

Type-II Seesaw Triplet Scalar Effects on Neutrino Trident Scattering arXiv:2204.05031 Yu Cheng, Xiao-Gang He, Zhong-Ly Huang and Ming-Wei Li

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Neutrino Trident Scattering (NTS) is a weak process:

A neutrino, scattering off a heavy nucleus, generates a pair of charged leptons.

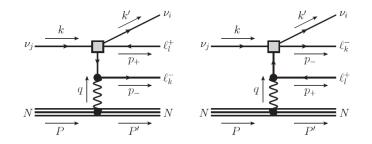


Figure 1: Diagrams for the $\nu_j \rightarrow \nu_i \ell_k^- \ell_l^+$.



- **1964** Trident scattering to examine the V-A theory. (Czyz, Sheppey, and Walecka)
- 1971 Momentum and angular distribution of the NTS. (Lovseth and Radomiski, Koike et al, Fujikawa)
- 1972 Trident scattering to examine the W-S theory. (Brown, Hobbs, Smith and Stanko)

 $\begin{array}{ll} \text{Measurements of } \nu_{\mu} \rightarrow \nu_{\mu} \mu^{+} \mu^{-}. \\ \textbf{1990} \quad \frac{\sigma_{\exp}}{\sigma_{\rm SM}} = 1.58^{+0.64}_{-0.64} & \text{CHARM-II. (Geiregat et al., 1990.)} \\ \textbf{1991} \quad \frac{\sigma_{\exp}}{\sigma_{\rm SM}} = 0.82^{+0.28}_{-0.28} & \text{CCFR. (S. R. Mishra et al., 1991.)} \\ \textbf{1999} \quad \frac{\sigma_{\exp}}{\sigma_{\rm SM}} = 0.72^{+1.73}_{-0.72} & \text{NuTeV. (T. Adams et al., 2000)} \\ \text{The average value } \frac{\sigma_{\exp}}{\sigma_{\rm SM}} = 0.95^{+0.25}_{-0.25} \end{array}$

Revival

2014 Trident scattering to constrain new physics. (Altmannshofer, Gori, Pospelov, and Yavin)

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In SM, the interaction terms between μ and ν_{μ} are

$$\mathcal{L}_{W} = \frac{g}{2\sqrt{2}} \left(\bar{\nu}_{\mu} \gamma^{\alpha} \left(1 - \gamma^{5} \right) \mu W_{\alpha}^{+} + \bar{\mu} \gamma^{\alpha} \left(1 - \gamma^{5} \right) \nu_{\mu} W_{\alpha}^{-} \right)$$
$$\mathcal{L}_{Z} = \frac{g}{4\cos\theta_{W}} \left(\bar{\nu}_{\mu} \gamma^{\alpha} \left(1 - \gamma^{5} \right) \nu_{\mu} + \bar{\mu} \gamma^{\alpha} \left(-1 + \gamma^{5} + 4\sin^{2}\theta_{W} \right) \mu \right) Z_{\alpha}$$
$$(1.1)$$

So in SM, a $\mu^+\mu^-$ pair can be generated by W and Z exchange from the operator

$$\frac{g^2}{8m_W^2}\bar{\nu}_{\mu}\gamma^{\alpha}\left(1-\gamma^5\right)\nu_{\mu}\bar{\mu}\gamma_{\alpha}\left(1-\gamma^5\right)\mu$$

$$\frac{g^2}{16m_W^2}\bar{\nu}_{\mu}\gamma^{\alpha}\left(1-\gamma^5\right)\nu_{\mu}\bar{\mu}\gamma_{\alpha}\left(-1+\gamma^5+4\sin^2\theta_W\right)\mu$$
(1.2)

where the momentum transfer is much smaller than m_W .

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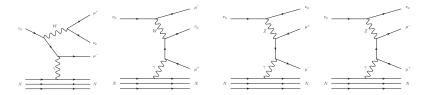


Figure 2: W contribution.

Figure 3: Z contribution.

In SM, the weak interaction contribution would be

$$\frac{g^2}{16m_W^2}\bar{\nu}_{\mu}\gamma^{\alpha}\left(1-\gamma^5\right)\nu_{\mu}\bar{\mu}\gamma_{\alpha}\left(1-\gamma^5+4\sin^2\theta_W\right)\mu\tag{1.3}$$



In Type-II Seesaw, one heavy Higgs triplet Δ is added¹, which contributes to reduce the cross section of NTS.

Table 1: Quantum number of the particles concerned with type-II seesaw. $Q = I_3 + Y$.

		$SU(2)_L$	$U(1)_Y$
L_{L_i}	$egin{pmatrix} u_{eL} \\ e_L \end{pmatrix}$	2	$-\frac{1}{2}$
Φ	$\begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix}$	2	$\frac{1}{2}$
Δ	$\begin{pmatrix} \Delta^+/\sqrt{2} & \Delta^{++} \\ \Delta^0 & -\Delta^+/\sqrt{2} \end{pmatrix}$	3	1

¹Konetschny and Kummer, 1977; Magg and Wetterich, 1980; Schechter and Valle, 1980; Cheng and Li, 1980; Lazarides et al., 1981; Mohapatra and Senjanovic, 1981.



The extra Yukawa coupling term

$$-\mathcal{L}_{M_{\nu}} = f_{ij} L_{L_i}^T C i \sigma^2 \Delta L_{L_j} + \text{h.c.}$$
(1.4)

1 lepton number violated by 2.

2 generate Majorana mass matrix of neutrino after SSB.





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2 generate Majorana mass matrix of neutrino after SSB.

Covariant derivative,

$$D_{\mu}\Phi = \partial_{\mu}\Phi - i\frac{g}{2}W^{a}_{\mu}\sigma^{a}\Phi - i\frac{g}{2}B_{\mu}\Phi$$

$$D_{\mu}\Delta = \partial_{\mu}\Delta - i\frac{g}{2}\left[W^{a}_{\mu}\sigma^{a}, \Delta\right] - ig'B_{\mu}\Delta.$$
(1.5)



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The Higgs sector is also extended

$$\mathcal{L}_{\text{Higgs}} = (D^{\mu}\Phi)^{\dagger} (D_{\mu}\Phi) + \text{Tr} \left[(D^{\mu}\Delta)^{\dagger} (D_{\mu}\Delta) \right] - V(\Phi, \Delta).$$
(1.6) where

$$V(\Phi, \Delta) = -m_{\Phi}^{2} \Phi^{\dagger} \Phi + M^{2} \operatorname{Tr} \Delta^{\dagger} \Delta + \left(\mu \Phi^{T} i \sigma^{2} \Delta^{\dagger} \Phi + \text{ h.c.} \right) + \frac{\lambda}{4} \left(\Phi^{\dagger} \Phi \right)^{2} + \lambda_{1} \Phi^{\dagger} \Phi \operatorname{Tr} \Delta^{\dagger} \Delta + \lambda_{2} \left(\operatorname{Tr} \Delta^{\dagger} \Delta \right)^{2} + \lambda_{3} \operatorname{Tr} \left(\Delta^{\dagger} \Delta \right)^{2} + \lambda_{4} \Phi^{\dagger} \Delta \Delta^{\dagger} \Phi.$$

$$(1.7)$$





After SSB,

$$\begin{split} \langle \Phi \rangle &\stackrel{SSB}{\Longrightarrow} \begin{pmatrix} 0 \\ v_d/\sqrt{2} \end{pmatrix} \quad \langle \Delta \rangle \stackrel{SSB}{\Longrightarrow} \begin{pmatrix} 0 & 0 \\ v_{\Delta}/\sqrt{2} & 0 \end{pmatrix} \\ m_{\Phi}^2 &= \frac{\lambda}{4} v_d^2 + \frac{\lambda_1 + \lambda_4}{2} v_{\Delta}^2 - \sqrt{2} \mu v_{\Delta} \\ M^2 &= -\frac{\lambda_1 + \lambda_4}{2} v_d^2 - (\lambda_2 + \lambda_3) v_{\Delta}^2 + \frac{\mu v_d^2}{\sqrt{2} v_{\Delta}} \sim m_{\Delta}^2 \equiv \frac{\mu v_d^2}{\sqrt{2} v_{\Delta}} \\ (1.8) \end{split}$$

we can generate the mass of gauge bosons and Higgs bosons

$$m_W^2 = \frac{g^2(v_d^2 + 2v_\Delta^2)}{4} \qquad m_Z^2 = \frac{(g^2 + g'^2)(v_d^2 + 4v_\Delta^2)}{4}.$$
 (1.9)



To show the mass of Higgs bosons, we write Φ and Δ in more explicit form

$$\Phi = \begin{pmatrix} \phi^+ \\ (v_d + h_d + iz_d)/\sqrt{2} \end{pmatrix}, \quad \Delta = \begin{pmatrix} \Delta^+/\sqrt{2} & \Delta^{++} \\ (v_\Delta + h_t + iz_t)/\sqrt{2} & -\Delta^+/\sqrt{2} \end{pmatrix}$$
(1.10)

Then from $\mathcal{L}_{\text{Higgs}}$, we can get the mass of Δ^{++} easily

$$m_{\Delta^{++}}^2 = \frac{\mu v_d^2}{\sqrt{2}v_{\Delta}} - \frac{\lambda_4}{2}v_d^2 - \lambda_3 v_{\Delta}^2 \approx \frac{\mu v_d^2}{\sqrt{2}v_{\Delta}} = m_{\Delta}^2.$$
(1.11)





While Δ^+ would mix with ϕ^+ , ie.

$$\mathcal{L}_{\text{Higgs}} \supset \left(\phi^{+} \Delta^{+}\right) M_{\text{charged}} \begin{pmatrix} \phi^{+} \\ \Delta^{+} \end{pmatrix},$$

$$M_{\text{charged}} = \begin{pmatrix} \sqrt{2}\mu v_{\Delta} - \frac{\lambda_{4}v_{\Delta}^{2}}{2} & -\mu v_{d} + \frac{\sqrt{2}}{4}\lambda_{4}v_{\Delta}v_{d} \\ -\mu v_{d} + \frac{\sqrt{2}}{4}\lambda_{4}v_{\Delta}v_{d} & \frac{\mu v_{d}^{2}}{\sqrt{2}v_{\Delta}} - \lambda_{4}v_{d}^{2} \end{pmatrix}$$

$$(1.12)$$

where det $(M_{\text{charged}})=0$ and $\text{Tr}(M_{\text{charged}})\neq 0$.





So there's a Goldstone boson

$$G^{\pm} \equiv \cos\beta\phi^{\pm} + \sin\beta\Delta^{\pm}, \qquad (1.13)$$

where $\tan \beta = \frac{\sqrt{2}v_{\Delta}}{v_d}$. It would be eaten up by W^{\pm}_{μ} boson.

And a massive physical field H^{\pm} is left,

$$H^{\pm} \equiv -\sin\beta\phi^{\pm} + \cos\beta\Delta^{\pm}, \quad m_{H^{\pm}}^2 = \frac{\left(2\sqrt{2\mu} - \lambda_4 v_{\Delta}\right)}{4v_{\Delta}} \left(v_d^2 + 2v_{\Delta}^2\right).$$
(1.14)

Because of the small mixing angle β , we can always treat H^+ as Δ^+ approximatively, eg.

$$m_{\Delta^+}^2 = m_{H^+}^2 \approx \frac{\mu v_d^2}{\sqrt{2}v_\Delta} = m_\Delta^2.$$

(1.15)



Acquire the neutrino mass matrix from the added Yukawa coupling terms $\mathcal{L}_{M_{\nu}}$,

$$-\mathcal{L}_{M_{\nu}} = f_{ij} L_{L_i}^T C i \sigma^2 \Delta L_{L_j} + \text{h.c.}$$

= $-f_{ij} \bar{\ell}_i^c P_L \ell_j \Delta^{++} - \sqrt{2} f_{ij} \bar{\nu}_i^c P_L \ell_j \Delta^{+} + f_{ij} \bar{\nu}_i^c P_L \nu_j \Delta^0 + \text{h.c.}$,
(1.16)

the last term would induce a Majorana-type mass matrix, ie.

$$m_{ij} = (M_{\nu})_{ij} = \sqrt{2} f_{ij} \nu_{\Delta} = \frac{\mu v_d^2}{m_{\Delta}^2} f_{ij},$$
 (1.17)

so the neutrino masses are suppressed by the mass of added Higgs boson m_{Δ} .

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$$-\mathcal{L}_{M_{\nu}} = -f_{ij}\bar{\ell}_i^c P_L \ell_j \Delta^{++} - \sqrt{2}f_{ij}\bar{\nu}_i^c P_L \ell_j \Delta^+ + f_{ij}\bar{\nu}_i^c P_L \nu_j \Delta^0 + \text{h.c.}$$

In type-II seesaw model, a $\mu^+\mu^-$ pair can be generated by Δ^+ ,

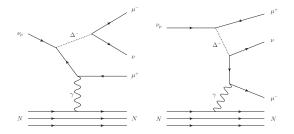


Figure 4: Type-II Seesaw contribution to NTS.

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Trident and Type-II See

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The effective operator would be

$$\frac{2f_{ij}f_{kl}^*}{m_\Delta^2}\bar{\nu}_i^c P_L\ell_j\bar{\ell}_l P_L\nu_k^c = \frac{m_{ij}m_{kl}^*}{2m_\Delta^2 v_\Delta^2}\bar{\nu}_k\gamma^\mu P_L\nu_i\bar{\ell}_l\gamma_\mu P_L\ell_j \qquad (2.1)$$

Notice that we should sum over all flavor neutrinos in final states. Then we get the final modification

$$\frac{\sigma}{\sigma_{SM}} = \frac{1}{\left(1 + 4s_w^2\right)^2 + 1} \left(\left(1 + 4s_w^2 - \frac{2m_W^2}{g^2} \frac{|m_{\mu\mu}|^2}{m_\Delta^2 v_\Delta^2} \right)^2 + \left(1 - \frac{2m_W^2}{g^2} \frac{|m_{\mu\mu}|^2}{m_\Delta^2 v_\Delta^2} \right)^2 + 2 \left(\frac{2m_W^2 |m_{\mu\mu}|^2}{g^2 m_\Delta^2 v_\Delta^2} \right)^2 \left(\frac{|m_{e\mu}|^2 + |m_{\tau\mu}|^2}{|m_{\mu\mu}|^2} \right) \right),$$
(2.2)

where $\frac{m_W^2}{g^2}$ could be replaced by the electroweak vev v because

$$\frac{4m_W^2}{g^2} = v_d^2 + 2v_\Delta^2 = v^2 = \frac{1}{\sqrt{2}G_F} = (246 \text{GeV})^2$$

(2.3)



Treat the ratio as a quadratic function of $\frac{1}{m_{\Lambda}^2 v_{\Lambda}^2}$

$$\frac{\sigma}{\sigma_{\rm SM}} = a \left(\frac{1}{m_{\Delta}^2 v_{\Delta}^2}\right)^2 + b \left(\frac{1}{m_{\Delta}^2 v_{\Delta}^2}\right) + 1, \qquad (2.4)$$

where

$$a = \frac{\left(v^2 |m_{\mu\mu}|^2\right)^2 \left(1 + \frac{|m_{e\mu}|^2 + |m_{\tau\mu}|^2}{|m_{\mu\mu}|^2}\right)}{2\left(\left(1 + 4s_w^2\right)^2 + 1\right)}.$$

$$b = -\frac{v^2 |m_{\mu\mu}|^2 \left(2 + 4s_w^2\right)}{\left(1 + 4s_w^2\right)^2 + 1}.$$
(2.5)

When degenerating to SM, we can see that

$$m_{\Delta}^2 v_{\Delta}^2 \to \infty, \qquad \frac{\sigma}{\sigma_{SM}} \to 1,$$

which means that the type-II contribution to trident scattering is negligible if $m_{\Delta}^2 v_{\Delta}^2$ is very large.



With the property of quadratic function, we can draw the figure,

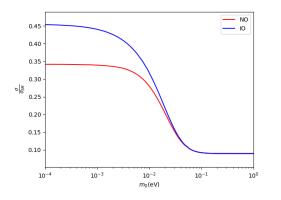


Figure 5: The rough lower bound of the ratio

 $\sigma_{\rm SM}$



Use upper bound of
$$\frac{1}{m_{\Delta}^2 v_{\Delta}^2}$$
 to constrain $\frac{\sigma}{\sigma_{SM}}$ better.
1 $\ell_i^- \to \ell_j^+ \ell_k^- \ell_l^-$
2 $\ell_i^- \to \ell_j^- \gamma$



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$$-\mathcal{L}_{M_{\nu}} = -f_{ij}\bar{\ell}_i^c\ell_j\Delta^{++} - \sqrt{2}f_{ij}\bar{\nu}_i^c\ell_j\Delta^{+} + f_{ij}\bar{\nu}_i^c\nu_j\Delta^{0} + \text{h.c.}$$

 $\ell_i^- \to \ell_j^+ \ell_k^- \ell_l^-$ arises at tree level and is mediated by Δ^{++} .

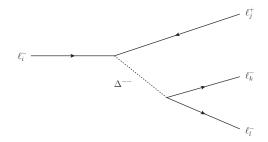


Figure 6: The diagram of $\ell_i^- \to \ell_i^+ \ell_k^- \ell_l^-$.





Its rate is found to be

$$\Gamma\left(\ell_{i}^{-} \to \ell_{j}^{+}\ell_{k}^{-}\ell_{l}^{-}\right) = \frac{1}{2\left(1+\delta_{kl}\right)} \frac{m_{\ell_{i}}^{5}}{192\pi^{3}} \left|\frac{f_{ij}f_{kl}}{m_{\Delta}^{2}}\right|^{2}.$$
 (2.6)

With the upper bound of branching ratio, we can give the lower bound of $m_{\Delta}^2 v_{\Delta}^2$.

Process	$Branching^2$	Constraint
$\mu^- \to e^+ e^- e^-$	1.0×10^{-12}	$m_{\Delta} v_{\Delta} > \left (M_{\nu})_{\mu e} (M_{\nu})_{ee} \right ^{1/2} \times 145 \text{TeV}$
$\tau^- \to \mu^+ e^- e^-$	1.5×10^{-8}	$m_{\Delta} v_{\Delta} > \left (M_{\nu})_{\tau\mu} (M_{\nu})_{ee} \right ^{1/2} \times 8.6 \text{TeV}$ $m_{\Delta} v_{\Delta} > \left (M_{\nu})_{\tau\mu} (M_{\nu})_{\mu e} \right ^{1/2} \times 8.8 \text{TeV}$
$\tau^- \to \mu^+ \mu^- e^-$	2.7×10^{-8}	$m_{\Delta} v_{\Delta} > \left (M_{\nu})_{\tau \mu} (M_{\nu})_{\mu e} \right ^{1/2} \times 8.8 \text{TeV}$
$\tau^- \to \mu^+ \mu^- \mu^-$	2.1×10^{-8}	$m_{\Delta} v_{\Delta} > \left (M_{\nu})_{\tau \mu} (M_{\nu})_{\mu \mu} \right ^{1/2} \times 7.9 \text{TeV}$

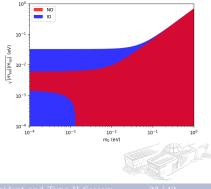
²R. Primulando, J. Julio and P. Uttayarat, 2019.



These constrains are weak. eg, for $\mu^- \rightarrow e^+ e^- e^-$,

Figure 7: The range of $|m_{\mu e}m_{ee}|^{1/2}$.

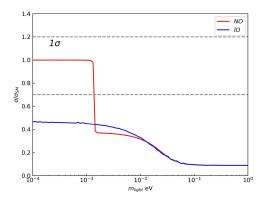
The lower bound of $m_{\Delta}^2 v_{\Delta}^2$ is significant only when lightest neutrino mass m_0 is smaller that 10^{-3} .





So the range of ratio is still very large.

Figure 8: The lower bound of the ratio $\frac{\sigma}{\sigma_{\rm SM}}$ with the constrain from $\mu^- \to e^+ e^- e^-$.





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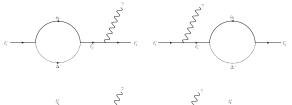


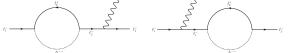
$$-\mathcal{L}_{M_{\nu}} = -f_{ij}\bar{\ell}_i^c \ell_j \Delta^{++} - \sqrt{2}f_{ij}\bar{\nu}_i^c \ell_j \Delta^+ + f_{ij}\bar{\nu}_i^c \nu_j \Delta^0 + \text{h.c.} .$$

 $\ell_i^- \to \ell_j^- \gamma$ occurs through one-loop penguin diagrams, mediated either by Δ^+ or Δ^{++} ,









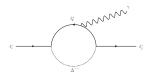


Figure 9: The diagram of $\ell_i^- \to \ell_j^- \gamma$.



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Neglected all internal lepton masses, we can get decay rate,

$$\Gamma\left(\ell_{i}^{-} \to \ell_{j}^{-} \gamma\right) = \frac{m_{\ell_{i}}^{5} \alpha_{em}}{(192\pi^{2})^{2}} \left|f^{\dagger}f\right|_{ij}^{2} \left(\frac{1}{m_{\Delta^{+}}^{2}} + \frac{8}{m_{\Delta^{++}}^{2}}\right)^{2} = \frac{m_{\ell_{i}}^{5} \alpha_{em}}{(192\pi^{2})^{2}} \left(\frac{9 \left|f^{\dagger}f\right|_{ij}}{m_{\Delta}^{2}}\right)^{2}.$$
(2.7)

Lower bound of $m_{\Delta}^2 v_{\Delta}^2$ could be

Process	$Branching^3$	Constraint
$\mu^- \to e^- \gamma$	4.2×10^{-13}	$m_{\Delta} v_{\Delta} > \sqrt{9 \left M_{\nu}^{\dagger} M_{\nu} \right _{\mu e}} \times 15.3 \text{TeV}$
$\tau^- \to e^- \gamma$	$3.3 imes 10^{-8}$	$m_{\Delta} v_{\Delta} > \sqrt{9 \left M_{\nu}^{\dagger} M_{\nu} \right _{\mu e}} \times 0.6 \text{TeV}$
$\tau^- \to \mu^- \gamma$	4.4×10^{-8}	$m_{\Delta} v_{\Delta} > \sqrt{9 \left M_{\nu}^{\dagger} M_{\nu} \right _{\mu e}} \times 0.56 \text{TeV}$

³R. Primulando, J. Julio and P. Uttayarat, 2019.



These constrains are much stronger, eg,

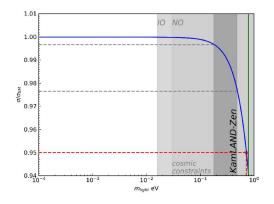
$$0.041 \text{eV} < 3\sqrt{\left|M_{\nu}^{\dagger}M_{\nu}\right|_{\mu e}} < 0.054 \text{eV}.$$
 (2.8)





So the range of ratio is very small.

Figure 10: The lower bound of the ratio $\frac{\sigma}{\sigma_{\rm SM}}$ with the constrain from $\mu^- \to e^- \gamma$.

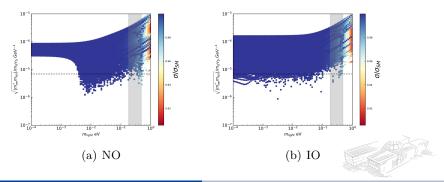






To combine the constraints from $\mu^- \to e^+e^-e^-$ and $\mu^- \to e^-\gamma$, **1** v_Δ in $(6.3 \sim 20) \text{eV}\left(\frac{100 \text{GeV}}{m_\Delta}\right) \Rightarrow \Gamma\left(\mu \to e\gamma\right)$ is satisfied. **2** scanning the parameter in 3σ region. **3** calculate the $\Gamma\left(\mu \to e^+e^-e^-\right)$ and $\frac{\sigma}{\sigma_{SM}}$.

Figure 11: Combine two constrains.



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The recent new measurement W mass by CDF,

$$m_W^{\text{CDF}} = 80,433.5 \pm 9.4 \text{MeV},$$

 7σ level above the SM prediction,

$$m_W^{\rm SM} = 80,357 \pm 6 {\rm MeV}.$$

Keeping other SM parameters unchanged as before,

$$\rho \equiv \frac{m_W^2}{m_Z^2 \cos^2 \theta_W} = 1.0019.$$





In type-II seesaw model, the mass of W and Z are changed as

$$m_W^2 = \frac{g^2(v_d^2 + 2v_{\Delta}^2)}{4} \qquad m_Z^2 = \frac{(g^2 + g'^2)(v_d^2 + 4v_{\Delta}^2)}{4},$$

so we can easily acquire

$$\rho = 1 + \Delta \rho = 1 - \frac{2v_{\Delta}^2}{v_d^2 + 4v_{\Delta}^2}.$$
 (3.1)

It's evidently smaller than 1, which is contradict to what we discussed before.



Considering

$$\frac{G_F}{\sqrt{2}} = \frac{g^2}{8m_W^2},$$
(3.2)

we get that $v_d^2+2v_\Delta^2=v^2=(246{\rm GeV})^2.$ Meanwhile ,

$$\rho = 1 - \frac{2v_{\Delta}^2}{v_d^2 + 4v_{\Delta}^2} = 1 - \frac{2v_{\Delta}^2}{v^2 + 2v_{\Delta}^2} = \frac{1}{1 + 2\left(\frac{v_{\Delta}}{v}\right)^2}$$
(3.3)

Significant effect on ρ needs v_{Δ} to be of order a GeV or so.

As the modification is in the wrong direction, so we want v_Δ to be little small

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Notice the constrains we discussed before,

$$m_{\Delta} v_{\Delta} > \sqrt{9 \left| M_{\nu}^{\dagger} M_{\nu} \right|_{\mu e}} \times 15.3 \text{TeV}$$

$$\Rightarrow v_{\Delta} > (6.25 - 8.39) \text{eV} \left(\frac{100 \text{GeV}}{m_{\Delta}} \right), \qquad (3.4)$$

which is compatible with experimental lower limit⁴ of a few hundred GeV of m_{Δ} .

So v_{Δ} could be very small.



⁴Particle Data Group Collaboration, P. A. Zyla et al., 2020.



To make the model in right direction, we can inrtoduce some new scalar Higgs bosons, the mass of W and Z are changed as

$$m_W = g^2 \sum_i \frac{v_i^2}{2} (I_i(I_i+1) - Y_i^2), \quad m_Z = (g^2 + g'^2) \sum_i (Y_i^2 v_i^2)$$
(3.5)

so that

$$\rho = \frac{m_W^2}{m_Z^2 \cos^2 \theta_W} = \frac{\sum_i v_i^2 (I_i(I_i+1) - Y_i^2)}{2\sum_i (Y_i^2 v_i^2)}.$$
 (3.6)

If their hypercharge Y is 0, they would give a positive contribution to ρ .



For example⁵, we can introduce a triplet ξ whose Y = 0, then

$$\rho = \frac{\frac{v_d^2}{2} + v_\Delta^2 + 2v_\xi^2}{2(\frac{v_d^2}{4} + v_\Delta^2)} = 1 - \frac{2v_\Delta^2}{v_d^2 + 4v_\Delta^2} + \frac{4v_\xi^2}{v_d^2 + 4v_\Delta^2}.$$
 (3.7)

If the contribution of v_{Δ} is small enough to neglect, with $\Delta \rho = 0.0019$ we can estimate that

$$v_{\xi} \approx 5.36 \text{GeV}.$$
 (3.8)



⁵J.-Y. Cen, J.-H. Chen, X.-G. He, and J.-Y. Su, 2018.

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• Combining constraints from
$$0\nu\beta\beta$$
, $\frac{\sigma}{\sigma_{\rm SM}} > 0.977$.





1 Combining constraints from
$$0\nu\beta\beta$$
, $\frac{\sigma}{\sigma_{\rm SM}} > 0.977$.

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 - A challenge to experimental test.









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Their vevs should reach of order a GeV.



Introduction 0000000000000	$\begin{array}{c} \mathbf{Constrians \ on \ NTS} \\ \texttt{ooooooooooooooooooo} \end{array}$	W mass 0000000	$\begin{array}{c} \mathbf{Conclusion}\\ 000 \end{array}$	$\mathbf{Discuss}$
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Thanks



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Ming-wei Li

Trident and Type-II Sees

Introduction 0000000000000	Constrians on NTS	W mass 0000000	Conclusion 000	$\mathbf{Discuss}$
Question			Ó	李改道研究所 TSUNG-DAO LEE INSTITUTE

Questions?



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