

# Status of muon polarization monitor

SJTU-KEK Workshop

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# Outline

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▷ Overview of muon polarization monitor

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Detector structure search

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Thermal muon simulation

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Cosmic muon simulation

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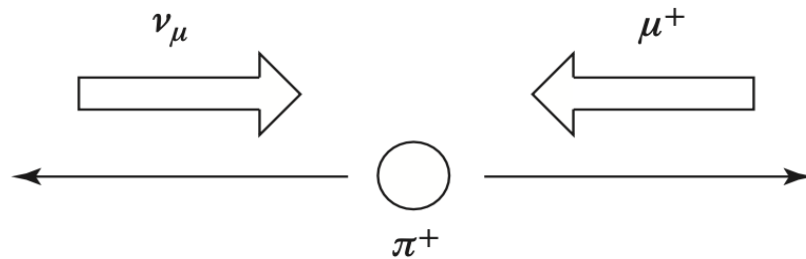
Conclusion and Future plan

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References

# How to detect muon polarization

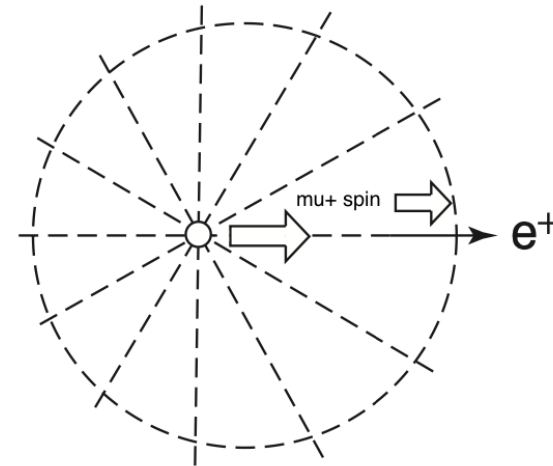
- Asymmetry angular distribution of  $e^+$  from  $\mu^+$  decay:



$$N_e(E_e) \propto [1 \mp A(E_e) \cos \theta_e]$$

$$A(E_e) = (E_e^{\max} - 2E_e) / (3E_e^{\max} - 2E_e)$$

upper sign:  $\mu^+$  decay,      lower sign:  $\mu^-$  decay

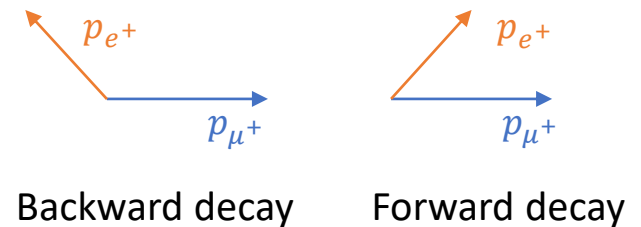


- By measuring the sign of  $e^+$  velocity's projection on  $\mu^+$  momentum, we can indirectly get the polarization fraction  $P$ :

- $N_F$  = Number of forward decay events

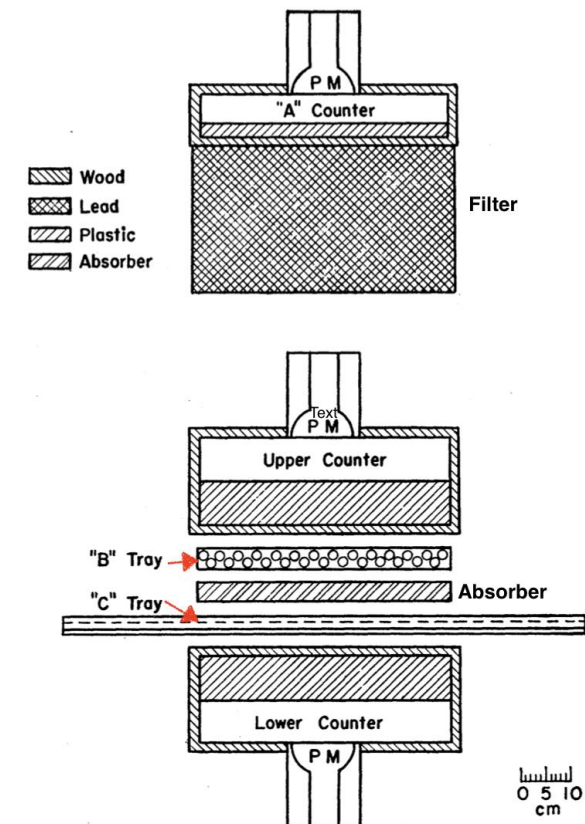
- $N_B$  = Number of backward decay events

- $R = N_F / N_B$ . In theory,  $R|_{P=1} < R|_{P=0.5} < R|_{P=0} = 1$



# Use cosmic muon to validate detector

- There is no muon source in Shanghai, therefore we plan to use muon from cosmic rays to calibrate detectors
- The polarization of cosmic muon is around 0.2-0.3
- The filter is used to select the energy range for a better yield
- Cu absorber: keep the polarization of cosmic muon
- Fe absorber: depolarize cosmic muon
- Design goal:
  - Select muons decaying in the absorber
  - Detect differences between Cu and Fe absorber
  - Less error
- An effective signal:  $ABC\bar{C}$



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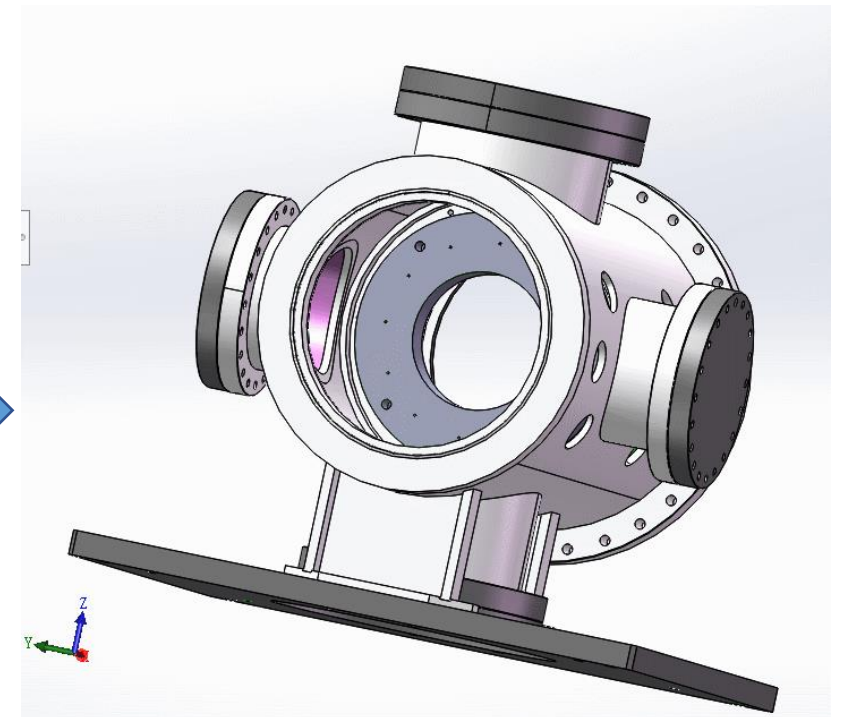
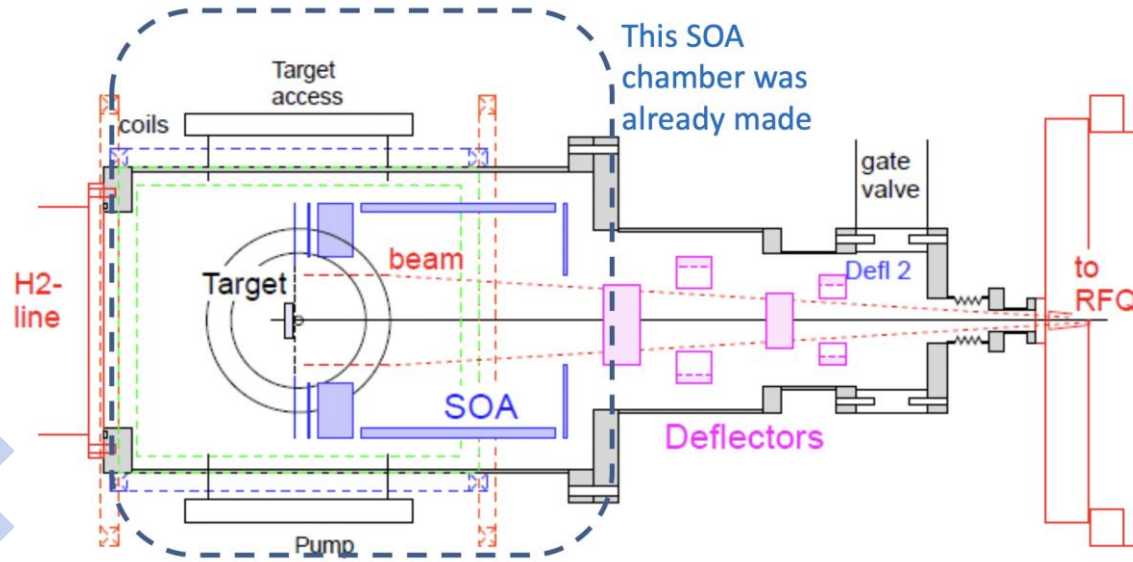
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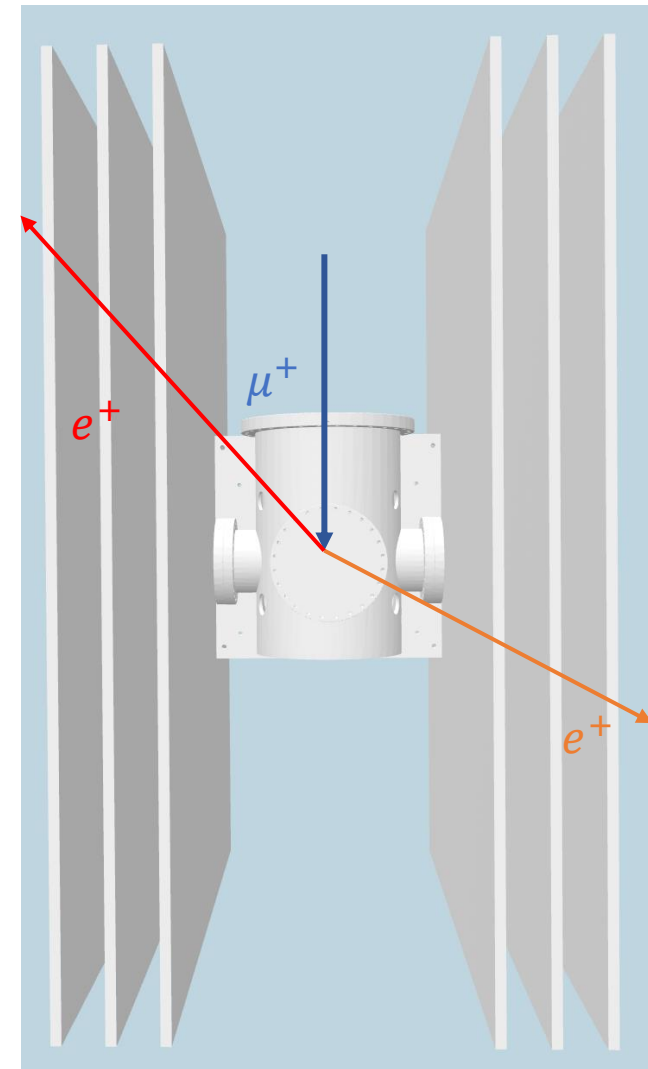
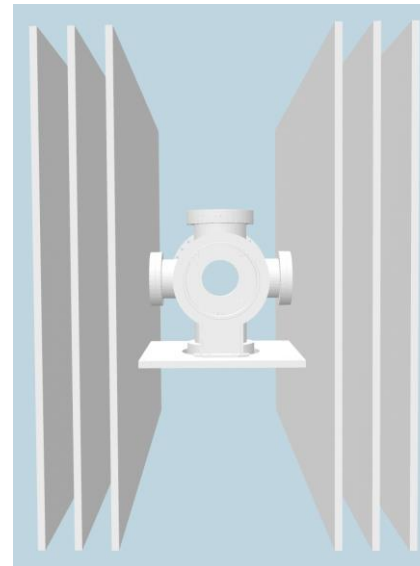
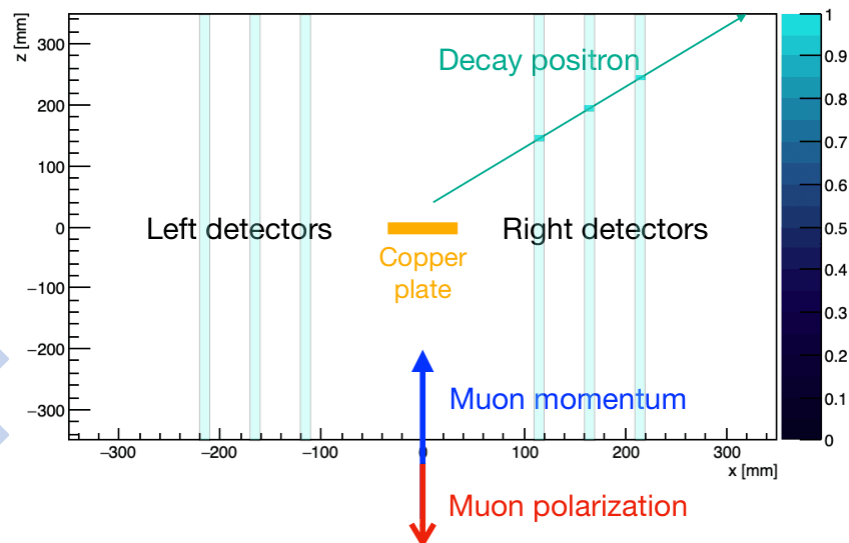
# Muon source chamber and its simplification

- Introduce chamber to simulation to study its absorbing and scattering influence on  $e^+$
- Use SolidWorks to defeature the chamber to a simpler one
- Reduce its size from 300 MB to 26 MB
- Similar performance, faster simulation speed



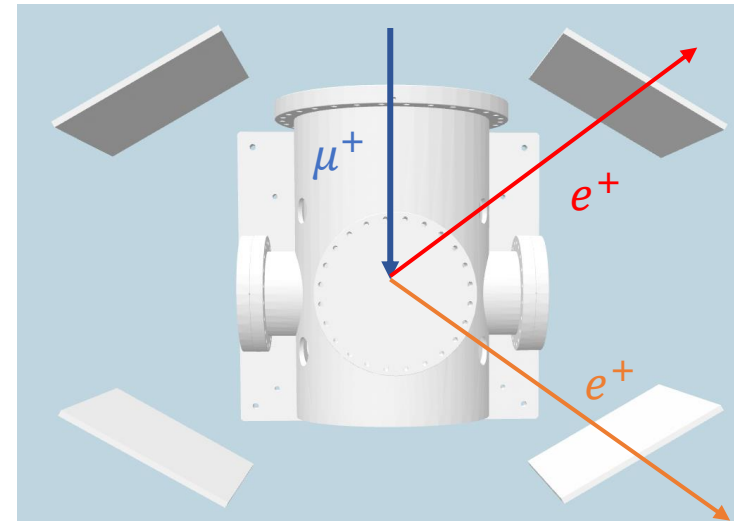
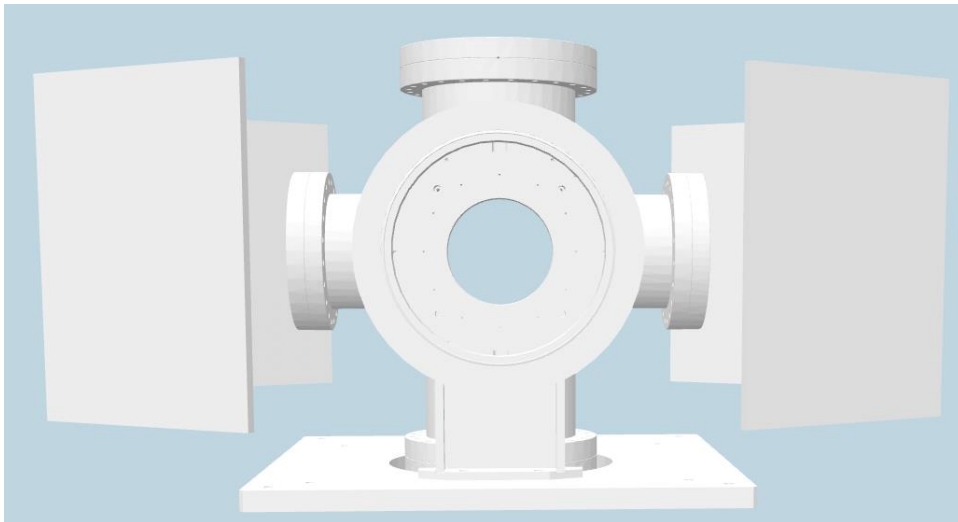
# Detector design: parallel detector

- 6 detectors, 2 sides, parallel to  $\mu^+$  momentum direction
- Positive gradient in momentum direction: Forward decay
- Negative gradient in momentum direction: Backward decay
- Each detector size: 1800 mm  $\times$  1400 mm  $\times$  10 mm



# Detector design: corner detector

- 4 detectors at each corner
- Hit signal on forward corner detector: Forward decay
- Hit signal on backward corner detector: Backward decay
- Each detector size: 300 mm × 400 mm × 10 mm





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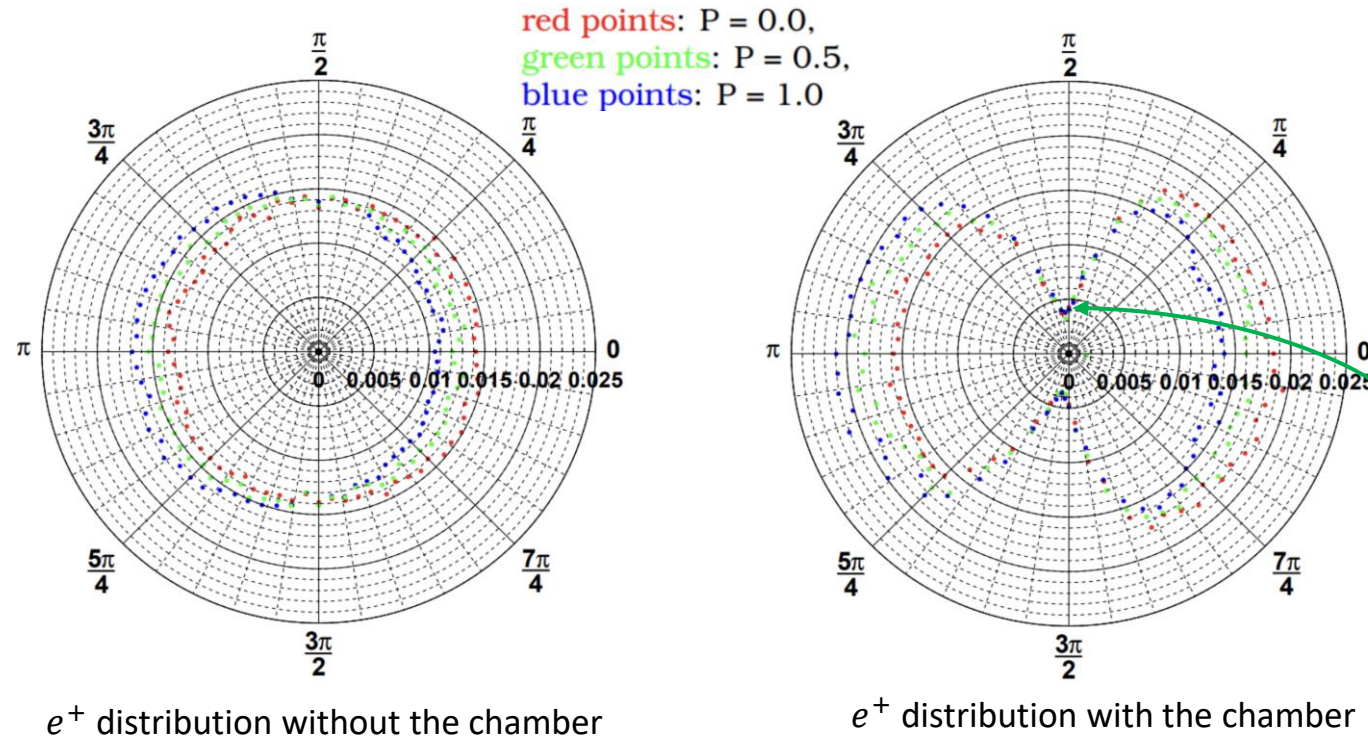
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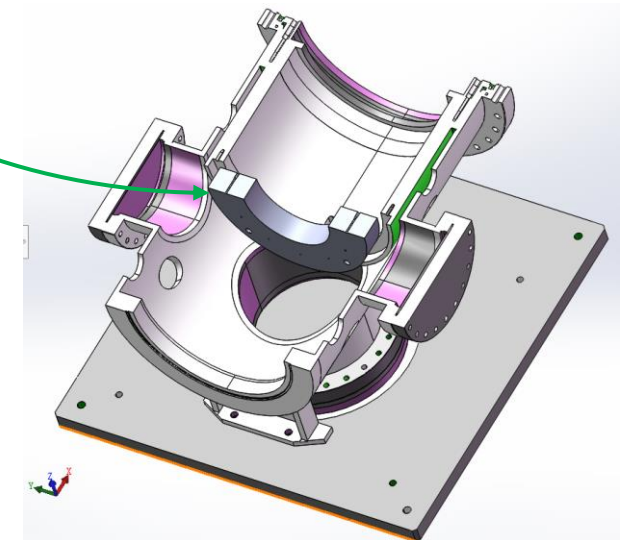
References

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# Angular distribution of $e^+$ from thermal muon

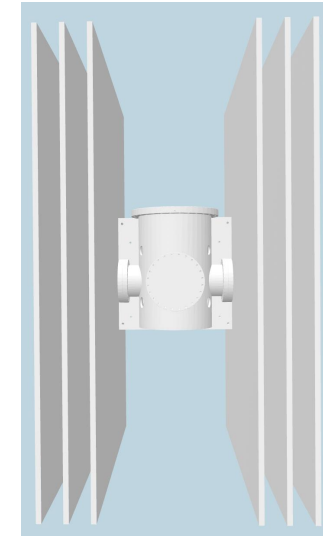


- After applying the chamber, asymmetry can still be detected outside the chamber
- The more  $e^+$  close to z angle, the stronger asymmetry it will show



# Parallel detector performance

- Event number:  $10^5$
- $F$  = Forward electron count;  $B$  = Backward electron count
- Asymmetry:  $A = \frac{F-B}{F+B}$
- Detector efficiency  $\epsilon = \text{Count number} / 10^5$

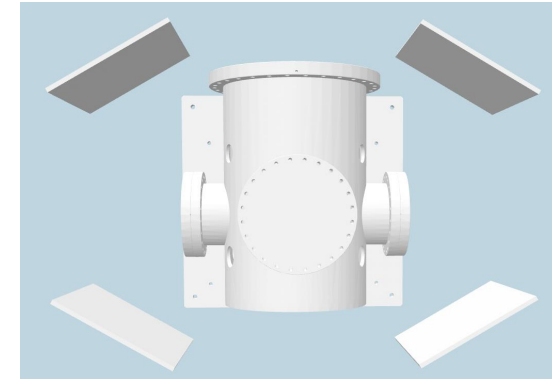


Polarization	$F$	$B$	$A$	$\epsilon$
0	6012	15535	-0.442	21.5%
0.5	5708	16805	-0.493	22.5%
1	5205	18056	-0.552	23.3%

- When  $\Delta P = 0.5$ ,  $\Delta A \approx 0.05$
- Costs high, but has lower SNR because of 3 layers
- Less influence on other parts of the apparatus

# Corner detector performance

- Event number:  $10^5$
- $F$  = Forward electron count;  $B$  = Backward electron count
- Asymmetry:  $A = \frac{F-B}{F+B}$
- Detector efficiency  $\epsilon = \frac{\text{Count number}}{10^5}$



Polarization	$F$	$B$	$A$	$\epsilon$
0	2160	6984	-0.53	9.1%
0.5	1886	7343	-0.59	9.2%
1	1507	8266	-0.69	9.8%

- When  $\Delta P = 0.5$ ,  $\Delta A \approx 0.09$
- Lower cost with better distinction performance, but lower efficiency
- Lower SNR since each corner only has 1 layer

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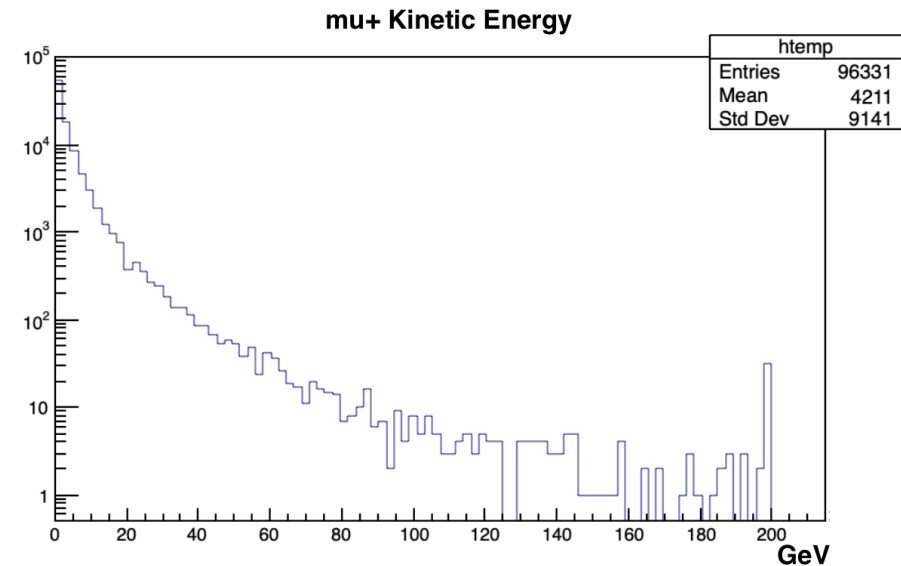
References

# Cosmic muon energy distribution

- Formula from reference[2]:

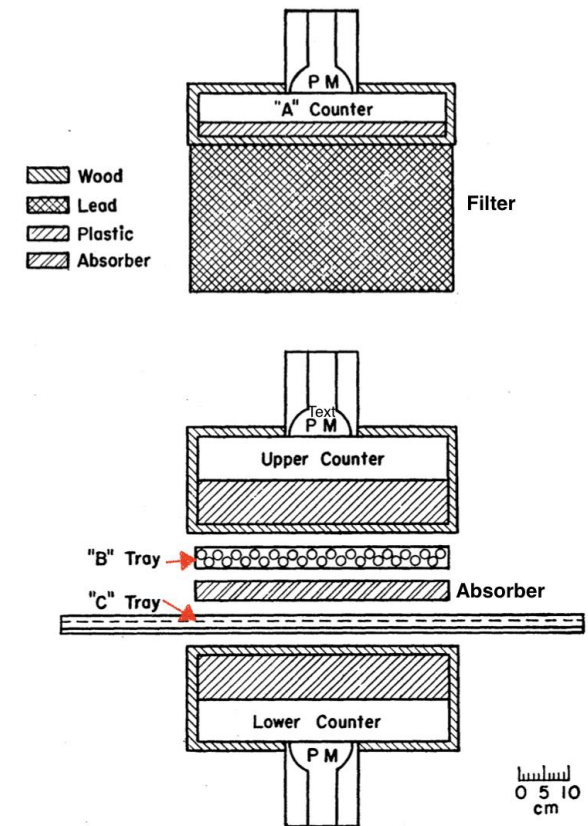
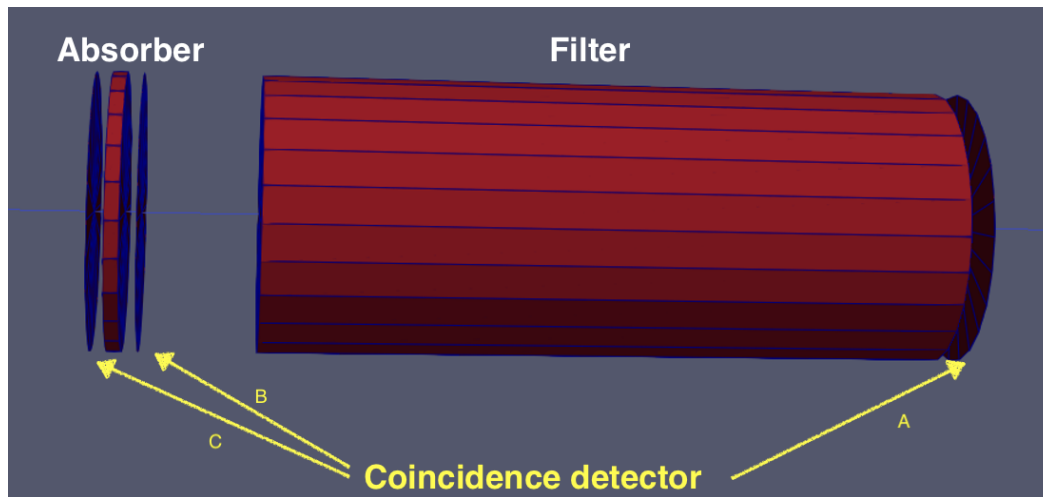
$$\frac{dI_\mu}{dE_\mu} = 0.14 \left[ \frac{E_\mu}{\text{GeV}} \left( 1 + \frac{3.64 \text{ GeV}}{E_\mu (\cos \theta^*)^{1.29}} \right) \right]^{-2.7} \times \left[ \frac{1}{1 + \frac{1.1 E_\mu \cos \theta^*}{115 \text{ GeV}}} + \frac{0.054}{1 + \frac{1.1 E_\mu \cos \theta^*}{850 \text{ GeV}}} \right]$$

- Currently set angle  $\cos \theta^* = 1$
- Simulation result:
  - Mean energy: 4.2 GeV, in agreement with the experiment
- This simulation is used as the muon source for further cosmic muon detector research



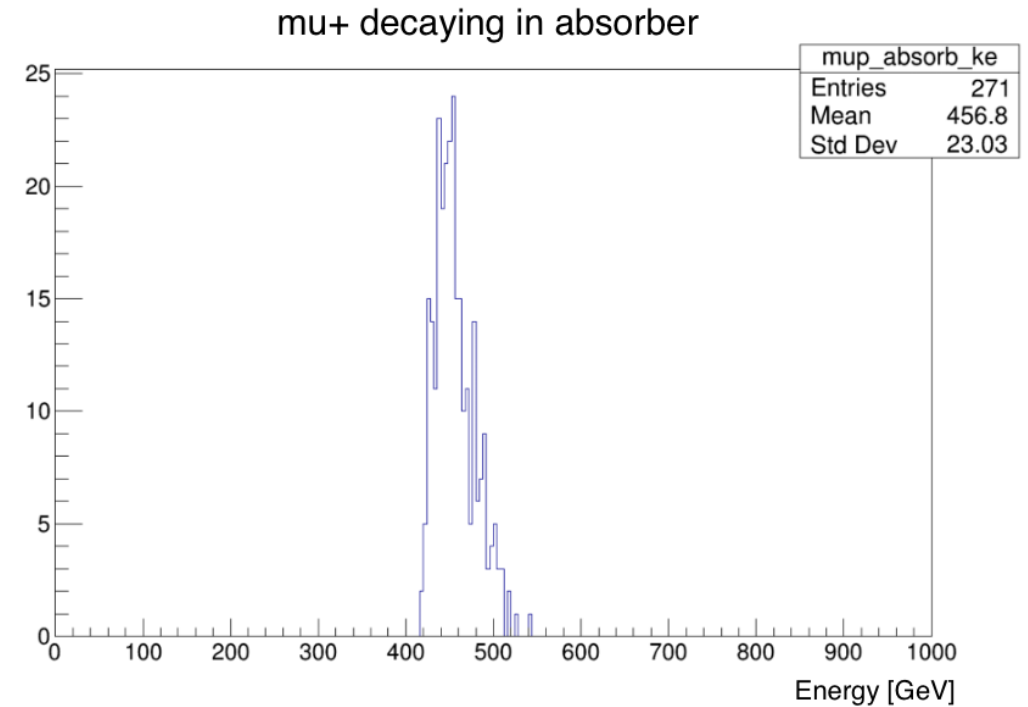
# Study on filter and absorber structure

- Use the cosmic muon source mentioned above
- Filter: lead, 300 mm
- Absorber: Cu, 63.5 mm (0.25 inch)
- A signal decay in absorber:  $ABC\bar{C}$



# Flux rate estimation

- The number of muons that decay in the absorber: 271 out of  $10^5$
- Energy range: 420 MeV – 520 MeV
- Assume absorber size: 50 cm × 50 cm
- Assume cosmic flux:  $200 \text{ m}^{-2} \text{ sec}^{-1}$
- Flux rate:
- $R = 200 \times 0.5 \times 0.5 \times \frac{271}{10^5} = 0.1355 \text{ sec}^{-1}$
- A signal will appear around each 7.38 seconds





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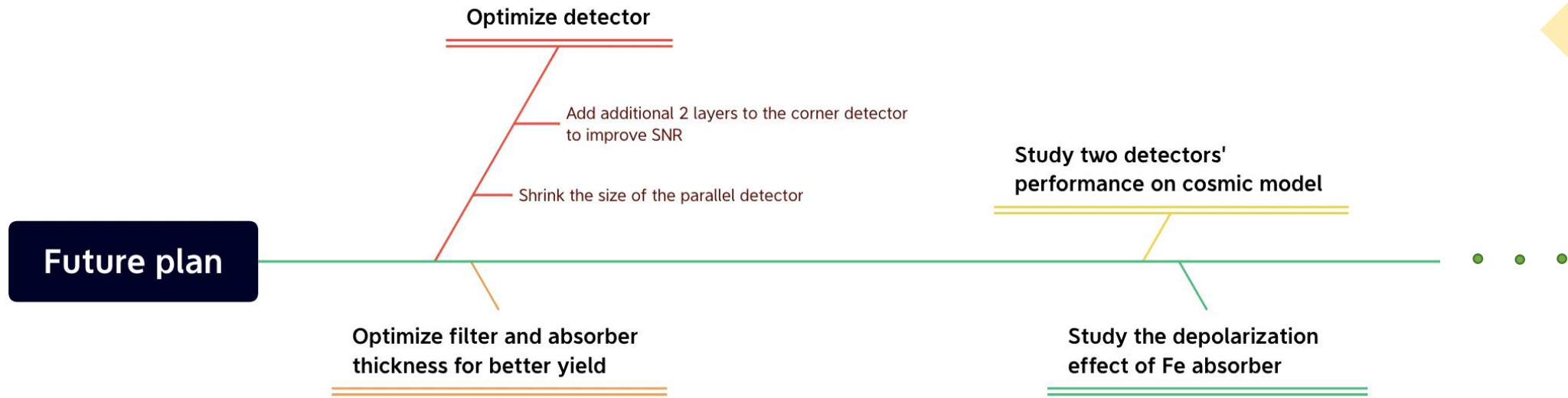
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References

# Conclusion

- ✓  $\mu^+ \rightarrow e^+$  decay asymmetry is the working principle of the detector
- ✓ We use cosmic muon to hit different absorbers to calibrate the polarization detector
- ✓  $e^+$  will still show asymmetry after adding the chamber
- ✓ The chamber will absorb most  $e^+$  near angle  $\pi/2$  and  $3\pi/2$ , therefore affecting the performance of the parallel detector
- ✓ Because of absorbing effect of the chamber, the corner detector performs better than the parallel detector with rather less cost
- ✓ Using given filter and absorber, we can obtain a  $\mu^+$  decaying in the absorber each 7.38 seconds

# Future plan



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▷ References

# References

- [1] Nagamine, K. (2003). Introductory Muon Science. Cambridge: Cambridge University Press  
doi:10.1017/CBO9780511470776
- [2] Johnson, C. Scott. (1961). Polarization of Cosmic-Ray Muons at Sea Level: PhysRev.122.1883
- [3] Guan, Mengyun and Chu et al. (2015). A parametrization of the cosmic-ray muon flux at sea-level  
doi:10.48550/ARXIV.1509.06176



Thanks!