Millicharged particles from proton bremsstrahlung in the atmosphere

Rundong Fang

Nanjing University

In collaboration with Zuowei Liu, Mingxuan Du, arXiv:2211.abcde

The 2022 Shanghai Particle Physics and Cosmology Symposium:
Neutrino and Dark Matter Physics (SPCS 2022)

2022.11.18

Content

- Motivation
- Production channels
- Earth attenuation and signals
- Constraints from SuperK
- Conclusion

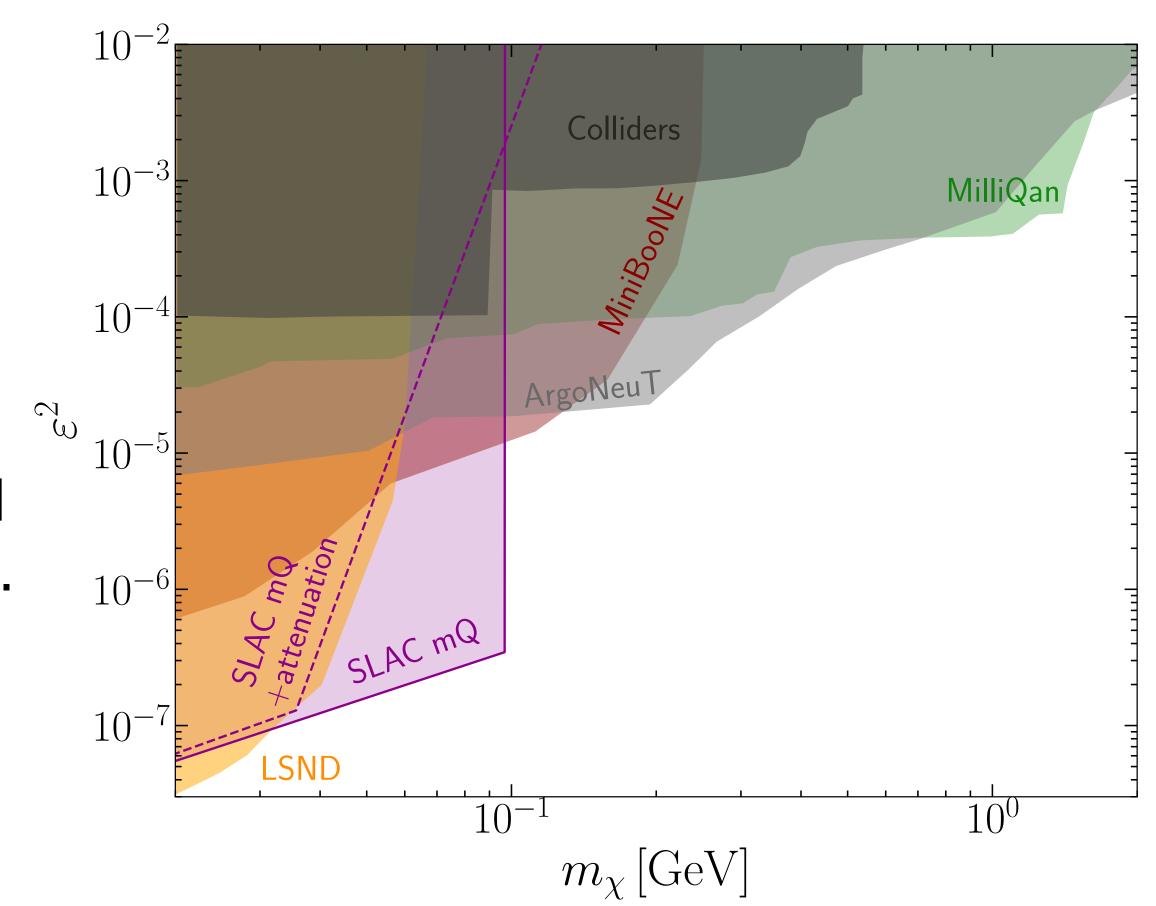
Motivation

 Millicharged particles (MCPs) are wellmotivated beyond the Standard Model particles with small electric charge:

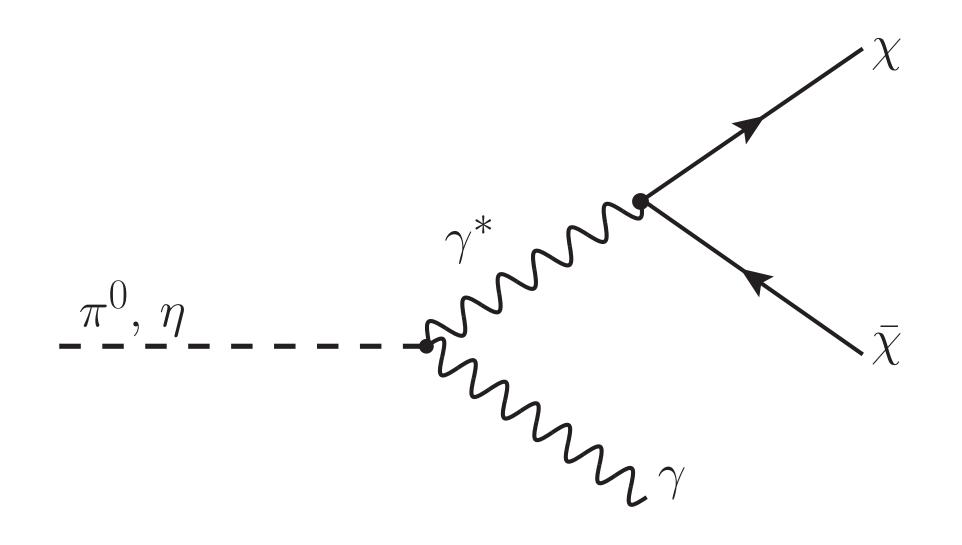
$$\mathcal{L}_{\rm int} = \varepsilon e A_{\mu} \bar{\chi} \gamma^{\mu} \chi$$

• The collision between cosmic ray protons and the atmosphere can copiously produce MCPs.

 MCPs produced in atmosphere can be detected by underground detectors.

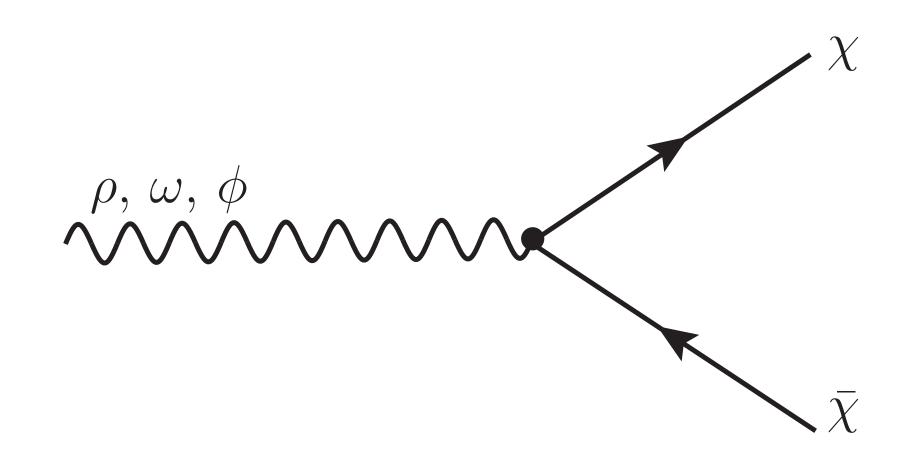


Meson Decay



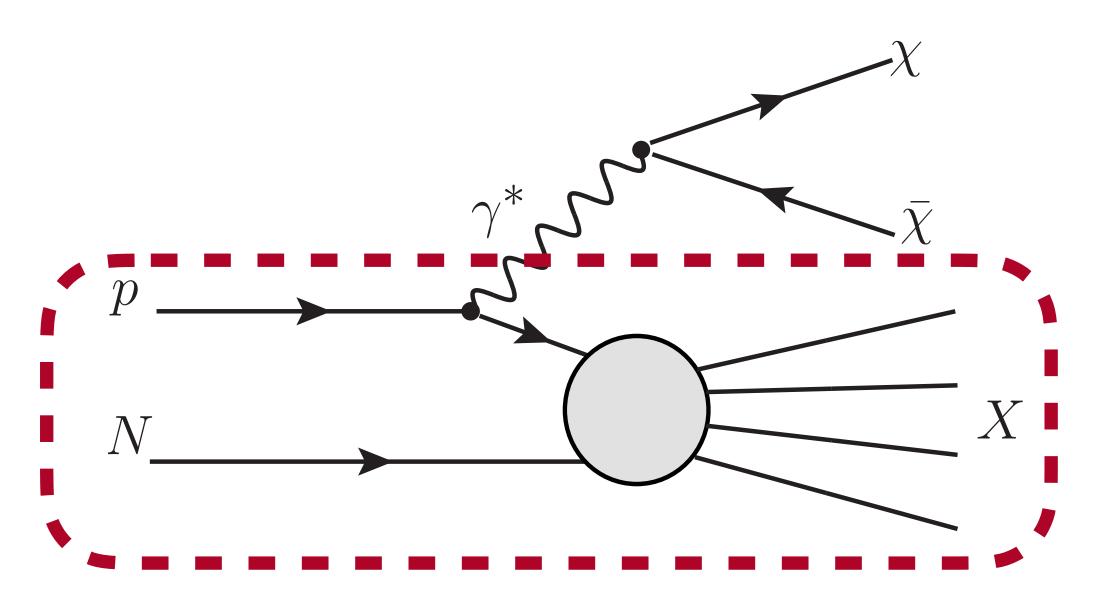
• In previous studies, the MCPs are produced via meson decay.

Plestid et al., 2022.11732 Kachelriess et al., 2104.06811 Arguelles et al., 2104.13924



 MCPs can be produced via the Dalitz decay of pseudo-scaler mesons and two body decay of vector mesons.

Proton Bremsstrahlung



We include a new process, proton bremsstrahlung, in the MCP production.

Factorization:

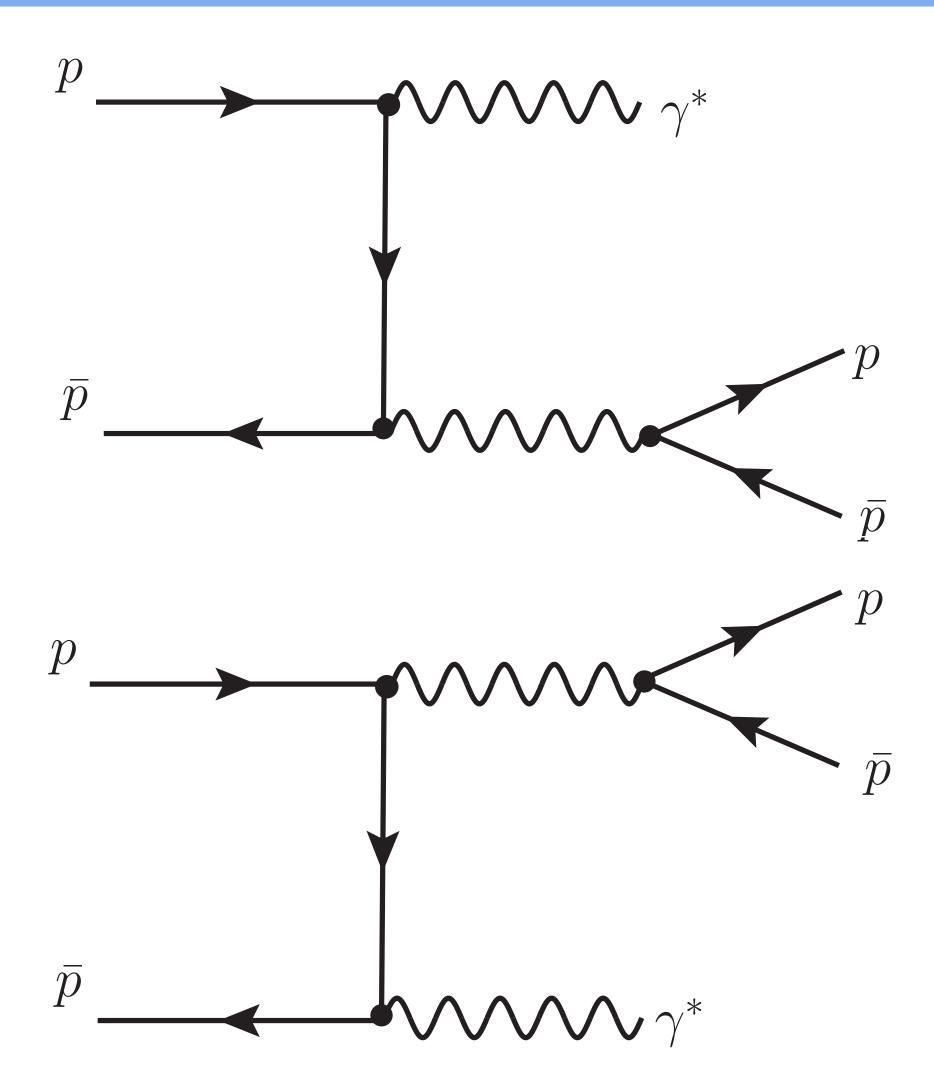
$$d\sigma_{pN\to\gamma^*X}(s) = d\mathcal{P}_{p\to\gamma^*p} \times \sigma_{pN\to X}(s')$$

The $d\mathcal{P}_{p\to\gamma^*p}$ is the splitting kernel.

The splitting kernel is usually given by the FWW approximation when all of the particles are highly relativistic.

Blumlein et al., 1311.3870

Splitting Kernel

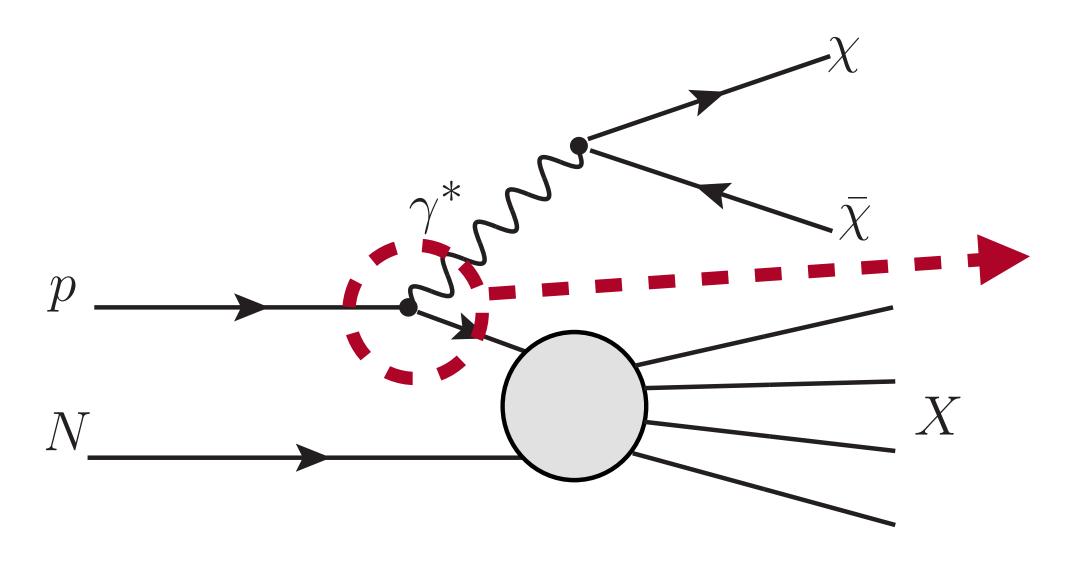


$$\frac{d\Phi_p}{dE_p d\Omega}(h_{\text{max}}) = \frac{0.74 \times 1.8 \times 10^4}{\text{m}^2 \text{ s sr GeV}} \left(\frac{E_p}{\text{GeV}}\right)^{-2.7}$$

The low energy protons are important

$$d\mathcal{P}_{p\to p\gamma^*} = \frac{d\sigma_{p\bar{p}\to\gamma^*p\bar{p}}(s)}{\sigma_{p\bar{p}\to p\bar{p}}(s')}$$

Form Factor



The radiation process is enhanced by the time-like form factor of proton.

$$F_{1}(q^{2}) = \sum_{V=\rho \, \rho' \, \rho'' \, \omega \, \omega' \, \omega''} \frac{f_{V} m_{V}^{2}}{m_{V}^{2} - q^{2} - i m_{V} \Gamma_{V}}$$

Faessler et al., 0910.5589

$$m_{\rho} \approx m_{\omega} \approx 0.77 \text{ GeV}$$

Cascade Equation

M. Thunman et al., hep-ph/9505417

• The MCP flux at the surface of Earth:

$$\frac{d\Phi_{\chi}^{s}}{dE_{\chi}^{s}} = \int \int dh dE_{p} \frac{d\Phi_{p}}{dE_{p}}(h) n_{T}(h) \sigma_{pT}(E_{p}) \sum_{i} \frac{dN_{\chi}^{i}}{dE_{\chi}}$$

• The proton flux at a given height:

$$\frac{d}{dh} \left(\frac{d\Phi_p}{dE_p}(h) \right) = \sigma_{pT}(E_p) n_T(h) \frac{d\Phi_p}{dE_p}(h)$$

Earth Attenuation

Gaisser et al., Cosmic Rays and Particle Physics

• MCPs lose energy when travel through Earth:

$$-\frac{dE}{dX} \approx \varepsilon^2 (a + bE)$$

MCP flux after earth attenuation

$$\frac{d^2\Phi_{\chi}^D(X)}{dE_{\chi}d\Omega} = e^{\varepsilon^2 bX} \frac{d^2\Phi_{\chi}^s}{dE_{\chi}^s d\Omega^s}$$

$$E_{\gamma}^{s} = (E_{\gamma} + a/b) \exp(\varepsilon^{2}bX) - a/b$$

Signal

• The signals of MCP come from the elastic scattering with electron in the detector.

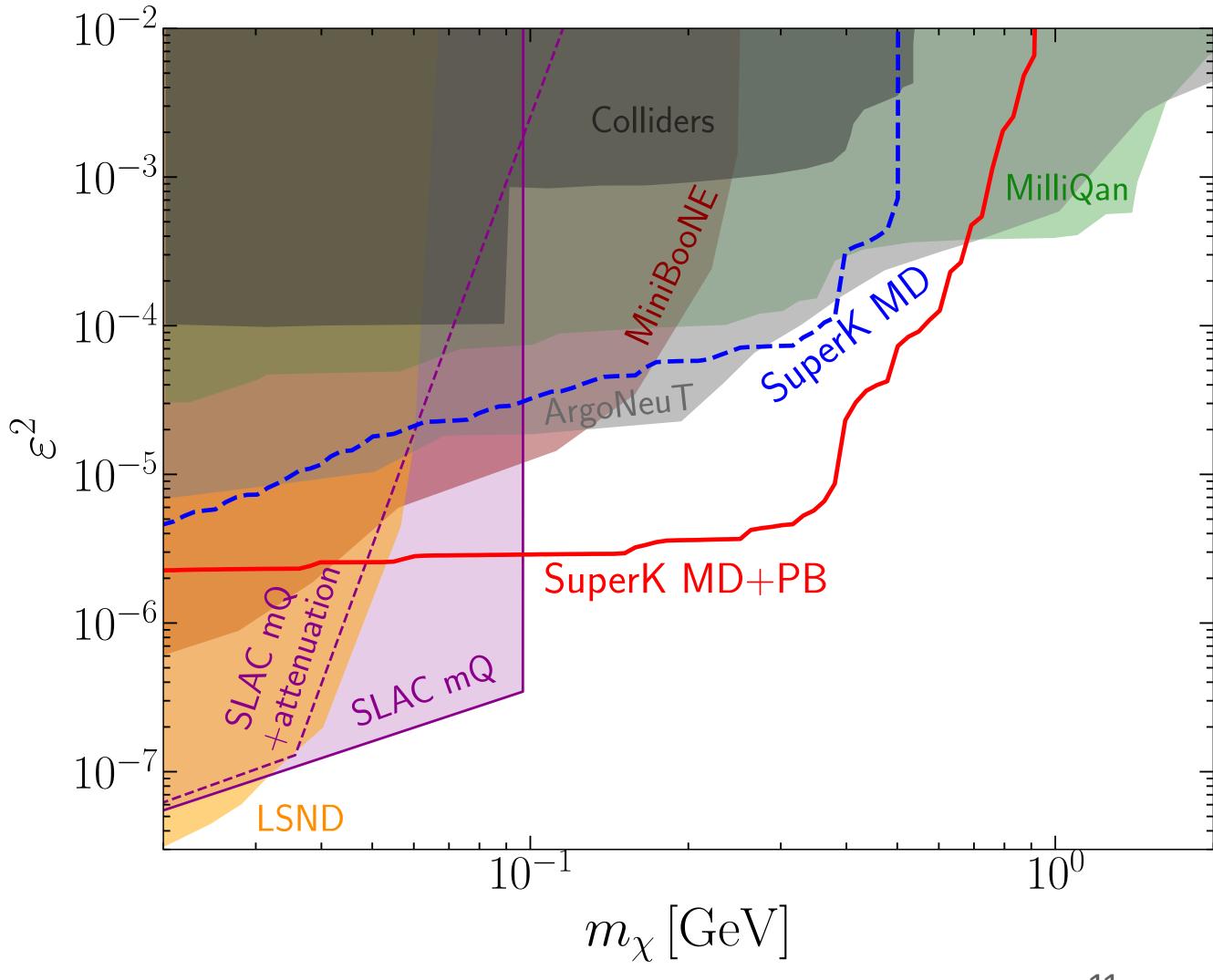
$$\frac{d\sigma}{dE_r} = \varepsilon^2 \alpha^2 \pi \frac{m_e \left(E_r^2 + 2E_\chi^2\right) - E_r \left(m_e \left(2E_\chi + m_e\right) + m_\chi^2\right)\right)}{E_r^2 m_e^2 \left(E_\chi^2 - m_\chi^2\right)}$$

• signals per bin:

$$S_{i} = 2\pi n_{e} \mathcal{E} \int dE_{r} f(E_{r}) \int dE_{\chi} \int d\cos\theta \frac{d^{2}\Phi_{\chi}^{D}}{dE_{\chi} d\cos\theta} \frac{d\sigma}{dE_{r}}$$

We use data and signal efficiency form SuperK phase I to phase III to set the constraint to MCPs.
 Arguelles et al., 2104.13924

SuperK constraints



Blue: meson decay

Red: meson decay plus proton

bremsstrahlung

Conclusion

- The MCPs can be copiously produced in the collision between cosmic ray protons and the atmosphere.
- We include a new production channel, proton bremsstrahlung, of MCPs which is enhanced by the resonance of the time-like form factor of proton in sub-GeV region.
- After including the PB process, the SuperK constraint for MCPs become much stronger than the meson decay processes.
- These enhancement will also appear in other pN collision experiments.

Thank you!

Detail Calculation of the Splitting Kernel

• The splitting kernel is calculated using the initial state radiation process $p\bar{p} \to p\bar{p}\gamma^*$, and in CM frame:

$$\frac{d^2 \mathcal{P}_{p \to \gamma^* p}}{dE_{\gamma^*}^0 d \cos \theta_{p,\gamma^*}^0} = \frac{1}{512\pi^4 E_p^0 E_{\bar{p}}^0 \left| v_p^0 - v_{\bar{p}}^0 \right|} \frac{\left| F_p(m_{\chi\bar{\chi}}^2) \right|^2}{\sigma_{2 \to 2}^s(s_1')} \int dE_{p'}^0 \int d\phi_{p',\gamma^*}^0 \left| \mathcal{M}_{2 \to 3}^s \right|^2$$

 Then we boost the splitting kernel back to the lab frame to calculate differential cross section:

$$\frac{d\sigma_{\text{PB}}}{dE_{\gamma^*}} = \int dE_{\gamma^*}^0 d\cos\theta_{p,\gamma^*}^0 \frac{d^2\mathcal{P}_{p\to\gamma^*p}}{dE_{\gamma^*}^0 d\cos\theta_{p,\gamma^*}^0} \sigma_{p,\text{air}}(s_2') \delta(E_{\gamma^*}' - E_{\gamma^*})$$

Proton Bremsstrahlung

$$d\sigma(P_1 + P_2 \to \gamma^* X \to \chi \bar{\chi} X) = d\sigma \left(P_1 + P_2 \to \gamma^* X \right) \times \frac{Q_{\chi}^2}{12\pi^2} \frac{dk_{\gamma^*}^2}{k_{\gamma^*}^2} \sqrt{1 - \frac{4m_{\chi}^2}{k_{\gamma^*}^2}} \left(1 + \frac{2m_{\chi}^2}{k_{\gamma^*}^2} \right).$$

S. N. Gninenko, et al. 1810.06856