Statistical properties of repeating FRBs, in comparison with earthquakes and solar flares

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Fast radio bursts trigger aftershocks resembling earthquakes, but not solar flares

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ABSTRACT

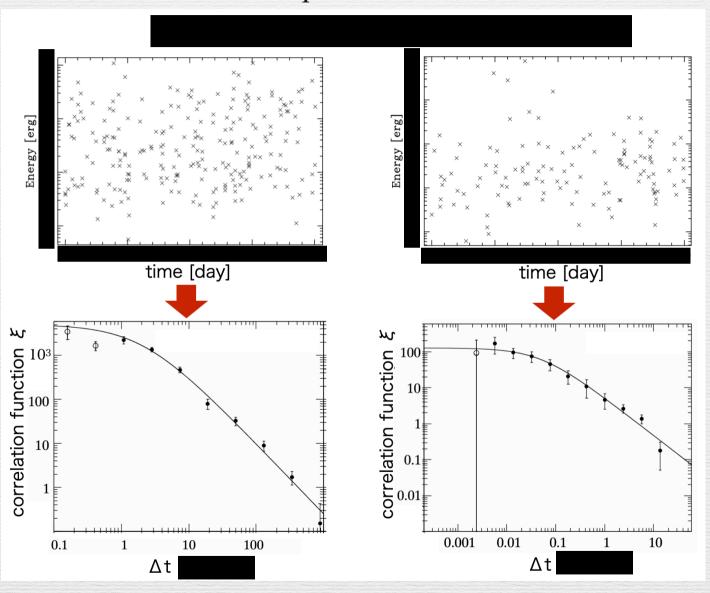
The production mechanism of repeating fast radio bursts (FRBs) is still a mystery, and correlations between burst occurrence times and energies may provide important clues to elucidate it. While time correlation studies of FRBs have been mainly performed using wait time distributions, here we report the results of a correlation function analysis of repeating FRBs in the 2D space of time and energy. We analyse nearly 7,000 bursts reported in the literature for the three most active sources of FRB 20121102A, 20201124A, and 20220912A, and find the following characteristics that are universal in the three sources. A clear power-law signal of the correlation function is seen, extending to the typical burst duration (\sim 10 msec) towards shorter time intervals (Δt). The correlation function indicates that every single burst has about a 10–60 per cent chance of producing an aftershock at a rate decaying by a power law as $\propto (\Delta t)^{-p}$ with p = 1.5–2.5, like the Omori–Utsu law of earthquakes. The correlated aftershock rate is stable regardless of source activity changes, and there is no correlation between emitted energy and Δt . We demonstrate that all these properties are quantitatively common to earthquakes, but different from solar flares in many aspects, by applying the same analysis method for the data on these phenomena. These results suggest that repeater FRBs are a phenomenon in which energy stored in rigid neutron star crusts is released by seismic activity. This may provide a new opportunity for future studies to explore the physical properties of the neutron star crust.

Key words: Sun: flares - stars: neutron - radio continuum: transients - fast radio bursts.

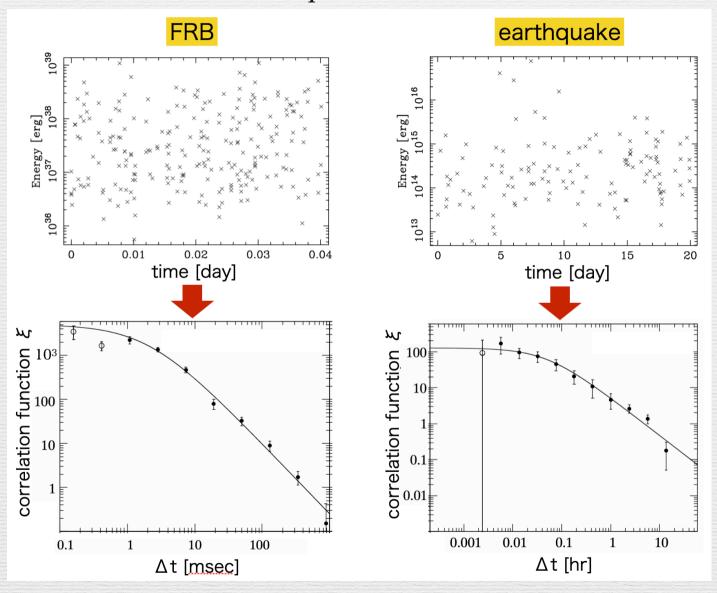
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FRB & earthquake… which is which?

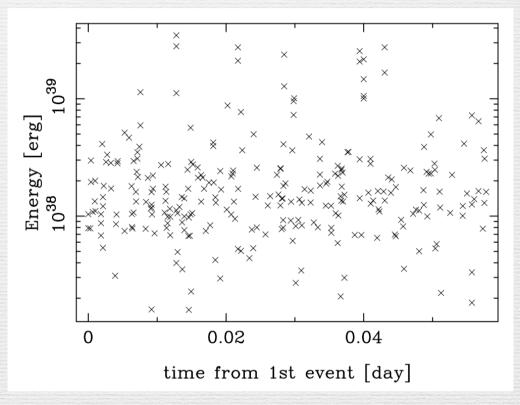


FRB & earthquake… which is which?



So many bursts from repeater FRBs!

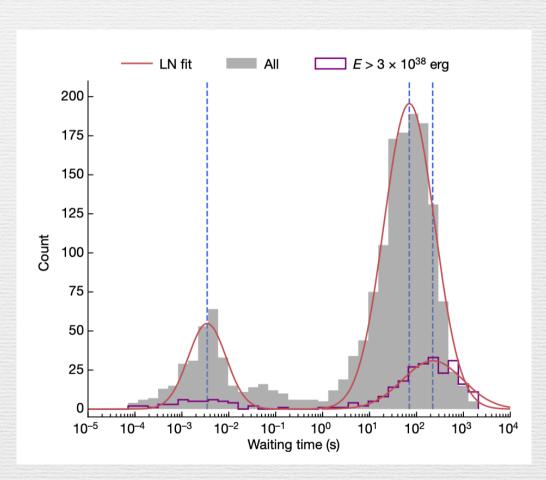
- Thousands of bursts from a few repeater FRB sources
- > 100 bursts in 1 hr
- Detailed studies on the statistical nature of these bursts now possible



A part of Jahns+'23 data for FRB 20121102A

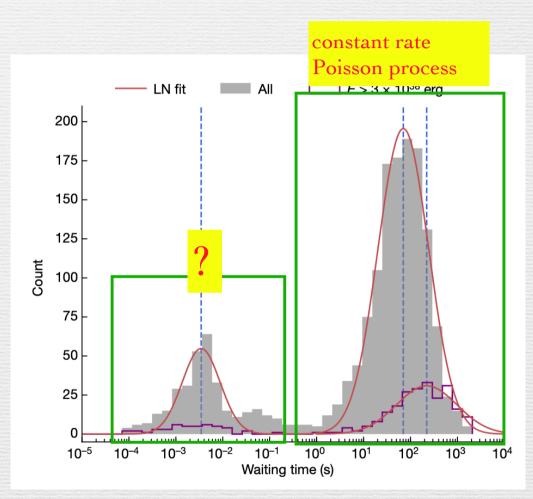
Statistical properties of repeating FRB occurrence time?

- bimodal wait-time distribution found universally for all repeater FRB sources
- The peak at longer wait times is consistent with a Poisson process with a constant event rate
- · The origin of short wait-time peak is unknown.
 - peaks at 1-10 msec, close to the duration of one FRB.
 - Related to radiative process/source activity?



Statistical properties of repeating FRB occurrence time?

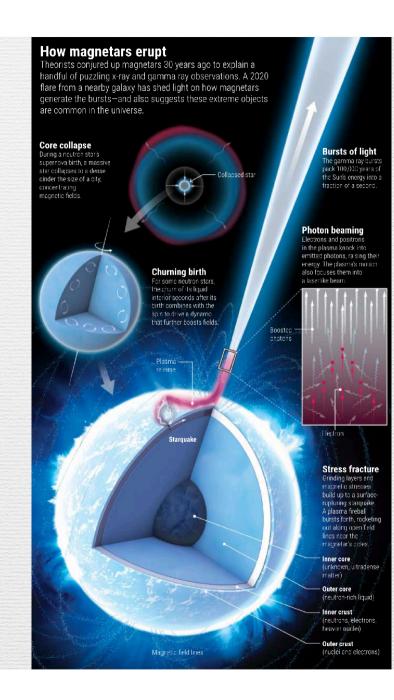
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Li+'21

FRBs vs. earthquakes and solar flares

- FRB statistical properties may be similar to earthquakes or solar flares
 - FRBs related to magnetars (e.g. SGR 1935+2154)
 - magnetar flares thought to occur by starquakes in the surface solid crust of a neutron star, induced by magnetic energy
 - similarity between magnetars, earthquakes, and solar flares investigated in the literature
 - e.g. the power-law energy distribution are common for magnetar bursts and earthquakes (Wadati-Gutenberg-Richter law of earthquakes)



Wadati? Gutenberg-Richter law?

- The power-law distribution of earthquake magnitudes (energies) often called "Gutenberg-Richter" (1944) law
 - $\log N(>M) \propto M^{-b}$, $dN/dE \propto E^{-1-2b/3}$, $b\sim 1$
 - \cdot M = magnitude, E = energy
- · But···
- ·和達清夫 (WADATI Kiyoo, 1902-1995, famous by Wadati-Benioff zone) found this law earlier in 1932
 - · Wadati, K. "On the frequency distribution of earthquakes." Journal of the Meteorological Society of Japan. Ser. II 10, 559-568 (1932) (in Japanese).

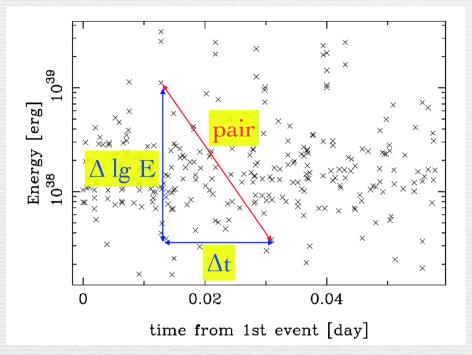
地震囘數の分布に就いて

氣象集誌第二輯第十卷第十號

(中央氣象臺彙報)



What we did: correlation function ξ in time-energy space



data for FRB 20121102A from Jahns+'23

• two-point correlation function ξ in the space of Δt and $\Delta \lg E$ (= $\Delta \log_{10} E$)

$$dn_p = (1 + \xi) \,\bar{n}_p \, d(\Delta t) \, d(\Delta \lg E) ,$$

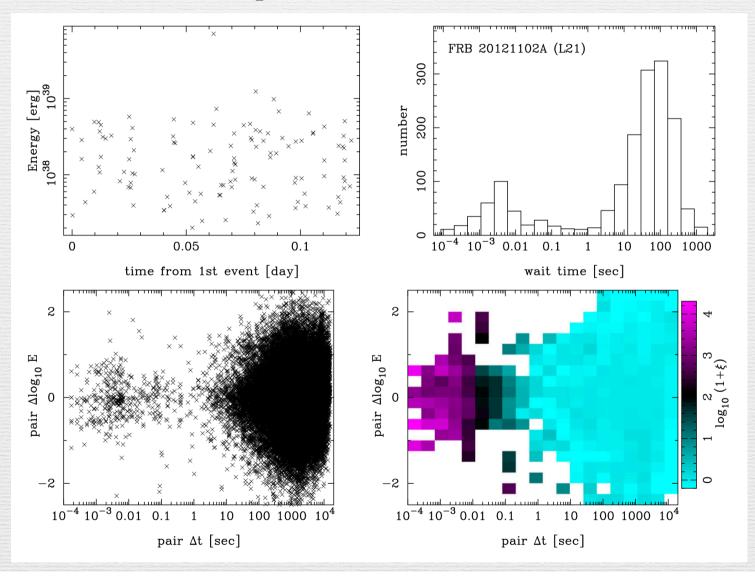
- \cdot ξ is the excess of pair counts compared with the case of no correlation (\bar{n}_p)
- random data (no correlation) is produced assuming "constant event rate" and "constant energy distribution" during one-day observation (~ a few hours)

7 FRB data sets for 3 sources

- nearly 7,000 events in total, from Arecibo & FAST
- · 3 sources (FRBs 20121102A, 20201124A, 20220912A)

Data set name	Telescope	Period	Days^a	$t_{ m obs}{}^b$	Events	$r_m{}^c$	$C_{ m best}{}^d$	$p_{ m best}$	$ au_{ m best}$	n^e
		(MJD)		(day)		(/day)	$C_{-1\sigma}$	$p_{-1\sigma}$	$ au_{-1\sigma}$	
							$C_{+1\sigma}$	$p_{+1\sigma}$	$ au_{+1\sigma}$	
FRB 20121102A (L21) (5)	FAST	58724.87-58776.88	39	1.76	1651	1500	5100	1.6	0.0020	0.2
							3100	1.4	0.0009	
							9700	1.8	0.0033	
FRB 20121102A (H22) (6)	Arecibo	57510.80-57666.42	18	0.733	475	870	490	9.1	0.28	0.1
							280	2.1	0.019	
							1200	∞	1.4	
FRB 20121102A (J23) (9)	Arecibo	58409.35-58450.28	8	0.272	1027	4900	770	2.3	0.012	0.4
					$(849)^{f}$		500	1.8	0.0063	
							1100	3.5	0.028	
FRB 20201124A (X22) (7)	FAST	59307.33-59360.18	45	3.13	1863	840	340	28.3	1.3	0.1
							250	4.5	0.13	
							500	∞	1.5	
FRB 20201124A (Z22 D3) (8)	FAST	59484.81-59484.86	1	0.040	232	5800	270	3.4	0.071	0.5
							83	1.5	0	
							∞	∞	1.7	
FRB 20201124A (Z22 D4) (8)	FAST	59485.78-59485.83	1	0.040	542	14000	54	4.2	0.19	0.5
							35	2.1	0.058	
							60	∞	1.9	
FRB 20220912A (Z23) (10)	FAST	59880.49-59935.39	17	0.32	1076	6900	70	5.7	0.26	0.3
							50	2.4	0.043	
							170	∞	1.8	

Example of ξ calculation

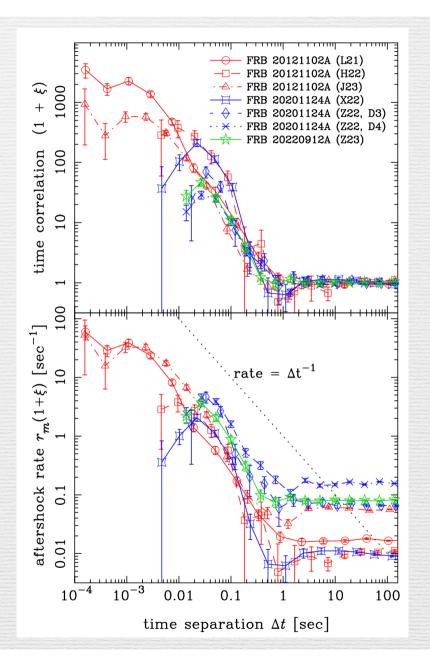


time correlation $\xi(\Delta t)$

- power-law signal at ∠t < 1 sec
- flat at ∠t ~ FRB duration(<10 msec)
 - · Note: different sub-burst treatments by different authors
- can be fit by $\xi \propto (\Delta t + \tau)^{-p}$
 - · similar to the Omori-Utsu law of aftershock rate of earthquakes
- "aftershock rate" after one event is given as $r_m (1 + \xi)$, where r_m is the mean event rate
 - expected number of aftershocks following one event:

$$n = \int r_m \, \xi(\Delta t) d(\Delta t)$$

- n = 0.1-0.5 for FRBs
- · multiple aftershocks to one event are rare
- \cdot stable against change of mean rate r_m or different sources



The Omori-Utsu law for earthquake aftershocks

- ・大森 房吉 (OMORI, Fusakichi, 1868-1923)
 - Omori law: Omori, F. "On the after-shocks of earthquakes." The journal of the College of Science, Imperial University, Japan 7, 111–200 (1894).
 - aftershock rate $\propto (\Delta t + \tau)^{-1}$
- ・宇津徳治 (UTSU, Tokuji, 1923-2004)
 - · modified Omori law (Omori-Utsu law) Utsu+ 1957, 1961
 - aftershock rate $\propto (\Delta t + \tau)^{-p}$



On the After-shocks of Earthquakes.

b

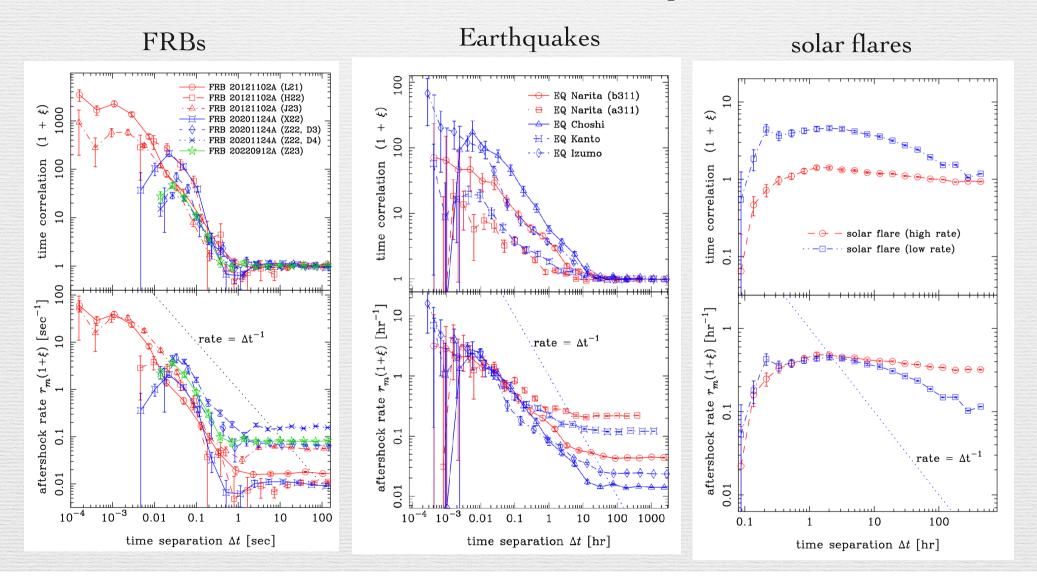
F. Omori, Rigakushi.

I. General Considerations.

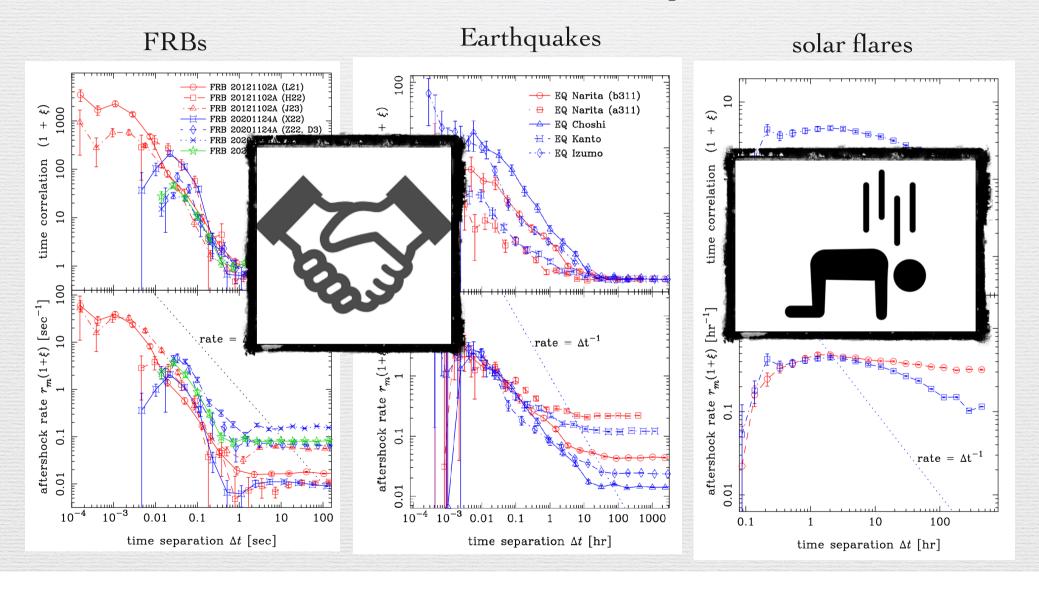
§ 1. A strong earthquake is almost invariably followed by weaker ones and when it is violent and destructive the number of minor shocks following it may amount to hundreds or even thousands. When after-shocks are not reported to have happened it is probably because they were deemed unimportant to record. Or it may be that the seat of origin of the earthquake being very deep or far out under the ocean-bed, the after-shocks did not reach the observer.

Complete records of after-shocks were obtained, I believe, for the first time in the cases of the three recent great earthquakes in Japan; namely, those of Kumamoto in 1889, of Mino and Owari in 1891,

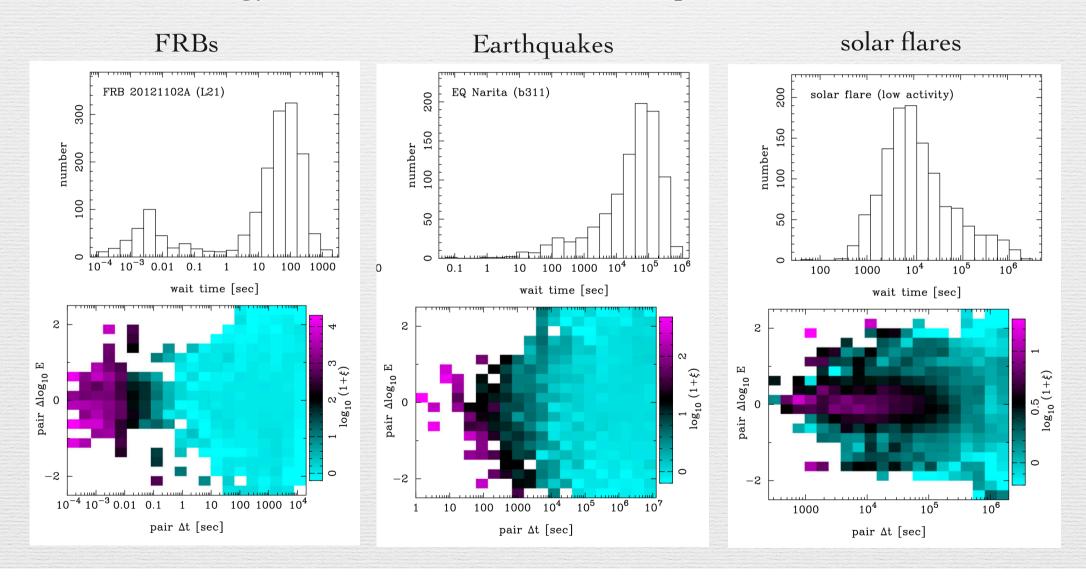
time correlation function: FRBs vs. earthquakes vs. solar flares



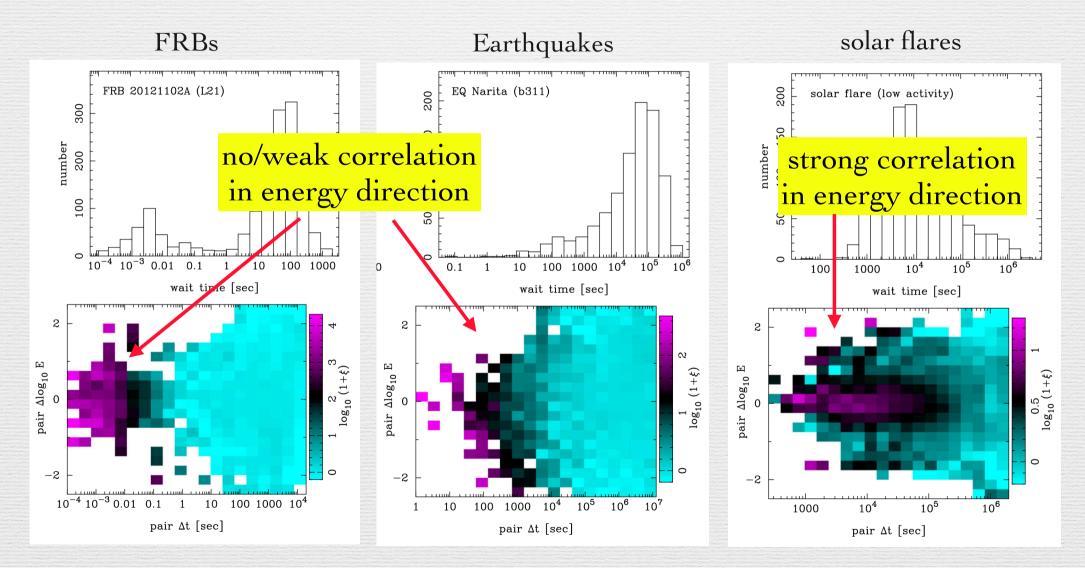
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time-energy correlation: FRBs vs. earthquakes vs. solar flares



time-energy correlation: FRBs vs. earthquakes vs. solar flares

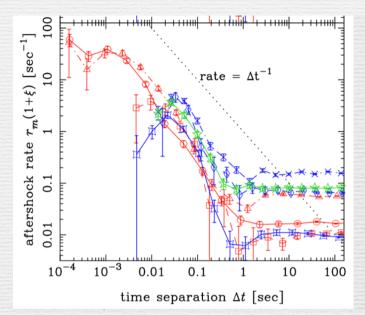


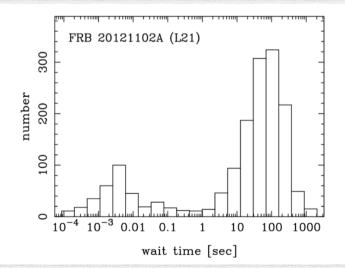
Difference between FRBs & earthquakes?

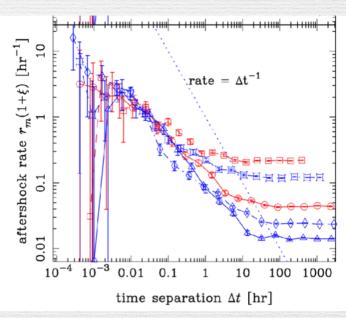
· index p of Omori-Utsu law

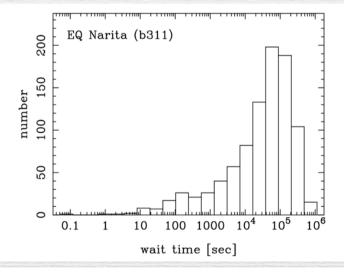
$$\cdot \propto (\Delta t + \tau)^{-p}$$

· wait-time distribution







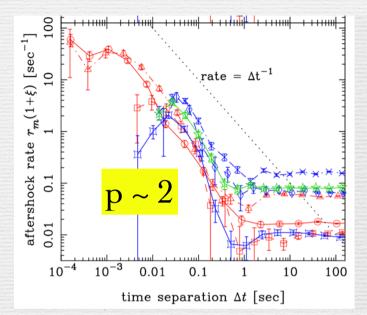


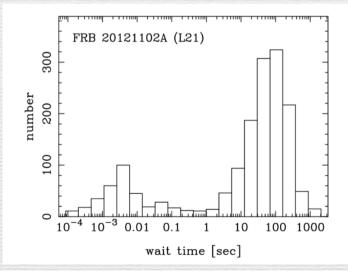
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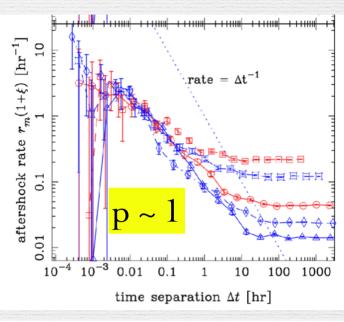
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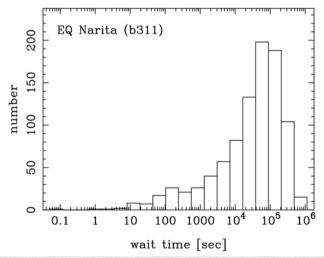
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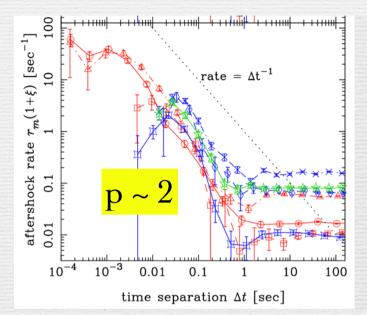


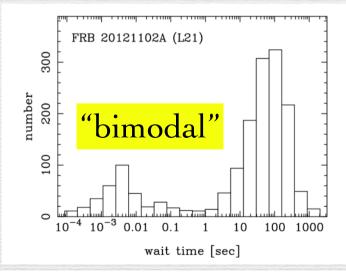
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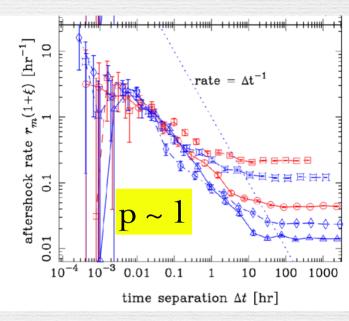
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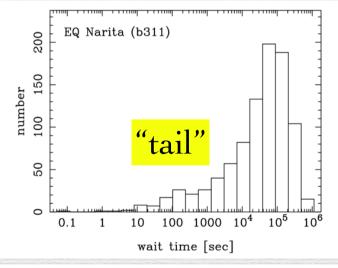
•
$$\propto (\Delta t + \tau)^{-p}$$

· wait-time distribution









Conclusions (1)

- FRBs are remarkably similar to earthquakes in time-energy correlation, with the universal laws on the aftershock statistics
 - 1. each event induces 0.1-0.5 aftershocks
 - 2. aftershock rate obeys the Omori-Utsu law $\propto (\Delta t + \tau)^{-p}$
 - 3. τ is close to the event duration (10 msec for FRBs, 1 min for earthquakes)
 - 4. even if the source activity changes, the aftershock rate remains stable
 - 5. almost no correlation between time and energy
- These features have been known for earthquakes as the ETAS (epidemic-type aftershock sequence) model
- the only difference is p ~ 2 (FRB) or ~ 1 (earthquakes)
- · In contrast, solar flares are NOT similar to FRBs/earthquakes
 - perhaps related to fluid surface of the Sun, compared with solid crusts at FRB/ Earth surfaces?

Conclusions (2)

- A natural interpretation: repeating FRBs are produced when the energy stored in solid neutron star crust is liberated by seismic activity
- Other FRB mechanism may not be excluded, but these aftershock properties must be explained in any FRB theory, putting strong constraints
- Future theoretical studies on FRB aftershock properties may give us new information about the neutron star crust / dense nuclear matter