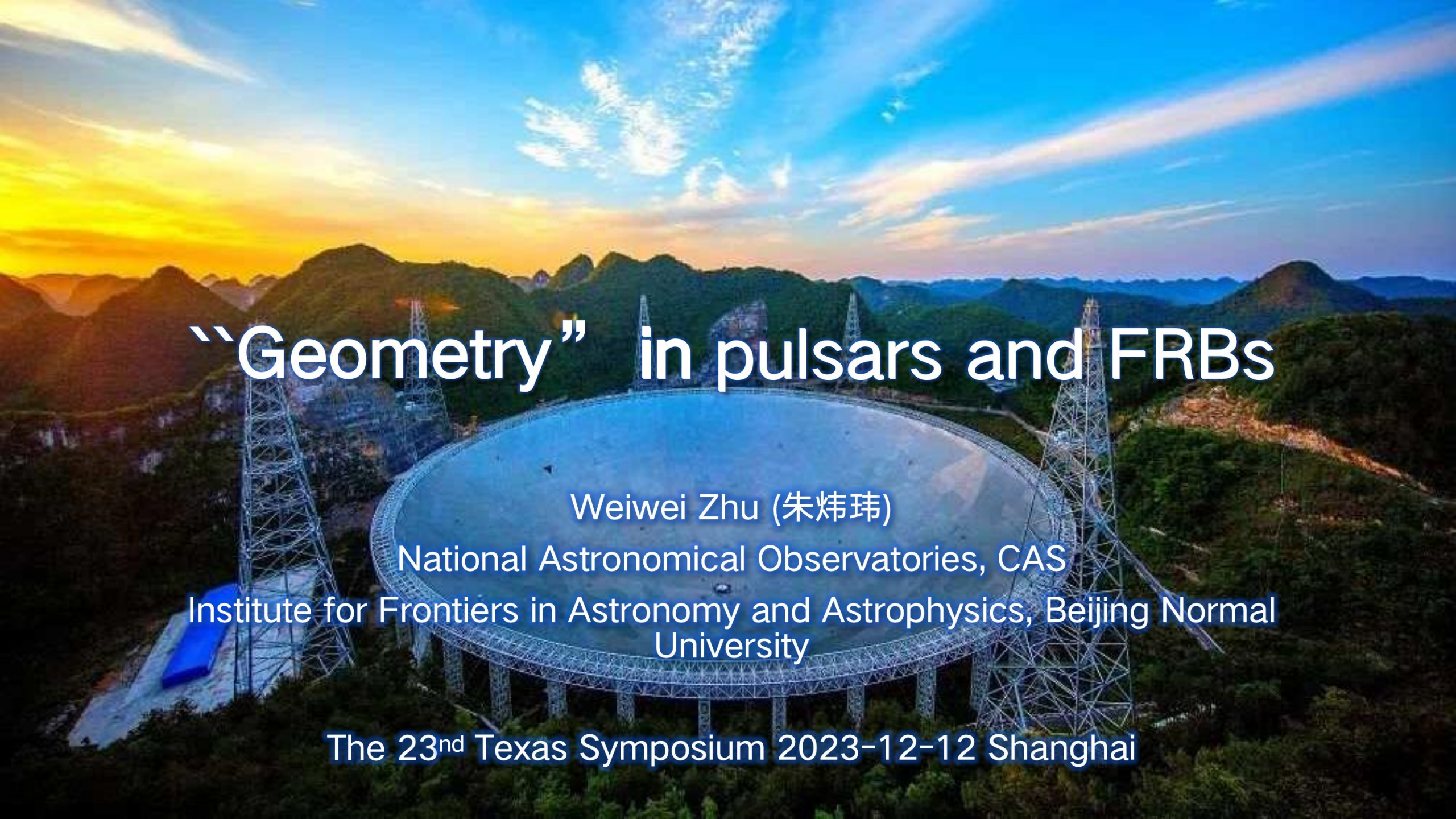


The background image shows a large, circular radio telescope dish, likely the FAST (Five-hundred-meter Aperture Spherical Telescope) in China, situated in a deep valley surrounded by lush green mountains. The sky is a vibrant mix of orange, yellow, and blue, indicating a sunset or sunrise. Several tall, lattice-like support towers are visible around the dish.

“Geometry” in pulsars and FRBs

Weiwei Zhu (朱炜玮)

National Astronomical Observatories, CAS

Institute for Frontiers in Astronomy and Astrophysics, Beijing Normal University

The 23rd Texas Symposium 2023-12-12 Shanghai

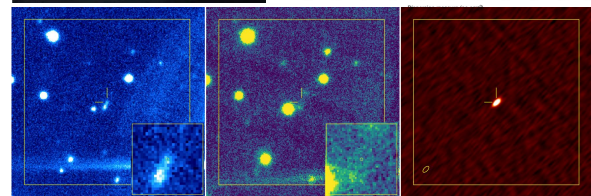
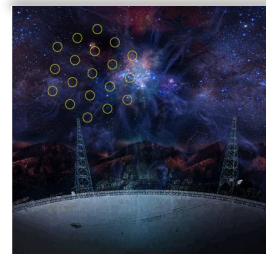


The FAST key science projects

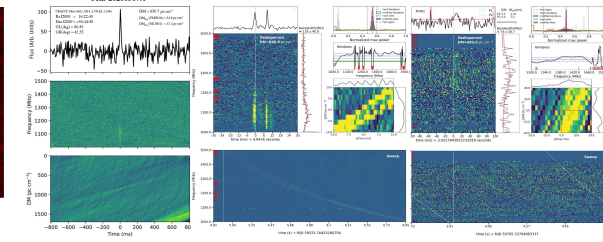
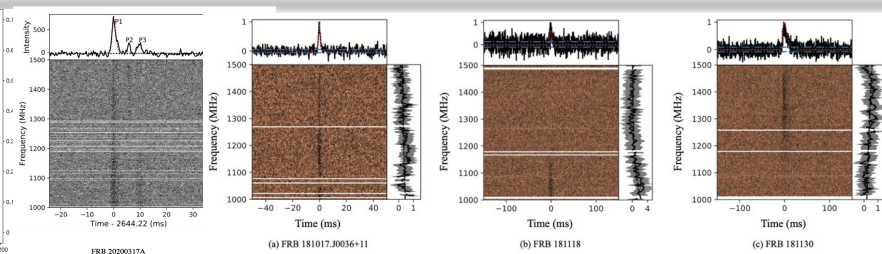
- ❑ The Commensal Radio Astronomy FAST Survey (**CRAFTS**; PI: Di Li)
- ❑ Fast Radio Burst Search and Multi-wavelength Observations (**FRB KSP**; PI: Bing Zhang, Weiwei Zhu)
- ❑ The Galactic Plane Pulsar Search (**GPPS**; PI: JinLin Han)
- ❑ M31 HI imaging and pulsar search (PI: Jie Wang)
- ❑ The Chinese Pulsar Timing Array (**CPTA**; PI: Kejia Lee)
- ❑ Pulsar Physics and Evolution (PI: Nina Wang)



CRAFTS pulsar discoveries



FRB 20190520B



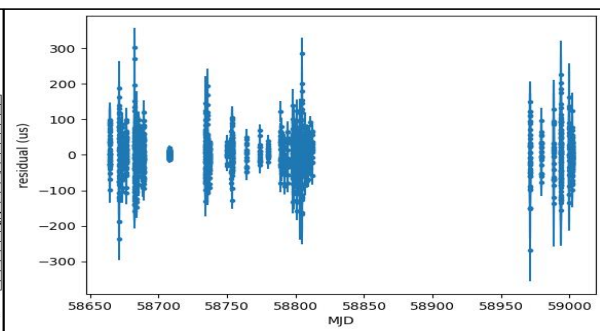
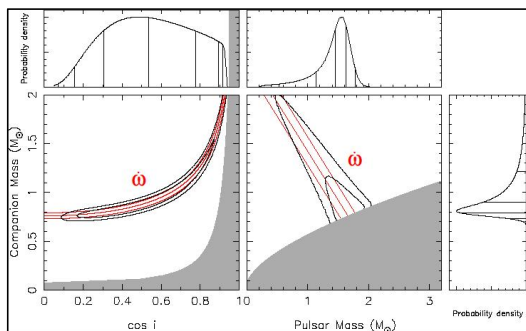
The CRAFTS survey have discovered:
180 pulsars

incl. 40 MSPs

20 binaries (3 DNS)

30 RRATs

8 FRBs (*Zhu et al. 2020; Niu et al. 2021; Niu et al. 2022 Nature; more in prep.*)



3 double neutron star systems

Wu et al. 2023

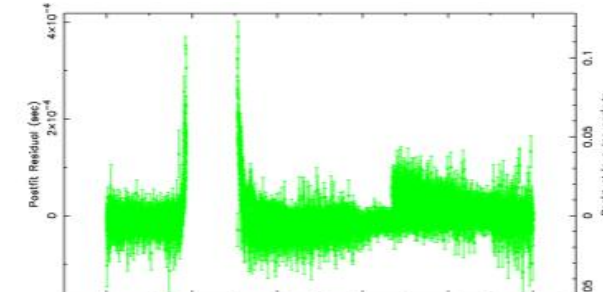
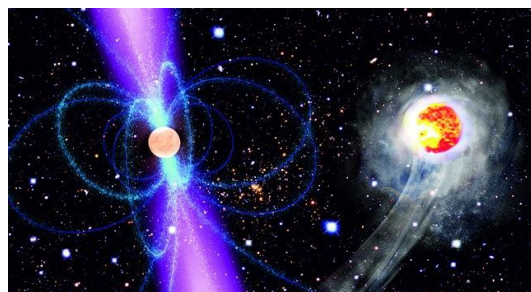
Zhao et al. 2023 in prep.

Miao et al. 2023 in prep.

MSP-black widow system with
interesting eclipses

Wang et al. MNRAS 2022

Miao et al. RAA 2023





The FAST FRB Key Science Project



FAST 快速射电暴优先及重大项目

FAST Fast Radio Burst (FRB) Key Science Project

- ~100 participants from 10s institutes
- Observers and theoreticians
- Multi wavelength & observatories



FAST



云南40米



Fermi



新疆26米



Swift

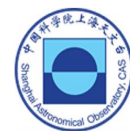


HMXT



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NATIONAL ASTRONOMICAL OBSERVATORIES, CAS



中国科学院上海天文台

Shanghai Astronomical Observatory, Chinese Academy of Sciences



中国科学院新疆天文台

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中国科学院紫金山天文台



中国科学院高能物理研究所

Institute of High Energy Physics Chinese Academy of Sciences



中国科学院国家授时中心

National Time Service Center, Chinese Academy of Sciences



南京大学

NANJING UNIVERSITY

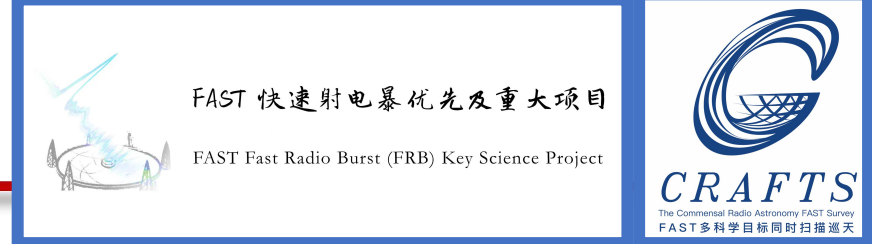


云南大学

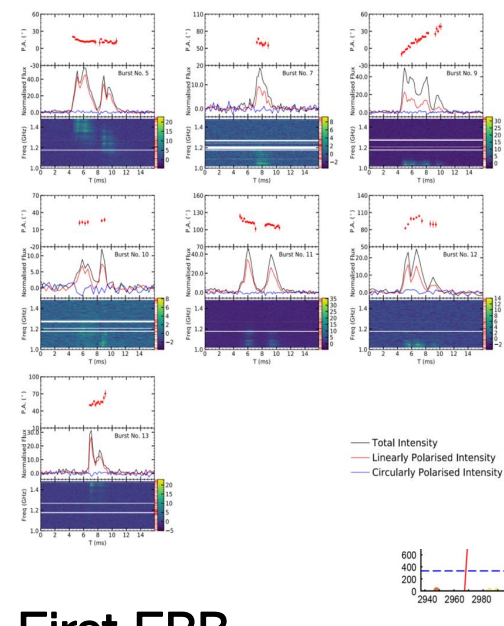




The FAST FRB Highlights

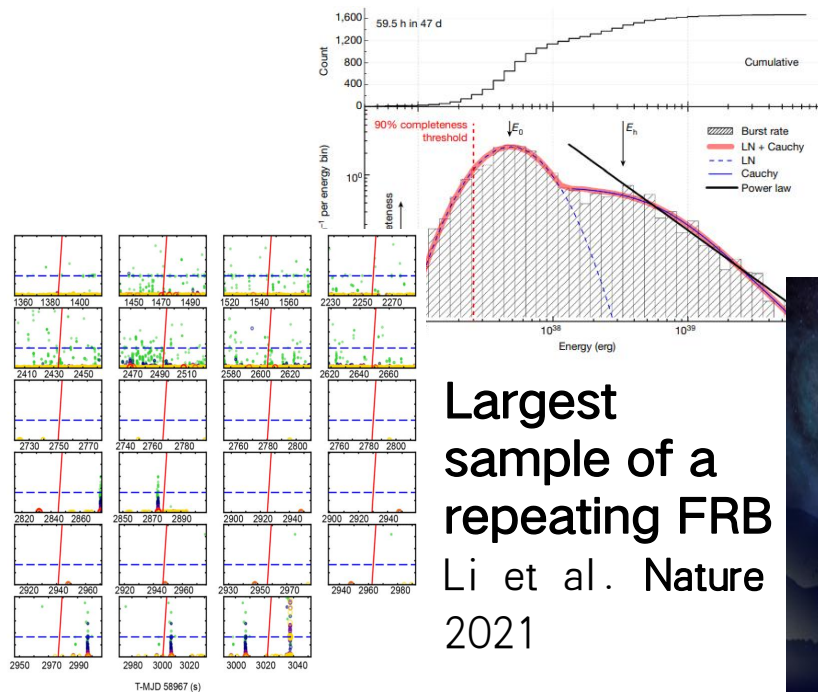


- 5 Nature, 1 Science since 2020

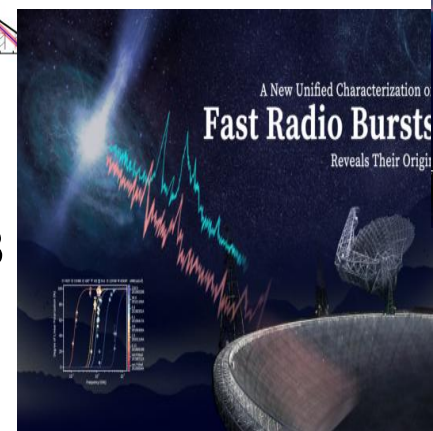


First FRB polarization swings
Luo et al. 2020 Nature

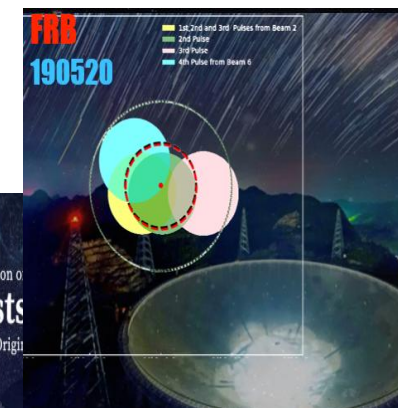
Stringent radio limit
On the FRB-magnetar
Lin et al. 2020 Nature



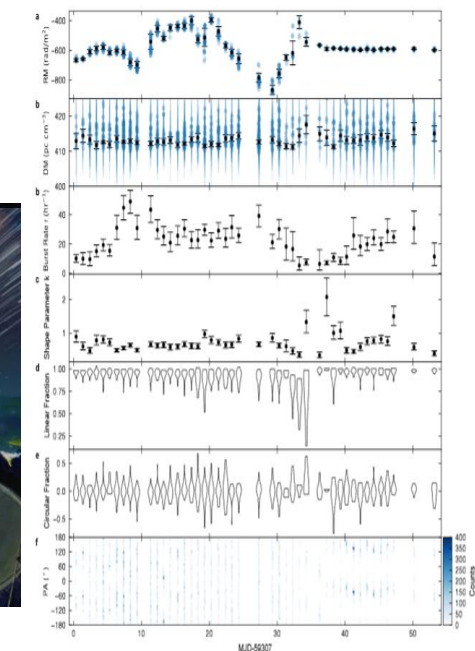
Largest sample of a repeating FRB
Li et al. Nature 2021



Key signature of FRB polarization
Feng et al. Science 2022



FRB190520, FAST 1st repeating FRB, special birth place and persistent radio source (PRS)
Niu et al. Nature 2022



Probing the complex magnetic environment of an repeating FRB
Xu Heng et al. Nature 2022

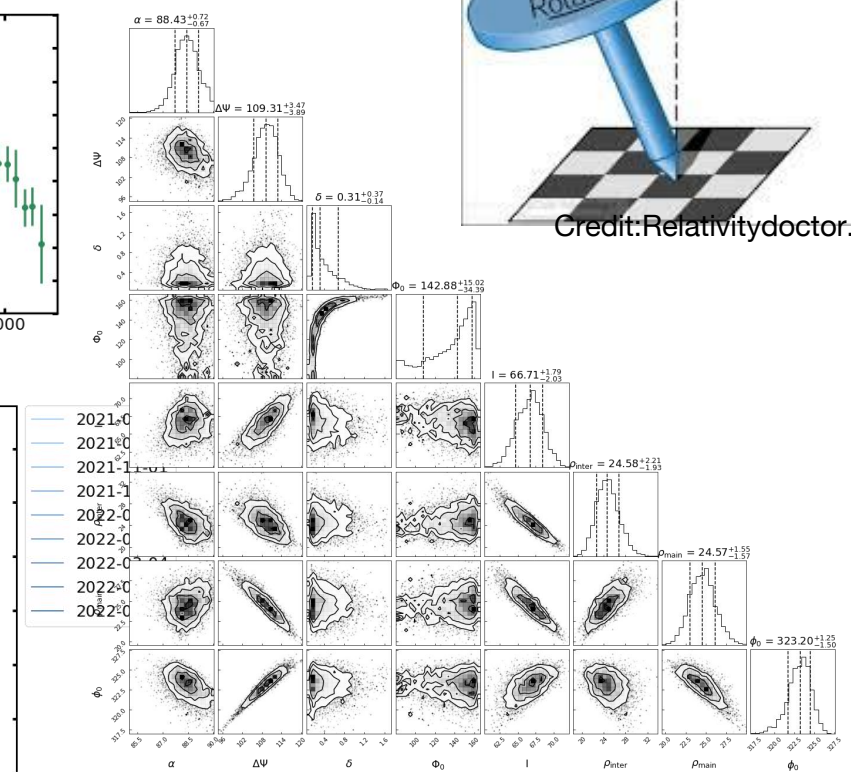
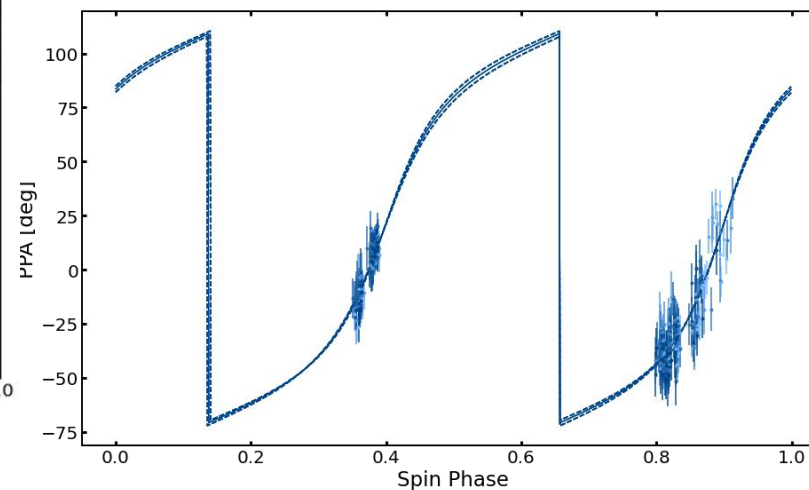
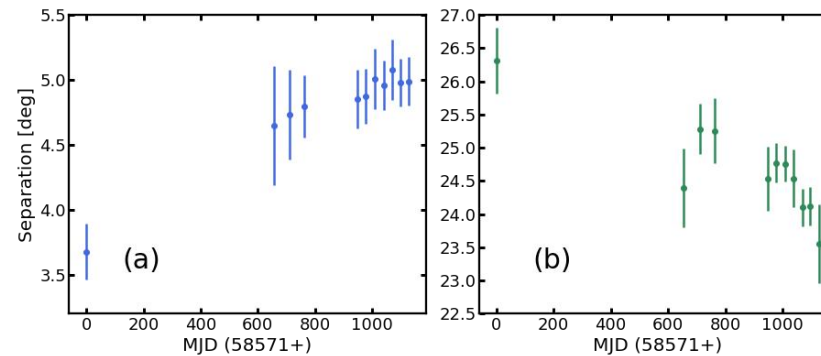
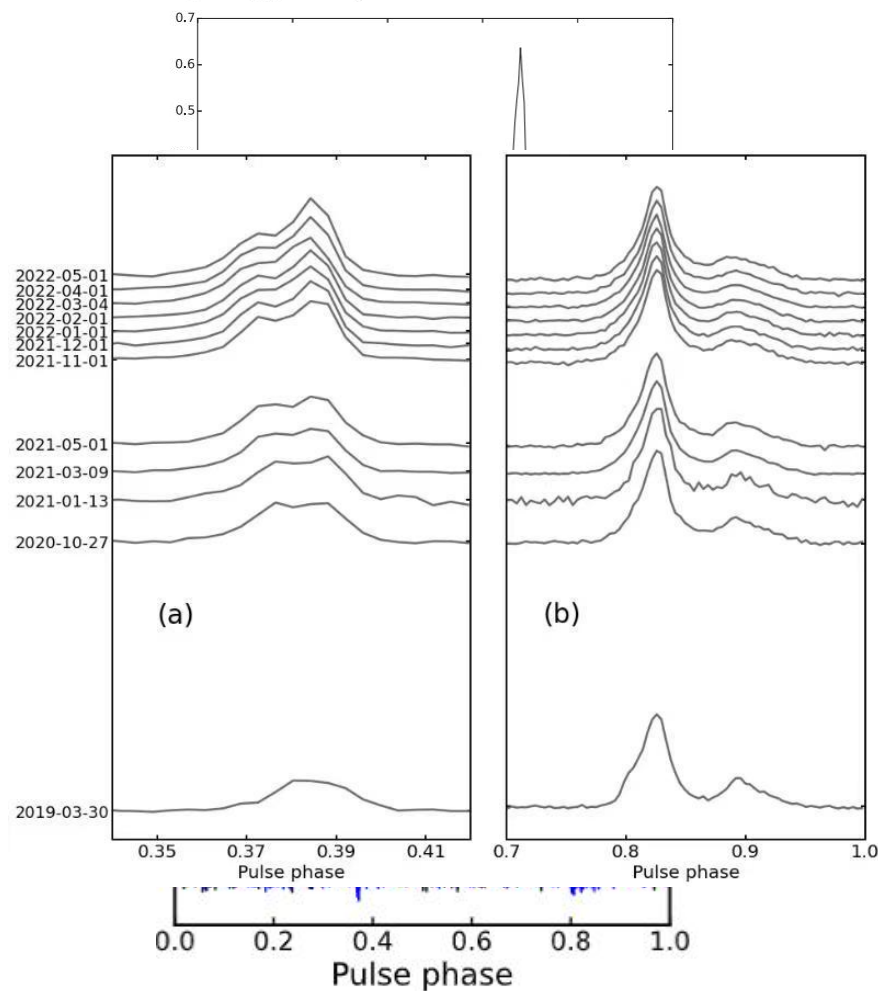
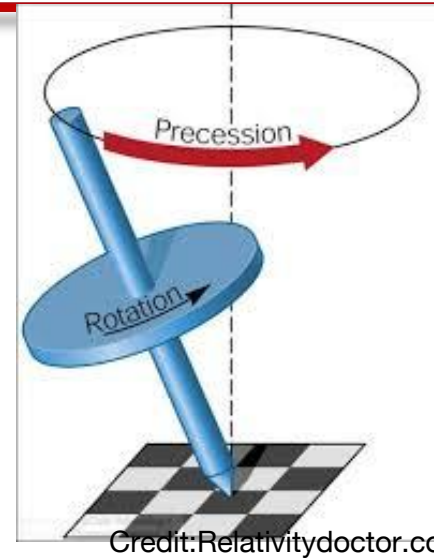
□ How do we get to the ``Geometry'' ?

1. Profiles (Polarizations)
2. Scintillation (arcs)
3. Spin (phases)

Relativistic spin precession in DNS

□ Profiles (Polarizations)

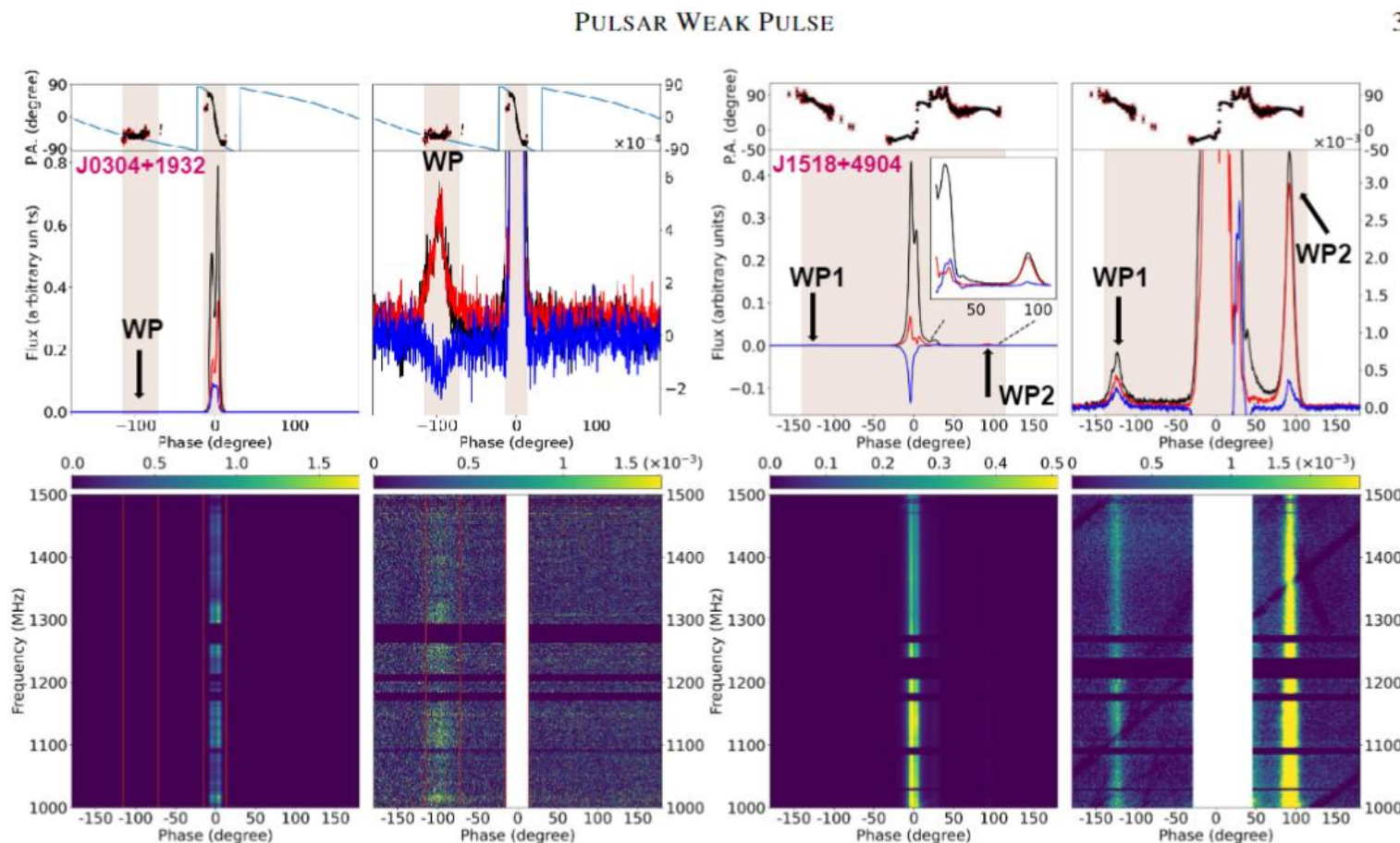
PSR J1946+2052
Closest DNS Pb=1.8 hr



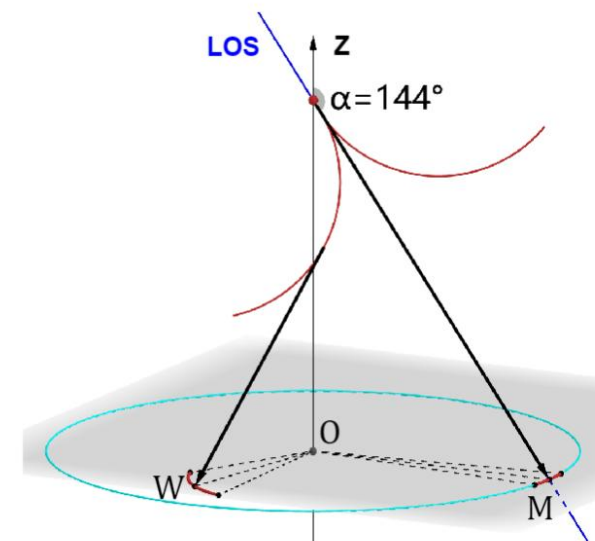
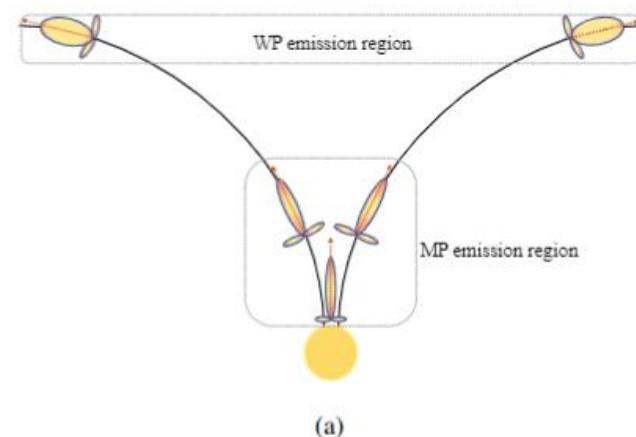
New weak pulsar emission components

□ Profiles (Polarizations)

□ Emissions from high magnetosphere altitudes



3



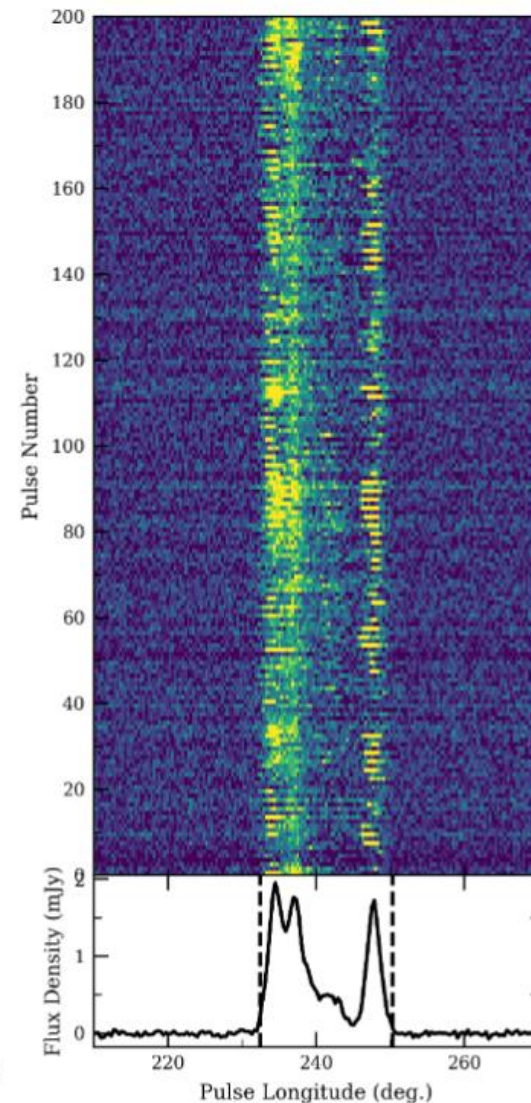
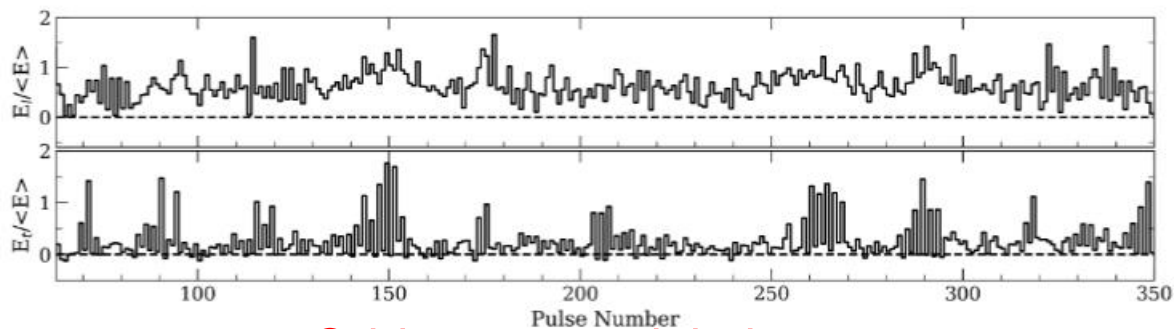
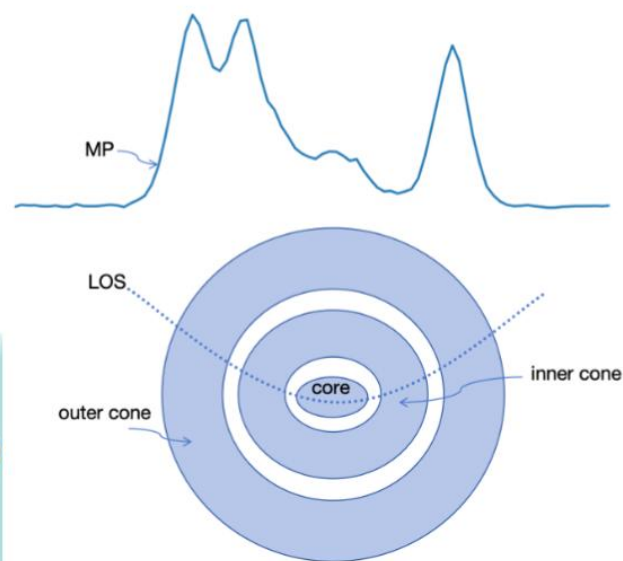
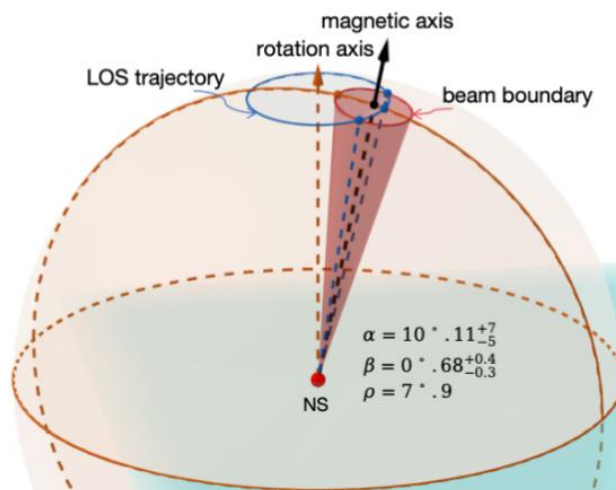
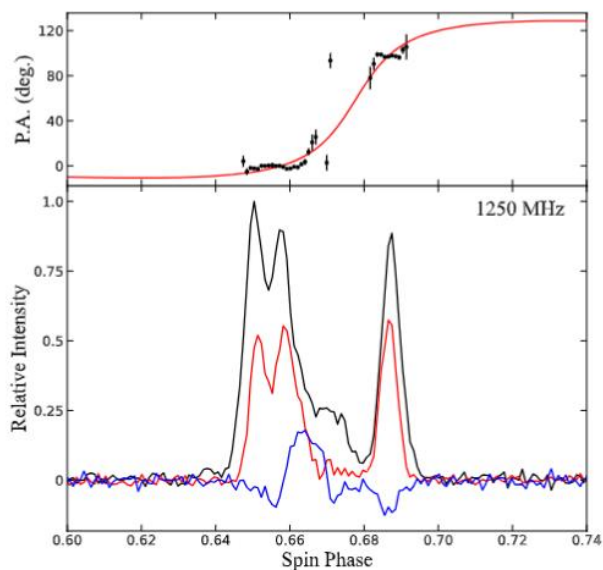
Yuan, Zhu, Kramer et al. ApJ 2023

A four components Pulsar

Profiles (Polarizations)

PSR J1239+0326

Double cone beam pattern



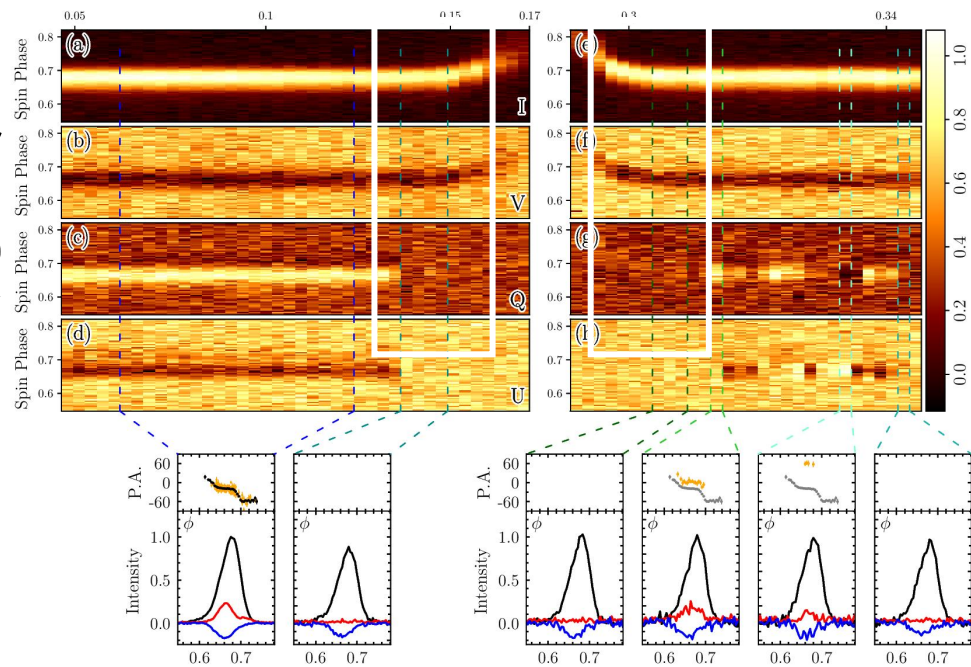
Rejep et al. in prep.

Odd-even modulation

Black Widow Pulsar Wind

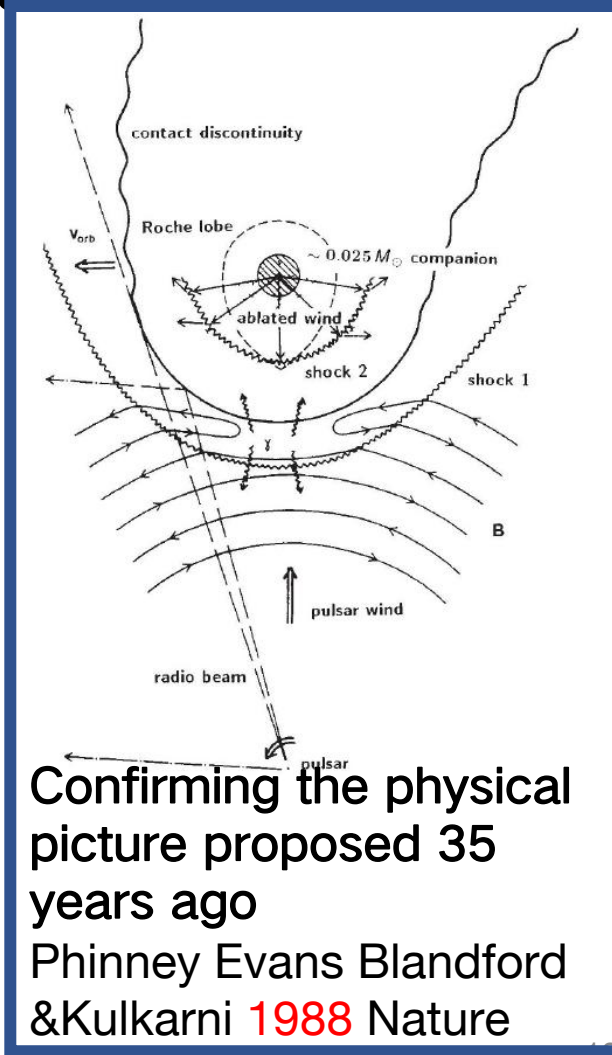
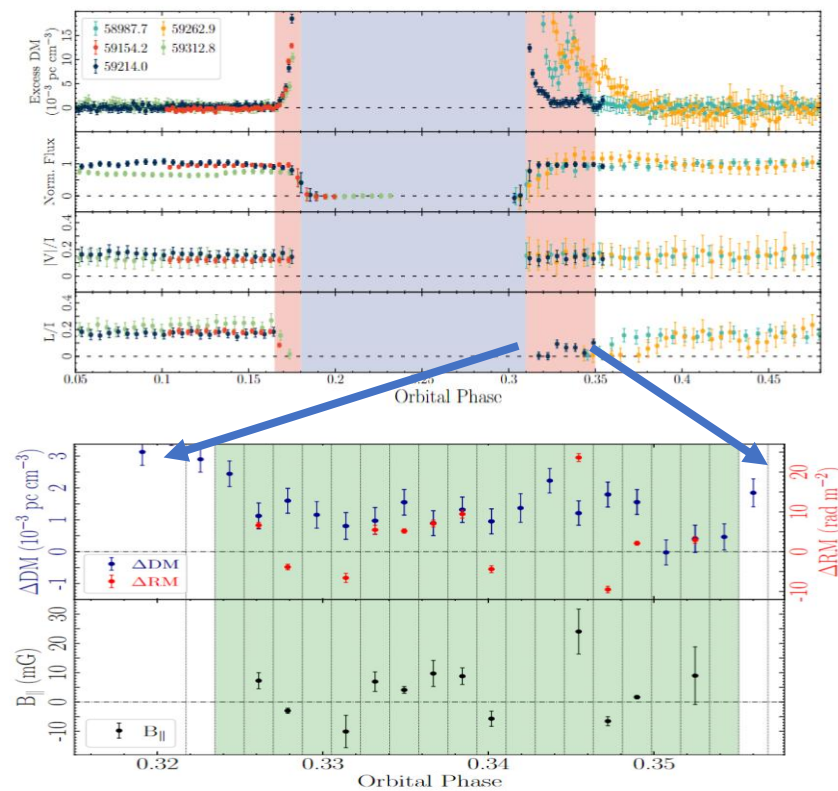
□ Profiles (Polarizations)

□ First evidence for Reciprocating Magnetic Fields in the Pulsar Wind from black-widow binary

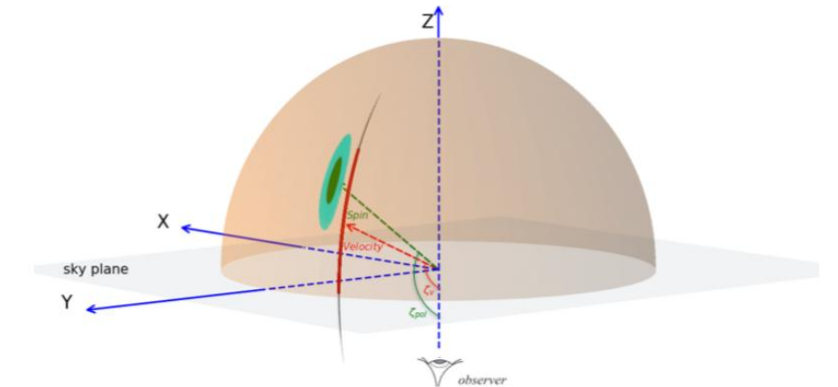
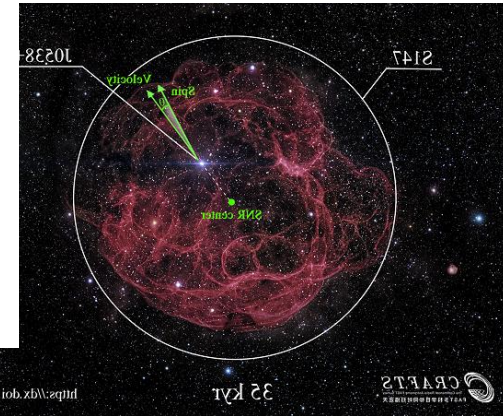
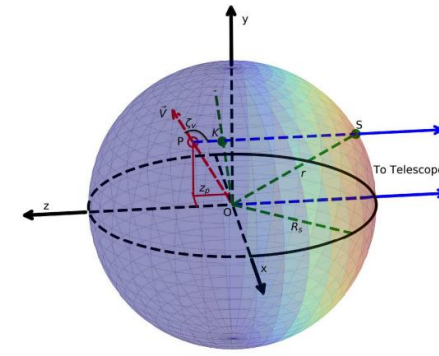
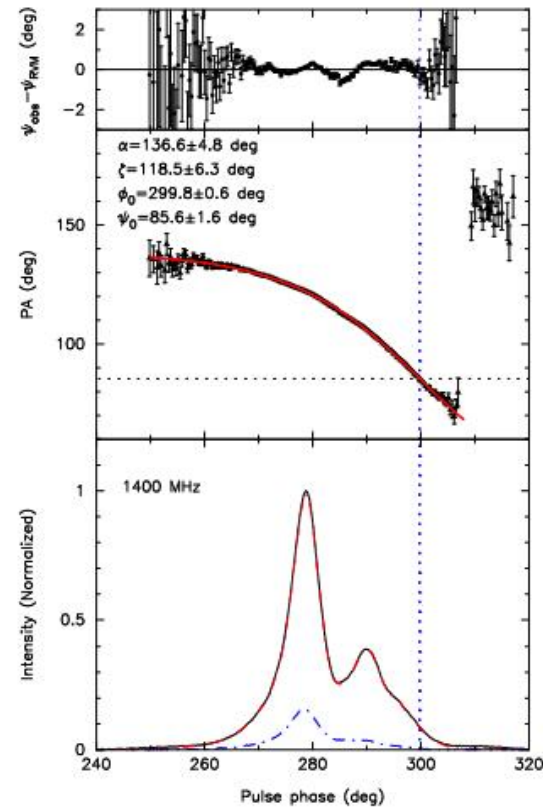
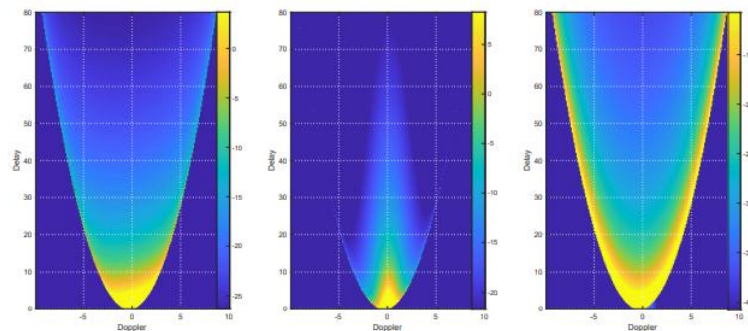
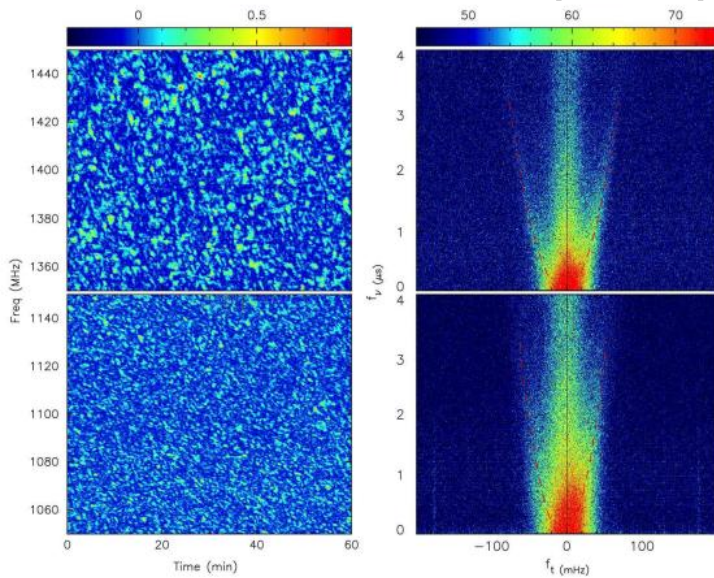


Miao, Blackmon, Zhu* et al. RAA 2023

In collaboration with authors from MPIfR, WVU, CSIRO etc.



□ Scintillation (arcs)



High altitude ISM structures

□ Scintillation (arcs)

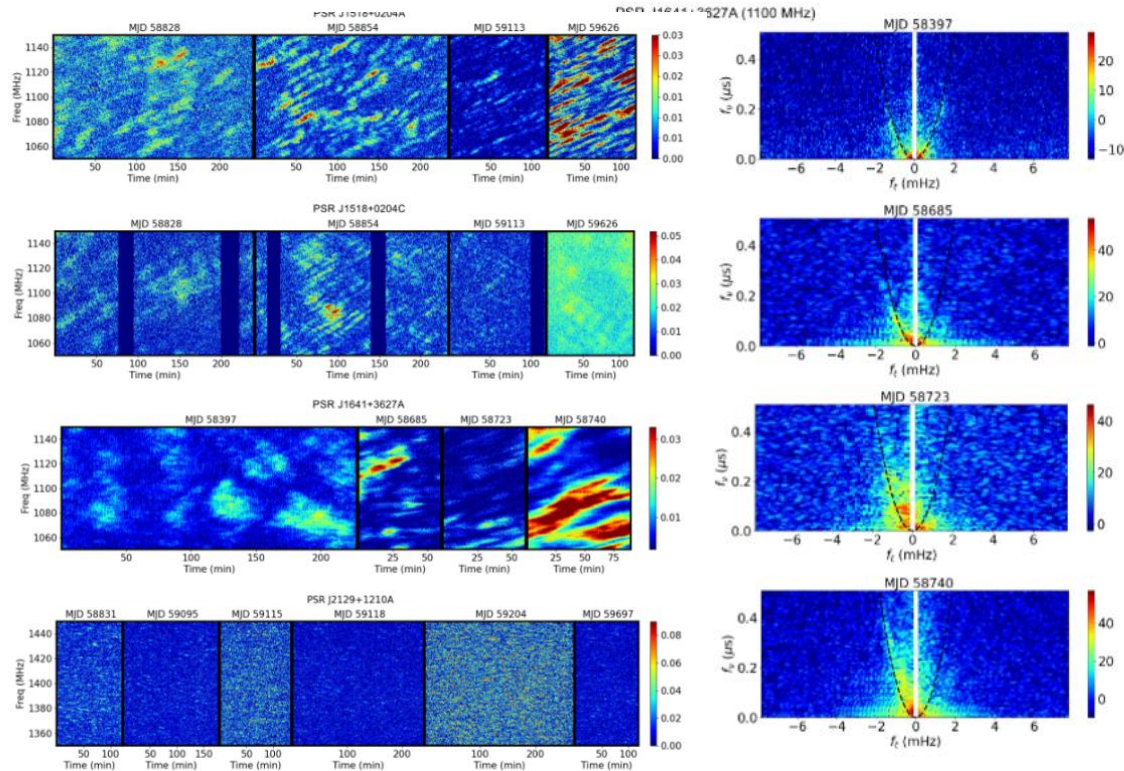
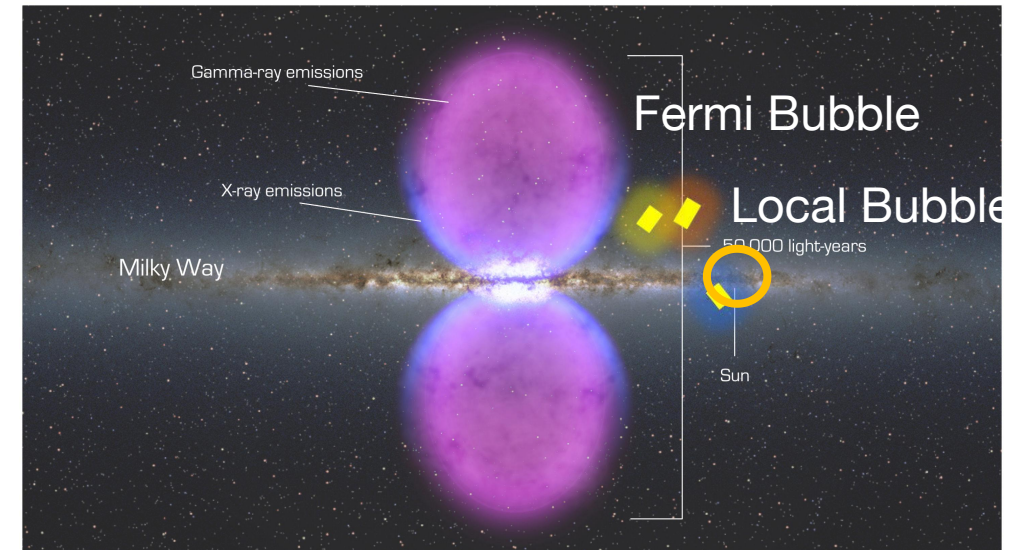
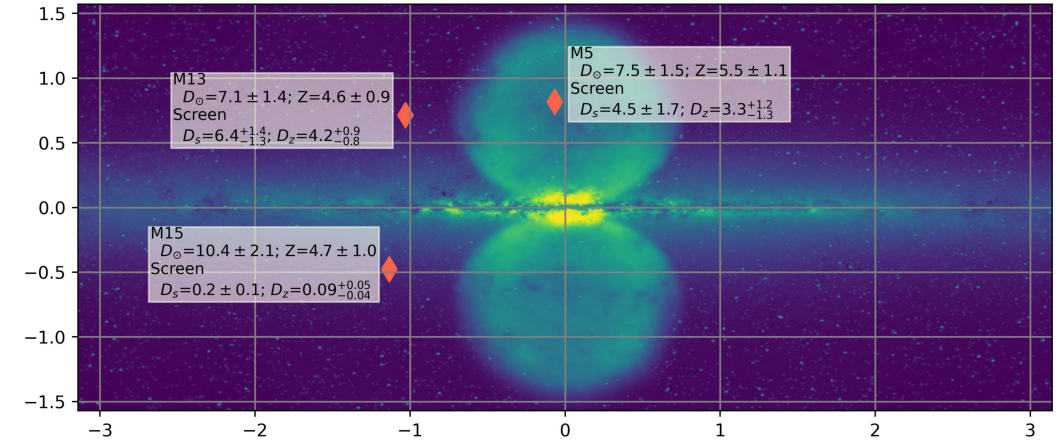
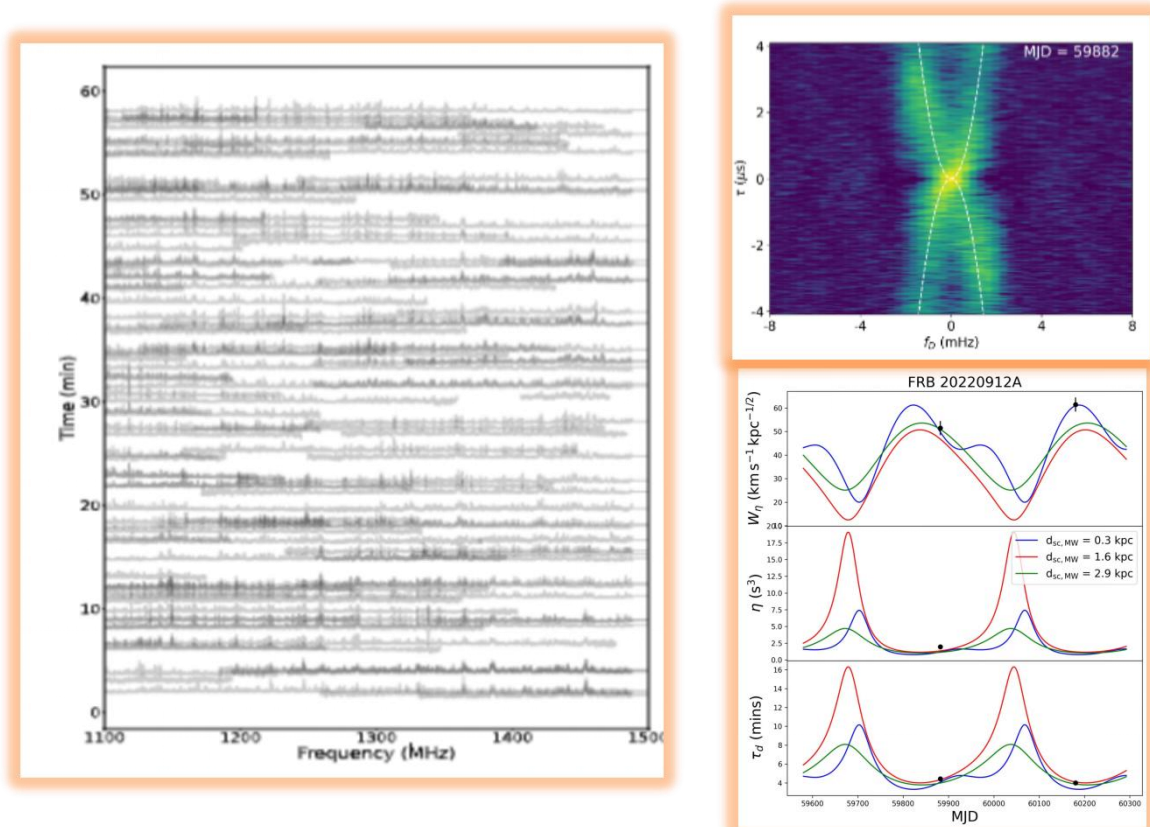


Figure A11 As in Figure A6, results of arc curvature fittings and secondary spectra for PSR J1641+3627A at 1100 MHz.



□ Scintillation (arcs)

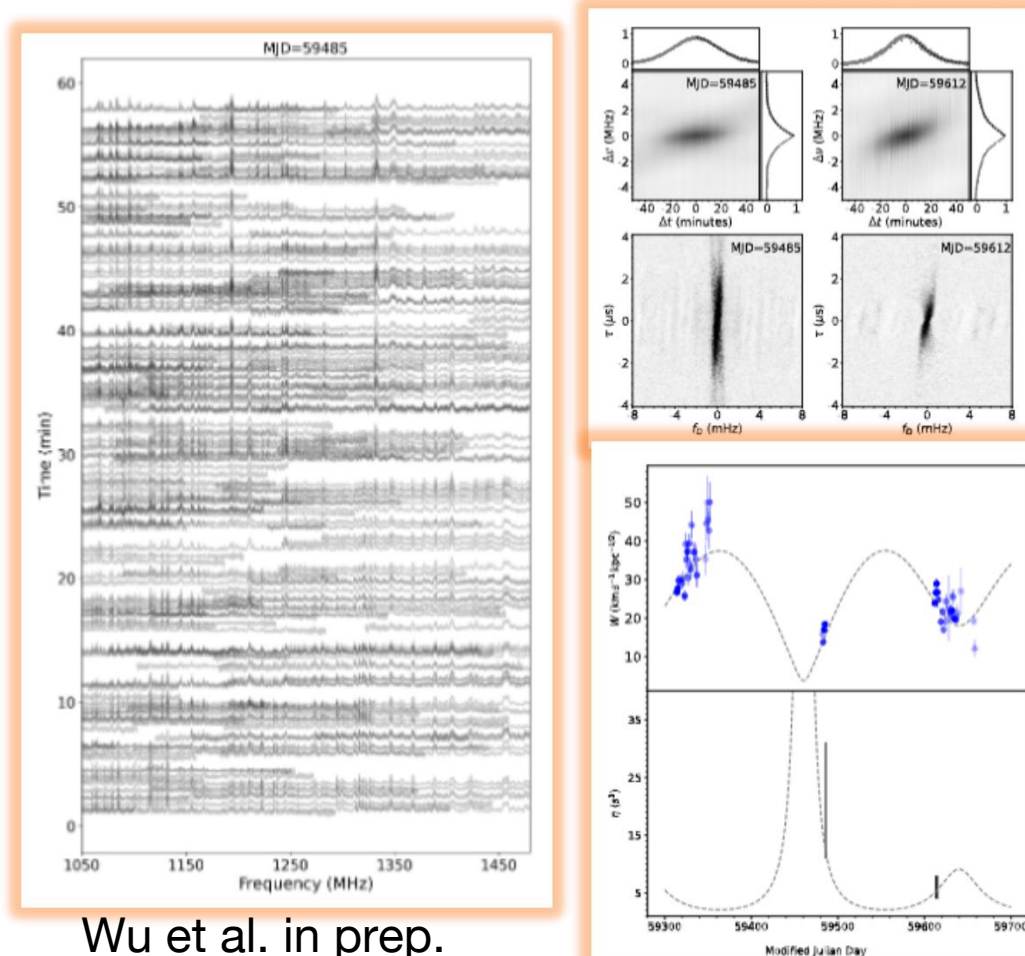
■ Scintillation arc and time scales of FRB20220912A



First detection of scintillation arc from FRBs

Wu, Main, Zhu et al. SCPMA in

■ Scintillation arc and time scales of FRB20221124A



Wu et al. in prep.

□ Spin (phases)

Research in Astronomy and Astrophysics

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* Volume 22 Number 12 December 2022 *

FAST Observations of an Extremely Active Episode of FRB 20201124A: I. Burst Morphology
D. J. Zhou, J. L. Han, B. Zhang, K. J. Lee, W. W. Zhu, D. Li, W. C. Jing, W. Y. Wang, Y. K. Zhang, J. C. Jiang et al.
2022, RAA, 22, 124001
[Abstract](#) | [PDF](#)

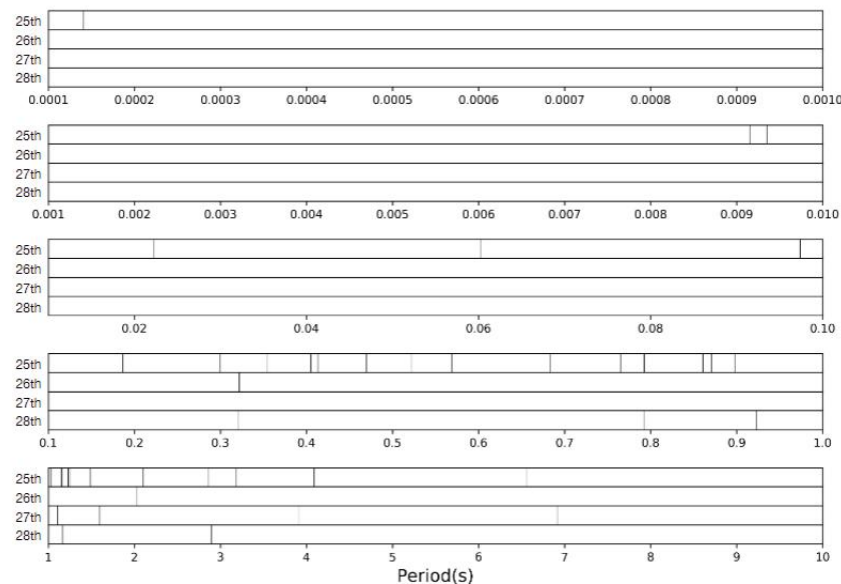
FAST Observations of an Extremely Active Episode of FRB 20201124A: II. Energy Distribution
Yang-Kun Zhang, Pei Wang, Yi Feng, Bing Zhang, Di Li, Chao-Wei Tsai, Chen-Hui Niu, Rui Luo, Ju-Mei Yao, Wei-Wei Zhu et al.
2022, RAA, 22, 124002
[Abstract](#) | [PDF](#)

FAST Observations of an Extremely Active Episode of FRB 20201124A: III. Polarimetry
Jin-Chen Jiang, Wei-Yang Wang, Heng Xu, Jiang-Wei Xu, Chun-Feng Zhang, Bo-Jun Wang, De-Jiang Zhou, Yong-Kun Zhang, Jia-Rui Niu, Ke-Jia Lee et al.
2022, RAA, 22, 124003
[Abstract](#) | [PDF](#)

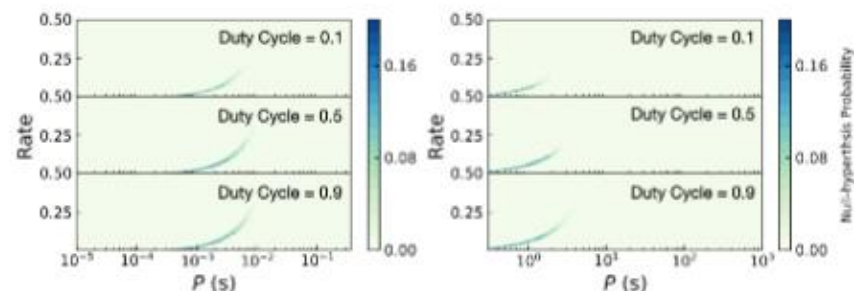
FAST Observations of an Extremely Active Episode of FRB 20201124A: IV. Spin Period Search
Jia-Rui Niu, Wei-Zhi Zhu, Bing Zhang, Mao Yuan, De-Jiang Zhou, Yong-Kun Zhang, Jin-Chen Jiang, J. L. Han, Di Li, Ke-Jia Lee et al.
2022, RAA, 22, 124004
[Abstract](#) | [PDF](#)

News and Views for Highlights

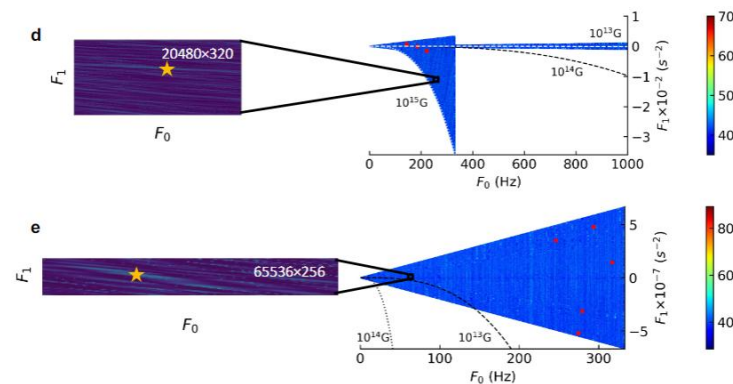
A hard stare at a bursting radio source with the world's largest telescope
Duncan Lorimer
2022, RAA, 22, No.12, N1
[Abstract](#) | [PDF](#)



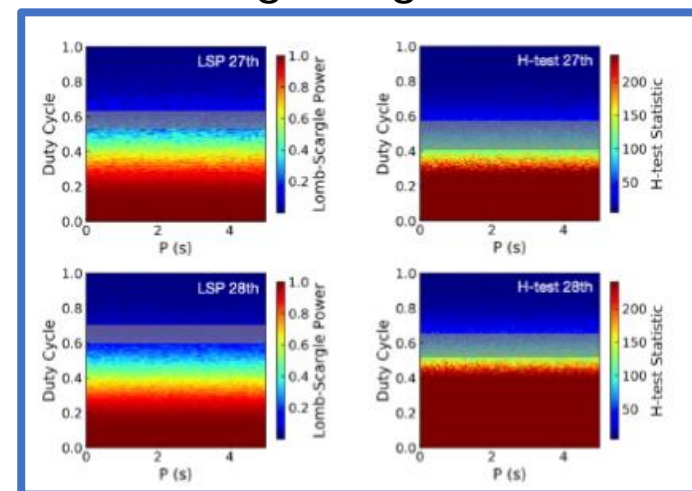
Compare Lomb-Scargle Periodogram from 4 days



Ranges of possible period



P-Pdot search for a large range
Cover B up to 10^{15} G
And a large range of binaries



Niu J. R. et al. 2022 RAA

FRB fine structures



□ Spin (phases)

- FRB20201124A most active episode
- Highest burst rate ~400 bursts/hr
- Burst fine structure study
- Many period-like fine structures $< 4\sigma$

Niu J. R. et al. 2022 RAA

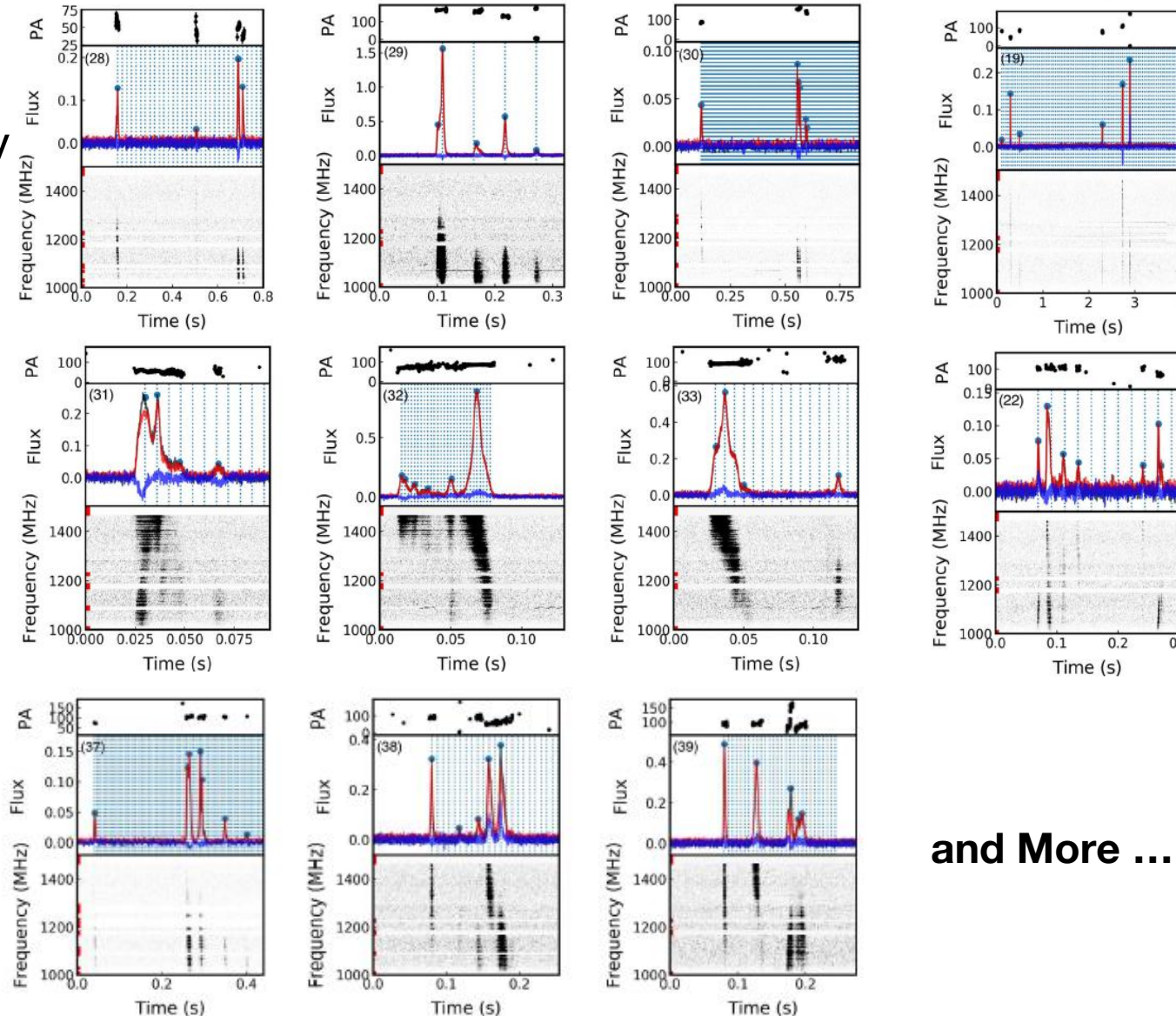


Table 6
Multi-components of FRB 20201124A

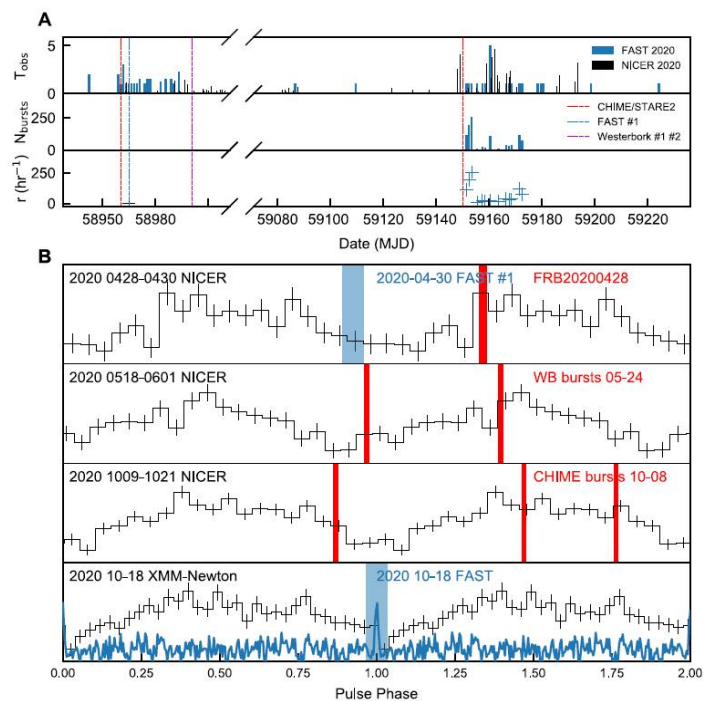
Index	N ^a	P (ms)	R	MJD
1	5	7.164	0.0240	59 484.816 06
2	7	2.638	0.0411	59 484.816 50
3	5	2.471	0.0685	59 484.821 73
4	4	11.996	0.0368	59 484.822 57
5	6	5.587	0.0415	59 484.823 57
6	5	3.801	0.0347	59 484.826 33
7	6	3.504	0.0716	59 484.830 85
8	4	5.637	0.0077	59 484.834 19
9	5	5.182	0.0337	59 484.840 94
10	8	13.521	0.0272	59 484.843 49
11	5	6.409	0.0308	59 484.847 19
12	4	8.077	0.0609	59 484.849 38
13	4	12.532	0.0532	59 484.851 48
14	4	9.736	0.0748	59 485.783 25
15	5	2.766	0.0198	59 485.783 38
16	5	4.013	0.0761	59 485.784 88
17	8	14.655	0.0238	59 485.785 61
18	4	9.308	0.0381	59 485.786 53
19	6	48.734	0.0058	59 485.789 12
20	5	4.869	0.0436	59 485.789 33
21	4	6.446	0.0931	59 485.790 68
22	7	21.804	0.0355	59 485.793 74
23	5	9.623	0.0456	59 485.794 32
24	5	9.678	0.0552	59 485.795 15
25	10	1.465	0.0471	59 485.796 12
26	5	4.255	0.0748	59 485.799 22
27	5	5.115	0.0209	59 485.799 49
28	4	20.35	0.0087	59 485.799 96
29	4	54.227	0.0292	59 485.800 02
30	6	4.915	0.0120	59 485.800 20
31	4	5.999	0.0629	59 485.801 31
32	6	2.311	0.0463	59 485.801 33
33	4	6.764	0.0444	59 485.801 61
34	4	21.211	0.0045	59 485.801 71
35	4	39.408	0.0172	59 485.803 09
36	5	12.762	0.0208	59 485.803 23
37	7	5.069	0.0078	59 485.803 79
38	5	7.867	0.0674	59 485.805 21
39	5	6.075	0.0679	59 485.807 17
40	6	21.629	0.0080	59 485.807 29
41	5	3.457	0.0429	59 485.807 64
42	4	4.967	0.0888	59 485.808 08
43	5	6.835	0.0148	59 485.809 27
44	4	8.007	0.0439	59 485.811 31
45	4	6.675	0.0213	59 485.811 93
46	4	4.045	0.0574	59 485.812 67
47	6	7.354	0.0171	59 485.812 70
48	4	3.547	0.0763	59 485.812 70
49	6	5.155	0.0141	59 485.817 02
50	5	33.551	0.0005	59 485.817 85
51	4	4.956	0.0169	59 485.818 90
52	7	5.918	0.0061	59 485.820 70
53	4	4.677	0.0258	59 485.821 38

Note.
^a Number of pulses.

and More ...

FAST detects pulsar emission from SGR J1935+2154, the FRB-magnetar

□ Spin (phases)



Radio-X-ray Phase

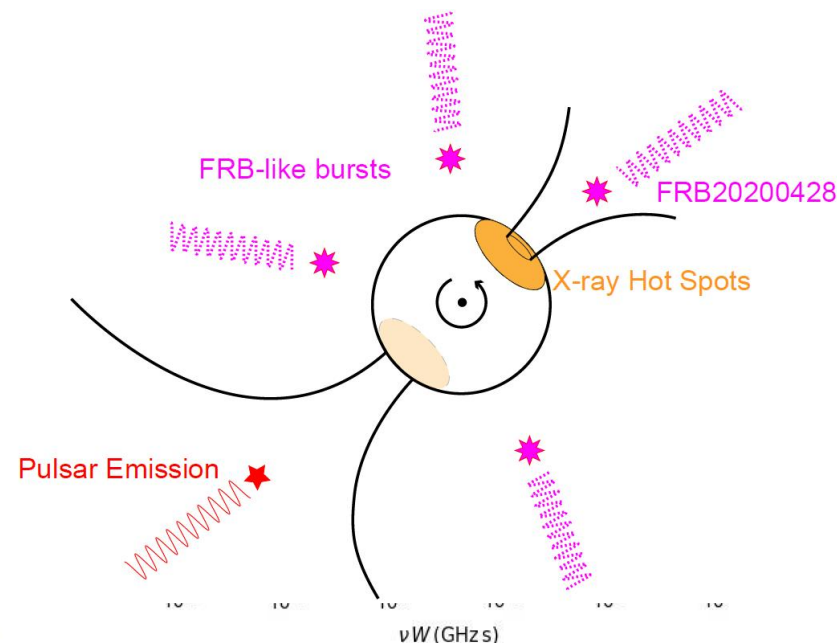
Table 1. Timing model parameters.

Parameter	Value
<i>Measured Parameters</i>	
Spin Frequency ν (s^{-1})	0.307899190(3)
Spin down rate $\dot{\nu}$ (s^{-2})	$-3.524(3) \times 10^{-12}$
<i>Fixed Parameters</i>	
Dispersion Measure (pc cm^{-3})	332.703 ^a
Reference epoch (MJD)	59131.377113
Right Ascension, α (J2000)	19:34:55.5978 ^b
Declination, δ (J2000)	21:53:47.7864 ^b
Solar System Ephemeris	DE436
<i>Derived Parameters</i>	
Dipole magnetic field, B (G)	3.52×10^{14}
Characteristic age, τ_c (yr)	1.38×10^3

^afixed to the weighted average DM value of single pulses; see method session for details.

^bposition taken from X-ray imaging [\[25\]](#) with an uncertainty of 0.4".

Timing results



Energy distribution
FRB bursts and pulsar emission

~800 pulses detected by FAST in Oct 2020

Zhu W.W.* Xu H. Zhou D. J. Zhang B.* Lee K. J.* et al. Science Advances 2023



More to come

Photo credit: Zhang C. F