

# Dark Matter Annihilation via Breit-Wigner Enhancement with Heavier Mediator

Y.C, Shao-Feng Ge, Jie Sheng and Tsutomu T. Yanagida, [arXiv: 2309.12043 [hep-ph]].

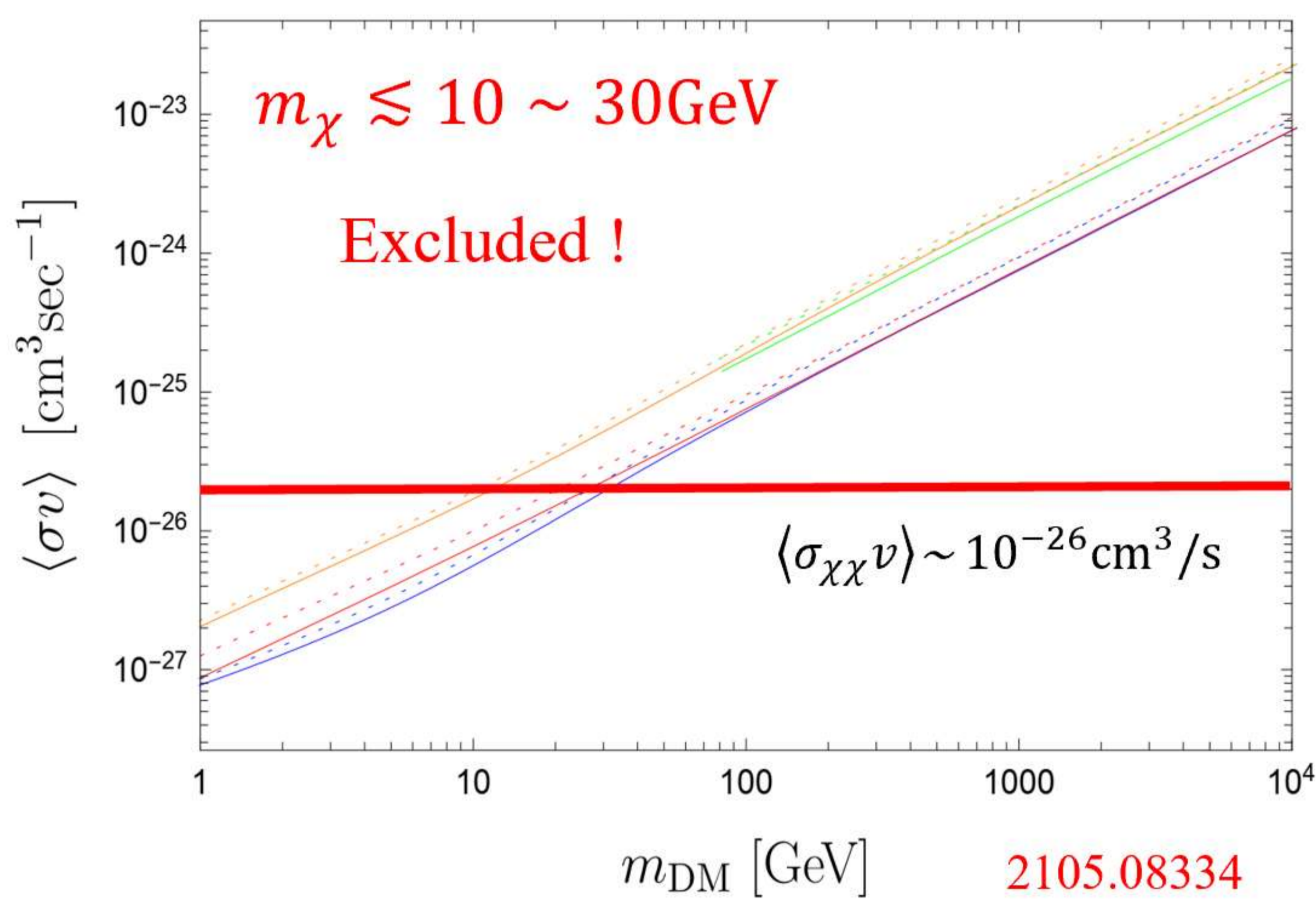


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**Research Interests:**

Dark Matter Physics, BSM Phenomenology.

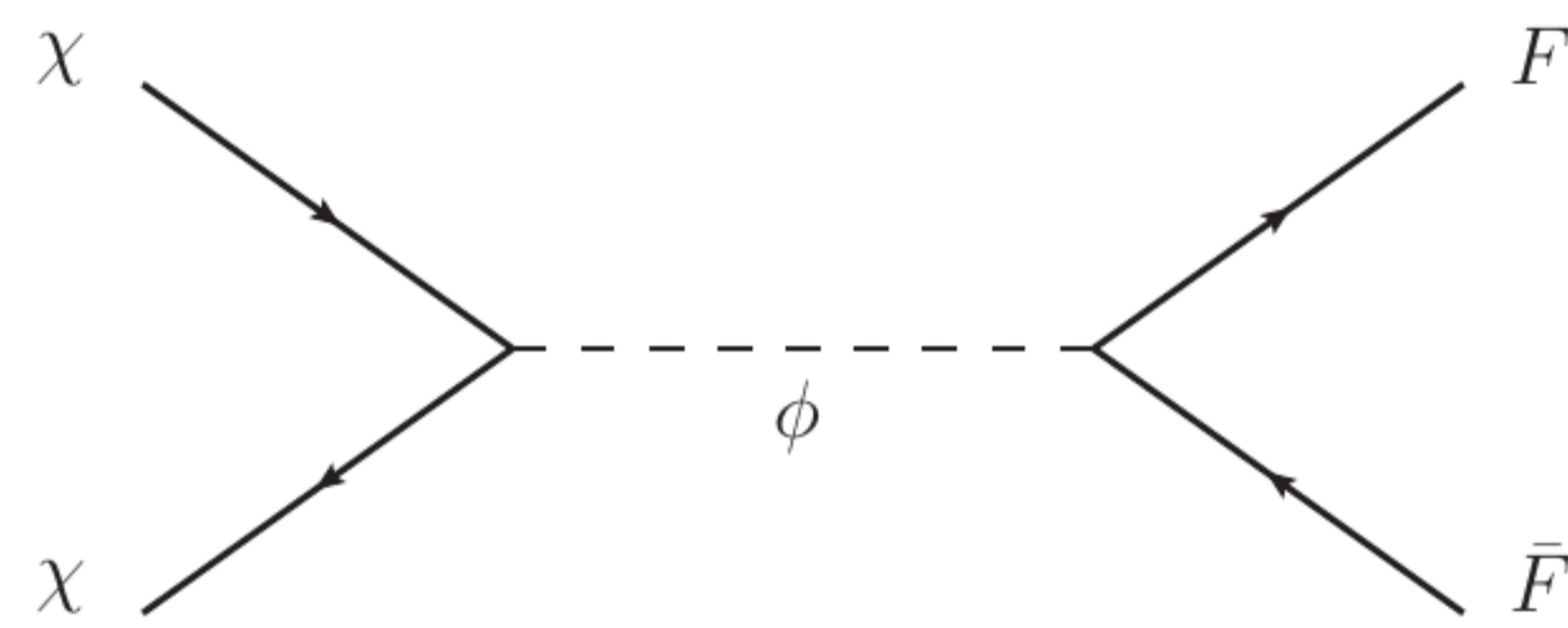
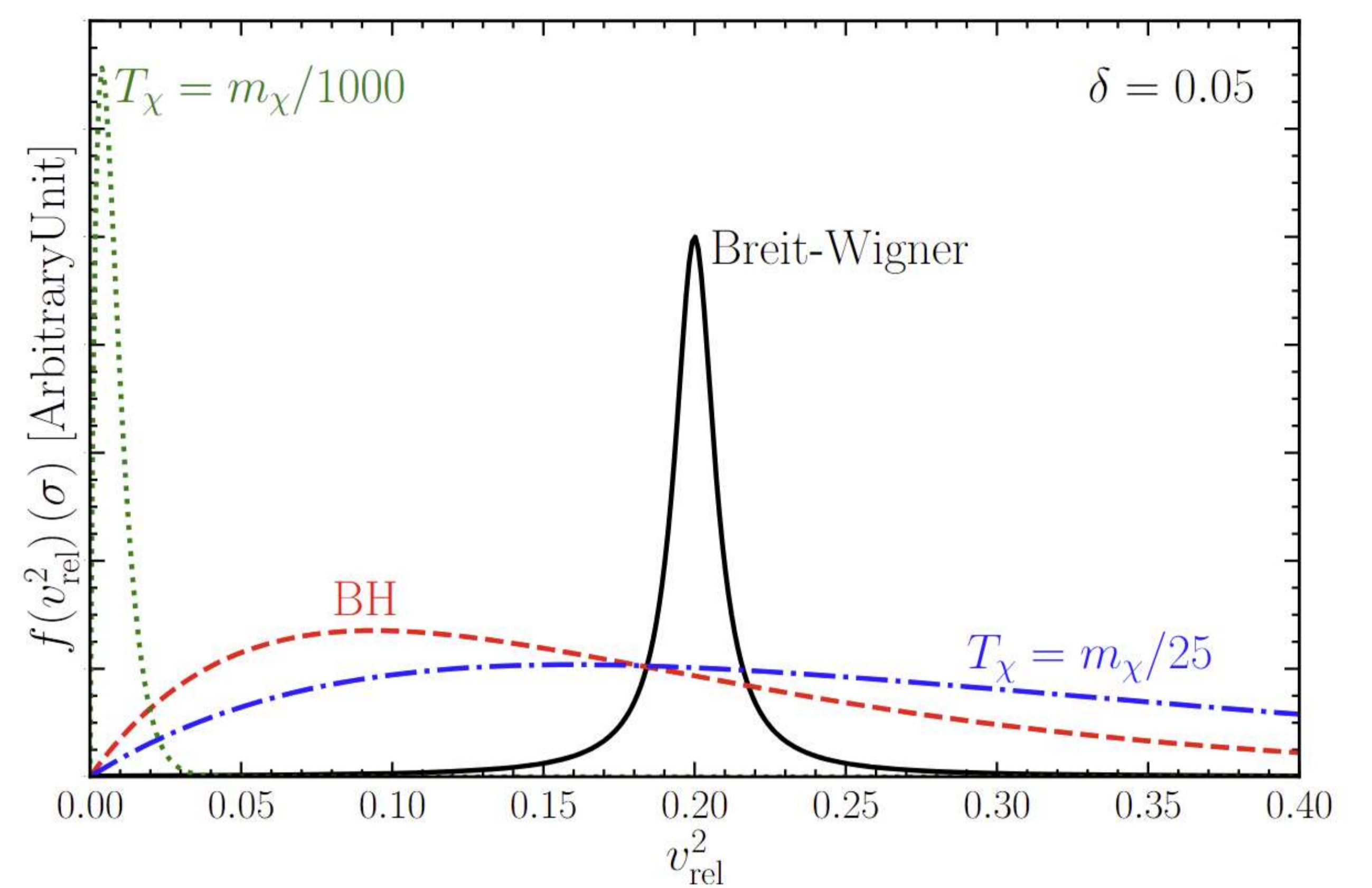
## 1. CMB constraint for Freeze out



**We need:**

- Freeze out:  $T \sim m_\chi/25$   
 $\langle\sigma_{\chi\chi}v\rangle \sim 10^{-26}\text{cm}^3/\text{s}$
- CMB:  $T \sim eV$   
 $\langle\sigma_{\chi\chi}v\rangle \ll 10^{-26}\text{cm}^3/\text{s}$

## 2. DM Production with BW Resonance



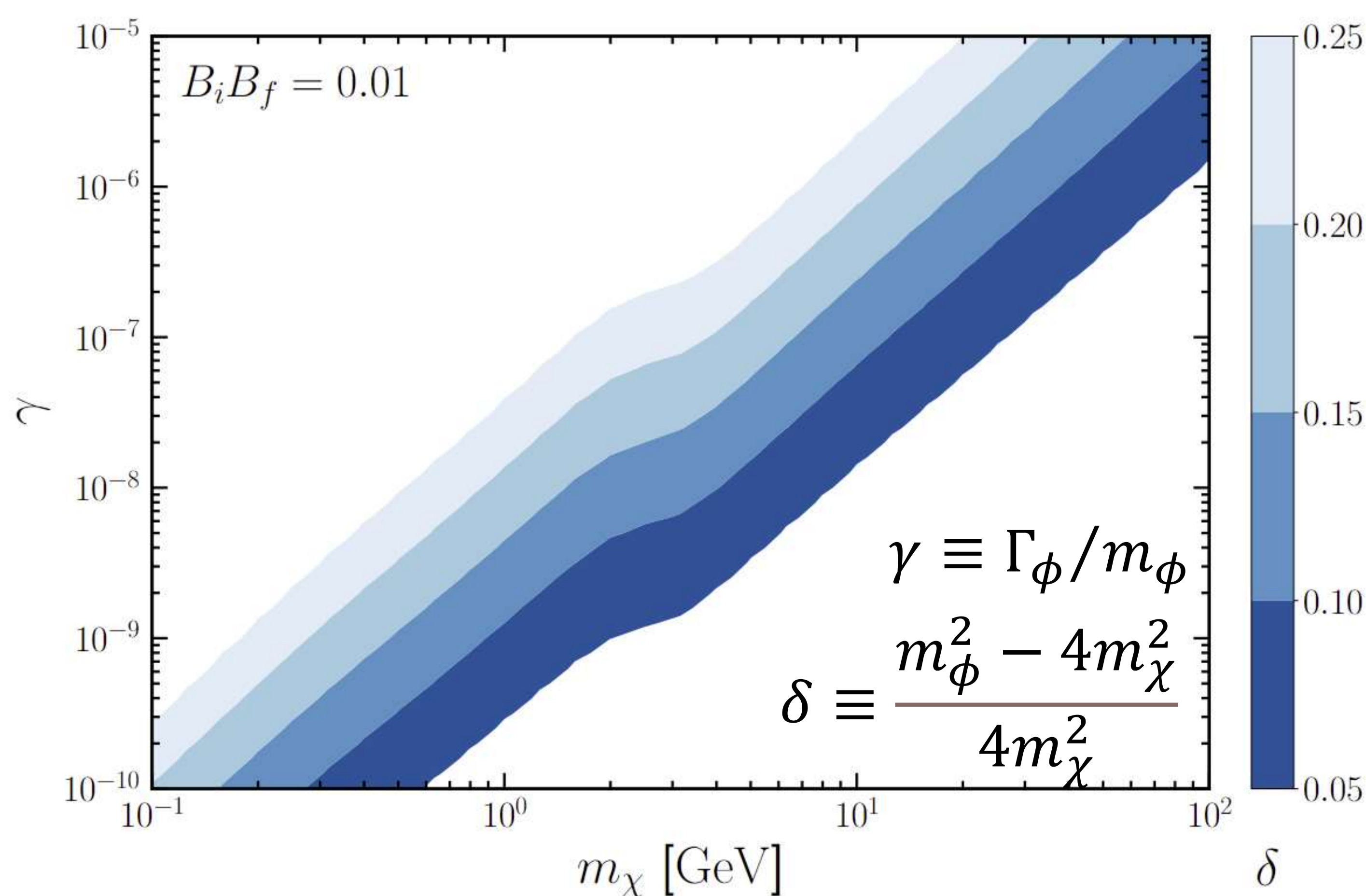
$$\sigma = \frac{16\pi\omega\beta_f}{s\bar{\beta}_i\bar{\beta}_f\beta_i} \frac{m_\phi^2\Gamma_\phi^2}{(s - m_\phi^2)^2 + m_\phi^2\Gamma_\phi^2} B_i B_f$$

## 3. Right-Handed Neutrino DM model

- **Seesaw mechanism**
  - $N_3$  Neutrino oscillation
  - $N_2$  Baryon asymmetry
  - $N_\chi$   $Z_2$  odd, Stable DM

$$\mathcal{L}_{\text{int}} = (y\phi N_\chi^T \epsilon N_\chi + \text{h.c.}) + \lambda m_\phi \phi H^\dagger H.$$

• **Relic Density:**



The parameter space to generate the correct relic density  $\Omega_\chi h^2 = 0.12$  for the Breit-Wigner resonance scenario with a heavy mediator.

## 4. Black hole acceleration



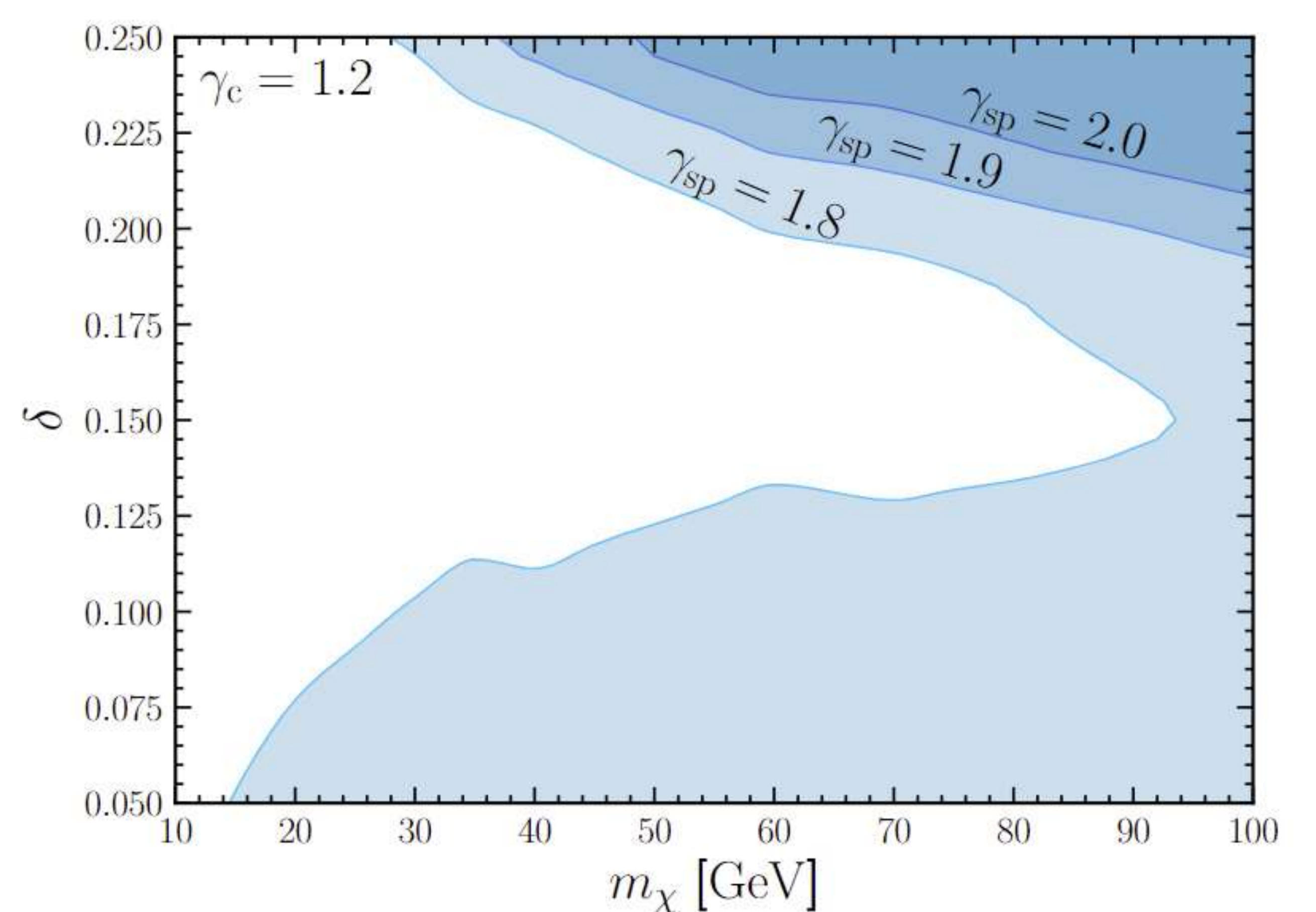
$$v^2 \sim \frac{GM}{r}$$

$$\downarrow$$

$$v(r) \sim \frac{1}{\sqrt{r}}$$

$$\rho \propto r^{-7/3}$$

• **Constraints from Fermi-LAT**



The allowed parameter space not exceeding the constraints from each data bin of the Fermi-LAT gamma-ray observation at 95 % C.L.