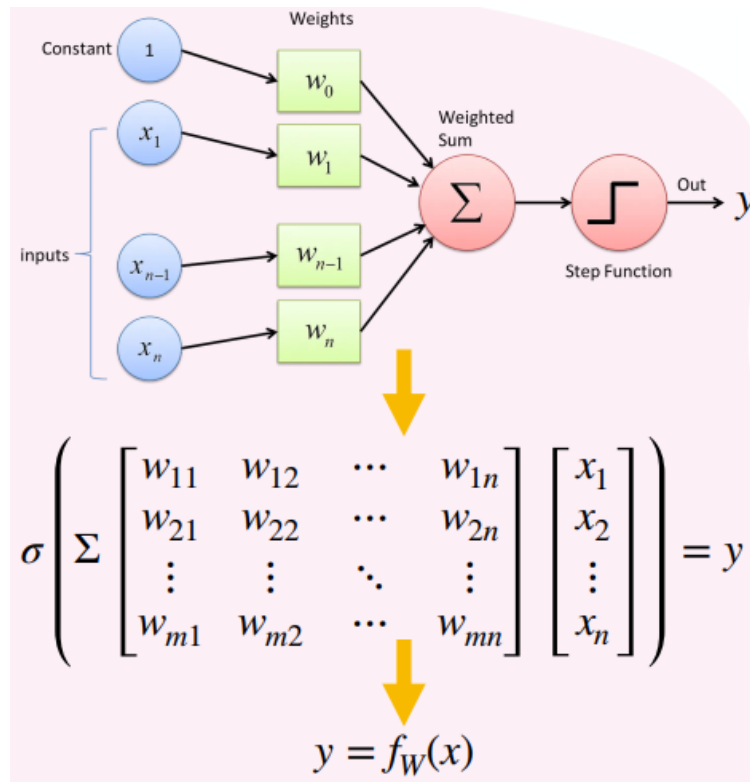


# Reconstruction Based on GNN

Cen Mo, Fan Hu, Fuyudi Zhang, Liang Li

- Neural network is only a **function** that maps input to output.
- Machine learning: use computer to find the **BEST function** for our tasks.



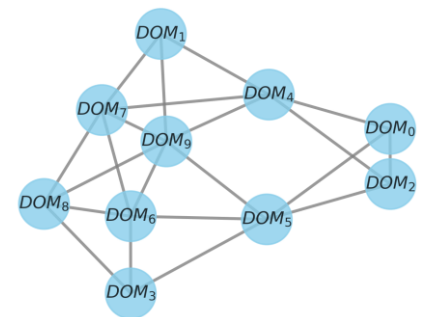
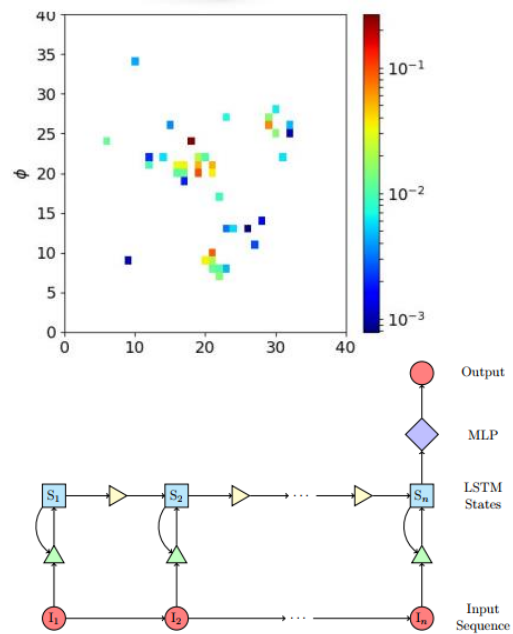
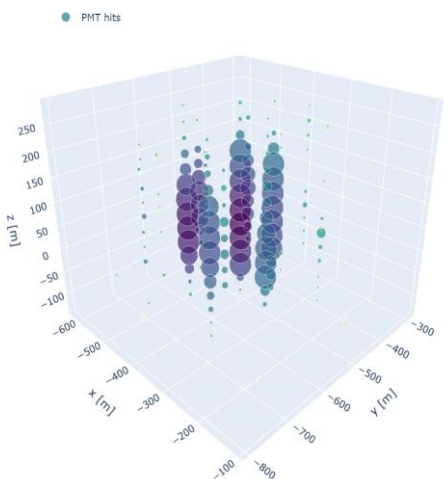
- Multilayer Perceptron (**MLP**)
- Boosted Decision Tree (**BDT**)
- Convolutional Neural Network (**CNN**)
- Recurrent Neural Network (**RNN**)
- Graph Neural Network (**GNN**)

**Vector** with fixed length:  $[DOM_1, DOM_2, \dots, DOM_{24220}]$  or Boosted Decision Tree

**Image**

**Ordered Sequence**

**Graph**



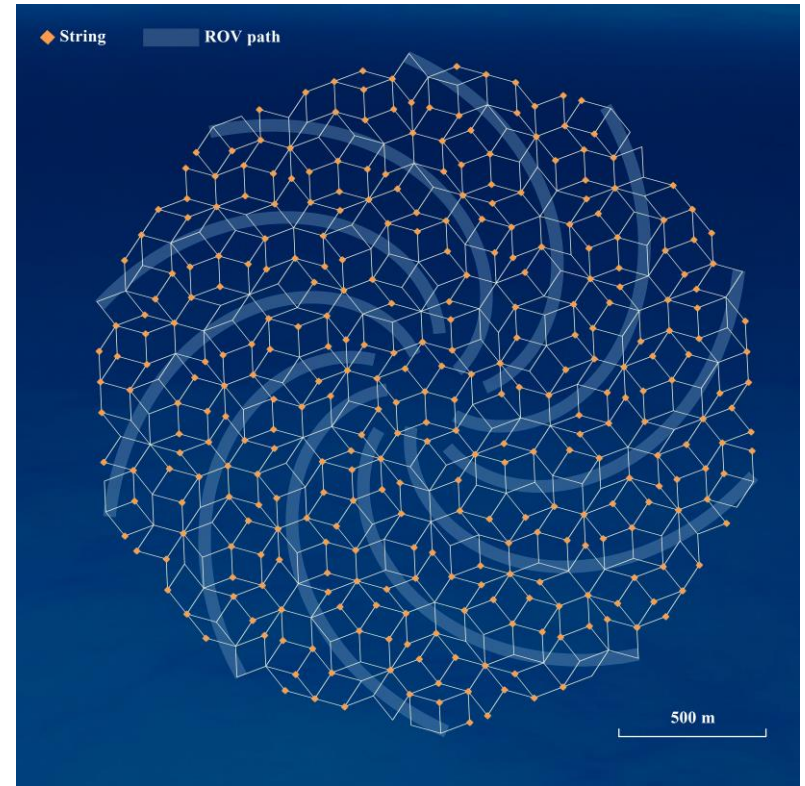
- Multilayer Perceptron **(MLP)**
- Boosted Decision Tree **(BDT)**
- Convolutional Neural Network **(CNN)**
- Recurrent Neural Network **(RNN)**
- Graph Neural Network **(GNN)**

## Neutrino telescope:

- Irregular detector geometry
- Sparse signal

Compared GNN and SSCNN ([arxiv:1706.01307](https://arxiv.org/abs/1706.01307)) performance:

- GNN **outperforms** SSCNN in terms of angular resolution in track-like events.

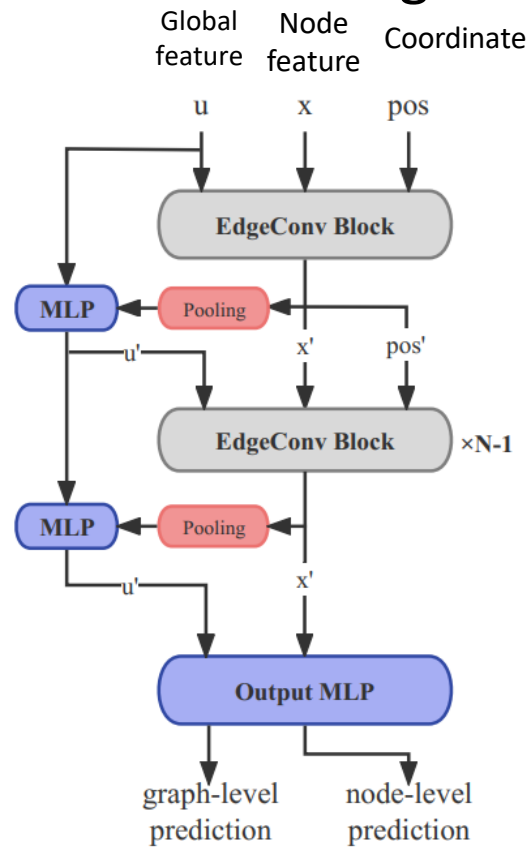


Top view of TRIDENT

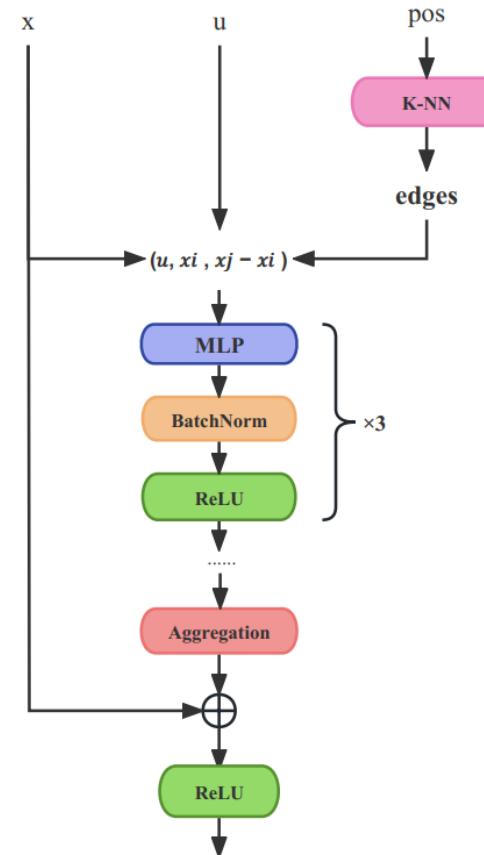
Use **point cloud** to represent neutrino events:

- Triggered DOMs → **Nodes** of point cloud
- Location of DOMs → Coordinate of nodes,  $pos_i$ .
- DOM-measured time and charge → Features of nodes,  $x_i$ .

- TridentNet is built based on EdgeConv block: modified from EdgeConv block used in ParticleNet ([arxiv:1902.08570](https://arxiv.org/abs/1902.08570)).
- Both **graph-level** and **node-level** target can be predicted.



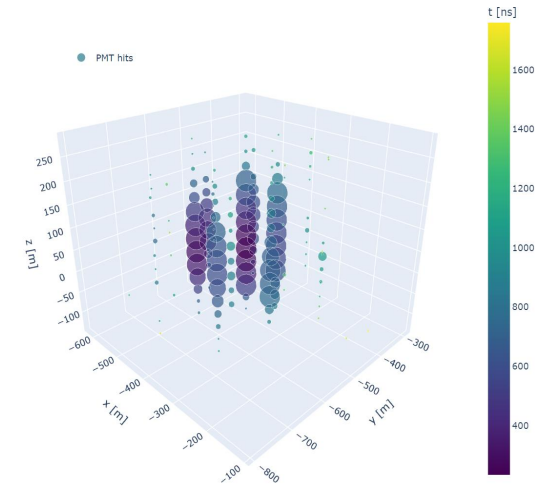
Network Structure



EdgeConv Block

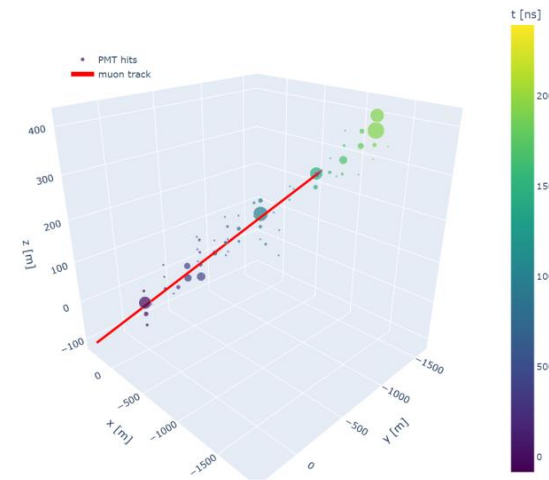
- Reconstruction of  $\nu_e$ /cascade events

- Direction Reconstruction
- Energy Reconstruction



- Reconstruction of  $\nu_\mu$ /track events

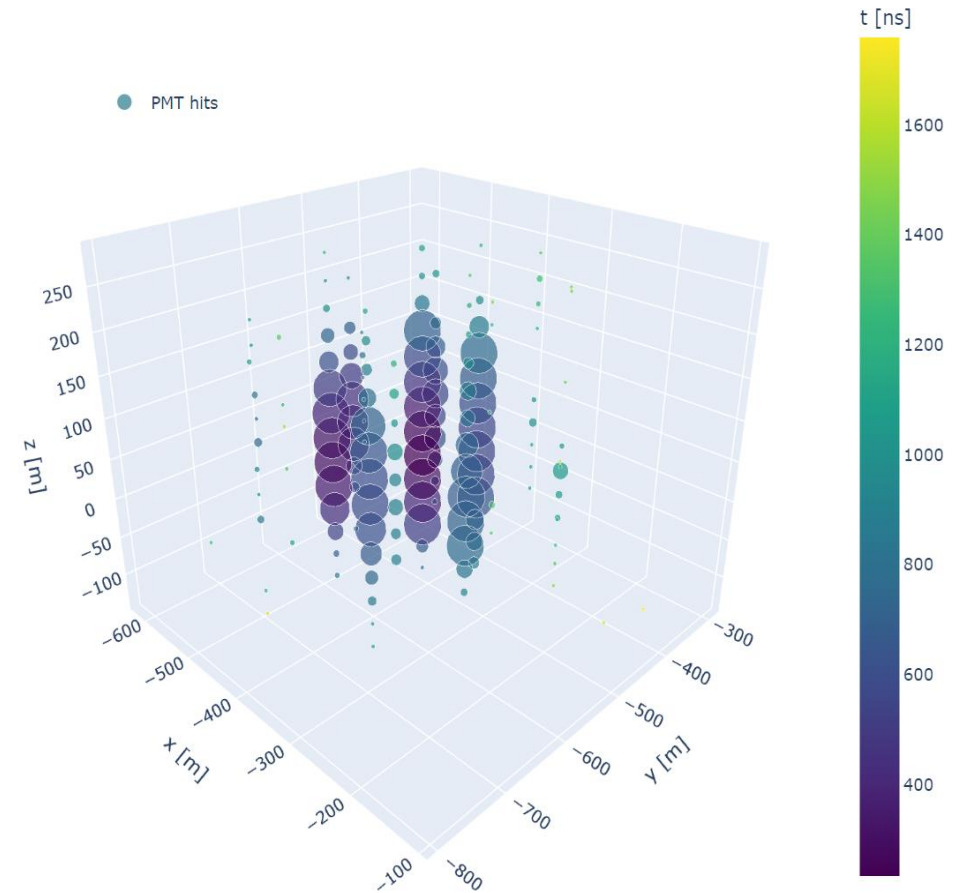
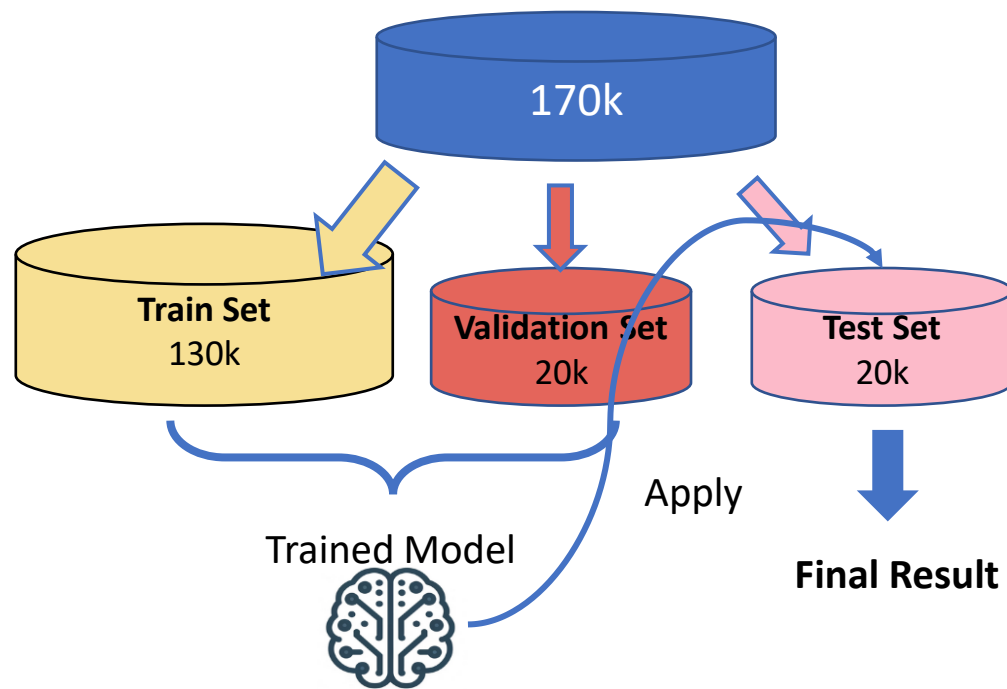
- Direction Reconstruction
- Energy Reconstruction



## $\nu_e$ direction reconstruction

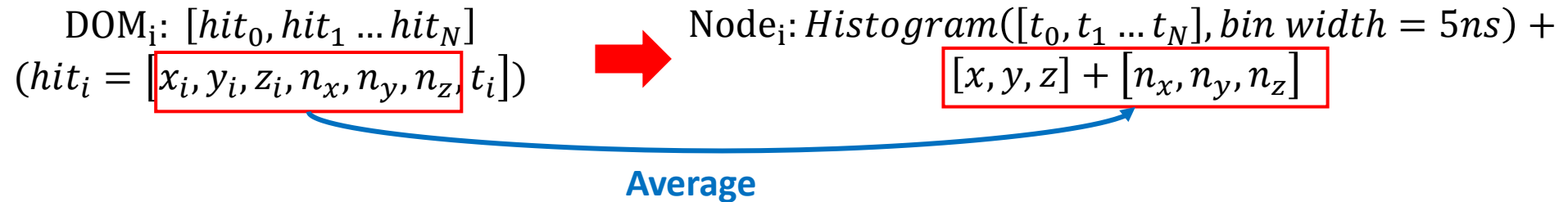
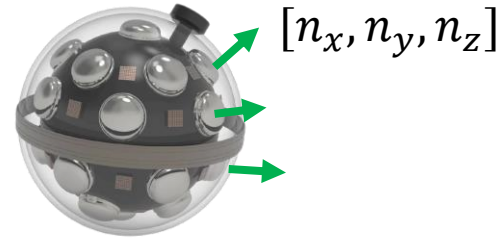
### Configuration for Direction Reconstruction

- $\nu_e$  energy: 100TeV
- Sample size:



## $\nu_e$ direction reconstruction

- Input feature of  $\text{DOM}_i$ :



- Histogram is further modified (as a **normalization**):

$$\text{Node}_i = [x, y, z, n_x, n_y, n_z, \ln(\text{Histogram} + 1)]$$

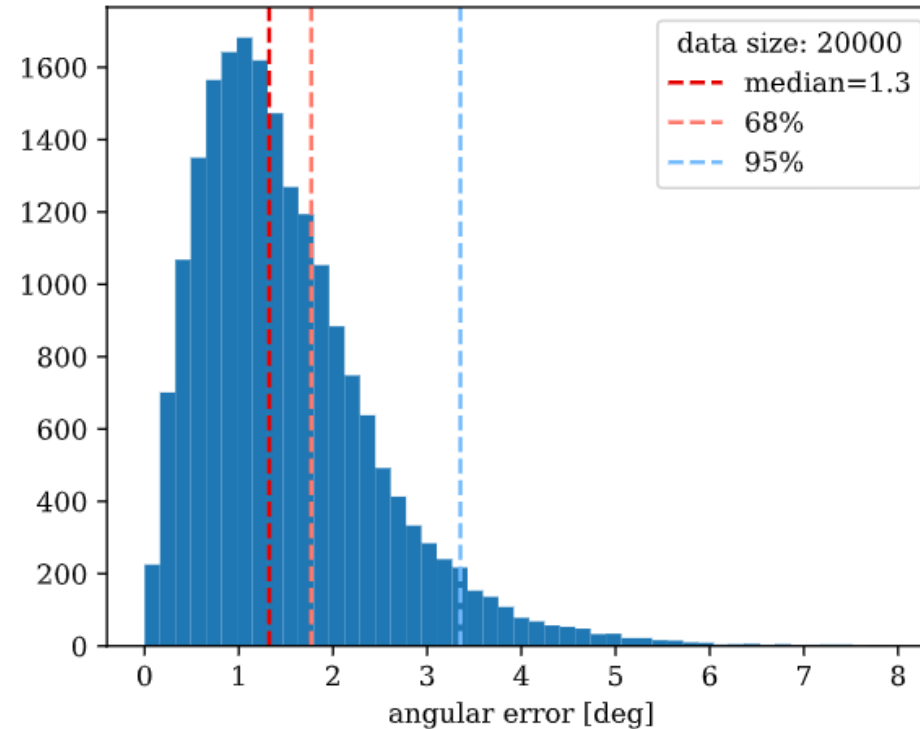
- Network is trained to predict  $\hat{n}_\nu$  with **MSE loss**:

$$\text{Loss} = \left| \frac{\overrightarrow{\text{output}}}{|\text{output}|} - \hat{n}_\nu \right|^2$$



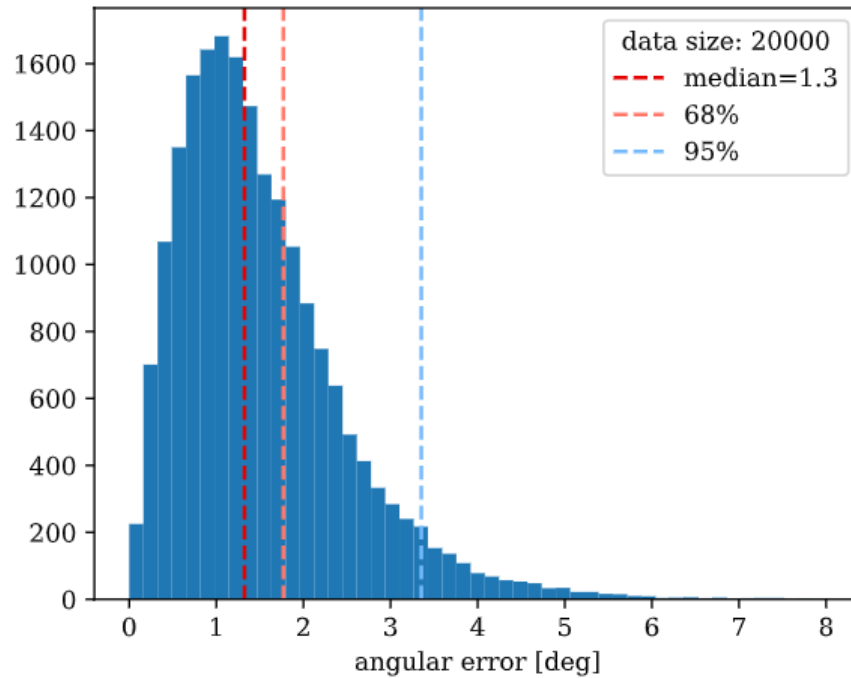
## $\nu_e$ direction reconstruction

- Resolution reaches **1.3 degrees**.

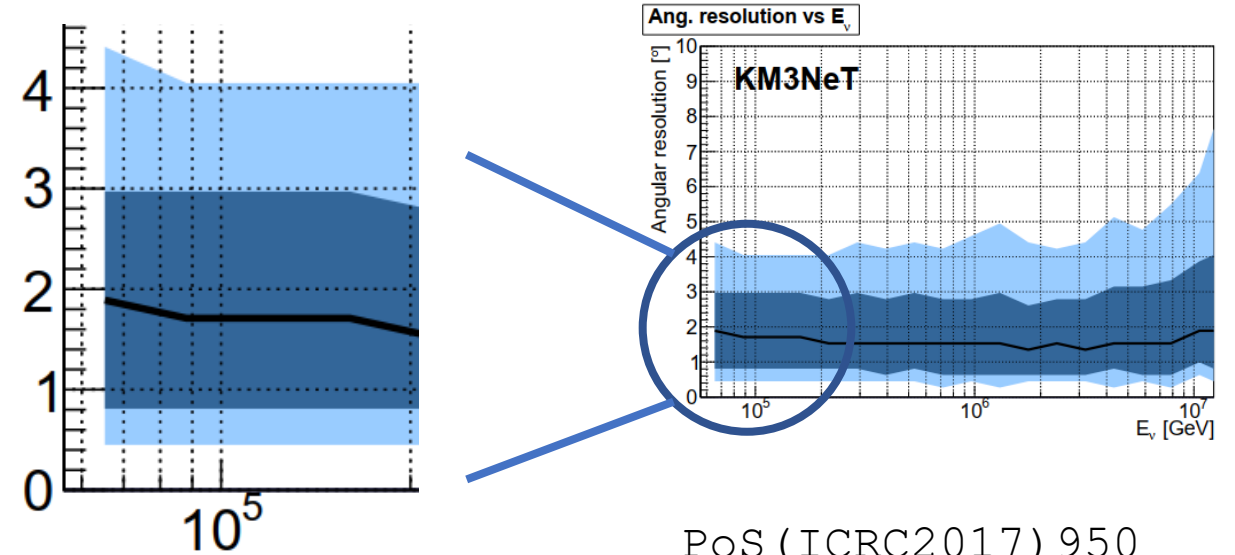


## $\nu_e$ direction reconstruction

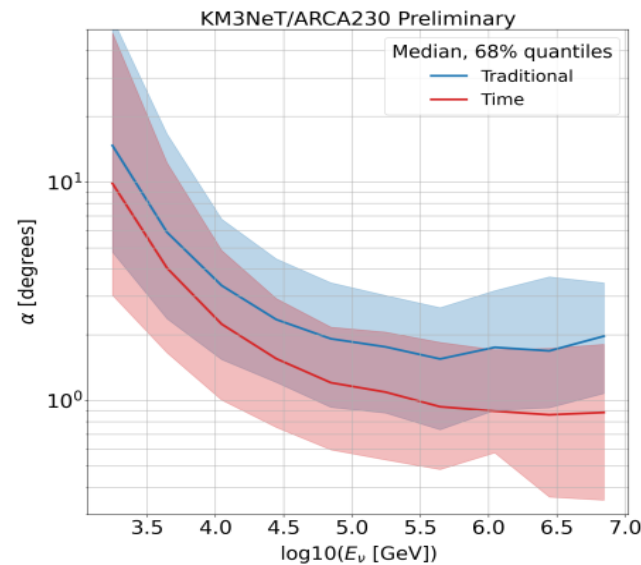
- Resolution reaches **1.3 degrees**.



## KM3NeT Results (likelihood)



PoS (ICRC2017) 950



PoS (ICRC2023) 1074  
**With pre-selection**

- $\nu_e$  energy reconstruction

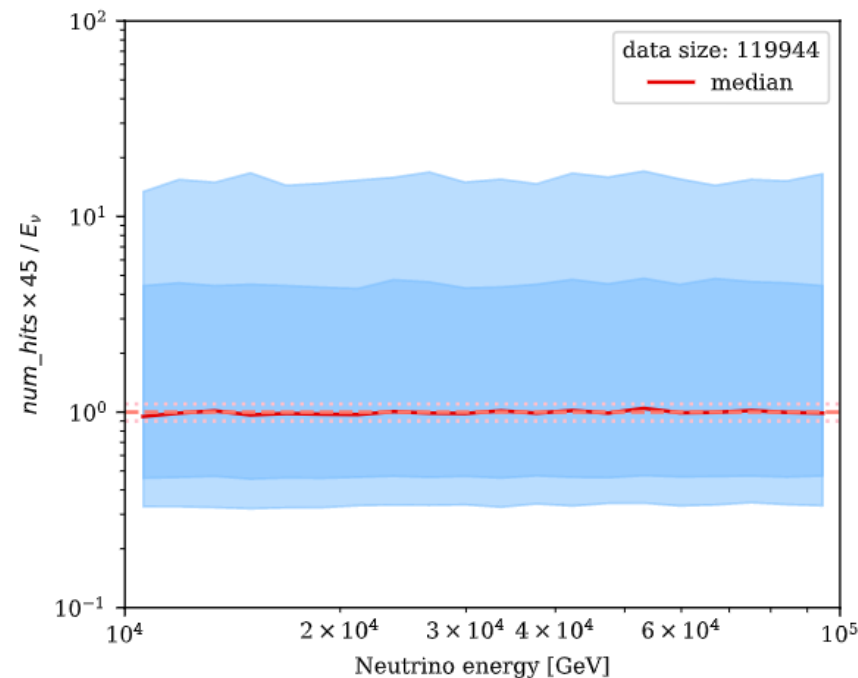
Configuration for Energy Reconstruction

- $\nu_e$  energy: 10TeV ~ 100TeV

- Sample size: 150k samples are splitted into:

*train* : *validation* : *test* = 120k : 15k : 15k

## Linearity between num\_hits & Energy

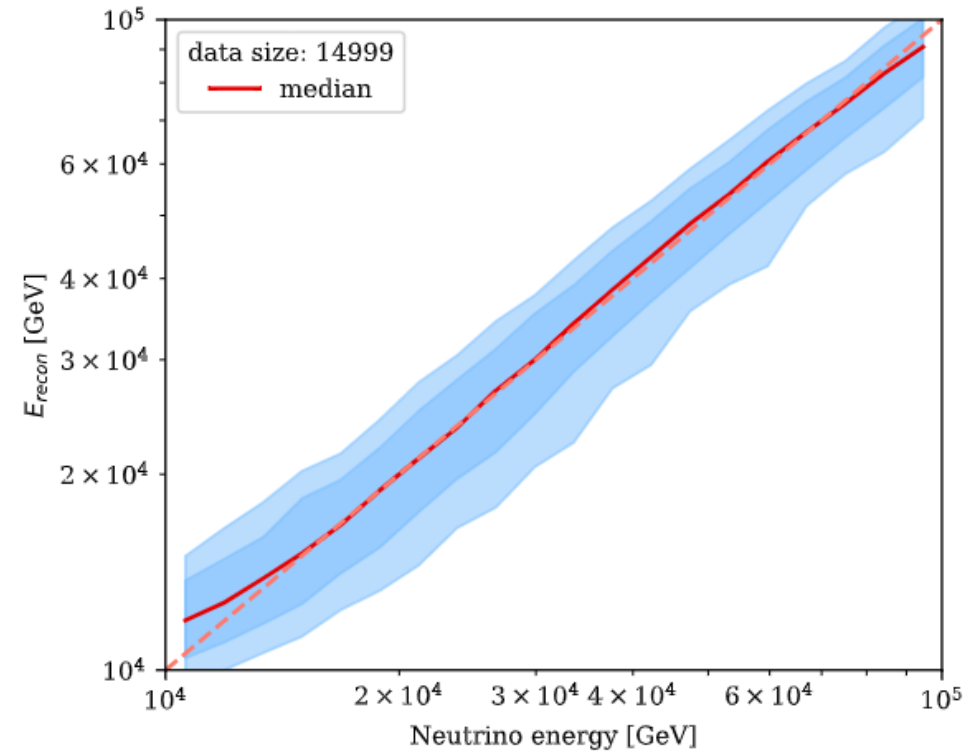
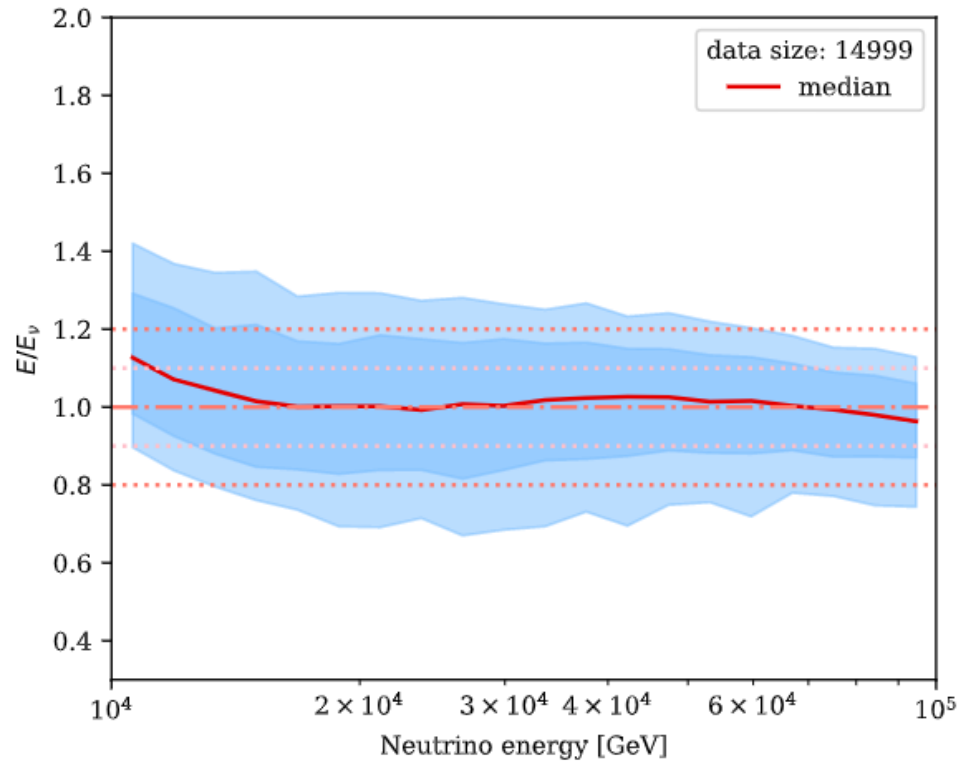


- Train GNN with:

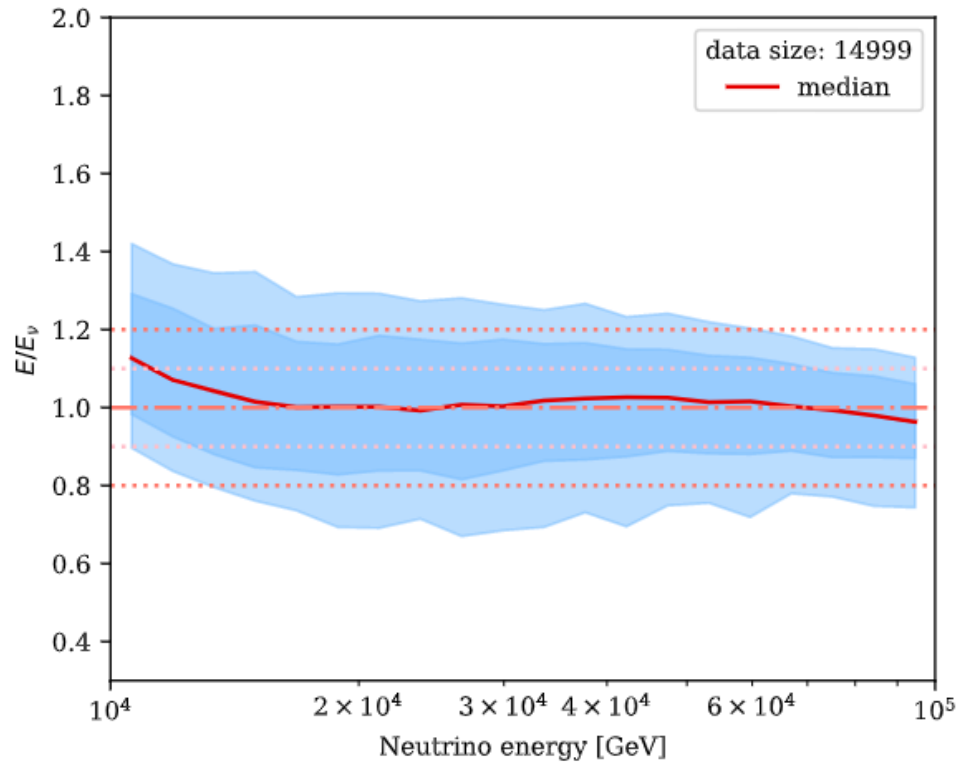
$$\log_{10} E = GNN(\text{graph}) + \log_{10}(\text{num\_hits} \times 45.78)$$

$$\text{Loss} = (\log_{10} E - \log_{10} E_{\text{truth}})^2$$

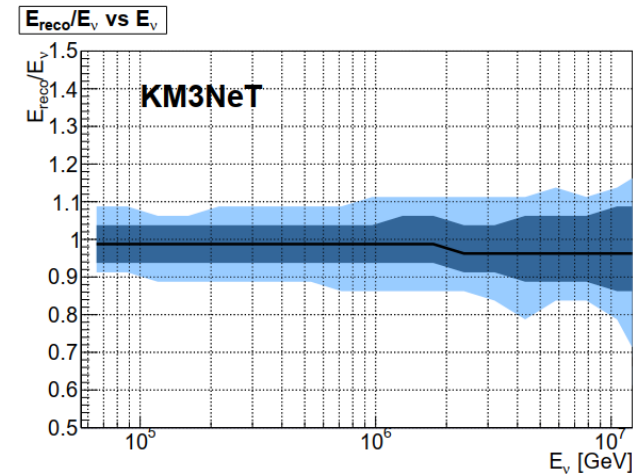
- Energy resolution is around 10% for high energy event.



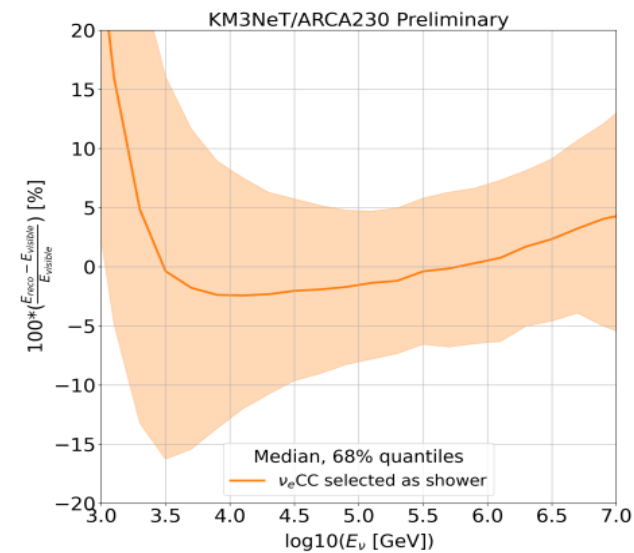
- Comparison



## KM3NeT Results (likelihood)



PoS (ICRC2017) 950



PoS (ICRC2023) 1074  
**With pre-selection**

- $\nu_\mu$  **Direction** reconstruction

Configuration for Energy Reconstruction

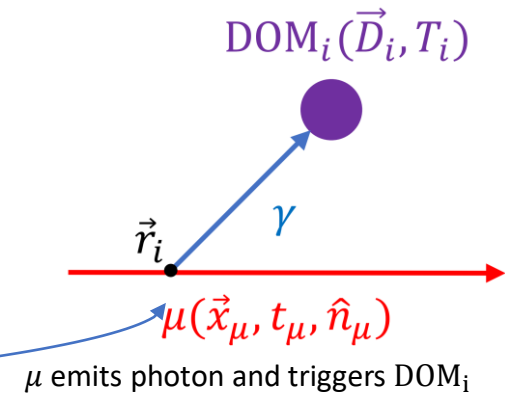
- $\nu_\mu$  energy: 1TeV ~ 1PeV

- Sample size:

*train* : *validation* : *test* = 900k : 70k : 100k

## $\nu_\mu$ Direction reconstruction

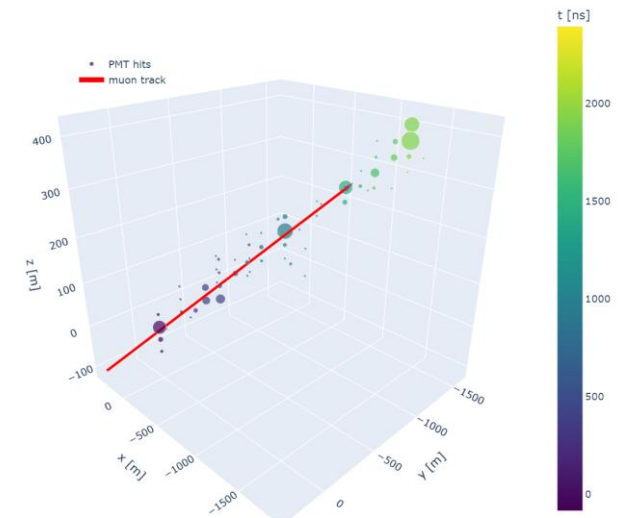
- **Input features:** location  $\vec{D}_i$ , first photon arrival time  $T_i$  and number of photo hits  $n_i$
- To make full use of the geometric feature of track-like events, the network is trained to predict  $\vec{r}_i$  for each  $\text{DOM}_i$ .



- **Loss function:** mean square error (MSE) with weight proportional to  $n_i$ :

$$\text{Loss} = \sum_i n_i \times |\overrightarrow{\text{output}}_i - \vec{r}_i|^2 / \sum_i n_i$$

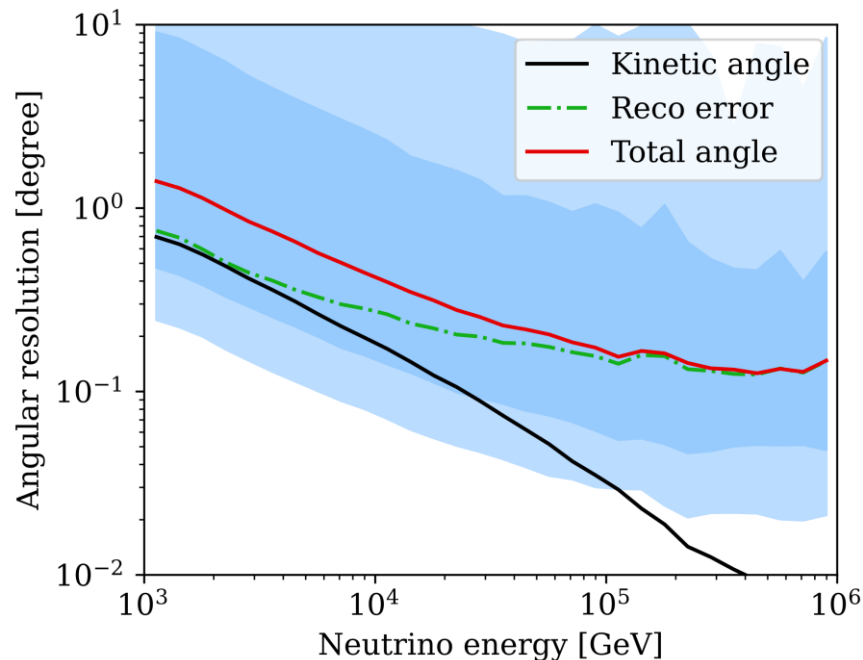
- Linear fit on the predicted  $\vec{r}'_i$  then reconstructs  $\hat{n}_\mu$ .



Track-like event display

## Direction reconstruction

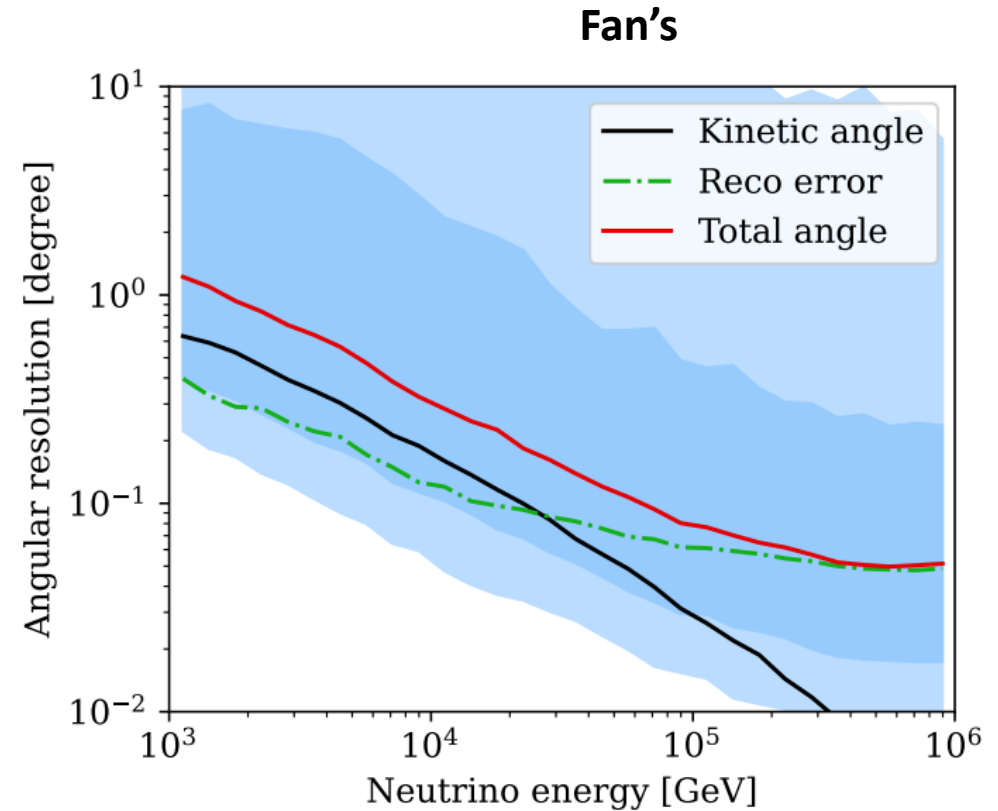
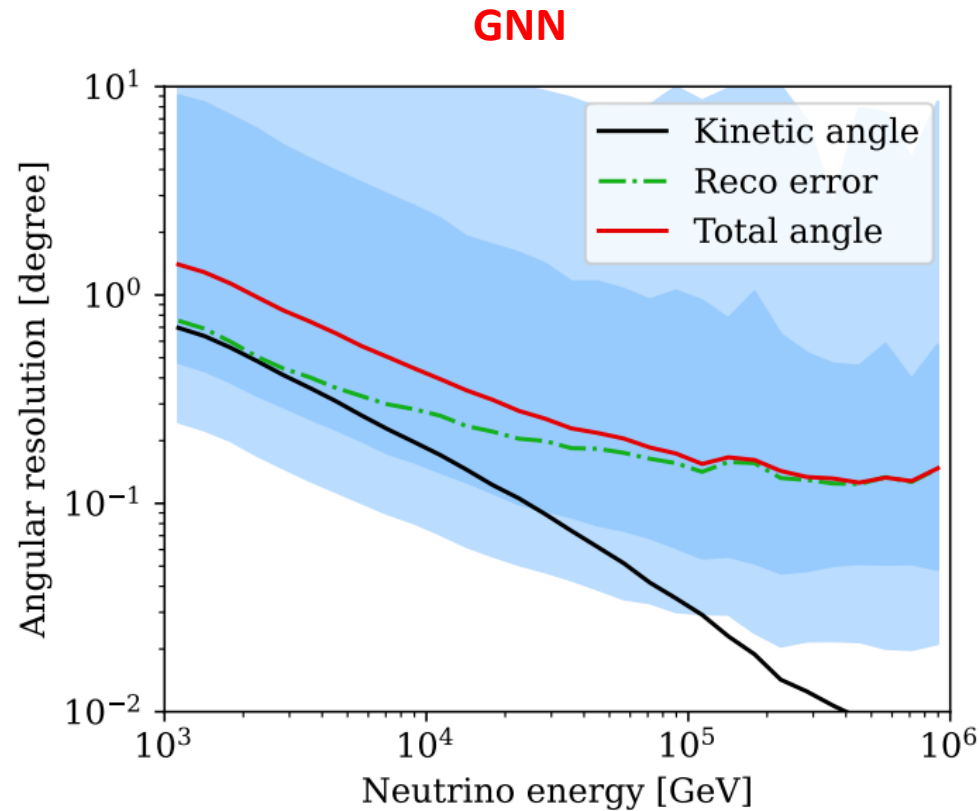
- Model is trained on events with track length > 500m.
- Median angular error decreases from 1 degree to **0.1 degree** as the energy of  $\nu_\mu$  increases – similar to the result of likelihood method.



- Kinetic angle =  $\langle \vec{n}_\mu, \vec{n}_\nu \rangle$
- Reco error =  $\langle \vec{n}_\mu, \vec{n}_{recon} \rangle$
- Total angle =  $\langle \vec{n}_\nu, \vec{n}_{recon} \rangle$



## Comparison



# Track-like Events Reconstruction

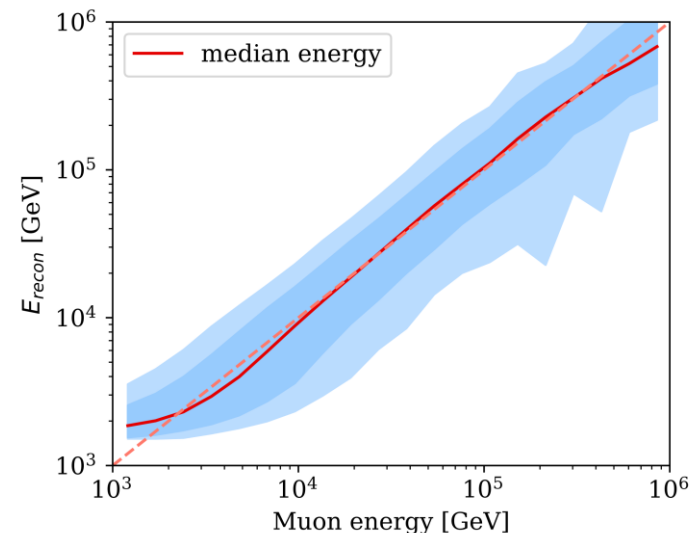
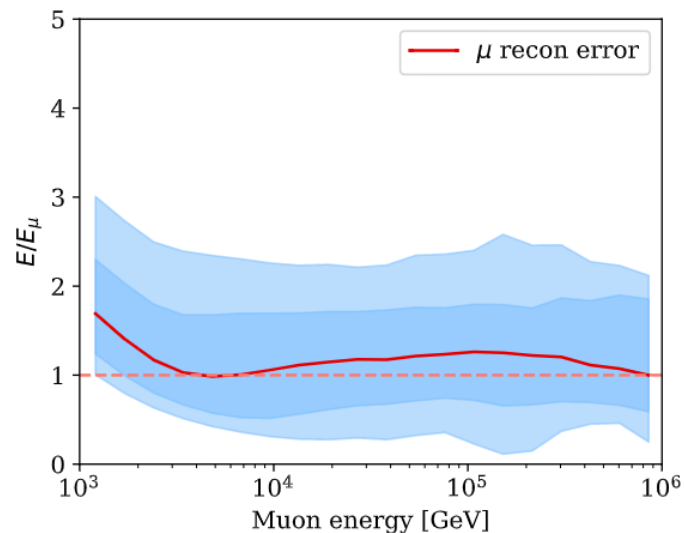
## Energy reconstruction

- Same input features as the direction reconstruction.
- Network is trained with **MSE loss** to predict  $\log_{10} E_\mu$ . Weight  $w = \log_{10} E_\mu - 2.5$  is applied:

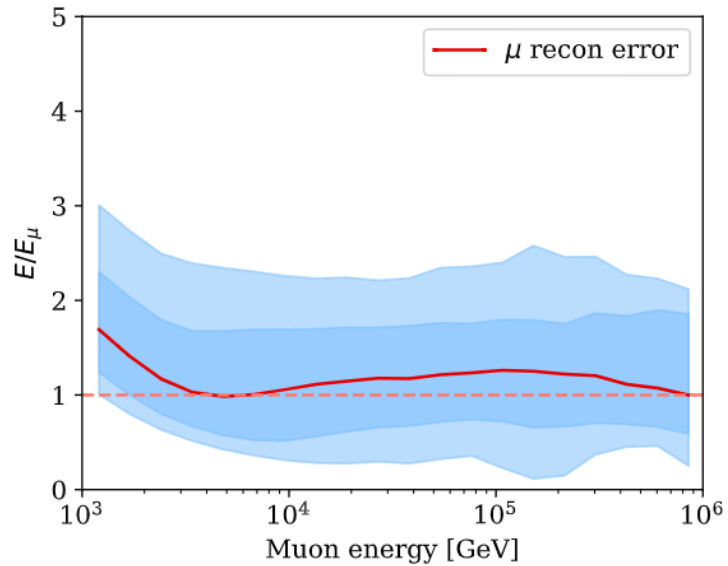
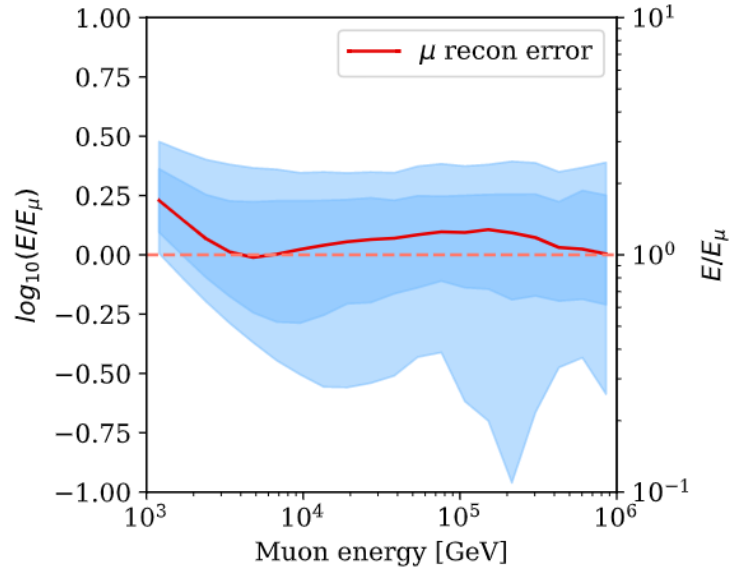
$$Loss = w(\text{output} - \log_{10} E_\mu)^2$$

- A shift term,  $b = 0.15$  is added to outputs of the model:

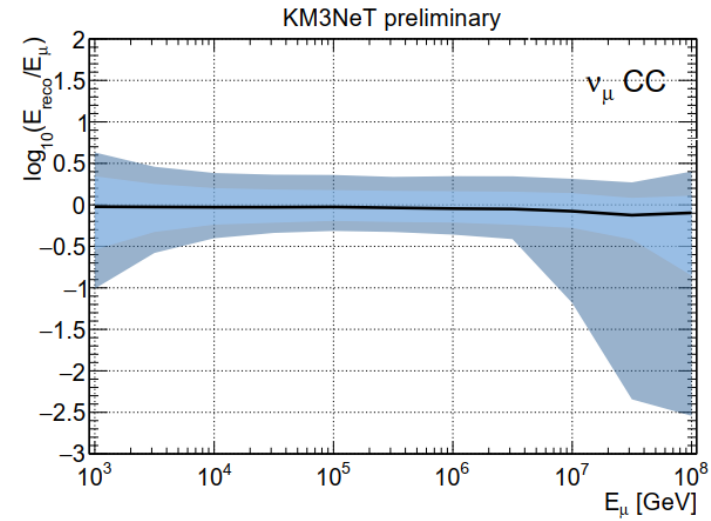
$$E_{recon} = 10^{\text{output}+b}$$



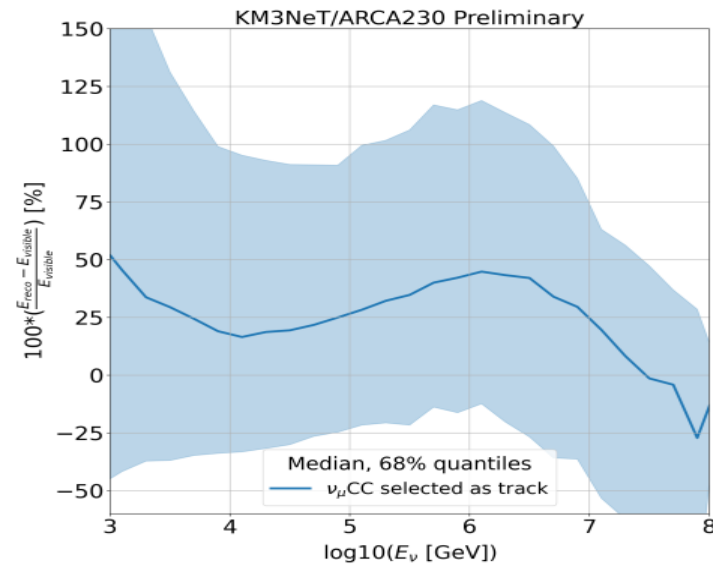
## Comparison



## KM3NeT Results (likelihood)



PoS (ICRC2017) 950



PoS (ICRC2023) 1074  
**With pre-selection**

Task	Resolution (GNN)	Resolution (KM3NeT)
<b>Cascade direction</b>	<b>1.3 degrees</b>	2 degrees (w/o time) ~1 degree (with time)
<b>Cascade energy</b>	<b>10% (high E)</b>	5~10%
<b>Track direction</b>	<b>~0.1 (high E)</b>	0.1 (high E)
<b>Track energy</b>	<b>100%</b>	100%

- Simulated neutrino events in TRIDENT are represented as point clouds and are reconstructed by TridentNet.
- GNN demonstrates high accuracy in reconstructing  $\nu_e$  and  $\nu_\mu$ .
  
- **Improvement** of neutrino reconstruction will be further studied.
- Future research will try to enhance the method's robustness against experimental **uncertainties and noise**.

# Thanks for listening!

Email: [mo\\_cen@sjtu.edu.cn](mailto:mo_cen@sjtu.edu.cn)

- **Learning rate:**

Initial learning Rate = 0.003

lr scheduler: `ReduceLROnPlateau(factor=0.5, patience=5)`

- **optimizer:**

`Adam(betas=(0.9, 0.999), eps=1e-8, weight_decay=0)`

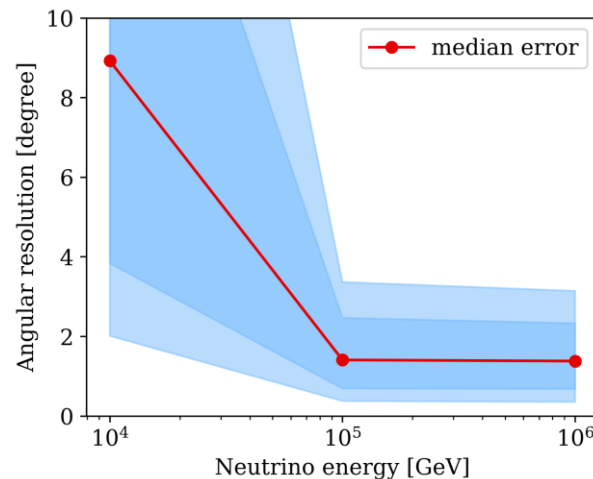
$$\log \mathcal{L} = \sum_{empties} \log[P_i^{nohit}] + \sum_{hitPMTs} \log[1 - P_i^{nohit}]$$

$$P_i^{nohit} = \exp[-\mu_{sig}(r_i, z_i, a_i, E_S) - R_{bg} \cdot T]$$

(equation 5.2) are considered. The expected number of hits from the shower  $\mu_{sig}$  is evaluated using interpolation of a three-dimensional histogram depending on  $r_i, z_i$  and  $a_i$  at a shower energy of 1 PeV (figure 8). The expected number of hits at different shower energies is calculated using the fact that the number of emitted photons scales linearly with the shower energy  $E_S$ .

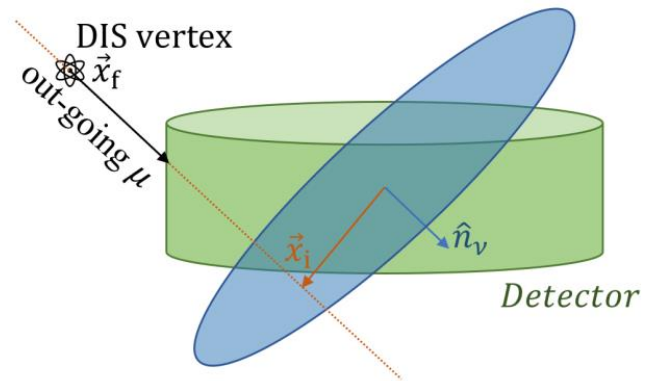
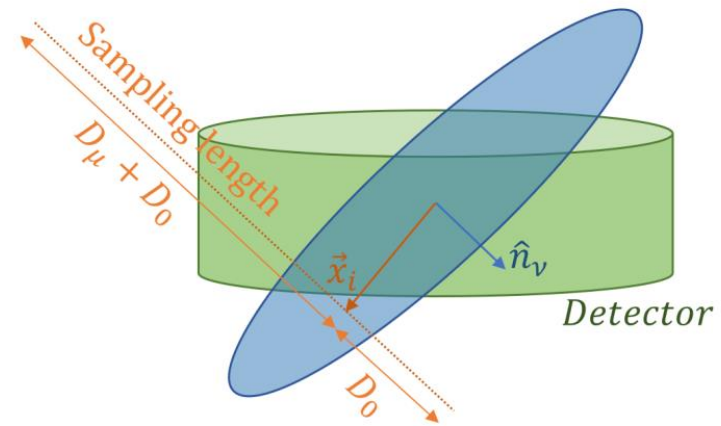
- Former GNN result on samples with other energy (by linear scaling num\_photons):

$$n' = n \times \frac{100TeV}{E_\nu}$$

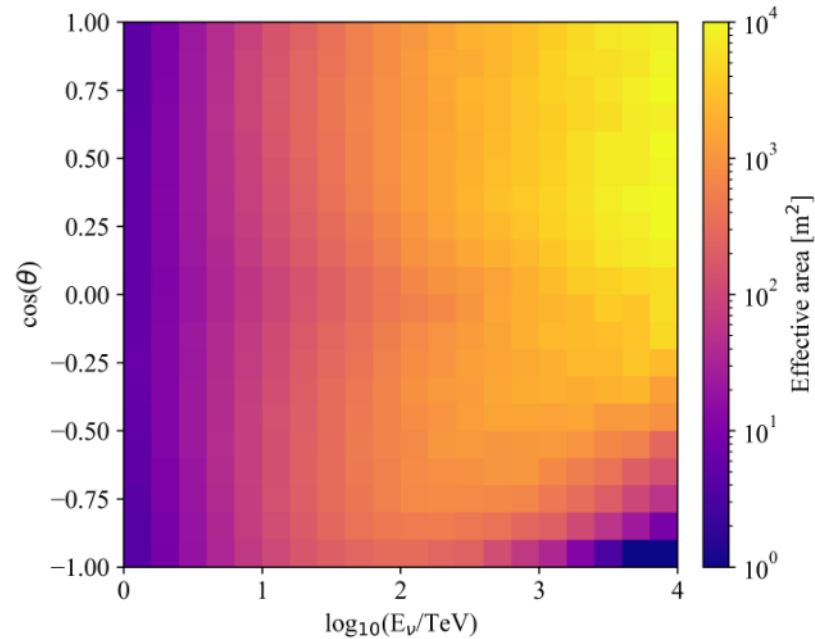




# $v_\mu$ Vertex Sampling



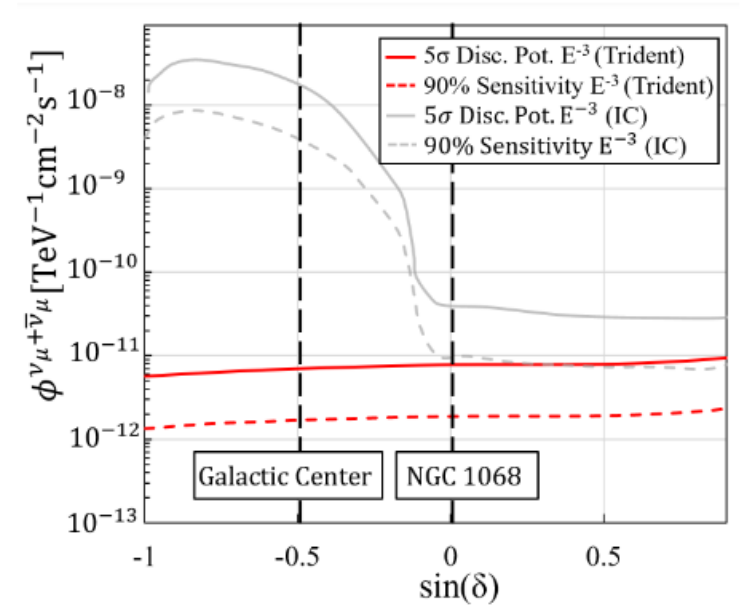
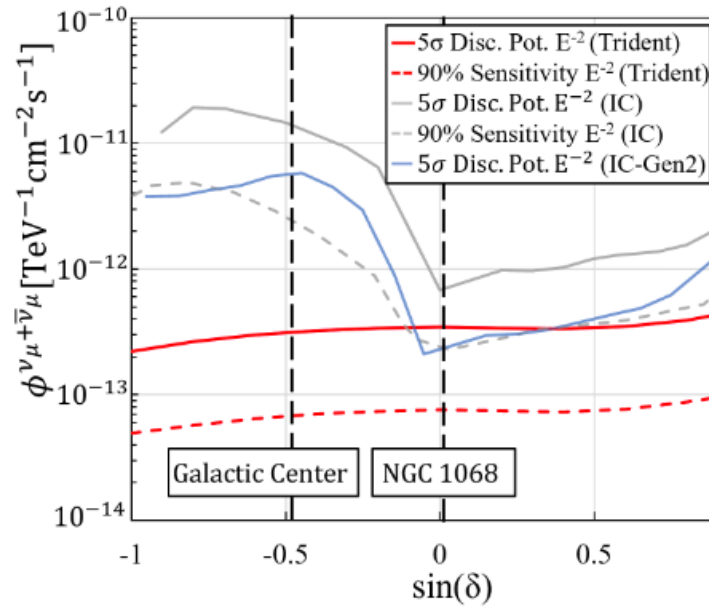
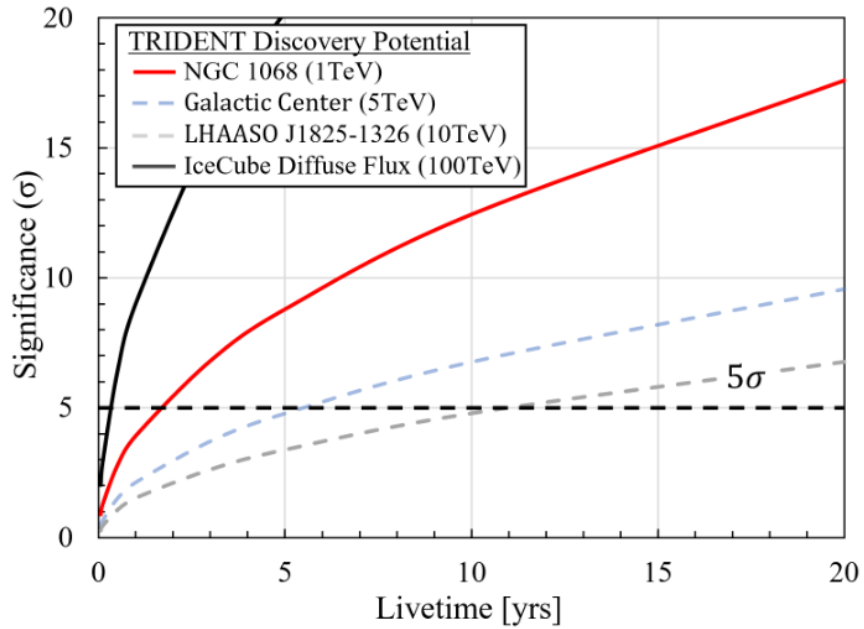
# Effective Area of $\nu_\mu$



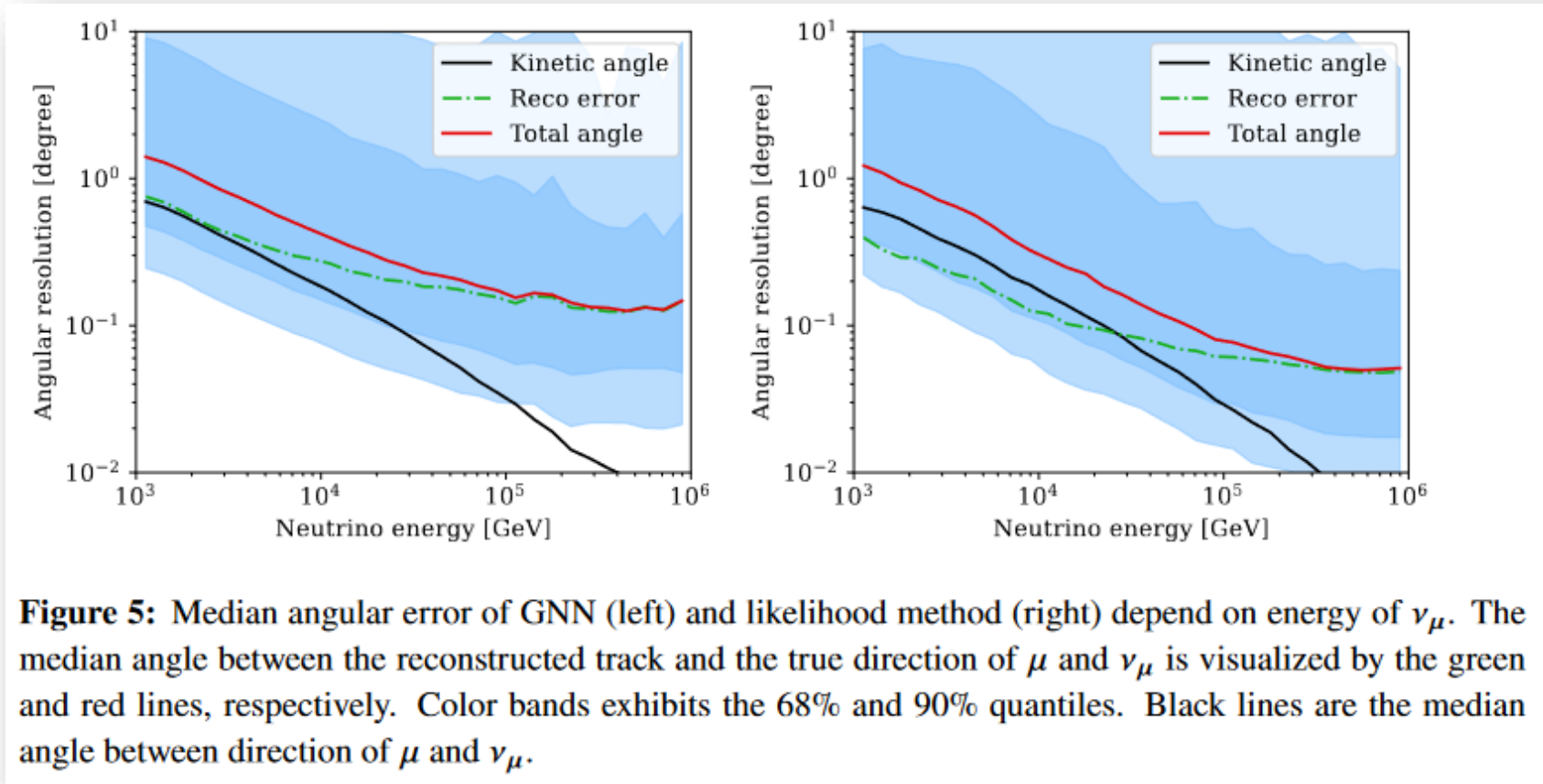
**Figure 15:** Effective areas at event reconstruction level for  $\nu_\mu$  track events as a function of primary neutrino energy and zenith angle in TRIDENT. At an energy of  $\sim 100$  TeV, the effective area for up-going events is expected to reach  $7 \times 10^2$   $\text{m}^2$ . Only events with angular error less than 6 degree are selected to evaluate the effective area.

# Significance & Sensitivity

[arXiv:2207.04519](https://arxiv.org/abs/2207.04519)



# Comparison with Likelihood Method



# Track-like Events Reconstruction

## Direction reconstruction

- Model is trained on events with track length > 500m.
- Median angular error decreases from 1 degree to **0.1 degree** as the energy of  $\nu_\mu$  increases.

