

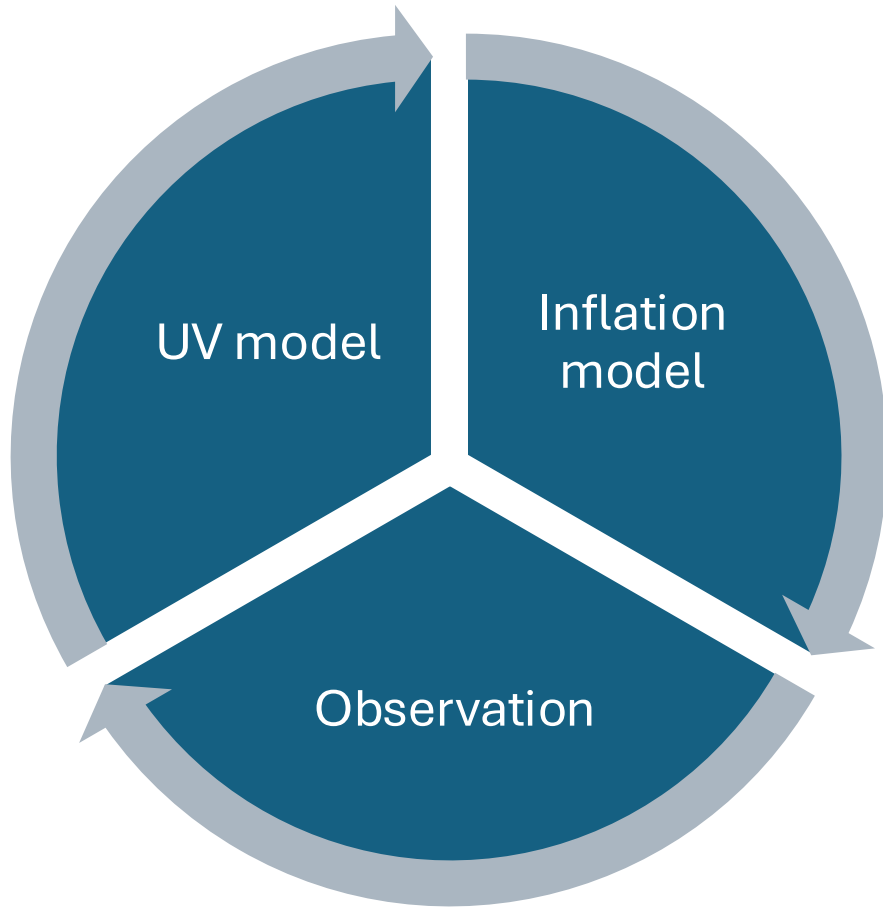
# Probing Particle Physics with the Cosmological Collider

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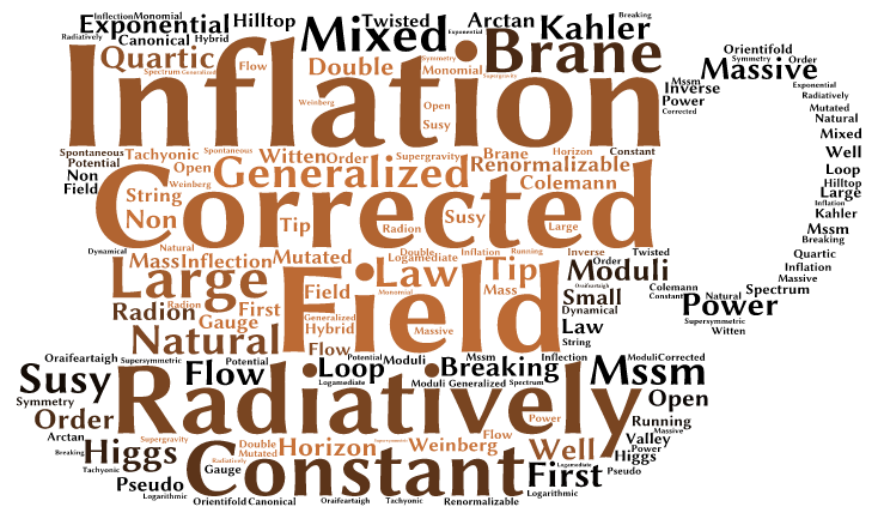
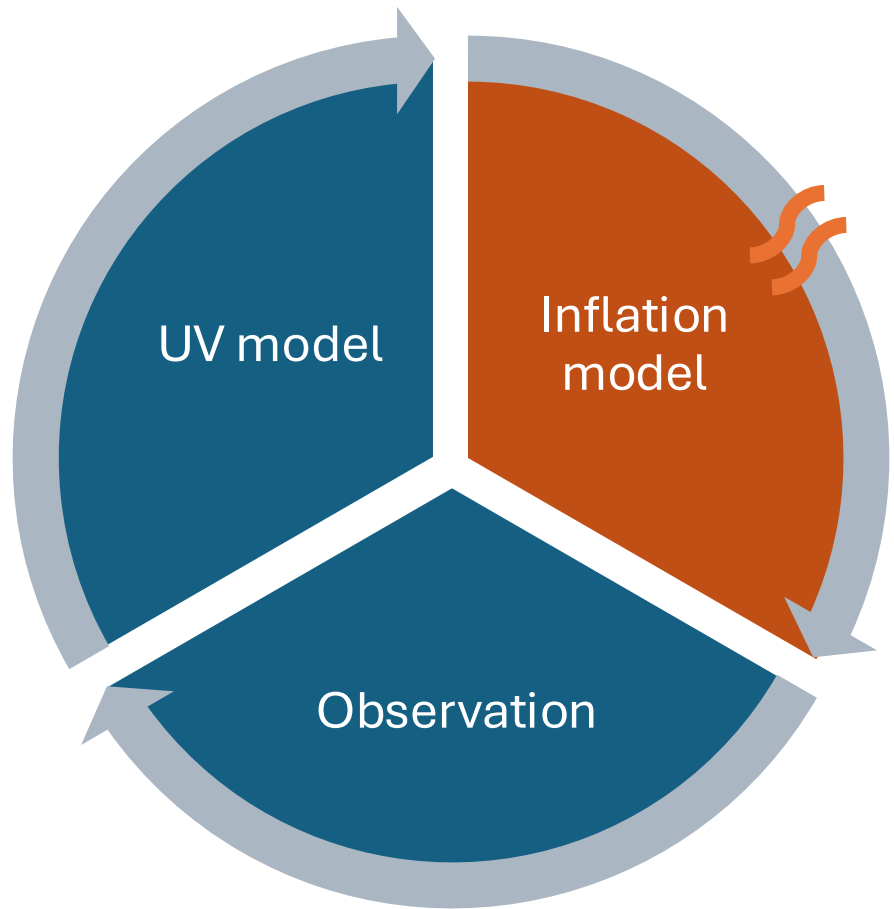
20 Aug 2024 @ Beihang University



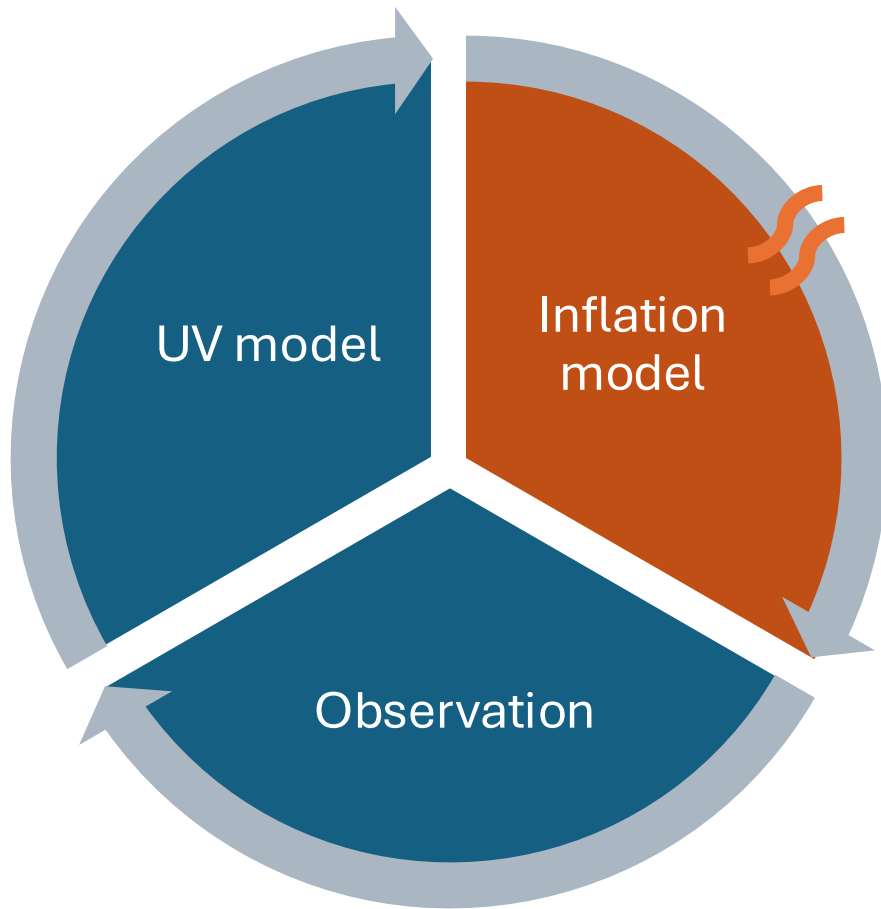
HEP from Inflation  
(Traditional)



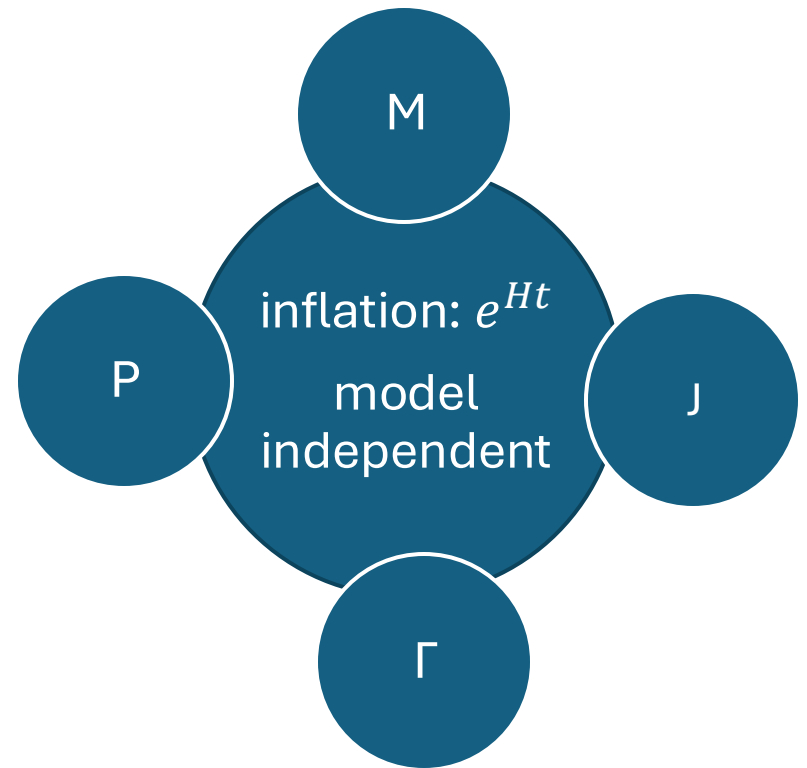
# HEP from Inflation (Traditional)



## HEP from Inflation (Traditional)



## Cosmological Collider



From correlation  $\langle \zeta_{\mathbf{k}_1} \zeta_{\mathbf{k}_2} \zeta_{\mathbf{k}_3} \cdots \rangle$ :

Extract mass:  $(k \text{ ratios})^{\pm i\mu}$  Chen, YW, 0911.3380

Extract angular momentum:  $P_S(\cos \theta)$  Arkani-Hamed, Maldacena, 1503.08043

Extract parity:  $\text{Im} \langle \cdots \rangle$  Liu, Tong, YW, Xianyu, 1909.01819

Extract width:  $(k \text{ ratios})^{-\frac{3}{2} - \alpha}$  Lu, Reece, Xianyu, 2108.11385

Example: How to extract mass?

# Local Particle Production on Vertices

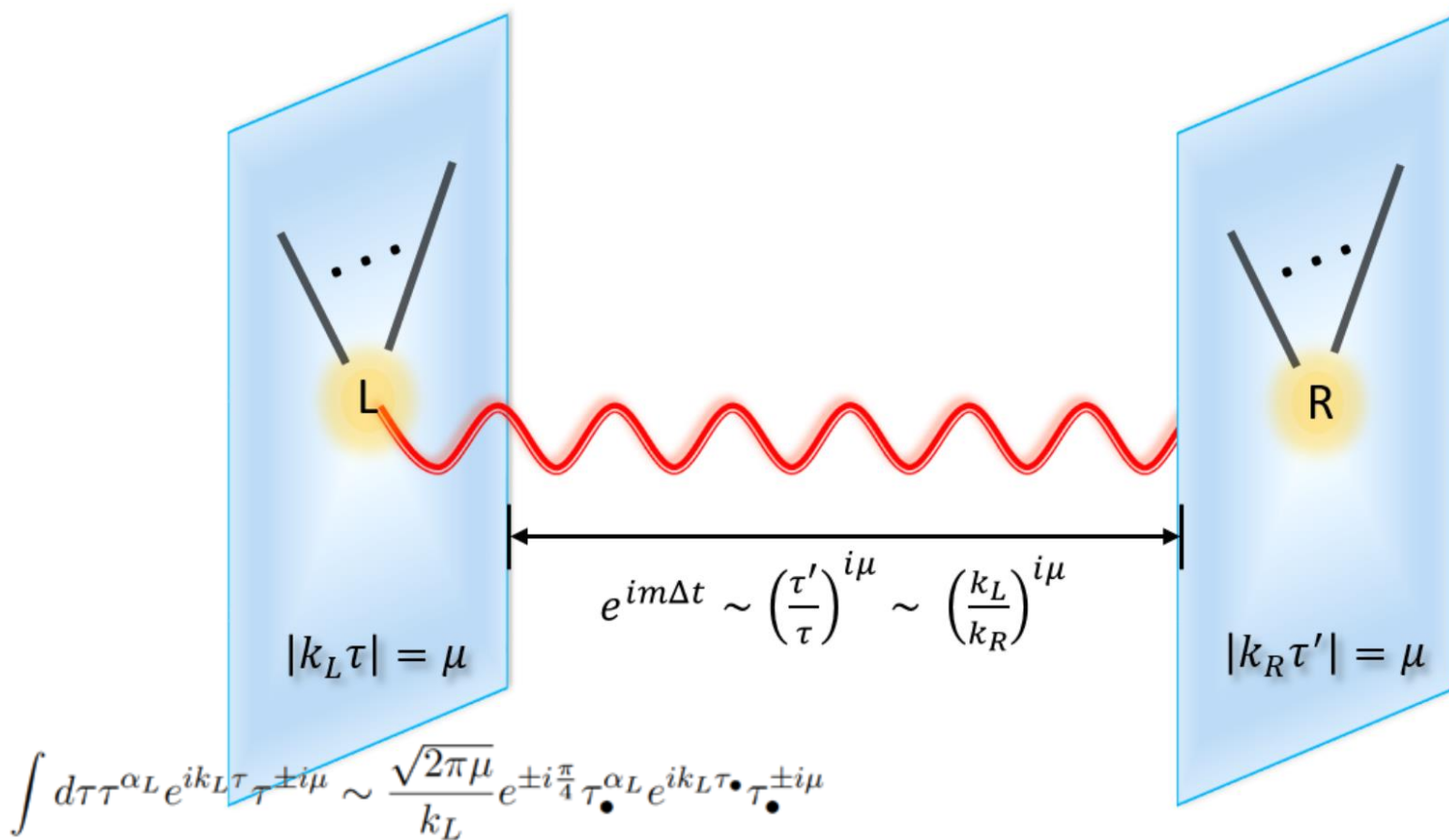


Figure: Tong, Zhu, YW, 2112.03448

See also Qin, Xianyu, 2208.13790

# Non-local Particle Production on Propagators

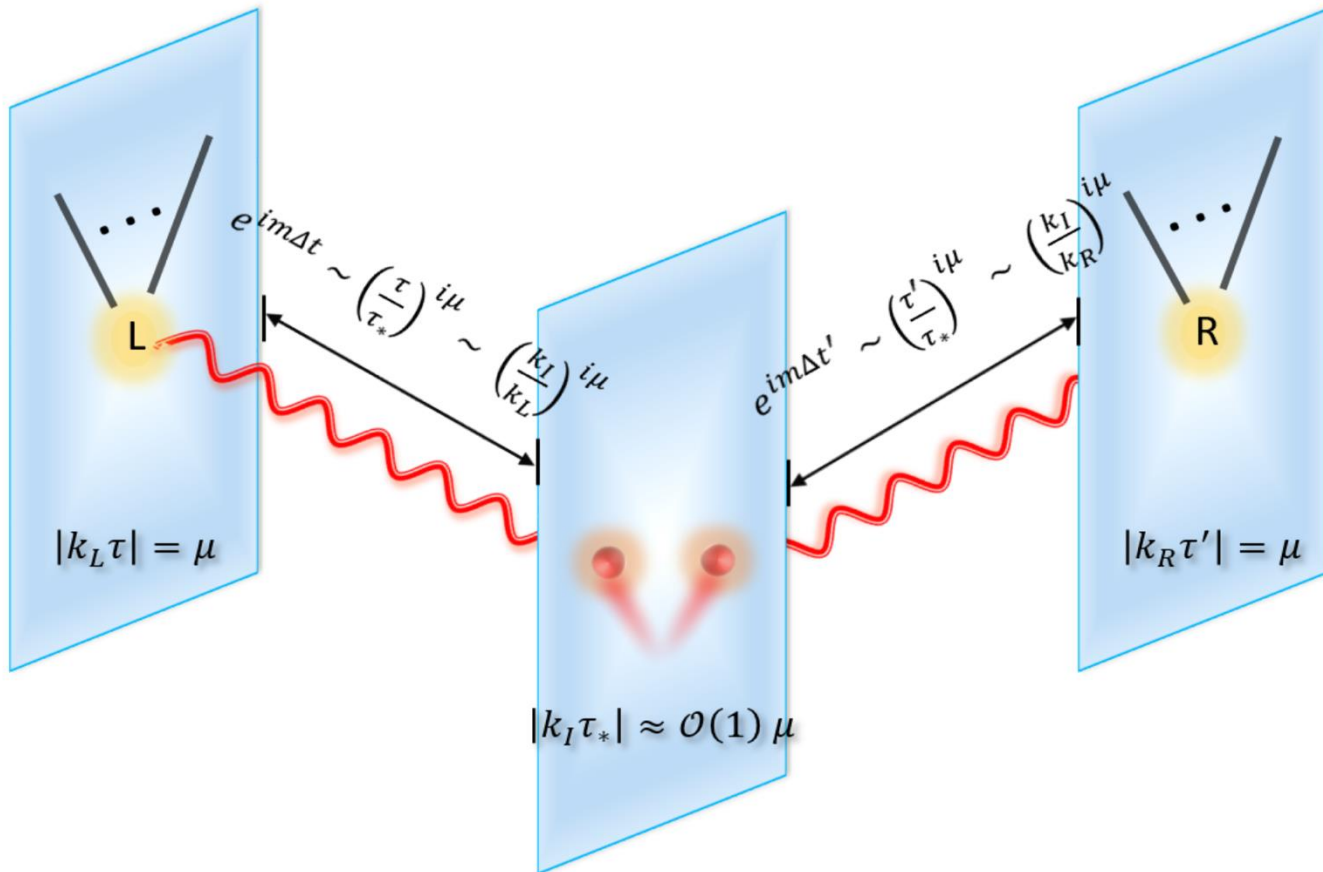


Figure: Tong, Zhu, YW, 2112.03448  
See also Qin, Xianyu, 2208.13790

## Imprints on observables:

- Curvature

Higgs: Lu, YW, Xianyu, 1907.07390

- Isocurvature

Curvaton: Kumar, Sundrum, 1908.11378

Axion: Lu, 2103.05958

- Primordial gravitational waves

YW, Zhu, 2001.03879

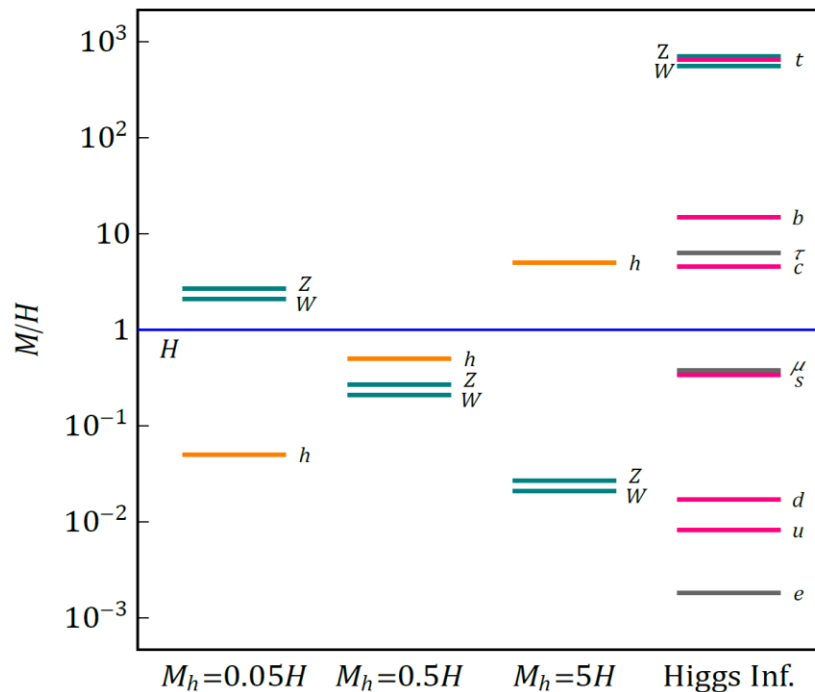


## Plan:

- Heavy particles (✓)
- What can we learn? ( $\leftarrow$  \_  $\leftarrow$ )
- Chemical potential
- Parity violation

# What can we learn?

- Cosmological collider signals
  - Individual particles: mass, spin, lifetime, parity
  - Particle physics models: Standard Model



$$\langle h^2 \rangle \sim H^2$$

$$\lambda h^4 \supset \lambda \langle h^2 \rangle h^2 \sim m_{\text{eff}}^2 h^2$$

Also: possible  $h^2 R \sim H^2 h^2$

## What can we learn?

- Cosmological collider signals
  - Individual particles: mass, spin, lifetime, parity
  - Particle physics models: Standard Model
  - Particle physics models: BSM

SUSY: Baumann, Green, 1109.0292

Neutrino seesaw: Chen, YW, Xianyu, 1805.02656

Grand unification: Kumar, Sundrum, 1811.11200

... ..

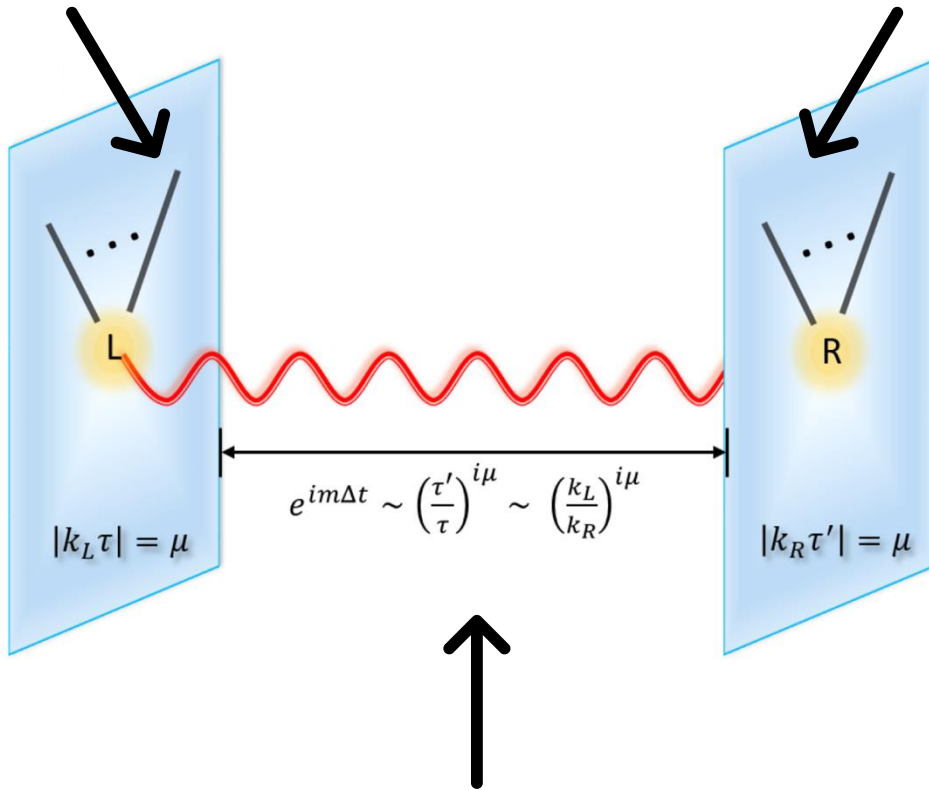
## What can we learn?

- Cosmological collider signals
  - Individual particles: mass, spin, lifetime, parity
  - Particle physics models: Standard Model
  - Particle physics models: BSM
  - A quantum “primordial standard clock”

Classical standard clocks: Chen, 1104.1323, 1106.1635

Quantum standard clocks: Chen, Namjoo, YW, 1509.03930

Light modes: tells conformal time  $\tau$



Heavy modes: tells physical time  $t$

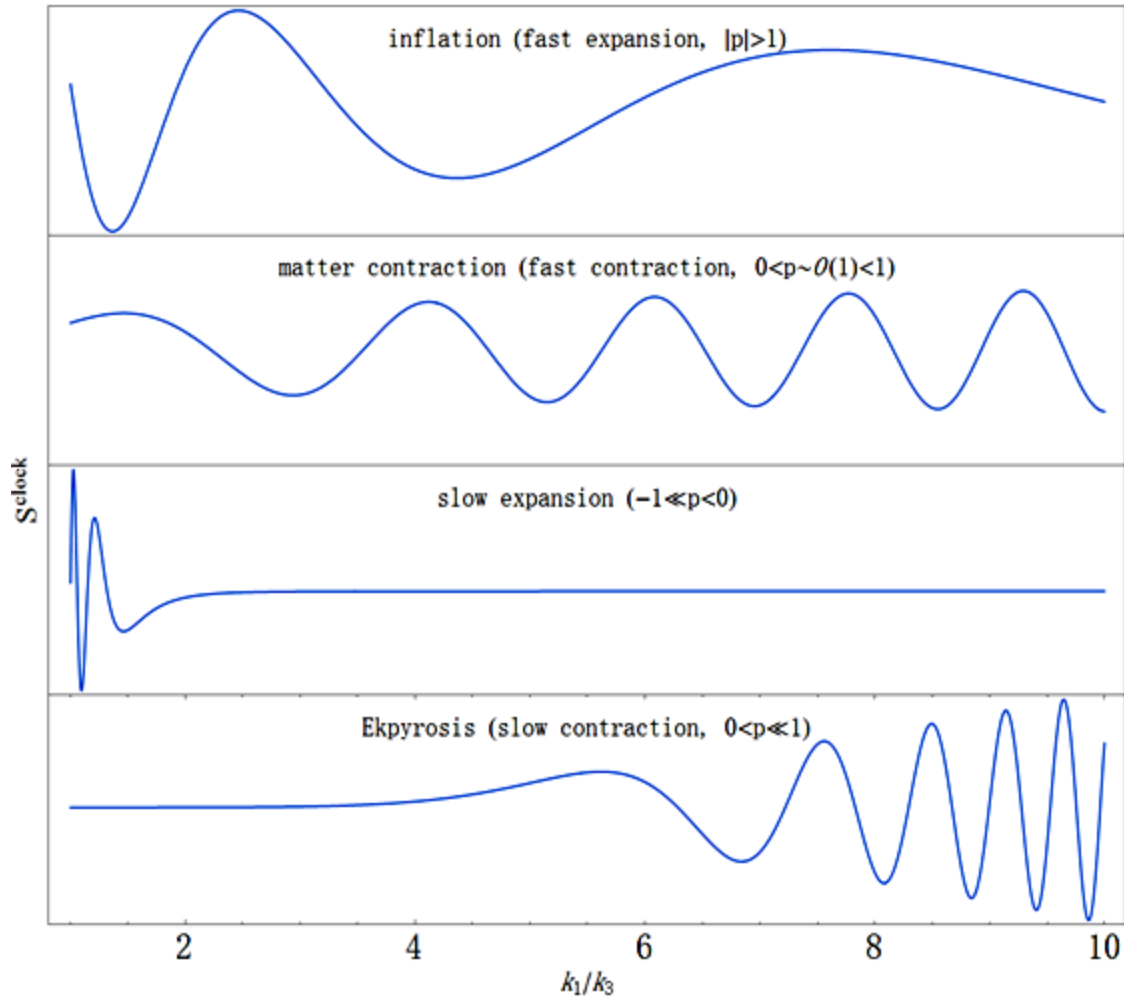
Correlation:

Can tell

$$a(t) = dt/d\tau$$

$$a(t) = a_0 \left( \frac{t}{t_0} \right)^p \int_{\tau_{\text{begin}}}^{\tau_{\text{end}}} d\tau g(t) e^{imt} e^{-iK\tau}$$

$$\rightarrow \sqrt{2\pi} g(t_*) \left( \frac{m}{|H_{k_0}|} \right)^{1/2} K^{-1} \left( \frac{K}{k_0} \right)^{1/2p} \exp \left[ -i \frac{p^2}{1-p} \frac{m}{H_{k_0}} \left( \frac{K}{k_0} \right)^{1/p} \mp i \frac{\pi}{4} \right]$$

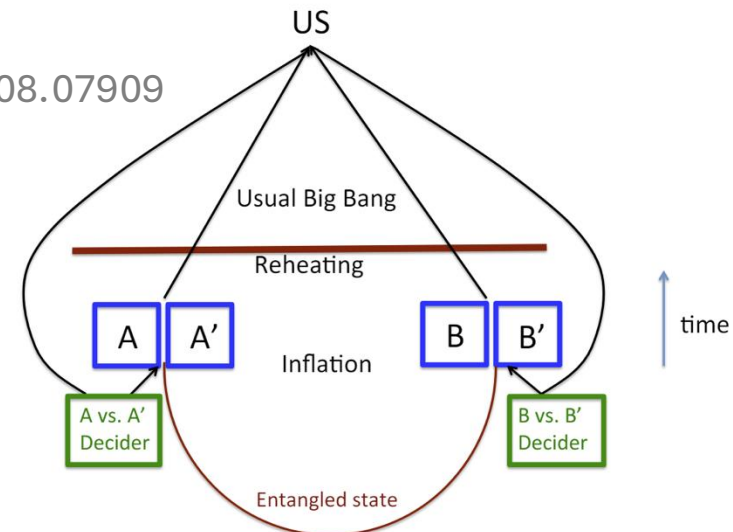


# What can we learn?

- Cosmological collider signals
  - Individual particles: mass, spin, lifetime, parity
  - Particle physics models: Standard Model
  - Particle physics models: BSM
  - A quantum “primordial standard clock”
- Probing quantum mechanics

Bell test: Maldacena, 1508.01082

Against decoherence: Liu, Sou, YW, 1608.07909



## What can we learn?

- Cosmological collider signals
  - Individual particles: mass, spin, lifetime, parity
  - Particle physics models: Standard Model
  - Particle physics models: BSM
  - A quantum “primordial standard clock”
  - Probing quantum mechanics
  - Testing modified gravity (?)

$$ds^2 = -(1 - 2\Phi)dt^2 + (1 - 2\Psi)dx^2 + \dots$$

- Distorted cosmological colliders (?)

$$f(\varphi)\sigma^2$$

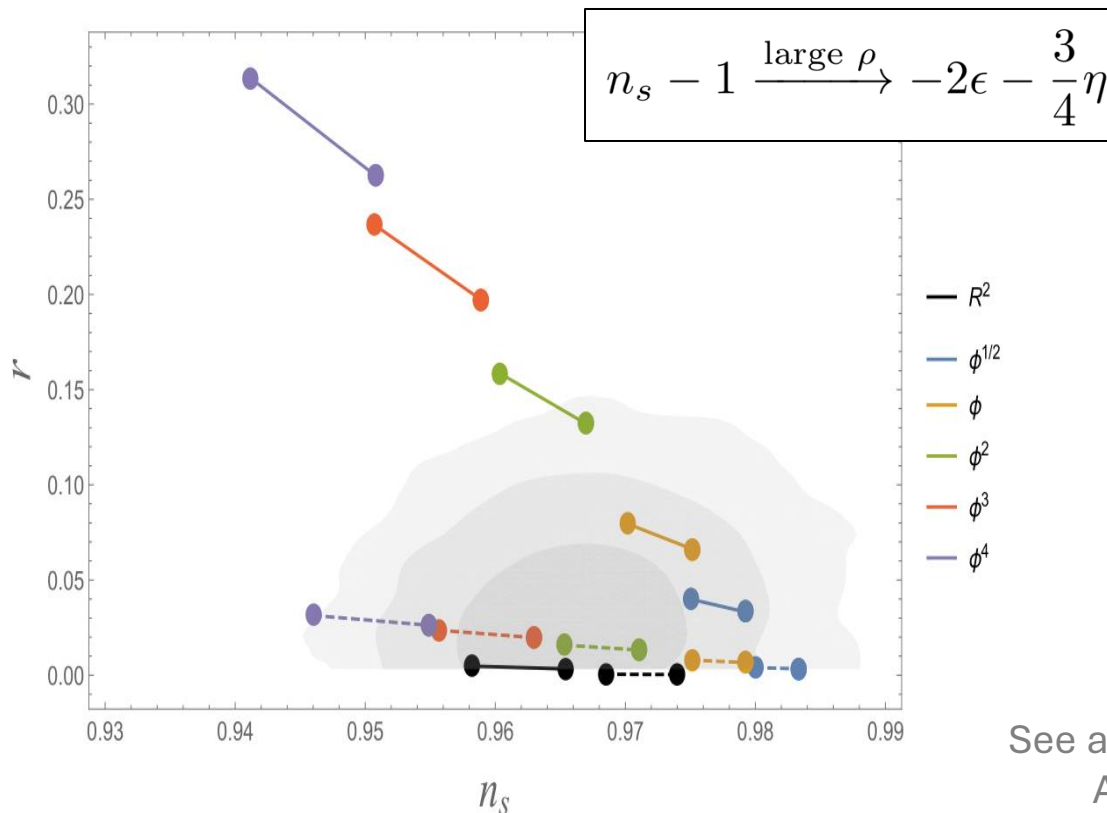
e.g., particle production: J,  $\Gamma$ , P are still valuable

For models like Romano, Sasaki, 0809.5142, Barnaby, Huang, 0909.0751



# What can we learn?

- Cosmological collider signals
- By-products:
  - Corrections on the  $n_s$ - $r$  diagram



$\delta\sigma$  is additional source

→ Enhance  $P_\zeta$

→ Lower  $H$

→ Lower  $r$

→ Saves  $m^2\varphi^2$

Jiang, YW, 1703.04477

Tong, YW, Zhou, 1708.01709

See also Achucarro, Atal, Welling 1503.07486

An, McAneny, Ridgway, Wise 1706.09971

## What can we learn?

- Cosmological collider signals
- By-products:
  - Corrections on the  $n_s$ - $r$  diagram
  - More careful abundance for WIMPZILLAs

Li, Nakama, Sou, YW, Zhou, 1903.08842

Li, Lu, YW, Zhou, 2002.01131

For WIMPZILLAs, see, e.g., Kolb, Chung, Riotto, hep-ph/9810361

# What can we learn?

## - Cosmological collider signals

Individual particles, SM, BSM

As a clock to measure the expansion (or even contraction?) history

Quantum effects, free from decoherence

Testing modified gravity (?)

Cosmological colliders with distorted mass (?)

## - By-products

$n_s$ - $r$  diagram; WIMPZILLAs

## - Assisted the advances of tools

bootstrap, Mellin-Barnes, spectral decomposition, dispersion relation...

Larger signals? Especially removing  $e^{-\pi m/H}$  for  $m \gg H$ ?

- Classical oscillation or periodic potential?  $\frac{H^2}{\dot{\phi}} \approx \frac{1}{3600}$

Chen 1104.1323; Chen, Namjoo, Wang: 1411.2349;

Flauger et al, 1606.00513; Chen, Ebadi, Kumar, 2205.01107

- Warm inflation? Tong, YW, Zhou, 1801.05688
- Distorted signals with varying mass?
- Chemical potential  
Chen, YW, Xianyu 1805.02656  
Wang, Xianyu 1910.12876  
Sou, Tong, YW, 2104.08772

## Plan:

- Heavy particles (✓)
- What can we learn? (✓)
- Chemical potential ( $\leftarrow$  \_  $\leftarrow$ )
- Parity violation

EFT power counting:

How to couple inflaton to other fields?

Nice to keep the inflaton shift symmetry:  $\partial_\mu \phi$ .

Leading: dim-5:  $\frac{1}{\Lambda} (\partial_\mu \phi) \bar{\psi} \gamma^5 \psi, \frac{1}{\Lambda} \phi F \tilde{F}$

(see also the scalar mixings & interactions  $\partial\phi\partial\sigma, \sigma(\partial\phi)^2$ )

Note:  $c_s$  is higher dimensional, though historically studied more

# Chemical potential: $\mu N \rightarrow \partial_\mu \phi J^\mu$

Long history in cosmology, e.g., Spontaneous Baryogenesis, Cohen, Kaplan, 1988

$$e^{-\pi m/H} \rightarrow \text{subsidy by } N \rightarrow e^{-\pi(m-\mu)/H}$$

Chen, YW, Xianyu, 1805.02656  
Wang, Xianyu, 1910.12876

## Various spins:

- **Scalar**  $\mathcal{L} = [(\partial_t + i\mu)\Phi^*][(\partial_t - i\mu)\Phi] - |\partial_i\Phi|^2 - m^2|\Phi|^2$

Can be rotated away by  $\Phi \rightarrow e^{i\mu t} \Phi$  (Wang, Xianyu, 1910.12876)

Symmetry breaking (Bodas, Kumar, Sundrum, 2010.04727)

- **Spinor**  $\frac{1}{\Lambda}(\partial_\mu\phi)\bar{\Psi}\gamma^5\gamma^\mu\Psi$  Chen, YW, Xianyu, 1805.02656; Wang, Xianyu, 1910.12876

- **Vector**  $-\phi F\tilde{F}/(4\Lambda)$  Lu, YW, Xianyu, 1907.07390; Wang, Xianyu, 1910.12876

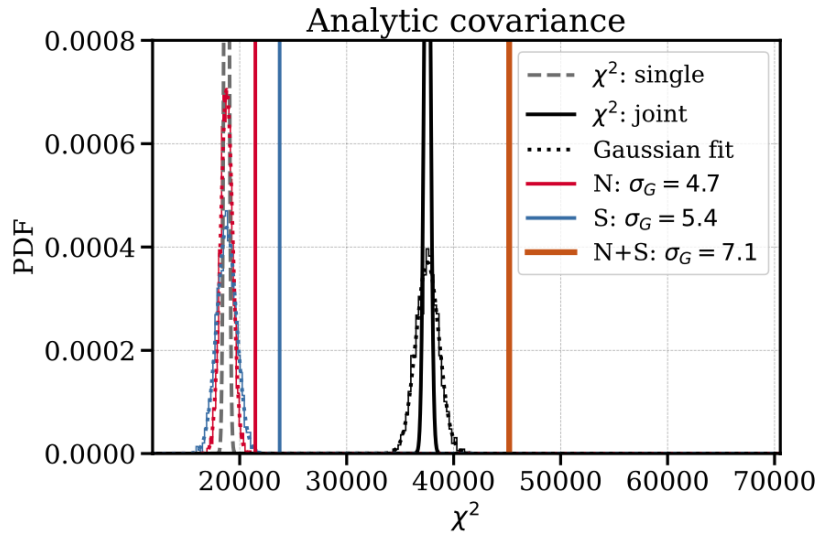
- **Tensor**  $\frac{1}{4\Lambda^3}\phi W_{\mu\nu\rho\sigma}\tilde{W}^{\mu\nu\rho\sigma}$  Wang, Xianyu, 1910.12876; Tong, Xianyu, 2203.06349

## Plan:

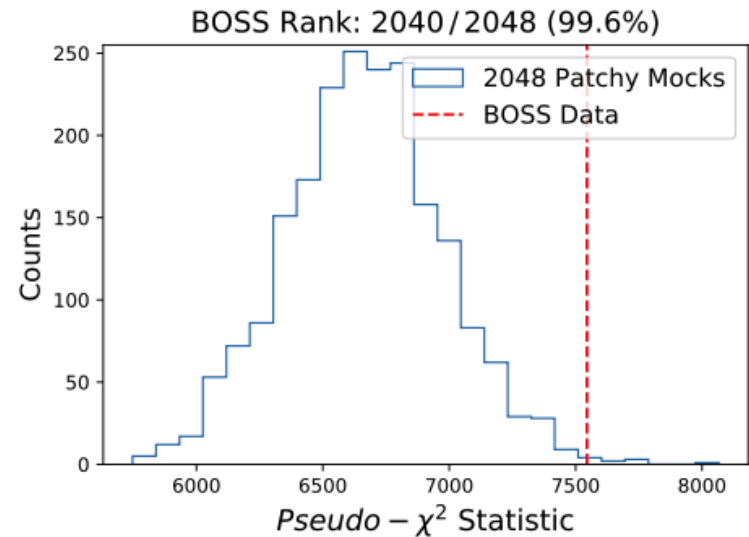
- Heavy particles (✓)
- What can we learn? (✓)
- Chemical potential (✓)
- Parity violation ( $\leftarrow$  \_  $\leftarrow$ )



# Parity violation in the trispectrum?



7.1 $\sigma$ : Hou, Slepian, Cahn, 2206.03625



2.9 $\sigma$ : evidence Philcox, 2206.04227



1.4 $\sigma$ : Philcox, Ereza, 2401.09523

(Improved estimation of covariance)

New method + attention = standard observable in new data?



## The way of Parity —

Chiral fermions, such as only left-handed neutrinos:

Cosmological collider primordial bispectrum: Chen, YW, Xianyu 1805.02656 => trispectrum?

Wenqi Yu, YW, to appear

## The way of CP —

CKM-like mixing matrices: Pinol, Aoki, Renaux-Petel, Yamaguchi 2112.05710 => trispectrum?

Axion-like couplings: Liu, Tong, YW, Xianyu, 1909.01819

How cosmological collider naturally violates parity?

Dimension counting of interactions,

preserving the shift symmetry of inflaton: starting from dim 5

- Scalar interaction:  $\sigma(\partial\varphi)^2$
- Vector interaction:  $\varphi F\tilde{F}$
- Spinor interaction:  $(\partial\varphi)\bar{\Psi}\gamma^5\gamma^\mu\Psi$

For vector and spinor:

Large signal (chemical potential) + parity violating

# Requirements for parity violating trispectrum

(1) Source of CP violation

(2) Four-point correlation needed

(if observing momenta only)

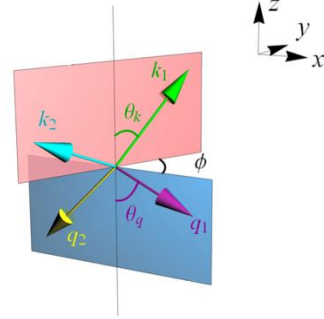
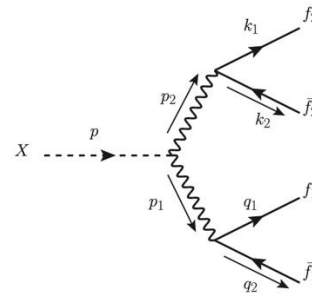
- Decay-plane correlation

- 3 sets of independent momenta

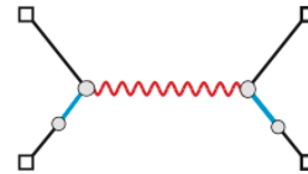
- For 3pt, 3D rot makes P trivial

(3) Difference between  $k_1$  &  $k_2$  needed

(4) Imaginary part in  $k$ -space  $\langle \zeta(\vec{k}_1) \dots \zeta(\vec{k}_4) \rangle - \langle \zeta(-\vec{k}_1) \dots \zeta(-\vec{k}_4) \rangle$



$$\epsilon^{ijk} (q_1 + q_2)_i (q_1 - q_2)_j (k_1 - k_2)_k$$



## No-go theorem: No parity violation if all below

- Massless
- Scale invariant
- IR finite
- BD initial condition
- Local
- Tree level

$$\begin{aligned}\tilde{\psi}_4 &\propto i\epsilon^{(3)}(ik_i)^{3+2m} \int_{-\infty}^0 d\tau a^{1-2m-n} \partial_\tau^n G_+^4 \\ &\propto \epsilon^{(3)}(k_i)^3 \times (k_j)^{2m} \int_{-\infty}^0 d\tau \tau^{2m+n-1} \partial_\tau^n u^{*4} \\ &\propto \epsilon^{(3)}(k_i)^3 \times (k_j)^{2m} \int_{-\infty}^0 d\tau \tau^{2m+n-1} \partial_\tau^n \left(1 - k \frac{\partial}{\partial k}\right)^4 e^{iK\tau} \\ &\propto \epsilon^{(3)}(k_i)^3 \times (k_j)^{2m} \left(1 - k \frac{\partial}{\partial k}\right)^4 \int_{-\infty}^0 d\tau \tau^{2m+n-1} (iK)^n e^{iK\tau} \\ &\propto \epsilon^{(3)}(k_i)^3 \times (k_j)^{2m} \left(1 - k \frac{\partial}{\partial k}\right)^4 (iK)^n (-1)^{n+1} (iK)^{-2m-n} \Gamma(2m+n) \\ &\propto \epsilon^{(3)}(k_i)^3 \times (k_j)^{2m} \left(1 - k \frac{\partial}{\partial k}\right)^4 K^{-2m} \in \mathbb{R} .\end{aligned}$$

Liu, Tong, YW, Xianyu, 1909.01819

More elegant proof from unitarity, etc., see

Cabass, Jazayeri, Pajer, Stefanyszyn, 2210.02907

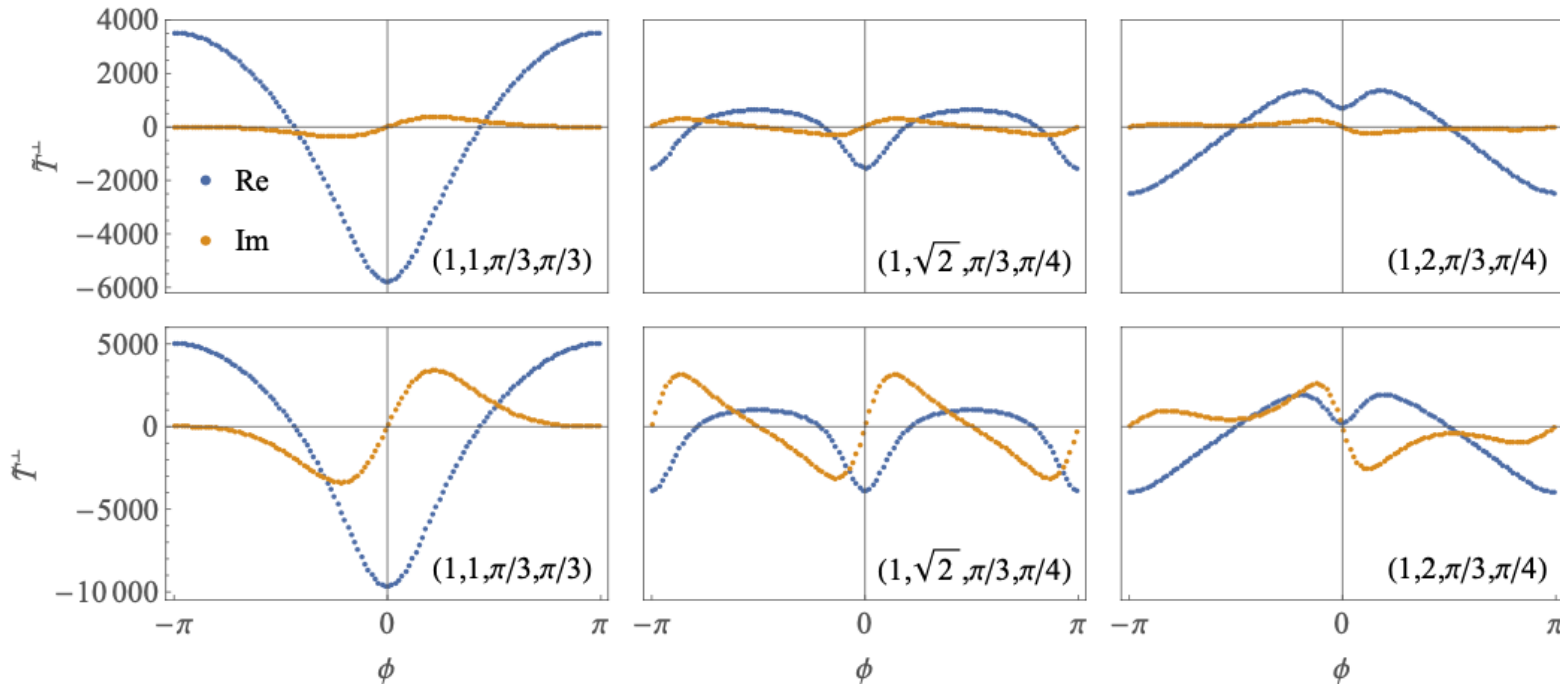
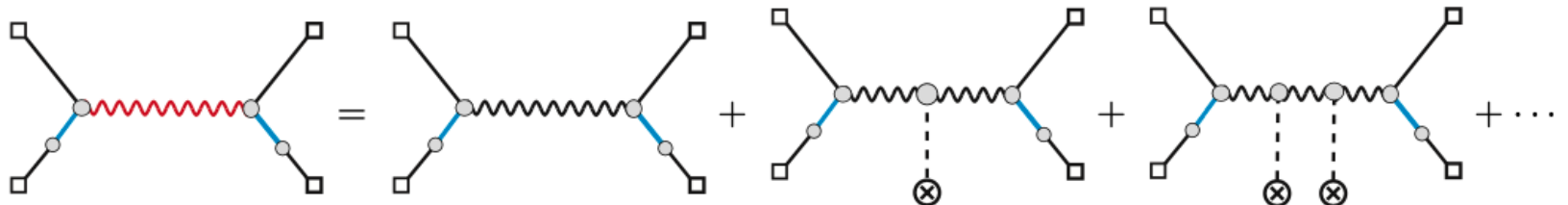
See also

Jazayeri, Renaux-Petel, Tong, Werth, Zhu, 2308.11315

Stefanyszyn, Tong, Zhu, 2309.07769

# Example: Z, Higgs and the inflaton

$$\Delta\mathcal{L}_1 = \frac{\rho_{1,Z}}{\dot{\phi}_0} h \partial_\mu \varphi Z^\mu, \quad \Delta\mathcal{L}_2 = -\rho_2 \dot{\phi} h, \quad \Delta\mathcal{L}_3 = -\frac{c_0}{4} \theta(t) Z_{\mu\nu} Z_{\rho\sigma} \mathcal{E}^{\mu\nu\rho\sigma}$$



# Thank you!

## Summary:

- Heavy particles

mass, spin, lifetime, parity

- What can we learn?

Signal: SM, BSM, dark matter, distorted (?)

Test: standard clock, quantum mechanics, modified gravity (?)

By-product: shift the  $n_s - r$  diagram

- Chemical potential

scalar (\*), spinor, vector, tensor

- Parity violation

observation, requirement, no-go theorem, vector example