

# Lecture 4:

## Variational principle for fluid dynamics

# Variational principle for perfect fluid

## Variational principle for the dynamics of the perfect fluid

- has a long history
- is nontrivial
- has many versions

## The geometric version we adopt here

- is close to the one used by Carter and by Arnold's school
- involves the use of “restricted variations”
- suitable for both relativistic and nonrelativistic hydro

.

# Action for a barotropic fluid

**Action** for barotropic fluid dynamics is a spacetime integral of fluid pressure

$$S[\pi_\nu] = \int d^4x P(\pi_\nu), \quad (67)$$

where pressure is understood as a function of fluid 4-momentum  $\pi_\nu$ .

**Admissible variations.** To obtain equations of motion one should vary action over  $\pi_\nu$  but only as

$$\delta\pi_\nu = \epsilon^\mu \partial_\mu \pi_\nu + \pi_\mu \partial_\nu \epsilon^\mu, \quad (68)$$

where  $\epsilon^\mu$  is an arbitrary 4-vector field.

This prescription is the variation as a Lie derivative of a one-form along  $\epsilon$ , that is  $\delta\pi = \mathcal{L}_\epsilon\pi$ .

# Equations of motion from action

## Exercise 22

Vary the action (67) as prescribed by (68) and obtain the equation of motion

$$J^\nu (\partial_\nu \pi_\mu - \partial_\mu \pi_\nu) + \pi_\mu \partial_\nu J^\nu = 0, \quad (69)$$

where we defined

$$J^\mu \equiv -\frac{\delta S}{\delta \pi_\mu} = -\frac{\partial P}{\partial \pi_\mu}. \quad (70)$$

## Exercise 23

Show that (69) leads to

$$\partial_\mu J^\mu = 0, \quad (71)$$

$$J^\nu (\partial_\nu \pi_\mu - \partial_\mu \pi_\nu) = 0. \quad (72)$$

*Hint: multiply (69) by  $J^\mu$ .*

# Admissible variations

Vary the action  $S[\pi]$  over admissible variations

$$\delta\pi_\nu = \epsilon^\mu \partial_\mu \pi_\nu + \pi_\mu \partial_\nu \epsilon^\mu, \quad \delta\pi = \mathcal{L}_\epsilon \pi$$

and obtain

$$\delta S = \int d^4x \frac{\delta S}{\delta \pi_\nu} (\epsilon^\mu \partial_\mu \pi_\nu + \pi_\mu \partial_\nu \epsilon^\mu) = 0.$$

Introducing “particle current”  $\mathcal{J}^\nu \equiv -\frac{\delta S}{\delta \pi_\nu}$

$$\mathcal{J}^\nu (\partial_\nu \pi_\mu - \partial_\mu \pi_\nu) + \pi_\mu \partial_\nu \mathcal{J}^\nu = 0.$$

Multiply by  $\mathcal{J}^\mu$

$$\mathcal{J}^\nu (\partial_\nu \pi_\mu - \partial_\mu \pi_\nu) = 0, \quad \partial_\nu \mathcal{J}^\nu = 0.$$

Notice, that the conservation of  $\partial_\nu \mathcal{J}^\nu = 0$  follows from diffeos.

## A few remarks

- For a given fluid we have  $P(\mu)$  – equation of state
- For a given space time symmetry we express  $\mu$  as a scalar made out of  $\pi_\mu$ .

$$\mu = -\pi_0 - \frac{\pi_i^2}{2}, \quad \text{- Galilean,} \quad (73)$$

$$\mu = \sqrt{-\pi_\mu^2}, \quad \text{- relativistic.} \quad (74)$$

- Equations are the same for  $S[\pi_\nu - A_\nu]$  as long as variation is done only over  $\pi_\mu$ .
- One can generalize the action by adding derivative terms. As long as  $S = S[\pi_\nu]$  the derivation holds.

# Conserving forms of the equations of motion

Variational principle gives equations of motion

$$\begin{aligned}\partial_\mu T_\nu^\mu &= 0, & T_\nu^\mu &= J^\mu \pi_\nu + \delta_\nu^\mu P, \\ \partial_\mu J^\mu &= 0, & J^\mu &= -\frac{\partial P}{\partial \pi_\mu},\end{aligned}$$

Identify:

- $T_\nu^\mu$  - energy-momentum tensor
- $J^\mu$  - charge/number current
- $\pi_\mu$  - specific momentum
- $P$  - pressure

# Energy-momentum tensor directly from the action

## Exercise 24\*

*Introduce metric  $g_{\mu\nu}$  into the action (67). Vary the obtained action over the metric and derive the form of the energy momentum tensor.*

*Hint: the metric enters into the measure of integration but also in a formula for a chemical potential in terms of  $\pi_\mu$ . It is easier to do this exercise in relativistic case.*

# Lecture 5:

## Anomalous action for fluid dynamics

# Adler-Bell-Jackiw Anomaly

Massless Dirac fermions in 3+1:

$$Z = \int D\psi e^{i \int d^4x \mathcal{L}}, \quad \mathcal{L} = \bar{\psi} i \not{\partial} \psi,$$

Global symmetries of classical Lagrangian  $\mathcal{L}$  with corresponding Noether currents (vector and axial)

$$\psi \rightarrow e^{i\lambda + i\gamma^5 \tilde{\lambda}} \psi, \quad j^\mu = \langle \bar{\psi} \gamma^\mu \psi \rangle, \quad j_A^\mu = \langle \bar{\psi} \gamma^\mu \gamma^5 \psi \rangle.$$

However, the complete quantum theory  $Z$  does NOT have both symmetries. For example, in the presence of the vector gauge field

$$\mathcal{L} = \bar{\psi} (i \not{\partial} - A) \psi,$$

we have

$$\partial_\mu j^\mu = 0, \quad \partial_\mu j_A^\mu = 2\mathbf{E} \cdot \mathbf{B} \frac{e^3}{4\pi^2 \hbar^2 c^2}.$$

## Mixed $U(1)_V \times U(1)_A$ anomaly

- 3+1 massless Dirac fermion has  $G = U(1)_V \times U(1)_A$  global symmetry
- Gauging this symmetry, i.e., introducing dynamic gauge fields coupled to  $G$ -currents makes theory inconsistent - anomaly
- In the presence of background gauge field the theory is consistent but charges do not conserve - mixed t'Hooft anomaly
- This anomaly can be interpreted as a system's coupling to a charge reservoir (5d topological insulator, 4d spectators ...)
- For any system with global symmetry one can ask whether there is an anomaly

# Motivation: role of chiral anomaly in fluids

- Quantum anomalies play important role in QFT with Dirac fermions
- Sometimes, Dirac fermions behave as fluids
- What are the effects of the anomalies on hydrodynamics?
- Pioneered by:  
Alekseev, Cheianov, Fröhlich; Son, Surowka; Haehl, Loganayagam, Rangamani; ...
- Hydrodynamics (bosonization) captures anomalies in 1D. What are anomalies in conventional 3D (Euler) fluids?
- Geometrical fluid dynamics:  
Lichnerowicz, Carter, Arnold, Marsden, Holm, ...
- Generalization of some of QHE physics to 4+1 dimensions with 3+1 dimensional boundary

# Geometric transport

The Lichnerowicz-Carter equation:

$$\mathcal{J}^\nu(\partial_\nu\pi_\mu - \partial_\mu\pi_\nu) = 0, \quad \text{or in form notation} \quad i_{\mathcal{J}}d\pi = 0.$$

Then

$$i_{\mathcal{J}}(d\pi \wedge d\pi) = 2(i_{\mathcal{J}}d\pi) \wedge d\pi = 0.$$

But  $d\pi \wedge d\pi$  is a 4-form and is proportional to the volume form and should be zero. Or we can conclude that  $d\pi$  has a rank 2 and, we have a fundamental property of barotropic fluid

$$d\pi \wedge d\pi = 0.$$

In components:

$$\partial_\mu(\epsilon^{\mu\nu\lambda\rho}\pi_\nu\partial_\lambda\pi_\rho) = 0.$$

# Chiral current

In external e/m field  $S[\pi] \rightarrow S[\pi - A]$  but still

$$\partial_\mu(\epsilon^{\mu\nu\lambda\rho}\pi_\nu\partial_\lambda\pi_\rho) = 0.$$

However, the helicity density is NOT gauge invariant. Introduce **the chiral current**:

$$J_A^\mu = \epsilon^{\mu\nu\lambda\rho}(\pi_\nu - A_\nu)\partial_\lambda(\pi_\rho + A_\rho) = \epsilon^{\mu\nu\lambda\rho}p_\nu(\partial_\nu p_\rho + F_{\nu\rho}), \quad p_\mu \equiv \pi_\mu - A_\mu.$$

The chiral current is gauge invariant but it is NOT conserved!

$$\partial_\mu J_A^\mu = 2\mathbf{E} \cdot \mathbf{B},$$

$$d(\pi - A) \wedge d(\pi + A) = -dA \wedge dA.$$

**Axial/chiral anomaly!**

AGA, P.B. Wiegmann, PRL 128, 054501 (2022) [arXiv:2110.11480]

## One-dimensional action (reminder)

In 1+1 dimensions with  $\pi_\mu = \partial_\mu \phi$  we had

$$S_{1+1} = \int d^2x P(\pi - A) + \tilde{A} \wedge (\pi - A) - \int_{M_3} \tilde{A} \wedge dA \quad (75)$$

and as a consequence

$$\partial j = [d\tilde{A}], \quad \partial \tilde{j} = -[dA], \quad (76)$$

$$j^\mu = -\partial P / \partial \pi_\mu, \quad \tilde{j} = [\pi - A]. \quad (77)$$

Can we introduce  $\tilde{A}$  into the 3+1 action similarly?

## 3+1 action with axial gauge field

In 3+1 dimensions

$$S_{3+1} = \int d^4x P(\pi - A) + \tilde{A} \wedge (\pi - A) \wedge d(\pi + A) - \int_{M_5} \tilde{A} \wedge dA \wedge dA. \quad (78)$$

Varying over  $\pi$  by diffeos (admissible variations) we derive

$$\partial j = -2[dA \wedge d\tilde{A}], \quad \partial \tilde{j} = -[dA \wedge dA], \quad (79)$$

$$j^\mu = -\partial P / \partial \pi_\mu + [(\pi - A) \wedge d\tilde{A}], \quad \tilde{j} = [(\pi - A) \wedge d(\pi + A)], \quad (80)$$

and

$$\partial_\nu T_\mu^\nu = F_{\mu\nu} j^\nu + \tilde{F}_{\mu\nu} \tilde{j}^\nu, \quad (81)$$

$$T_\mu^\nu = J^\nu p_\mu + \delta_\mu^\nu P. \quad (82)$$

These currents and anomalies are known as covariant current and covariant anomalies.

However, this is not the most general form of anomalies.

# Generalization

In 3+1 dimensions

$$S_{3+1}[\pi, \psi] = \int d^4x P(\pi - A) + \tilde{A} \wedge (\pi - A) \wedge d(\pi + A) + \psi \left[ d\pi \wedge d\pi + \alpha dA \wedge dA \right] \quad (83)$$

$$- \int_{M_5} \tilde{A} \wedge dA \wedge dA + \alpha \tilde{A} \wedge d\tilde{A} \wedge d\tilde{A}. \quad (84)$$

Varying over  $\pi$  and  $\psi$  **freely**, we obtain

$$\partial j = -2[dA \wedge d\tilde{A}], \quad \partial \tilde{j} = -[dA \wedge dA] - 3\alpha[d\tilde{A} \wedge d\tilde{A}]. \quad (85)$$

## Exercise 25\*

*Show that this action is axial and vector gauge invariant. Derive expressions for  $j, \tilde{j}, T$  for this action.*

*Hint: see arXiv:2403.12360*

## Further generalization

In the curved spacetime

$$\begin{aligned} S_{3+1}[\pi, \psi] = & \int d^4x \sqrt{g} P(\pi - A) + \tilde{A} \wedge (\pi - A) \wedge d(\pi + A) \\ & + \psi \left[ d\pi \wedge d\pi + \alpha dA \wedge dA + \beta \operatorname{tr}(R \wedge R) \right] \\ & - \int_{M_5} \tilde{A} \wedge dA \wedge dA + \alpha \tilde{A} \wedge d\tilde{A} \wedge d\tilde{A} + \beta \tilde{A} \wedge \operatorname{tr}(R \wedge R). \end{aligned} \quad (86)$$

Varying over  $\pi$  and  $\psi$  **freely**, we obtain mixed axial-gravitational anomaly

$$D_\nu j^\nu = -2[dA \wedge d\tilde{A}], \quad D_\nu \tilde{j}^\nu = -\left[ dA \wedge dA + 3\alpha[d\tilde{A} \wedge d\tilde{A} + \beta \operatorname{tr}(R \wedge R)] \right]. \quad (87)$$

### Exercise 26\*

*Derive expressions for  $j, \tilde{j}, T$  for this action.*

*Hint: see arXiv:2403.12360*

## Conclusion: five lectures in three sessions!

- 1 Lecture 1: Introduction to Euler fluid dynamics
- 2 Lecture 2: Introduction to chiral anomaly
- 3 Lecture 3: Covariant form of Euler equation
- 4 Lecture 4: Variational principle for fluid dynamics
- 5 Lecture 5: Anomalous action for fluid dynamics

- + many more details in [arXiv:2403.12360](https://arxiv.org/abs/2403.12360) and references therein
- + many generalizations producing nondissipative fluids with anomalies and assortments of conservation laws
- + dissipative versions using similar tricks but in Keldysh formalism (work in progress)