

Acceleration and Cooling Processes in Outburst Phase of the TeV Blazar W Com

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Curved broad-band spectral distributions of non-thermal sources like blazars are described well by a log-parabolic law where the spectra can be obtained for relativistic electrons by means of a statistical acceleration mechanism whose probability of acceleration depends on energy. In this work we used the curvature parameter of the synchrotron spectra (b) and the peak energy (E_p) to investigate the acceleration and cooling mechanism during outburst phase of the TeV blazar W com. We carried out a detailed time-resolved temporal and spectral study using multi-wavelength data taken by *Swift* and *XMM-Newton* satellites. Similar spectral variation of the main two emission components supports the SSC scenario. During the event, the source showed a significantly positive E_p - b relation, which, incorporating previous theoretical predictions, likely be associated with a magnetic field-driven stochastic process in the jet.

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