# First Order Electroweak Phase Transition from sub-GeV Physics and Possible Link to Flavor

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Based on: H.D., 2101.05319

# Introduction:

• A mostly "dark" Universe of unknown physics Dark Energy: ∼ 68% ; Dark Matter: ∼ 27%

- "Visible" Universe: ∼ 5%
- Known physics (atoms)
- 
- Matter dominates over anti-matter
- Lack of cosmic annihilation signals, CMB, . . .
- Baryon asymmetry established after inflation
- Need a *baryogenesis* mechanism





**PDG** 

• Baryogenesis generally requires three Sakharov conditions:

Sakharov, 1967

- (i) Baryon number violation
- (ii) C and CP violation
- (iii) Departure from equilibrium
- SM does not provide sufficient levels of (ii) and (iii)
- Baryon asymmetry strongly motivates physics beyond SM (BSM)
- Sakharov condition (iii): needed to avoid erasing the asymmetry
- SM could have supplied (iii) if EW Phase Transition (EWPT) was first order
- Yet the measured Higgs mass  $\approx$  125 GeV too large, implies this was not the case
- Motivates BSM in order to have an EW First Order Phase Transition (FOPT)
- It is typically assumed that EW FOPT requires

Critical temperature  $T_c$ , Higgs vev  $v$ 





JodyDole/GettyImages

• SM: thermal contributions of  $W, Z$  gauge bosons  $E.g.,$  Quiros, 1999

$$
\frac{v(T_c)}{T_c} = \frac{2m_W^3 + m_Z^3}{3\pi\lambda_H v^3} \approx 0.1
$$

• SM Higgs quartic coupling  $\lambda_H \approx 0.13$  for  $m_H \approx 125$  GeV

 $v = 246$  GeV,  $\langle H \rangle = v/\sqrt{2}$ 

• Since  $m_W, m_Z \propto v$ , for  $\lambda_H \to \sim \lambda_H/10$ , a FOPT with  $m_H \sim 40$  GeV

Assuming SM gauge couplings

- There have been numerous proposals for realizing an EW FOPT via BSM
- Adding new weak scale states with  $\mathcal{O}(1)$  coupling to the Higgs
	- Thermal effects suppressed for states with masses  $\gg \langle H \rangle$
- Addition of  $(H^{\dagger}H)^{(2+n)}/\Lambda^n$  operators;  $\Lambda \sim 1\,$  TeV
	- Allowing for deviations of  $\lambda_H$  to allow a FOPT
- Typically, expect states near weak scale and significant coupling to Higgs

See, e.g., Ramsey-Musolf, 1912.07189, and references therein

This talk<sup>\*</sup>: a different approach, using sub-GeV physics weakly coupled to Higgs

- Basic idea:  $\lambda_H$  starts small before EWPT, grows to its SM value afterwards
- Stability of our model near weak scale suggests small top Yukawa before EWPT
- Conjecture: vanishing Yukawa couplings for all SM fermions before EWSB
- Other works considering alternative approaches include
- Using weakly coupled intermediate mass ≥ 10 GeV scalars: Jeong, Jung, Shin, 1806.02591; Kozaczuk, Ramsey-Musolf, Shelton, 1911.10210
- Via axion-like particles: Jeong, Jung, Shin, 1811.03294

\* We will not specify a baryogenesis mechanism, but consider FOPT as a potential key ingredient

# **A Model:** H.D., 2101.05319

$$
V(\phi, H, \eta) = \frac{m_{0\phi}^2}{2} \phi^2 - (\mu_0^2 + 2\mu \phi) H^{\dagger} H + (\lambda_0 + 2\frac{\phi^2}{M^2}) (H^{\dagger} H)^2 + \frac{\kappa}{4} \phi^2 \eta^2
$$

- $\bullet$   $m_{0\phi}^2 >$  0,  $\mu_0^2 >$  0,  $\lambda_0 >$  0, and 0  $< \mu \ll m_H$
- Scale  $M \gg m_H$  descends from some ultraviolet (UV) dynamics
- Scalar  $\eta$  coupled to  $\phi$  with  $0 < \kappa \ll 1$ , to allow a thermal "slow-roll" for  $\phi$
- $\eta$  is initially massless
- Potential V consistent with a softly broken  $\mathbb{Z}_2$  acting on  $\phi$
- Vacuum solutions with  $\partial_h V = \partial_\phi V = \partial_\eta V = 0$ ; we will set  $\langle \eta \rangle = 0$

 $H\to (0,v+h)^T/$ √ 2, with  $\langle h \rangle = v = 246$  GeV

• Background value of  $\phi$  at  $T=0$ 

$$
\bar{\phi} = \frac{\mu v^2 M^2}{m_{0\phi}^2 M^2 + v^4}
$$

Induced by the tadpole term  $\propto \phi H^{\dagger}H$ 

• To get SM values,  $\mu_H \approx 89$  GeV and  $\lambda_H \approx 0.13$ , at  $T = 0$ 

$$
\mu_H^2 = \mu_0^2 + 2\mu \bar{\phi}
$$
 and  $\lambda_H = \lambda_0 + 2\frac{\bar{\phi}^2}{M^2}$  (A)

- Parameterize  $\lambda_0 = \varepsilon \lambda_H$ , with  $\varepsilon \lesssim 0.1$
- Higgs and  $\phi$  mix after EWSB, governed by mixing angle

$$
\theta \approx \frac{2v^2}{m_H^2} \left( \frac{2\bar{\phi}v}{M^2} - \frac{\mu}{v} \right)
$$

 $\bullet$  For  $|\theta|\ll 1$ ,  $m_H\approx$ √  $\overline{2}\,\mu_H$  (approximately SM) and lightest scalar mass

$$
m_{\phi}^2 \approx m_{0\phi}^2 + \frac{v^4}{M^2} - \theta^2 \, m_H^2
$$

• We choose benchmark values; fix  $\bar{\phi}$  from (A) above

$$
\varepsilon = 0.1 \quad \text{and} \quad M = 4.0 \times 10^3 \text{ TeV}
$$

• With  $\bar{\phi}$  fixed  $\Rightarrow m_{0\phi}$  and  $\theta$  functions of  $\mu$ 

• (I) Stable potential:  $m_{0\phi}^2 > 0$  (II) triggering symmetry breaking:  $\mu_0^2 > 0$ 

(I) and (II) 
$$
\Rightarrow \frac{\overline{\phi} v^2}{M^2} < \mu < \frac{\mu_H^2}{2\overline{\phi}}
$$
 (B)

• For benchmark values:  $\vert$  5 MeV  $\lesssim m_\phi \lesssim 10$  MeV

#### Thermal Evolution

- Degenerate minimum at critical temperature  $T_c \approx \frac{T_0}{\sqrt{1-\frac{C^2}{C^2}}}$  $1-E^2/(\lambda_0 D)$
- Barrier disappears at  $T = T_0$

$$
T_0^2 = \frac{m_{0H}^2 - 8Bv_0^2}{4D} \; ; \; B = \frac{3(2m_W^4 + m_Z^4 - 4m_t^4)}{64\pi^2 v^4}
$$
\n
$$
D = \frac{2m_W^2 + m_Z^2 + 2m_t^2}{8v^2} \; ; \; E = \frac{2m_W^3 + m_Z^3}{6\pi v^3}
$$

- <u>Initial values</u> ( $\langle \phi \rangle = 0$ ):  $m_{0H} =$ √  $\overline{2} \mu_0$ , and  $v_0^2 = \mu_0^2$  $_0^2/\lambda _0$ 

- $m_t = 0$  (vanishing top Yukawa) chosen here, for vacuum stability as explained later
- $T_c = 0$  not sensible at upper limit in  $(B)$ ; choose a slightly smaller upper limit

3.7 MeV  $\leq \mu \leq 4.0$  MeV  $\Rightarrow$  38 GeV  $\leq T_c \leq 90$  GeV and 37 GeV  $\leq T_0 \leq 87$  GeV

- Underlying assumption: FOPT is achieved for  $\phi \approx 0$
- Thermal effects mediated by  $\eta$  that are  $\propto \kappa$  pin  $\phi$  near origin



 $\phi/M$  at  $T_c$  (solid) and  $T_0$  (dashed), for benchmark parameters and  $\kappa = 10^{-4}$ 

- Numerical inspection:  $\phi/M$  remains small down to  $T \sim 10$  GeV
- One can show that  $\eta$  and  $\phi$  in thermal equilibrium

 $\bullet$   $m_\eta\sim \kappa^{1/2}\bar{\phi}\sim 10$  TeV;  $\eta$  can decay fast, *e.g.* via  $\eta F_{\mu\nu}F^{\mu\nu}/M_\eta$ , with  $M_\eta\lesssim 10^{18}$  GeV

# UV Stability

- FOPT through small  $\lambda_0 \sim 0.01$  can lead to instability
- This is a quantum (running) effect from  $y_t \sim 1$  top Yukawa

$$
16\pi^2 \frac{d\lambda_0}{dt} = 24\lambda_0^2 + \lambda_0(12y_t^2 - 9g^2 - 3g'^2) - 6y_t^4 + \frac{9}{8}g^4 + \frac{3}{8}g'^4 + \frac{3}{4}g^2g'^2
$$

- $\{g, g'\}$  are  $\{SU(2)_L, U(1)_Y\}$  gauge couplings, respectively
- $t \equiv \ln Q/Q_0$ , with  $Q_0 = 100$  GeV reference scale
- With our benchmark choices  $\lambda_0$  goes negative ("instability") for  $Q \gtrsim 170$  GeV
- Too close to  $T_c \sim 100$  GeV to allow for a reasonable theory
- If  $y_t \ll 1$  initially one could avoid this problem
- After FOPT, dynamics of  $\phi$  could allow for  $y_t$  to rise to SM value

• We introduce

$$
O_t = \frac{\phi}{M_t} H^* \epsilon \, \overline{Q}_L t_R + \text{H.C.} \qquad [Q_L = (t, b)_L; t_R \, \mathbb{Z}_2 \text{ odd}]
$$

$$
\Rightarrow y_t^{\text{SM}}(Q_0) = y_{0t}(Q_0) + \frac{\overline{\phi}}{M_t}
$$

• Initial dim-4 top Yukawa  $y_{0t} = 0$ ; for  $y_t^{\text{SM}}$  $_{t}^{\mathsf{SM}}(Q_{0})\approx 0.92\,\rightarrow\,M_{t}\approx 1.1\times 10^{3}\,$  TeV



Running of  $\lambda_0$ , for  $\lambda_0(Q_0) = 0.1\lambda_H$ , with  $Q_0 = 100$  GeV

# Flavor from  $\phi$

- $\mathbb{Z}_2$  odd  $t_R$  not fundamentally different from other fields in our model
- We thus extend  $\phi$ -induced Yukawa coupling to all fermions; schematically

$$
O_f = \frac{\phi}{M_f} H \,\bar{F}_L f_R + \text{H.C.}
$$

F:  $SU(2)_L$  doublet containing flavor f

- In general, one needs a Yukawa matrix defined by flavor-dependent coefficients

• Assuming vanishing initial Yukawa couplings

$$
y_f^{\text{SM}}(Q_0) = \frac{\bar{\phi}}{M_f} \quad \Rightarrow \quad M_f = \frac{y_t^{\text{SM}}(Q_0)}{y_f^{\text{SM}}(Q_0)} M_t
$$

- After mass matrix diagonalization
- Direct coupling of the fermion f to  $\phi$ :  $\xi_f = \langle H \rangle / M_f$

$$
\begin{array}{c|c|c|c|c|c} \hline \varepsilon & \kappa & M \ (\text{TeV}) & M_t \ (\text{TeV}) \\ \hline 0.1 & 10^{-4} & 4.0 \times 10^3 & 1.1 \times 10^3 \end{array}
$$

Benchmark input parameters

- Softly broken  $\mathbb{Z}_2$ : loop-induced Yuakwa couplings
- For top quark, with our reference parameters

$$
\delta y_t(Q_0) \sim \frac{\mu}{16\pi^2 M_t} \ln(M_t/Q_0) \sim 10^{-10},
$$

- Completely negligible, does not affect stability
- Smaller for lighter fermions



⊕:  $\mathbb{Z}_2$  soft-breaking interaction  $\propto \mu$ ; ■: dimension-5  $O_f$  operator

## Couplings of  $\phi$  to Fermions

• Including mixing, the total coupling  $\bar{\xi}_f \phi \bar{f}_L f_R + \text{H.C.}$  given by

$$
\left| \bar{\xi}_f \equiv \xi_f + \theta \, y_f^{\text{SM}} \right|
$$

- One can show that in our model  $\theta y_f \sim \xi_f$
- For electron Yukawa we find  $M_e \approx 3.3 \times 10^{11}$  GeV
- Assuming Dirac neutrinos with  $y^{\text{SM}}_{\nu}$  $_{\nu}^{\mathsf{SM}}(Q_0) \sim 10^{-12} \Rightarrow M_{\nu} \sim 4 \times 10^{18}$  GeV
- Close to, but below, Planck mass  $M_P \sim 10^{19}$  GeV
- All flavor can arise from  $\phi$  dynamics coupled to scales  $\sim 10^3$  TeV  $M_P$
- Alternatively, the UV scale  $\sim 10^3$  TeV, with hierarchic  $\phi$  couplings to fermions
- For benchmark values,  $\phi$  decay length typically macroscopic

$$
c\,\tau_\phi \approx \frac{8\pi}{\bar\xi_e^2\,m_\phi} \approx 8.4\times 10^3\,\,\text{km} \left(\frac{8.4\times 10^{-20}}{\bar\xi_e^2}\right)\!\!\left(\frac{7\,\,\text{MeV}}{m_\phi}\right) \qquad (m_\phi < 2 m_\mu)
$$

c speed of light and  $y_e \approx 2.9 \times 10^{-6}$  SM electron Yukawa coupling

• Generally expect  $\phi$  to escape as missing energy in laboratory experiments

• For purposes of illustration, let  $m_{\phi} = 7$  MeV (typical for our parameters)



- Allowed region between dashed lines
- G. Krnjaic, 2015
- Revised SN bounds

Dev, Mohapatra, Zhang, 2020

- BBN sets lower limit, dotted line
- We get  $\bar{\xi}_t \approx 2.6 \times 10^{-4}$ ; allowed by known bounds See, e.g., Foroughi-Abari, Ritz, 2020



From G. Krnjaic, Phys. Rev. D 94, 073009 (2016), 1512.04119

# Low Energy Tests

- $\bullet$   $\phi$  can appear as missing energy in rare meson decays
- Example: KOTO, measuring  $K_L \rightarrow \pi^0 \bar{\nu} \nu$ , to  $\sim 10 \%$
- SM prediction for branching fraction:  $3.4 \pm 0.6 \times 10^{-11}$ Cirigliano, Ecker, Neufeld, Pich, Portoles, 2011; Egana-Ugrinovic, Homiller, Meade, 2019
- A few anomalous events were observed by KOTO Shinohara (KOTO), 2020
- $\phi$ -top coupling in the 1-loop penguin diagram for  $K_L \rightarrow \pi^0 \phi$  can mimic the events
- $\phi$ -top coupling  $\sim 2 \times 10^{-4}$  and  $m_{\phi} \lesssim 50$  MeV could yield a  $2\sigma$  explanation Egana-Ugrinovic, Homiller, Meade, 2019; Kitahara, Okui, Perez, Soreq, Tobioka, 2019

• KOTO has since concluded events are consistent with background Ahn et al. (KOTO), 2020

- Future KOTO measurements can potentially probe  $\bar{\xi}_t \lesssim 10^{-5}$
- Generically, also expect gravitational waves from the FOPT
- Possibly detectable by e.g. LISA See, for example, Caprini et al., 2019

# Summary

• FOPT, a possible key element in weak scale baryogenesis, not feasible in SM with  $m_H \approx 125$  GeV

• Typically need to extend SM with particles near weak scale with  $\mathcal{O}(1)$  couplings to get FOPT

• We propose that a smaller initial Higgs quartic could result in a FOPT

• Proposal: dynamics of a weakly coupled sub-GeV scalar can yield the late time SM Higgs potential

• Stability of our scenario suggests top Yukawa coupling initially small or vanishing

• Natural extension to all fermions: flavor turns on after EWSB, via operators suppressed by scales  $\sim 10^3$  TeV (top) to near Planck mass (Dirac neutrinos)

• The sub-GeV scalar is typically long-lived, can show up as missing energy in rare meson decays

- KOTO, for example, can probe this scenario in  $K_L \to \pi^0 +$  "missing energy"
- Generically also expect primordial gravitational waves from weak scale FOPT
- Potential signal for future observatories, like LISA