

[1] JHEP 07 (2019) 050,
[2] Phys.Rev.D 110 (2024) 6, 063535, [3] JHEP 05 (2024) 281,
[4] JCAP 08 (2025) 059, and [5] JCAP 11 (2025) 013.

Bridging the MeV-Gap for Light Higgs Portal Dark Matter

Yue-Lin Sming Tsai

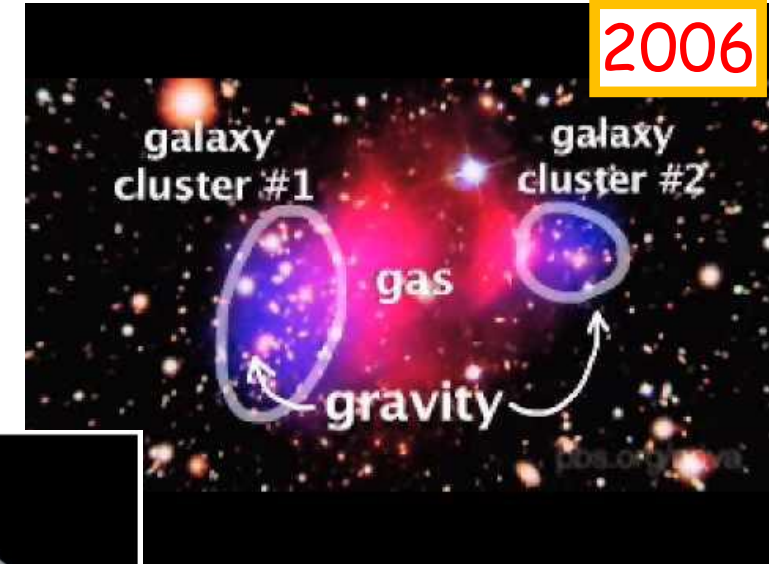
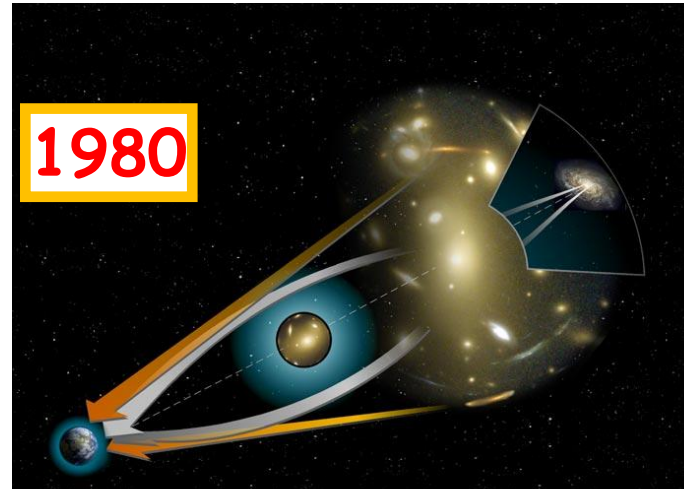
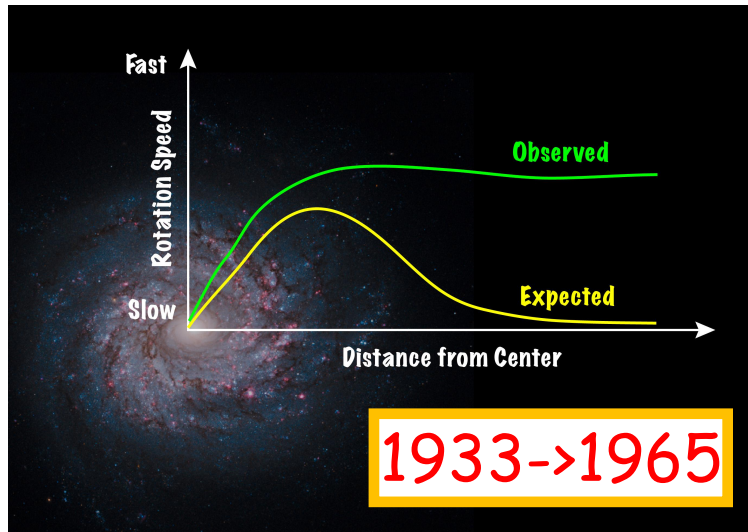
(Purple Mountain Observatory)

2025.12.21@HiggsPotential 2025

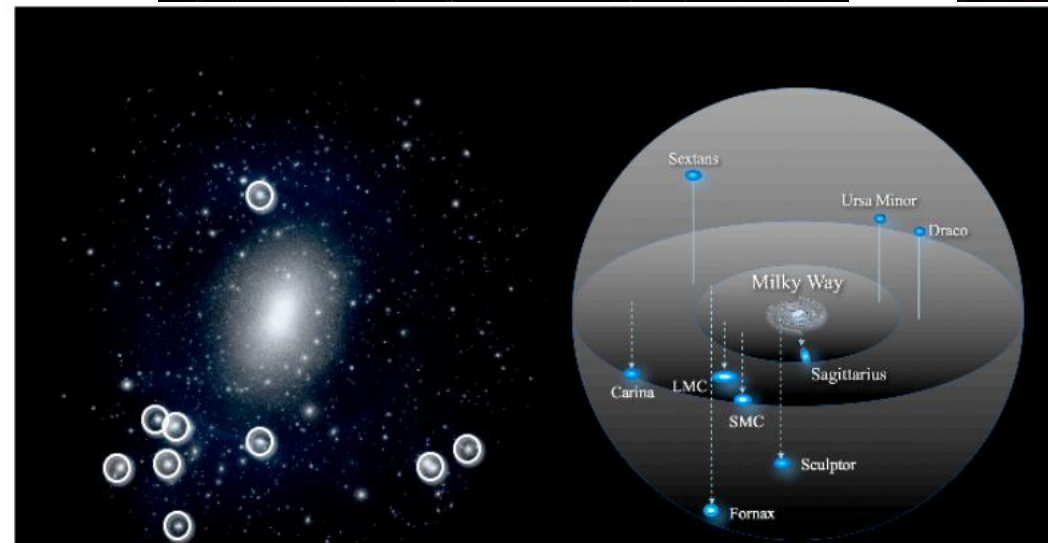
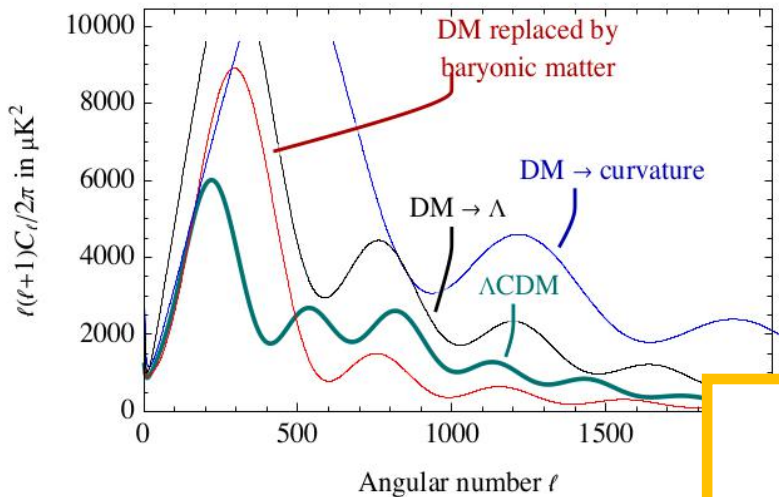
Outline

- ☐ Motivations.
- ☐ Minimal dark matter Models.
- ☐ Parameter space to be detected in gamma-ray telescope (VLAST).
- ☐ Results and summary.

Dark Matter Problems



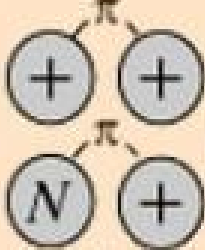
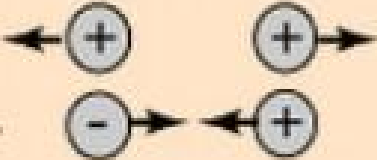

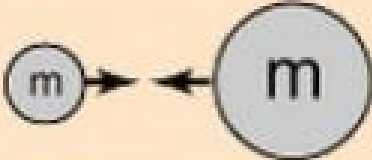
CMB power spectrum



IF GR is correct,
it will be difficult
to explain the
universe without
DM assumption.

Diversity problem:
More and more dSphs were found!
Some of them are DM-rich but some are DM-poor.

Fundamental Forces

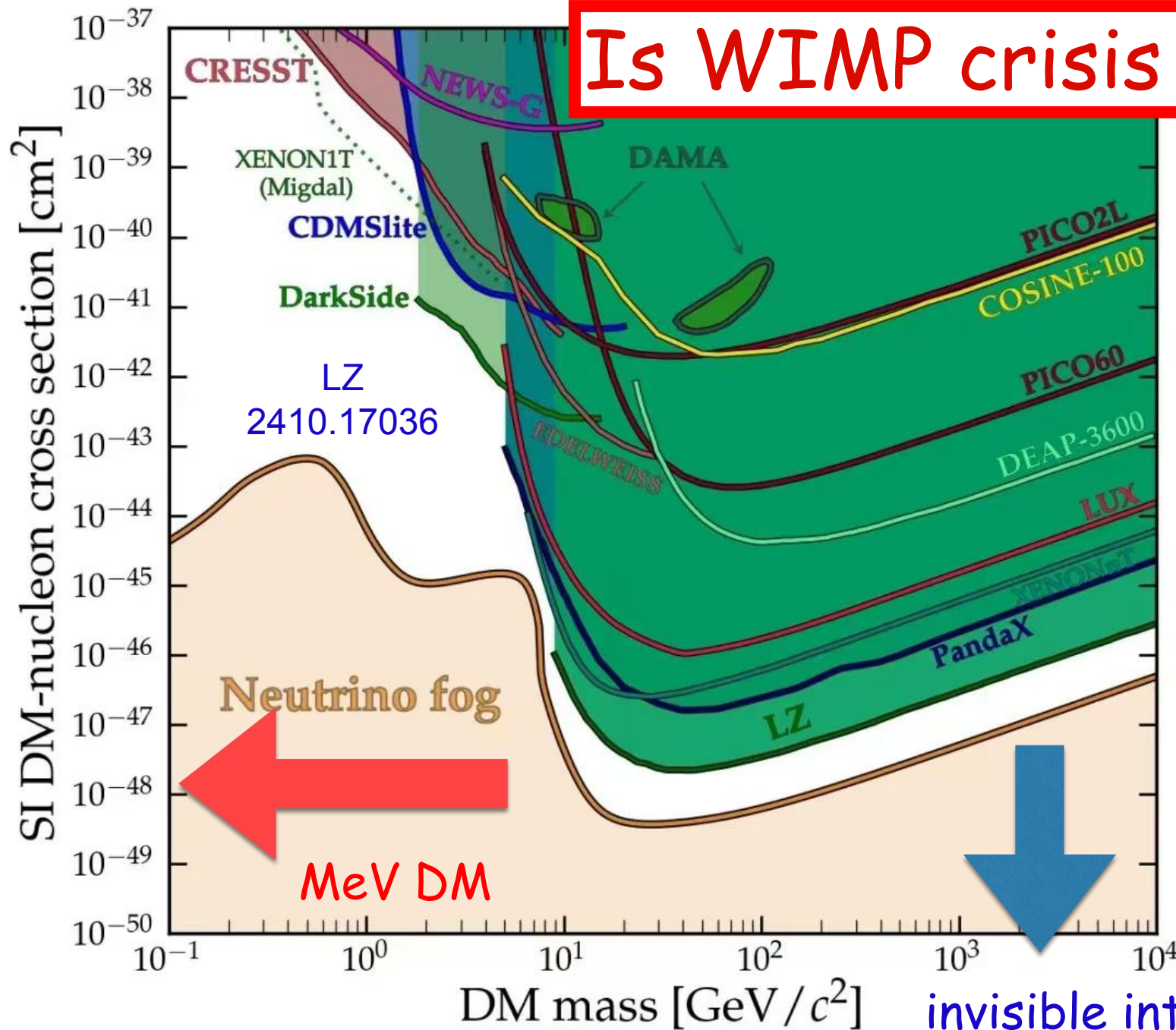
<i>Strong</i>		Force which holds nucleus together	Strength 1	Range (m) 10^{-15} (diameter of a medium sized nucleus)	Particle gluons, π (nucleons)
<i>Electro-magnetic</i>			Strength $\frac{1}{137}$	Range (m) Infinite	Particle photon mass = 0 spin = 1
<i>Weak</i>		neutrino interaction induces beta decay	Strength 10^{-6}	Range (m) 10^{-18} (0.1% of the diameter of a proton)	Particle Intermediate vector bosons W^+ , W^- , Z_0 , mass > 80 GeV spin = 1
<i>Gravity</i>			Strength 6×10^{-39}	Range (m) Infinite	Particle graviton ? mass = 0 spin = 2

How is possible that no interaction between $1e-6$ and $1e-39$?

If new interaction greater than Gravity...

What is the portal?

Is WIMP crisis or Human panic?



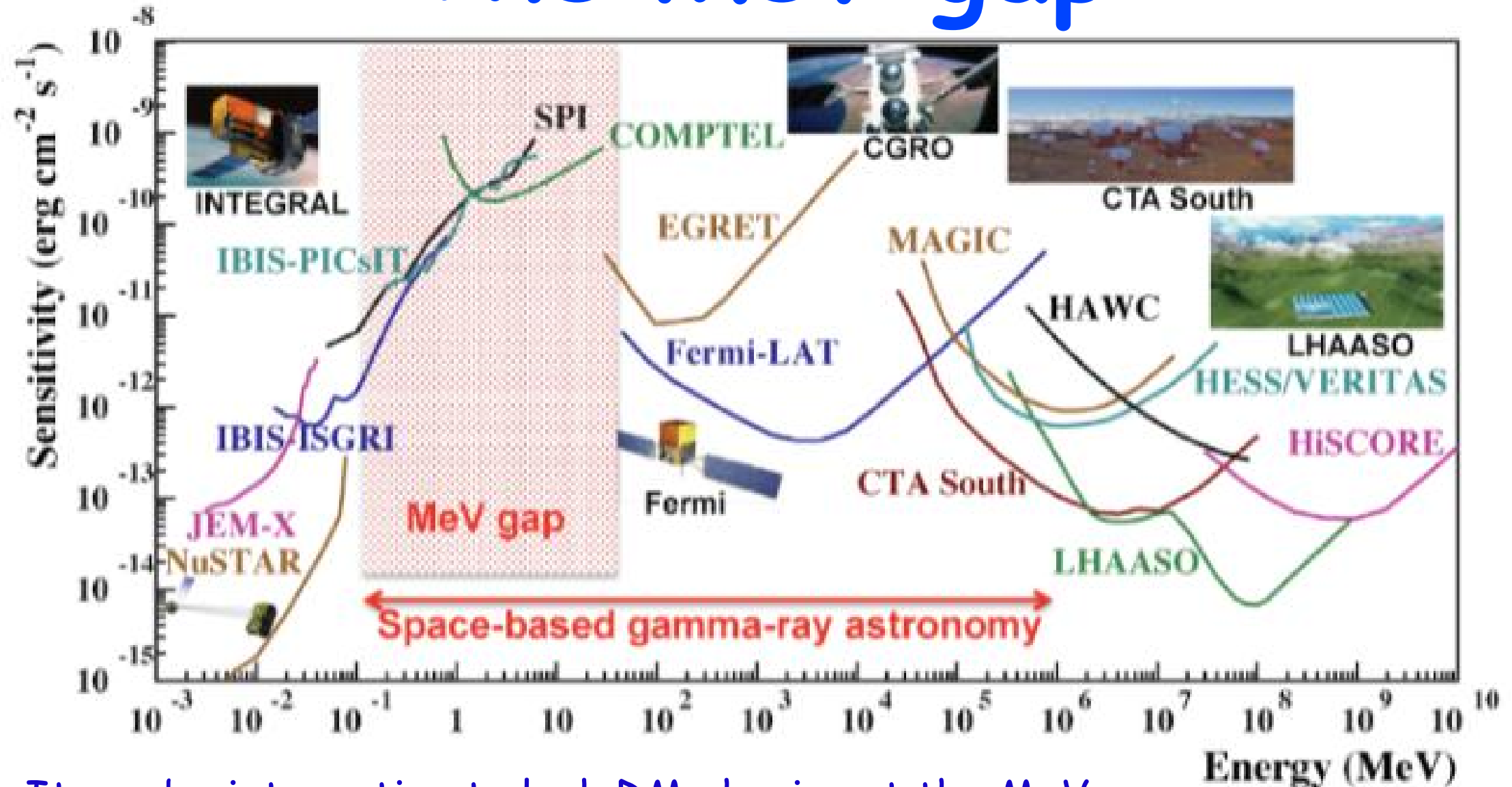
Lower energy,
higher exposure, or
wrong DM density?

- p-wave
- Resonance
- Forbidden DM
- Coannihilation
- Secluded DM

Velocity dependent annihilations!

invisible interactions

The MeV gap



It can be interesting to look DM physics at the MeV gap.

Challenges of MeV dark matter

The light DM mass region

Can we go to the region below GeV?

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Cosmological Lower Bound on Heavy-Neutrino Masses

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(Received 13 May 1977)

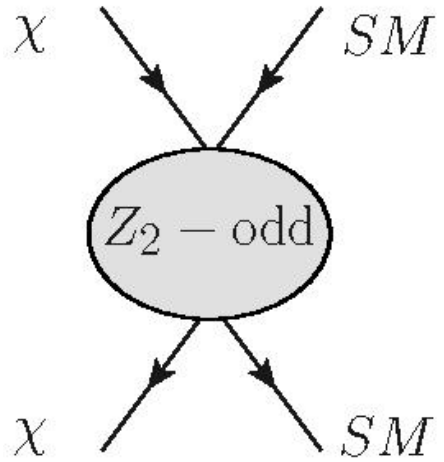
If only a DM introduced...

$g = \text{Weak coupling}$

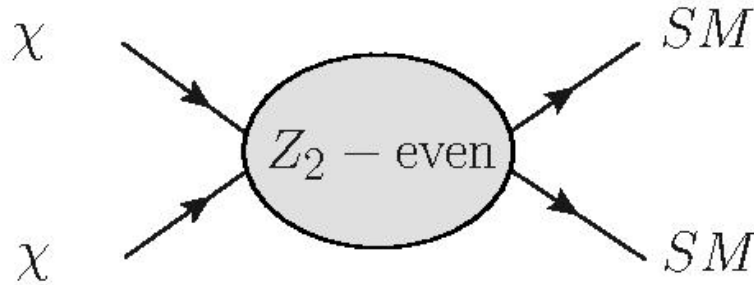
The present cosmic mass density of possible stable neutral heavy leptons is calculated in a standard cosmological model. In order for this density not to exceed the upper limit of $2 \times 10^{-29} \text{ g/cm}^3$, the lepton mass would have to be *greater* than a lower bound of the order of 2 GeV.

Unless, a new light mediator is introduced!

Simplicity and Light mediator

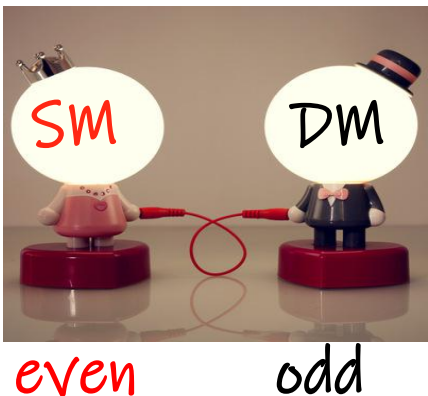


t-channel
annihilation



s-channel
annihilation

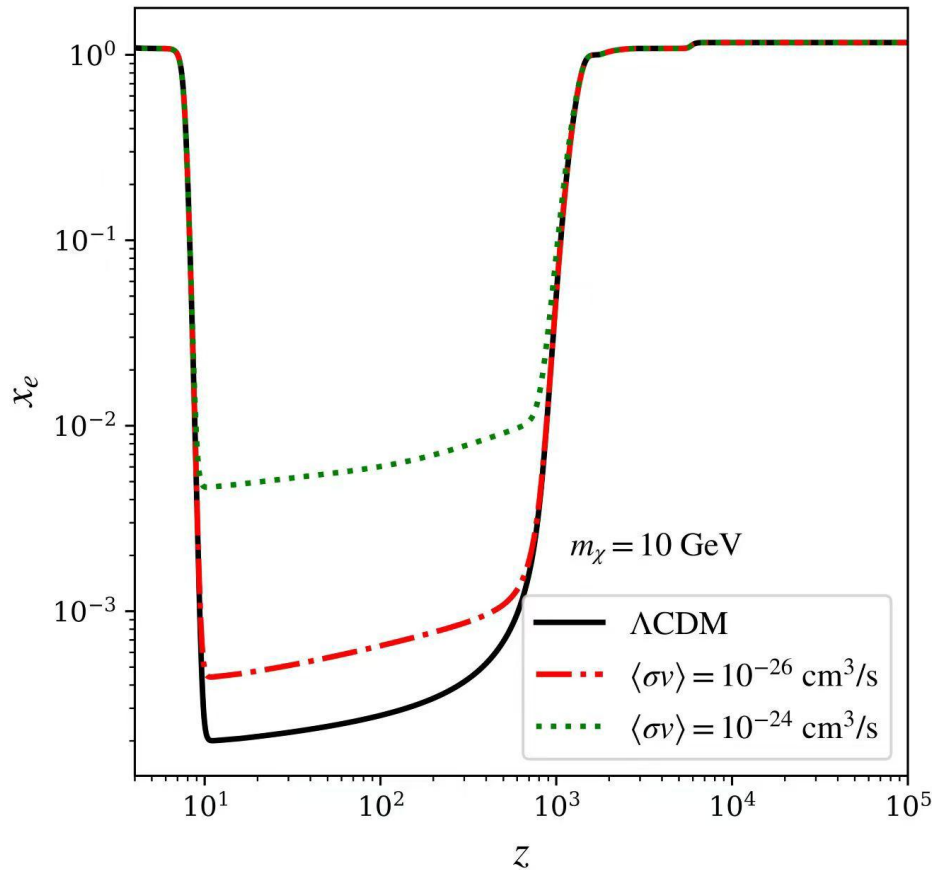
- ① Z_2 odd scalar mediator (like squark) + SM fermion. LEP mass limit for charged mediator is heavier than 100 GeV.
- ② Z_2 odd fermion mediator (like Chargino) + SM gauge boson. Invisible decay gives a severe limit.



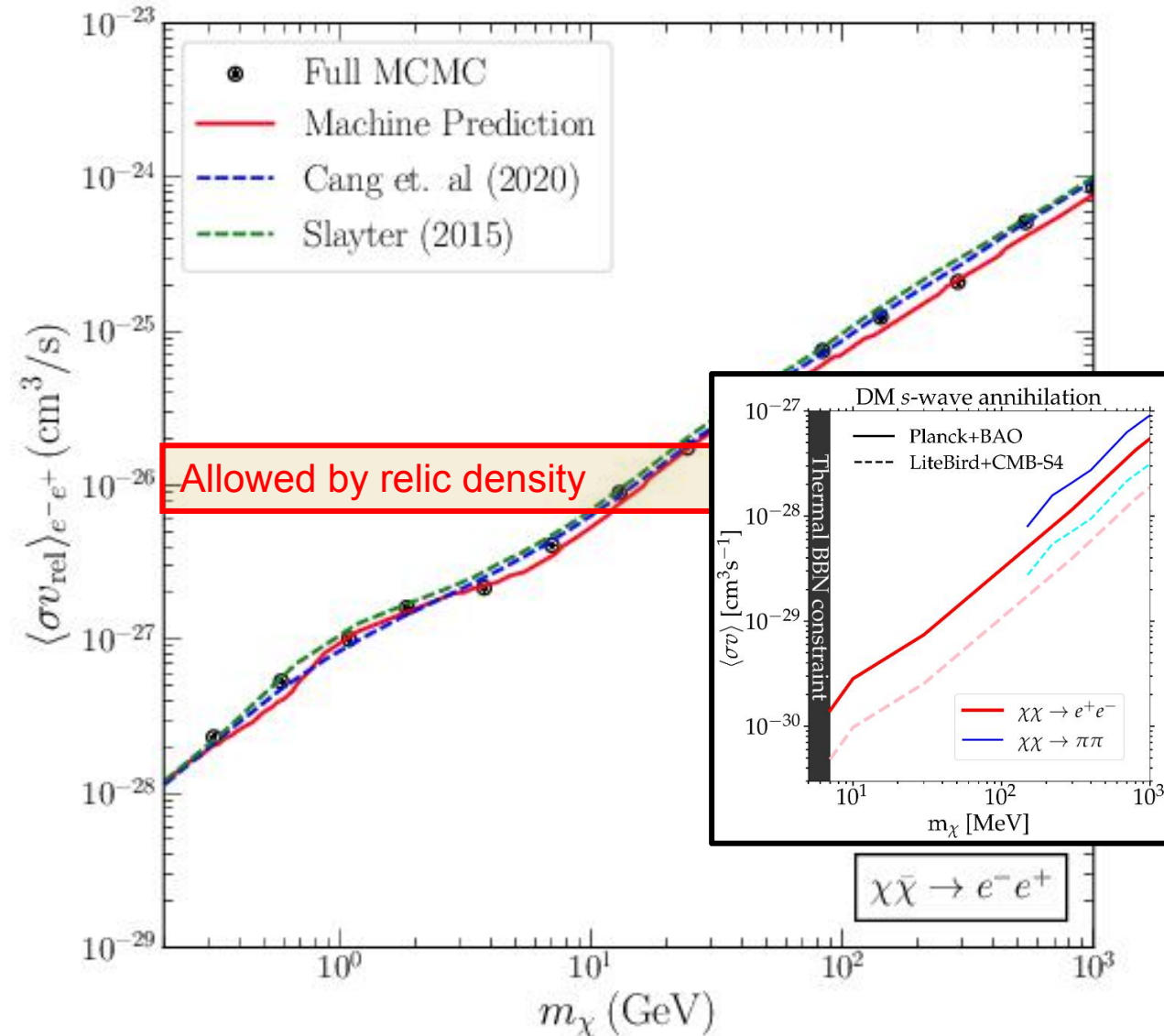
Therefore, an MeV mediator of the the DM annihilation to SM pair via t-channel **CANNOT** be Z_2 -odd.

Higgs portal or dark photon portal?

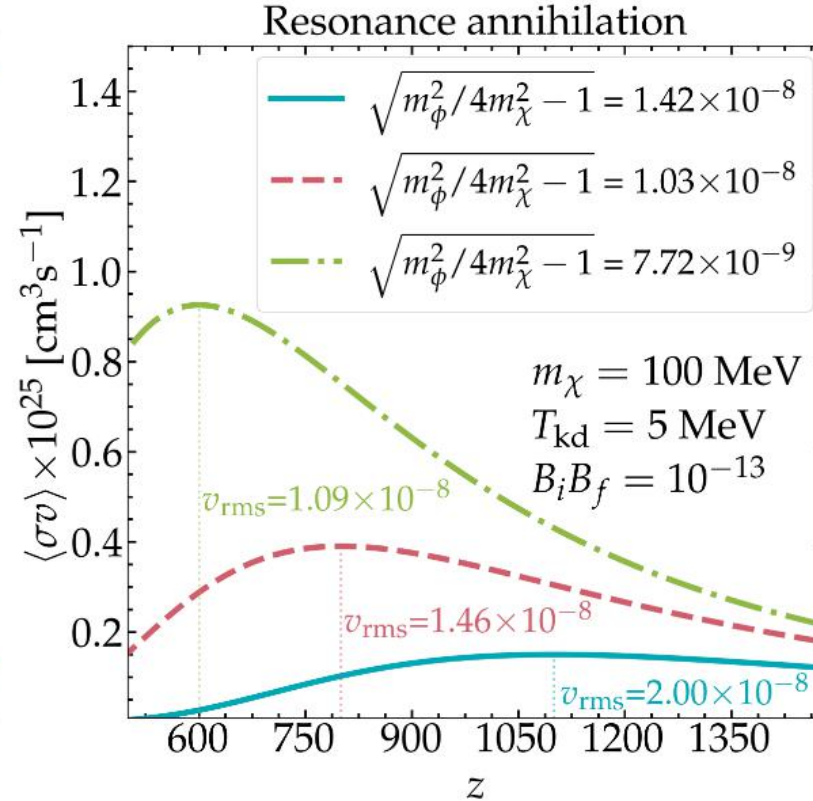
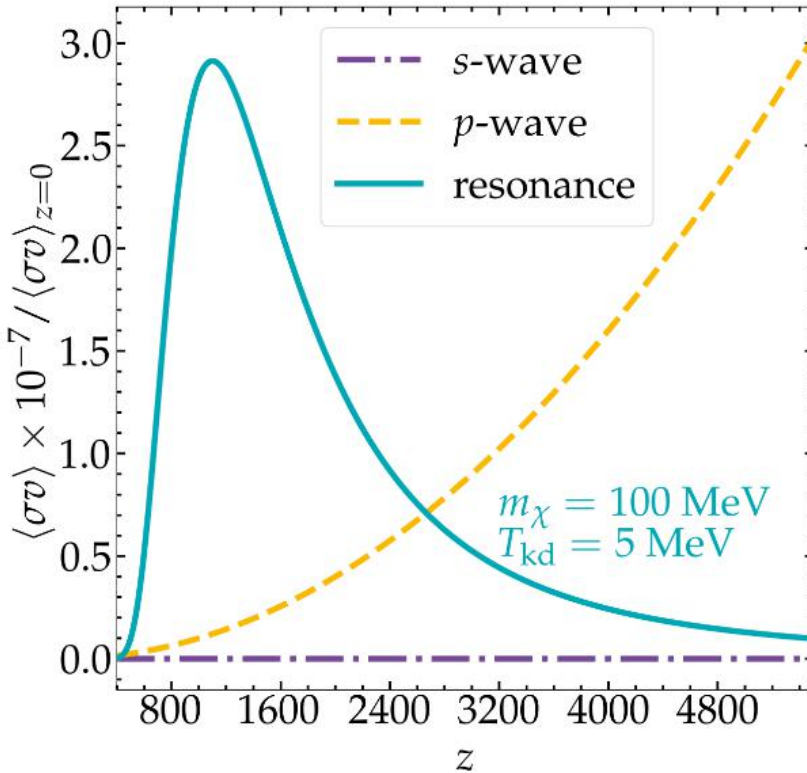
CMB constraints



- DM annihilation injects energy into gas to ionize and heat Hydrogen.
- For sub-GeV DM with correct relic density, s-wave may be excluded?



CMB constraints



- The velocity is $\sim 10^{-8} c$ during recombination epoch.
- p -wave annihilation cross-section is suppressed by the velocity.
- It requires an extreme fine-tuning for resonance annihilation in order to be testable in CMB data.

s-wave

p-wave

resonance

Cross-section

$$\langle \sigma v_{\text{rel}} \rangle \propto a_s$$

$$\langle \sigma v_{\text{rel}} \rangle \equiv b \langle v_\chi^2 \rangle$$

$$\langle \sigma v_{\text{rel}} \rangle = \frac{576}{\sqrt{3}} \frac{\pi^{3/2}}{m_\phi^2} \frac{\gamma}{v_{\text{rms}}^3} e^{-3\xi/v_{\text{rms}}^2} B_i B_f$$

Free parameters

$$\langle \sigma v_{\text{rel}} \rangle, m_\chi$$

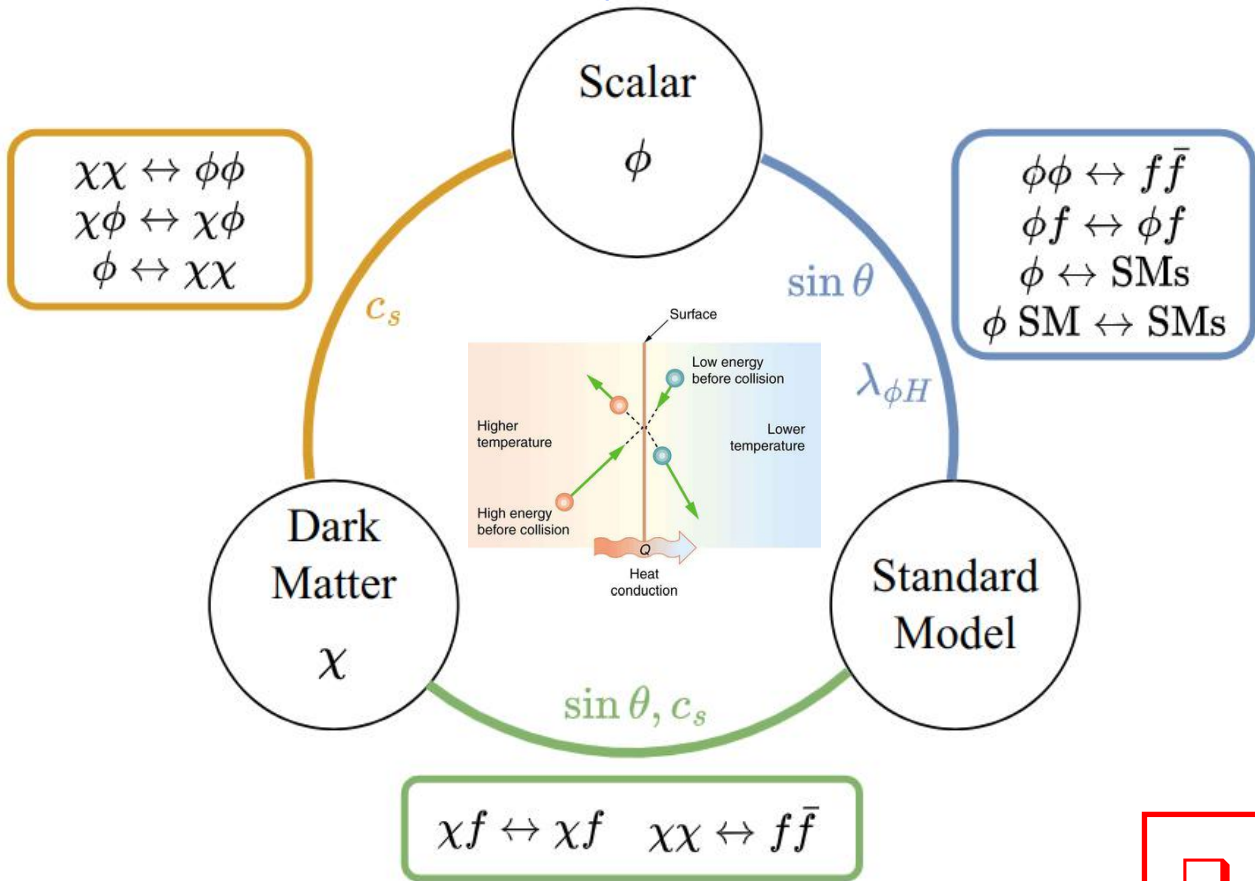
$$b, T_{\text{kd}}, m_\chi$$

$$B_i B_f, T_{\text{kd}}, m_\chi$$

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Thermal dark matter

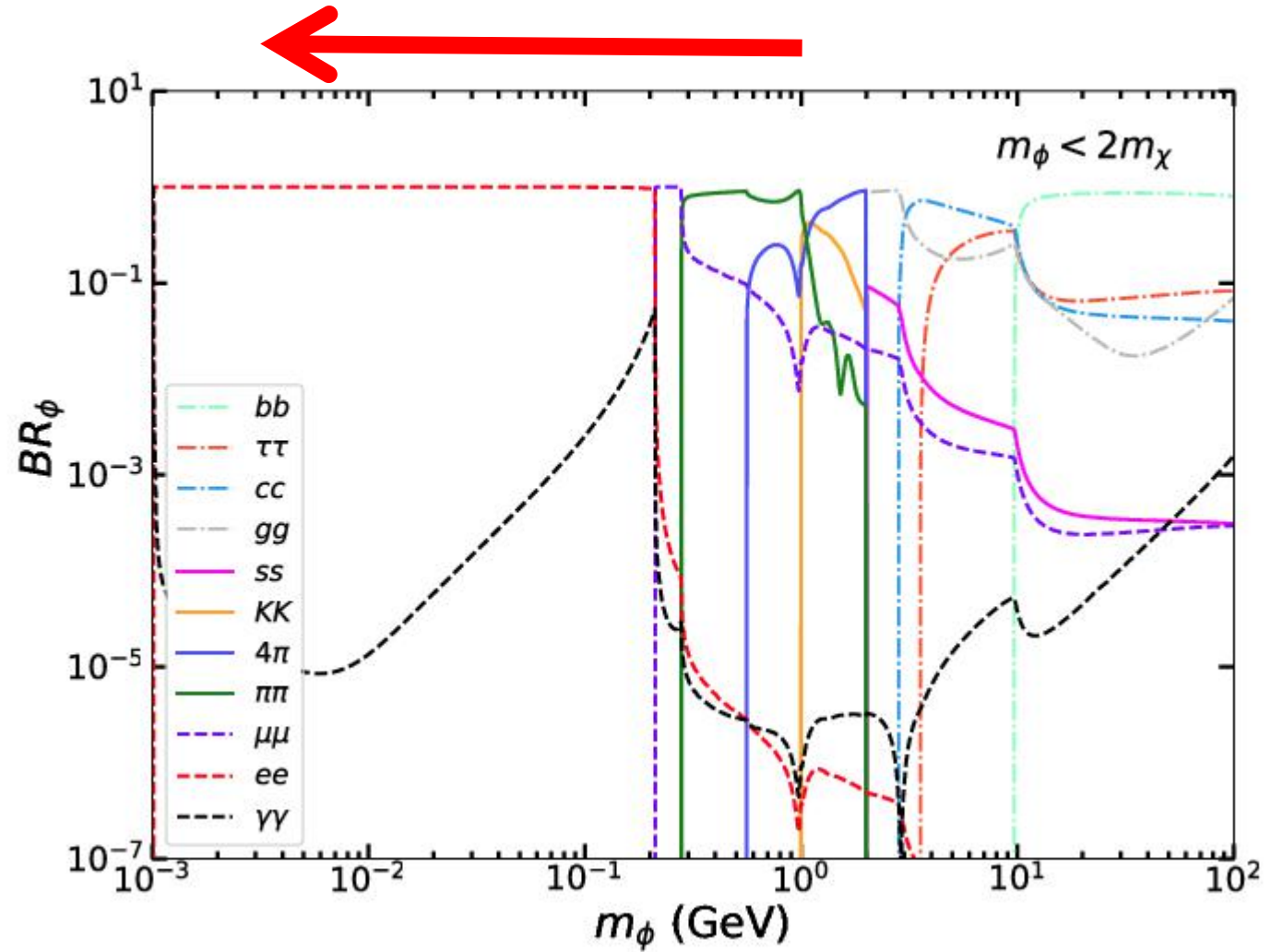


	Likelihood	Constraints
Relic abundance	Gaussian	$\Omega_{\chi}^{\text{exp}} h^2 = 0.1193 \pm 0.0014$ [90]; $\sigma_{\text{sys}} = 10\% \times \Omega_{\chi}^{\text{th}} h^2$.
Equilibrium	Conditions	either $(\Gamma_{\chi\text{SM}}^{\text{FO}} \geq H_{\text{FO}})$, or $(\Gamma_{\phi\text{SM}}^{\text{FO}} \geq H_{\text{FO}} \text{ and } \Gamma_{\chi\phi}^{\text{FO}} \geq H_{\text{FO}})$
DM direct detection	Half Gaussian	$9 \text{ GeV} < m_{\phi} < 10 \text{ TeV}$ (LZ [91]), $3.5 \text{ GeV} < m_{\phi} < 9 \text{ GeV}$ (PANDAX-4T [16]), $60 \text{ MeV} < m_{\phi} < 5 \text{ GeV}$ (DarkSide [92]).
ΔN_{eff}	Half Gaussian	$\Delta N_{\text{eff}} < 0.17$ for 95% C.L. [90]
BBN	Conditions	if $(m_{\phi} \geq 2m_{\pi})$ then $\tau_{\phi} \leq 1 \text{ s}$ [93], if $(m_{\phi} \leq 2m_{\pi})$ then $\tau_{\phi} \leq 10^5 \text{ s}$ [94].

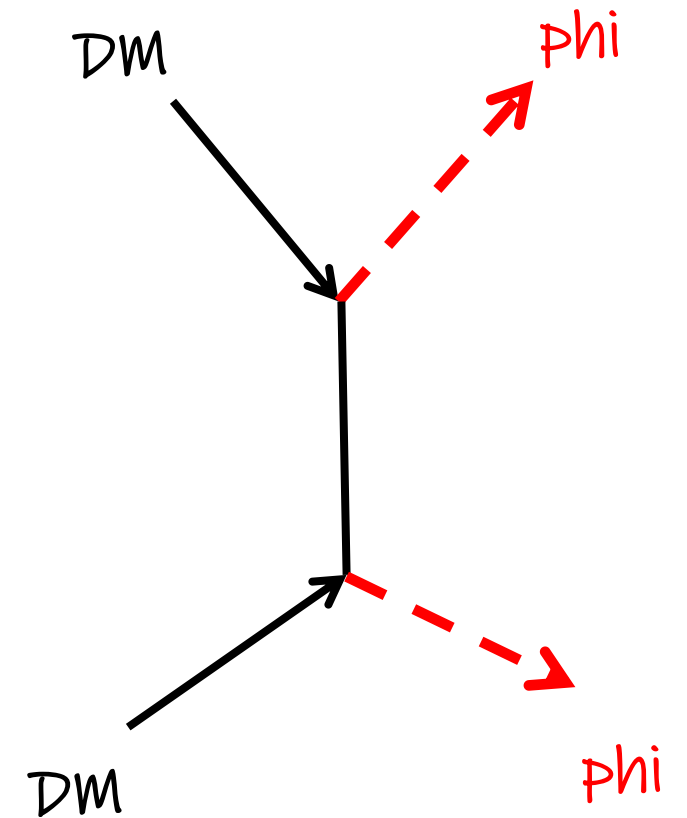
Heat transfer can be via the
green or orange+blue.

- ❑ Must be frequent momentum exchange!
(Common Problems of MeV DM)
- ❑ Number density can be described by
 $n \sim \exp(-m/T)$!

Sub-GeV Leptophilic Dark Matter interactions



Z ₂ even mediator				
types	Lagrangian	$\langle\sigma v\rangle_{2\mu}$ $\simeq a + bv^2$	$\langle\sigma v\rangle_{4\mu}$ $\simeq a + bv^2$	DD
χ and ϕ	$\mathcal{L}_1 = (g_D \bar{\chi} \chi + g_f \bar{f} f) \phi$	$a = 0$	$a = 0$	Eq. (B1)
	$\mathcal{L}_2 = (g_D \bar{\chi} \chi + g_f \bar{f} i \gamma^5 f) \phi$	$a = 0$	$a = 0$	—
	$\mathcal{L}_3 = (g_D \bar{\chi} i \gamma^5 \chi + g_f \bar{f} f) \phi$	Case (i)	$a = 0$	Eq. (B2)
	$\mathcal{L}_4 = (g_D \bar{\chi} i \gamma^5 \chi + g_f \bar{f} i \gamma^5 f) \phi$	Case (i)	$a = 0$	—
χ and V_μ	$\mathcal{L}_5 = (g_D \bar{\chi} \gamma^\mu \gamma^5 \chi + g_f \bar{f} \gamma^\mu f) V_\mu$	$a = 0$	Case (A)	Eq. (B3)
	$\mathcal{L}_6 = (g_D \bar{\chi} \gamma^\mu \gamma^5 \chi + g_f \bar{f} \gamma^\mu \gamma^5 f) V_\mu$	Case (ii)	Case (A)	—
	$\mathcal{L}_7 = (g_D \bar{\chi} \gamma^\mu \chi + g_f \bar{f} \gamma^\mu f) V_\mu$	Case (i)	Case (C)	Eq. (B4)
	$\mathcal{L}_8 = (g_D \bar{\chi} \gamma^\mu \chi + g_f \bar{f} \gamma^\mu \gamma^5 f) V_\mu$	Case (i)	Case (C)	—
S and ϕ	$\mathcal{L}_9 = (M_{D\phi} S^\dagger S + g_f \bar{f} f) \phi$	Case (i)	Case (B)	Eq. (B5)
	$\mathcal{L}_{10} = (M_{D\phi} S^\dagger S + g_f \bar{f} i \gamma^5 f) \phi$	Case (i)	Case (B)	—
	$\mathcal{L}_{9'} = (g_D S^\dagger S \phi + g_f \bar{f} f) \phi$	—	$b = 0$	—
	$\mathcal{L}_{10'} = (g_D S^\dagger S \phi + g_f \bar{f} i \gamma^5 f) \phi$	—	$b = 0$	—
S and V_μ	$\mathcal{L}_{11} = (ig_D S^\dagger \partial_\mu S + g_D^2 S^\dagger S V_\mu + g_f \bar{f} \gamma_\mu f) V^\mu$	$a = 0$	Case (C)	Eq. (B6)
	$\mathcal{L}_{12} = (ig_D S^\dagger \overleftrightarrow{\partial}_\mu S + g_D^2 S^\dagger S V_\mu + g_f \bar{f} \gamma_\mu \gamma^5 f) V^\mu$	$a = 0$	Case (C)	—
X_μ and ϕ	$\mathcal{L}_{13} = (M_{D\phi} X^\mu X_\mu^\dagger + g_f \bar{f} f) \phi$	Case (i)	Case (D)	Eq. (B7)
	$\mathcal{L}_{14} = (M_{D\phi} X^\mu X_\mu^\dagger + g_f \bar{f} i \gamma^5 f) \phi$	Case (i)	Case (D)	—
	$\mathcal{L}_{13'} = (g_D X^\mu X_\mu^\dagger \phi + g_f \bar{f} f) \phi$	—	$b = 0$	—
	$\mathcal{L}_{14'} = (g_D X^\mu X_\mu^\dagger \phi + g_f \bar{f} i \gamma^5 f) \phi$	—	$b = 0$	—
X_μ and V_μ	$\mathcal{L}_{15} = ig_D \{X^{\mu\nu} X_\mu^\dagger V_\nu - X^{\mu\nu\dagger} X_\mu V_\nu + X_\mu X_\nu^\dagger V^{\mu\nu}\} + g_D^2 \{X_\mu^\dagger X^\mu V_\nu V^\nu - X_\mu^\dagger V^\mu X_\nu V^\nu\} + g_f \bar{f} \gamma^\mu f V_\mu$	$a = 0$	Case (C)	Eq. (B8)
	$\mathcal{L}_{16} = ig_D \{X^{\mu\nu} X_\mu^\dagger V_\nu - X^{\mu\nu\dagger} X_\mu V_\nu + X_\mu X_\nu^\dagger V^{\mu\nu}\} + g_D^2 \{X_\mu^\dagger X^\mu V_\nu V^\nu - X_\mu^\dagger V^\mu X_\nu V^\nu\} + g_f \bar{f} \gamma^\mu \gamma^5 f V_\mu$	$a = 0$	Case (C)	—



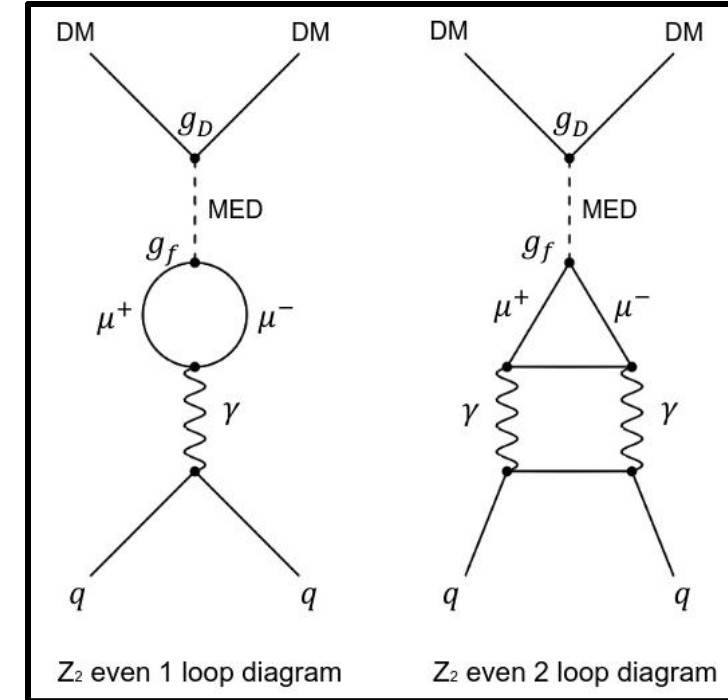
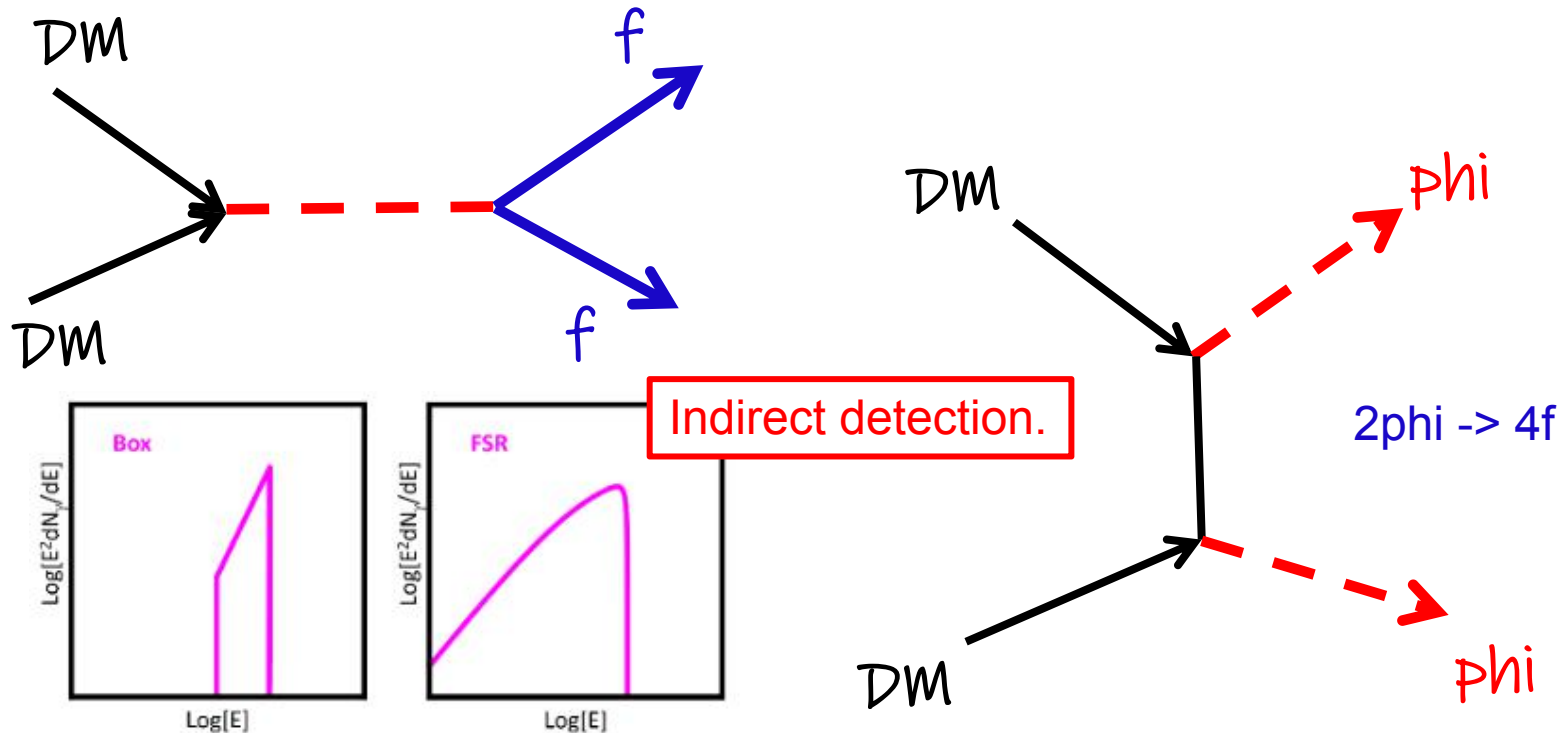
$$\mathcal{L}_{\text{int}} \supset -\frac{\cos\theta}{2}(c_s \phi \bar{\chi} \chi + c_p \phi \bar{\chi} i \gamma_5 \chi) + \frac{\sin\theta}{2}(c_s h \bar{\chi} \chi + c_p h \bar{\chi} i \gamma_5 \chi).$$

Abdughani, Fan, Lu, Tang and Tsai,
JHEP 07 (2022), 127

The joint contribution of $|c_s| \approx |c_p|$ leads to s-wave annihilation of $\chi\chi \rightarrow \phi\phi$

Minimum leptonphilic Lagrangian

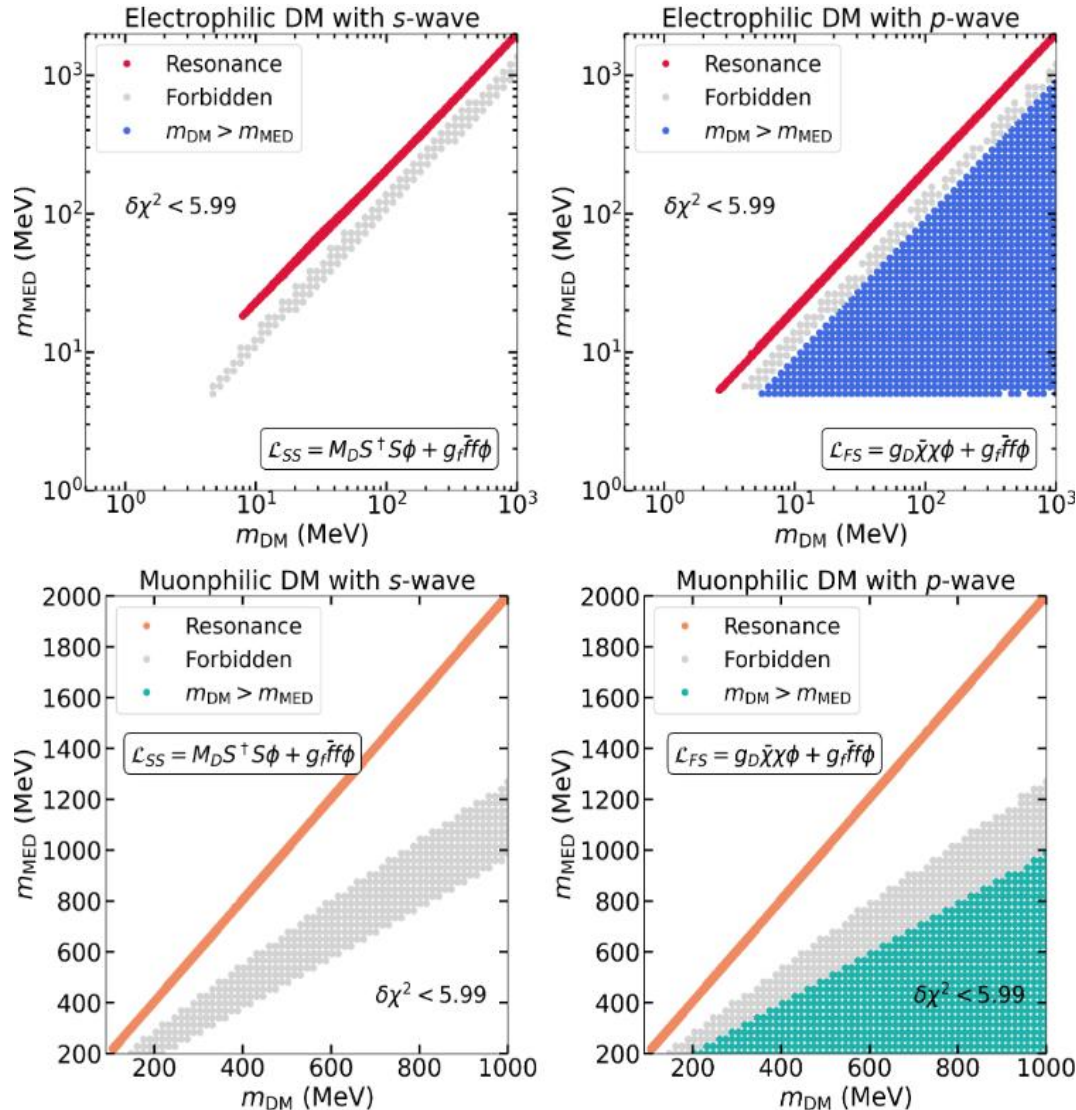
- (i) Scalar DM and scalar mediator (s-wave): $\mathcal{L}_{SS} = M_D S^\dagger S \phi + g_f \bar{f} f \phi$,
- (ii) Dirac DM and scalar mediator (p-wave): $\mathcal{L}_{FS} = g_D \bar{\chi} \chi \phi + g_f \bar{f} f \phi$.



Direct detection.

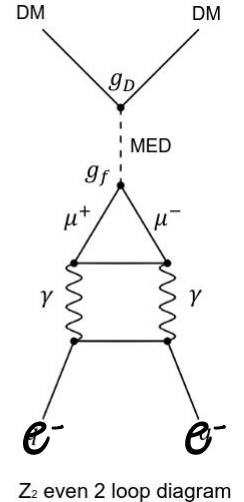
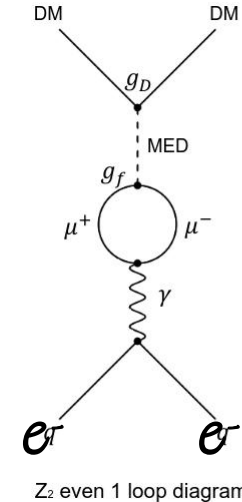
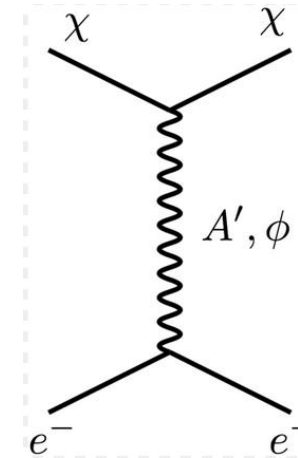
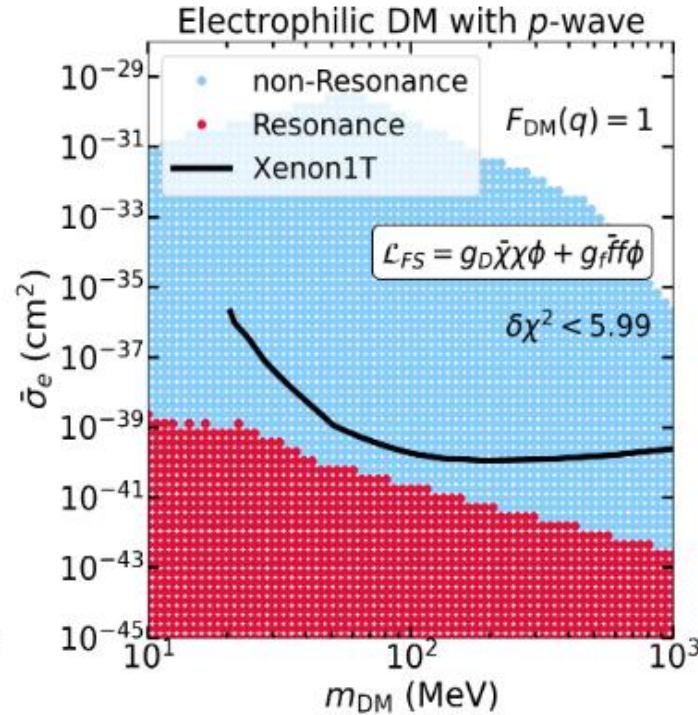
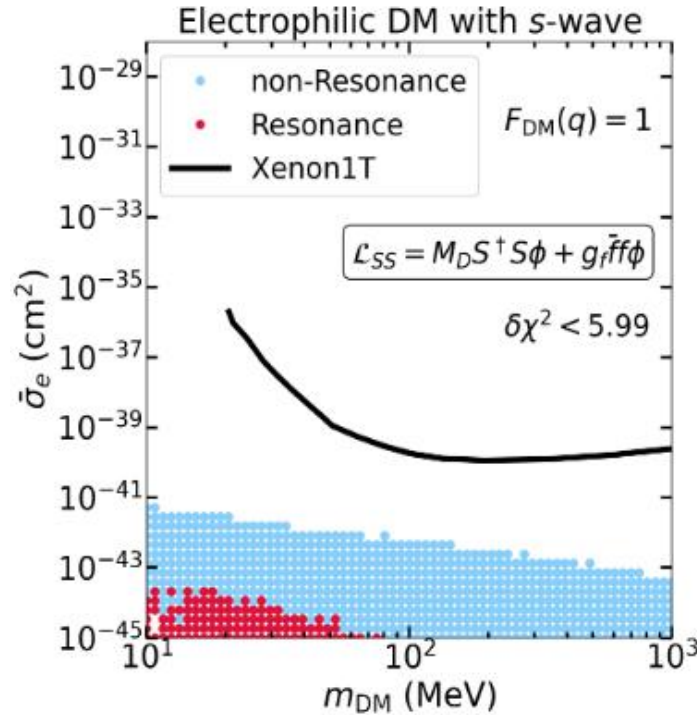
SM+ Z_2 -odd DM+ Z_2 -even mediator.

Allowed parameter space



- Only resonance, forbidden, and secluded annihilation mechanisms remain.
- Secluded DM with s-wave annihilation is completely excluded.
- The lower mass limits for DM vary between the three mechanisms.

DM-electron scattering



- In the electrophilic case, direct detection experiments impose strong constraints on the non-resonance parameter space due to tree-level DM-electron scattering.
- However, VLAST can effectively probe the resonance region that escapes these bounds.

Light thermal
dark matter in
minimal Higgs
portal model

Basic and minimum Lagrangian

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \frac{1}{2}\bar{\chi}(i\not{\partial} - m_{\chi})\chi + \frac{1}{2}(\partial\Phi)^2 - \frac{c_s}{2}\Phi\bar{\chi}\chi - \frac{c_p}{2}\Phi\bar{\chi}i\gamma_5\chi - V(\Phi, H),$$

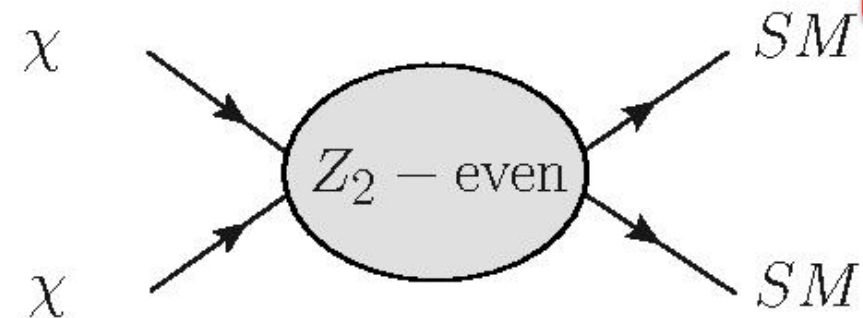
Majorana DM

SM singlet scalar

pseudo-scalar
interaction

Scalar interaction

Mixing between New
mediator and SM Higgs.



$$\mathcal{L}_{\text{int}} \supset -\frac{\cos\theta}{2}(c_s\phi\bar{\chi}\chi + c_p\phi\bar{\chi}i\gamma_5\chi) + \frac{\sin\theta}{2}(c_s h\bar{\chi}\chi + c_p h\bar{\chi}i\gamma_5\chi).$$

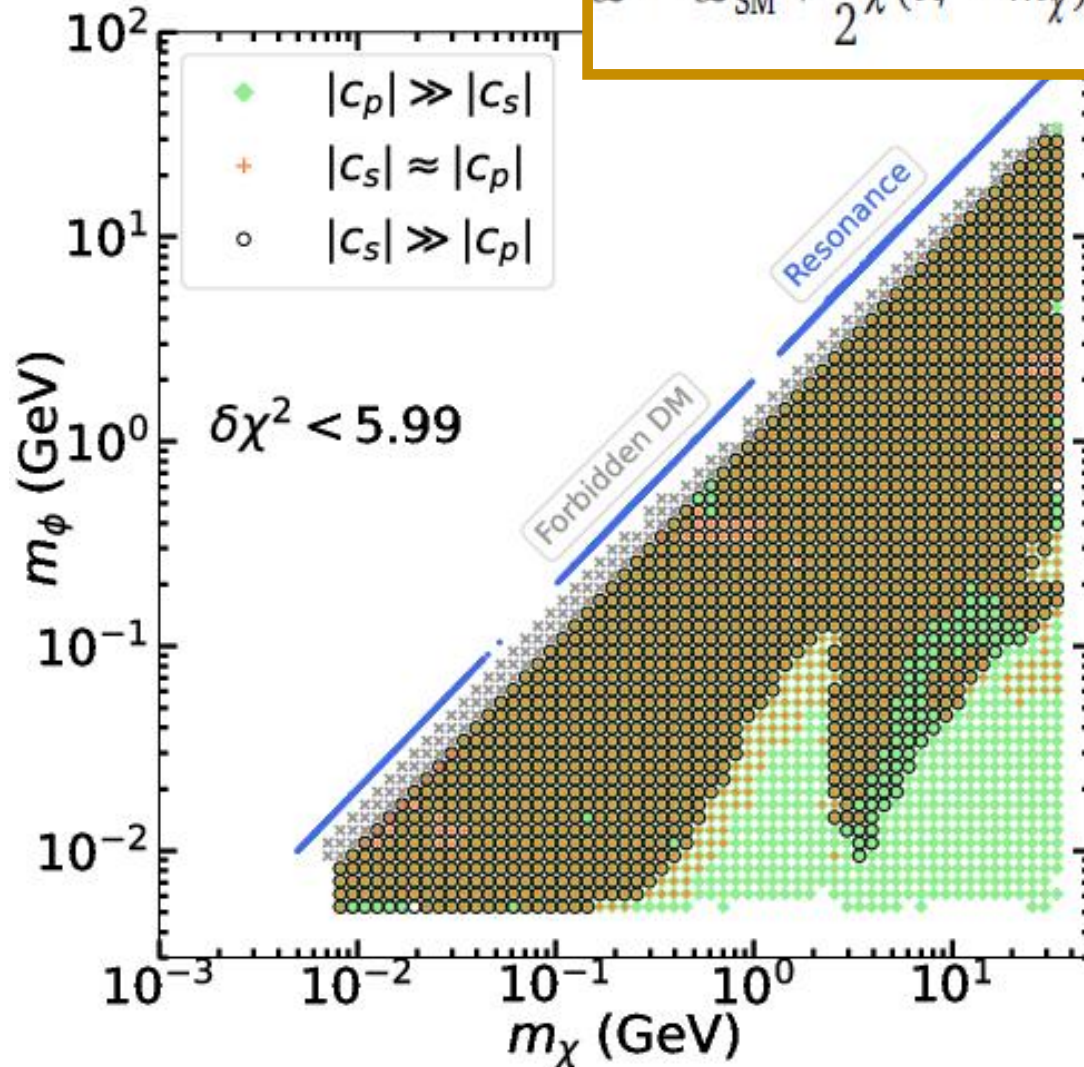
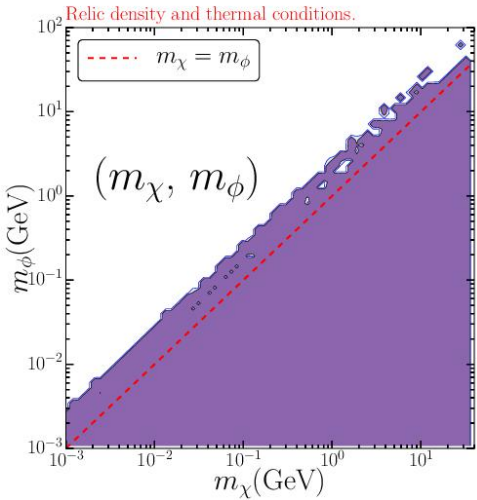
A minimum setup:

one SM singlet Majorana DM + one
SM singlet scalar mediator.

Possible parameter space

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \frac{1}{2} \bar{\chi} (i \not{\partial} - m_\chi) \chi + \frac{1}{2} (\partial \Phi)^2 - \frac{c_s}{2} \Phi \bar{\chi} \chi - \frac{c_p}{2} \Phi \bar{\chi} i \gamma_5 \chi - V(\Phi, H),$$

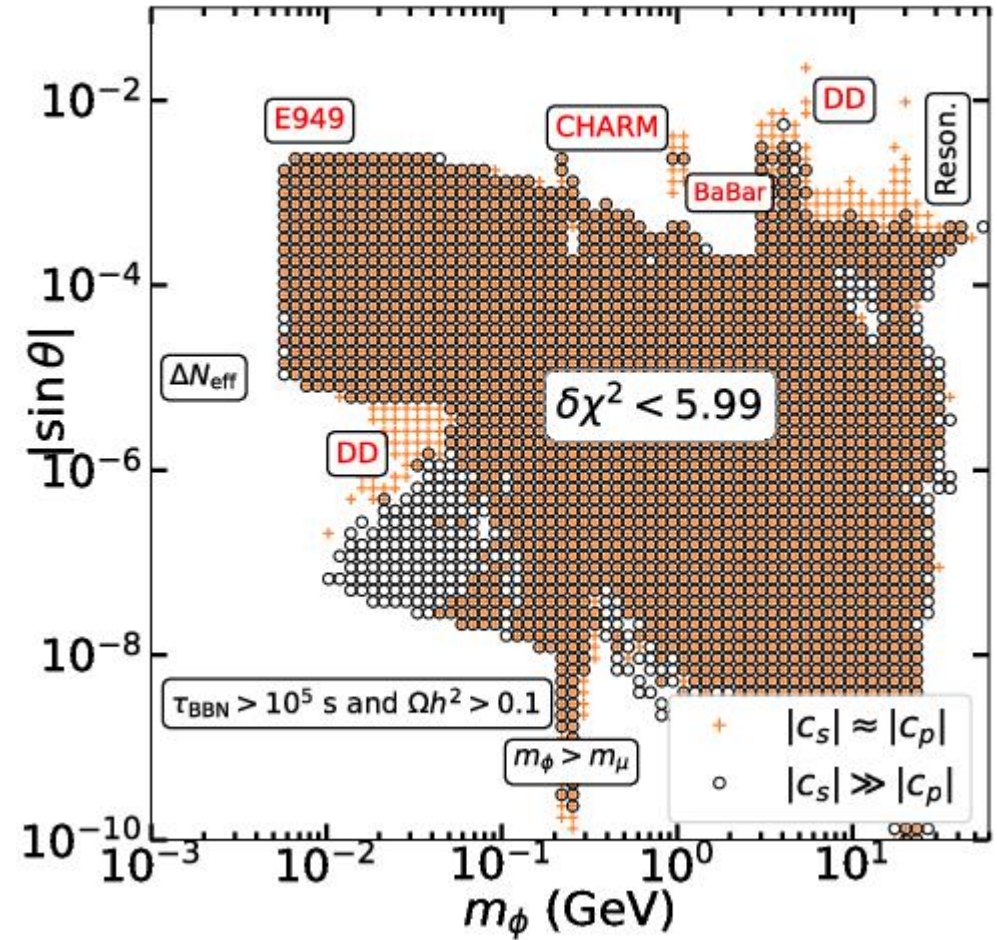
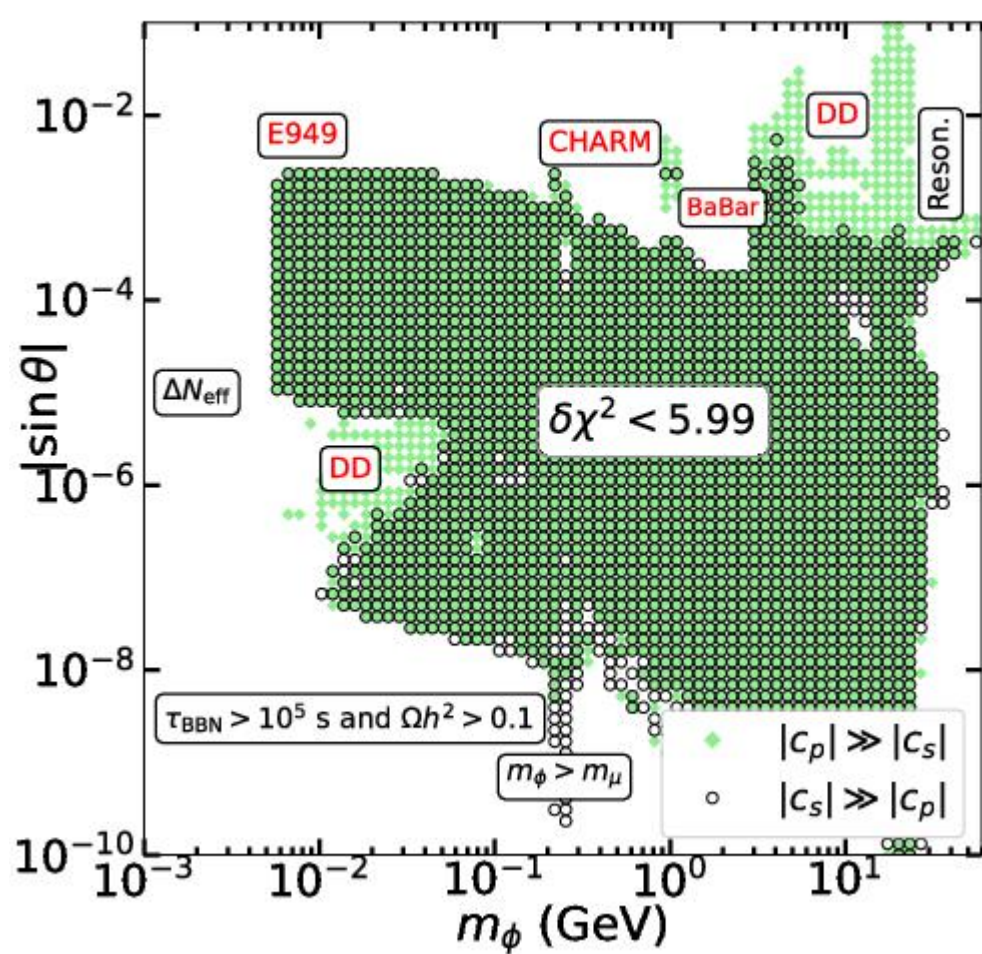
The unitarity, stability, and perturbative constraints.



$$\begin{aligned} 1 \text{ MeV} &\leq m_\chi \leq 30 \text{ GeV}, \\ -1 &\leq c_p \leq 1, \\ -1 &\leq c_s \leq 1, \\ 1 \text{ MeV} &\leq m_\phi \leq 60 \text{ GeV}, \\ -\pi/6 &\leq \theta \leq \pi/6, \\ -1 \text{ TeV}^2 &\leq \mu_\Phi^2 \leq 1 \text{ TeV}^2, \\ -1 \text{ TeV} &\leq \mu_3 \leq 1 \text{ TeV}, \\ -1 &\leq \lambda_\Phi \leq 1. \end{aligned}$$

	ϕ signature	Constraints
Higgs decay	Prompt*	See the upper limits of $\text{BR}(h \rightarrow \phi\phi)\text{BR}(\phi \rightarrow ll)^2$ from Fig. 12 of Ref. [99] and Fig. 7 of Ref. [100].
	Displaced*	See Ref. [101, 102]
	Long-lived*	$\text{BR}(h \rightarrow \text{inv.})_{\text{BSM}} \leq 0.145$ [103]
B decay	Prompt	$\text{BR}(B^\pm \rightarrow K^\pm \mu^- \mu^+) \lesssim 3 \times 10^{-7}$ [104]
	Displaced	(1) $\sin^2 \theta \gtrsim 2 \times 10^{-8}$ for the region $0.5 < m_\phi / \text{GeV} < 1.5$ and $1 < c\tau_\phi / \text{cm} < 20$ [105] (2) See Fig. 5 of Ref. [106] for details.
	Long-lived*	$P_p \text{BR}(B^\pm \rightarrow K^\pm \nu \bar{\nu}) = (2.3 \pm 0.7) \times 10^{-5}$ [107]
Kaon decay	Prompt	(1) $\text{BR}(K^+ \rightarrow \pi^+ \mu^- \mu^+) \leq 4 \times 10^{-8}$ [108] (2) $\text{BR}(K_L \rightarrow \pi^0 e^- e^+) \leq 2.8 \times 10^{-10}$ [109] (3) $\text{BR}(K_L \rightarrow \pi^0 \mu^- \mu^+) \leq 3 \times 10^{-10}$ [110]
	Displaced	CHARM detected events $\gtrsim 2.3$ [111]
	Long-lived*	(1) $\text{BR}(K_L \rightarrow \pi^0 \nu \bar{\nu}) \leq 3.0 \times 10^{-9}$ [112] (2) See $\text{BR}(K^+ \rightarrow \pi^+ \nu \bar{\nu})$ limits from Fig. 18 of Ref. [113] and Fig. 4 of Ref. [114] for details.

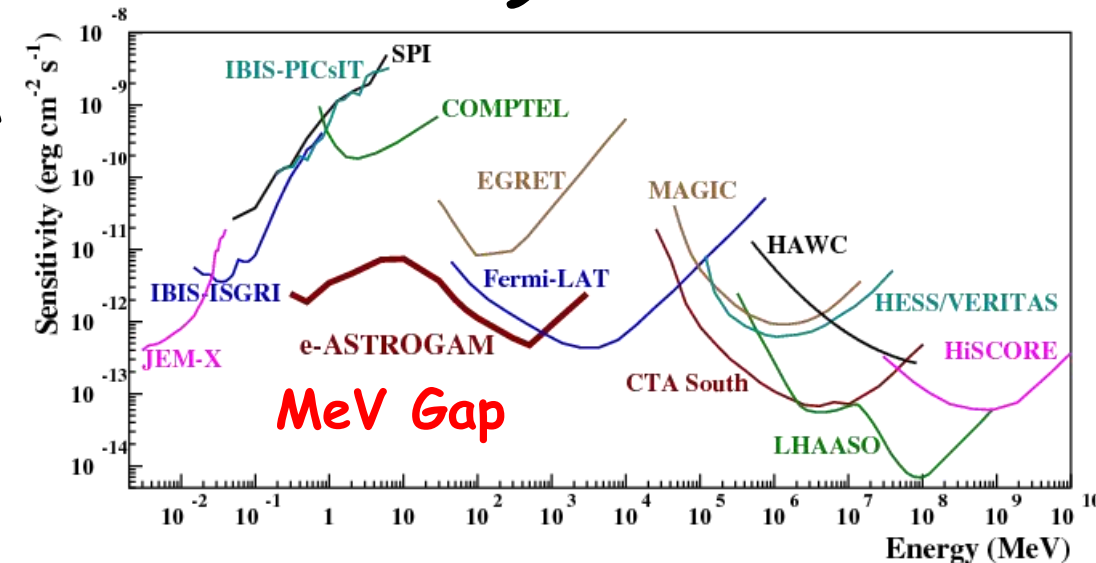
Possible parameter space



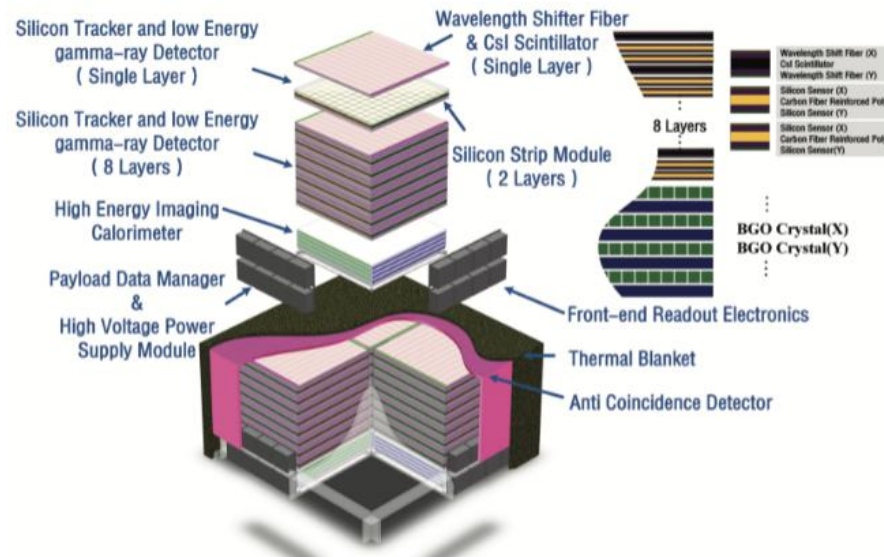
Parameter space is finite and we may be able to probe them ALL!

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VLAST-closing the MeV gap



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甚大面积伽马射线空间望远镜计划*

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(3 中国科学技术大学天文与空间科学学院 合肥 230026)

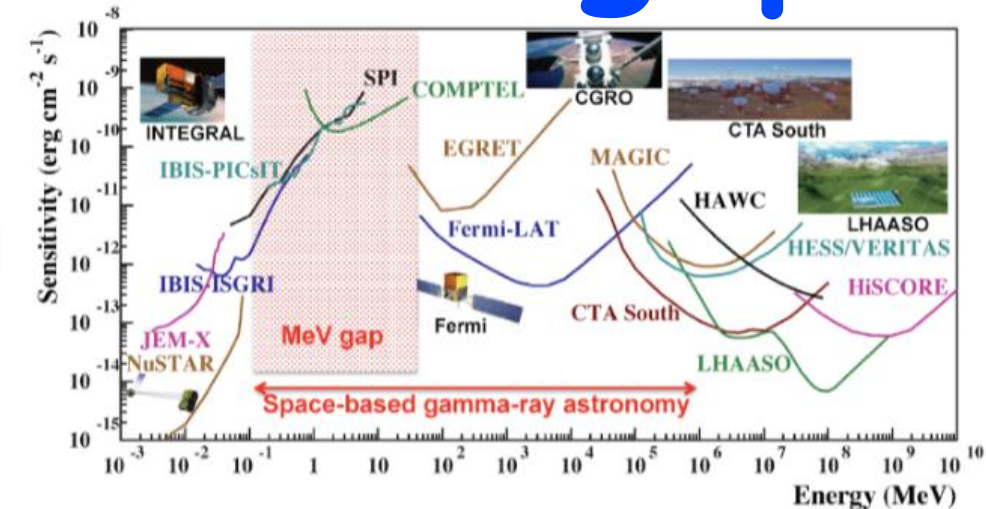
(4 核能与核电子技术国家重点实验室 中国科学技术大学 合肥 230026)

(5 中国科学技术大学近代物理系 合肥 230026)

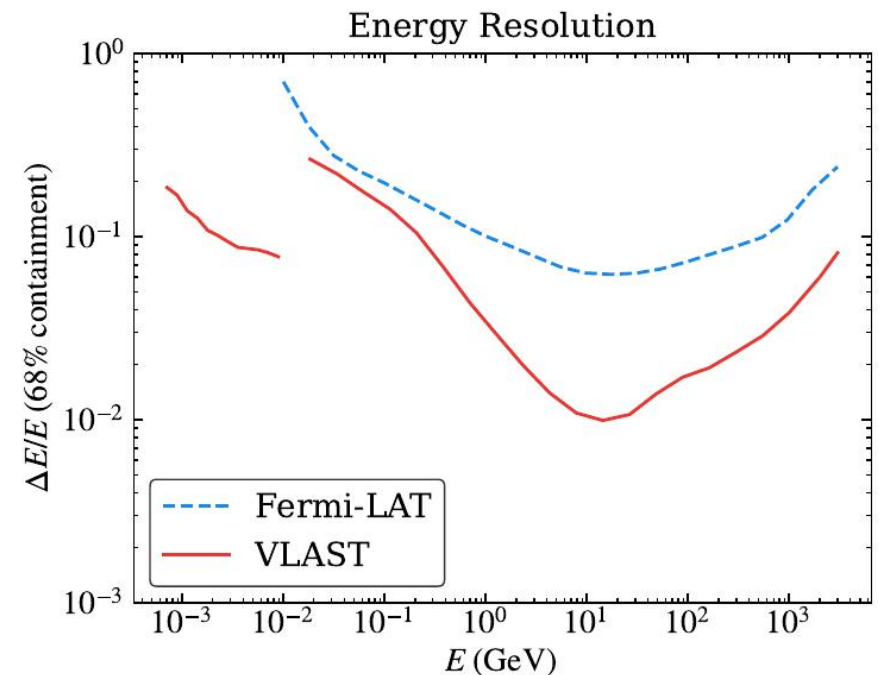
(6 中国科学院近代物理研究所 兰州 730000)

(7 临沂大学物理与电子工程学院 临沂 276000)

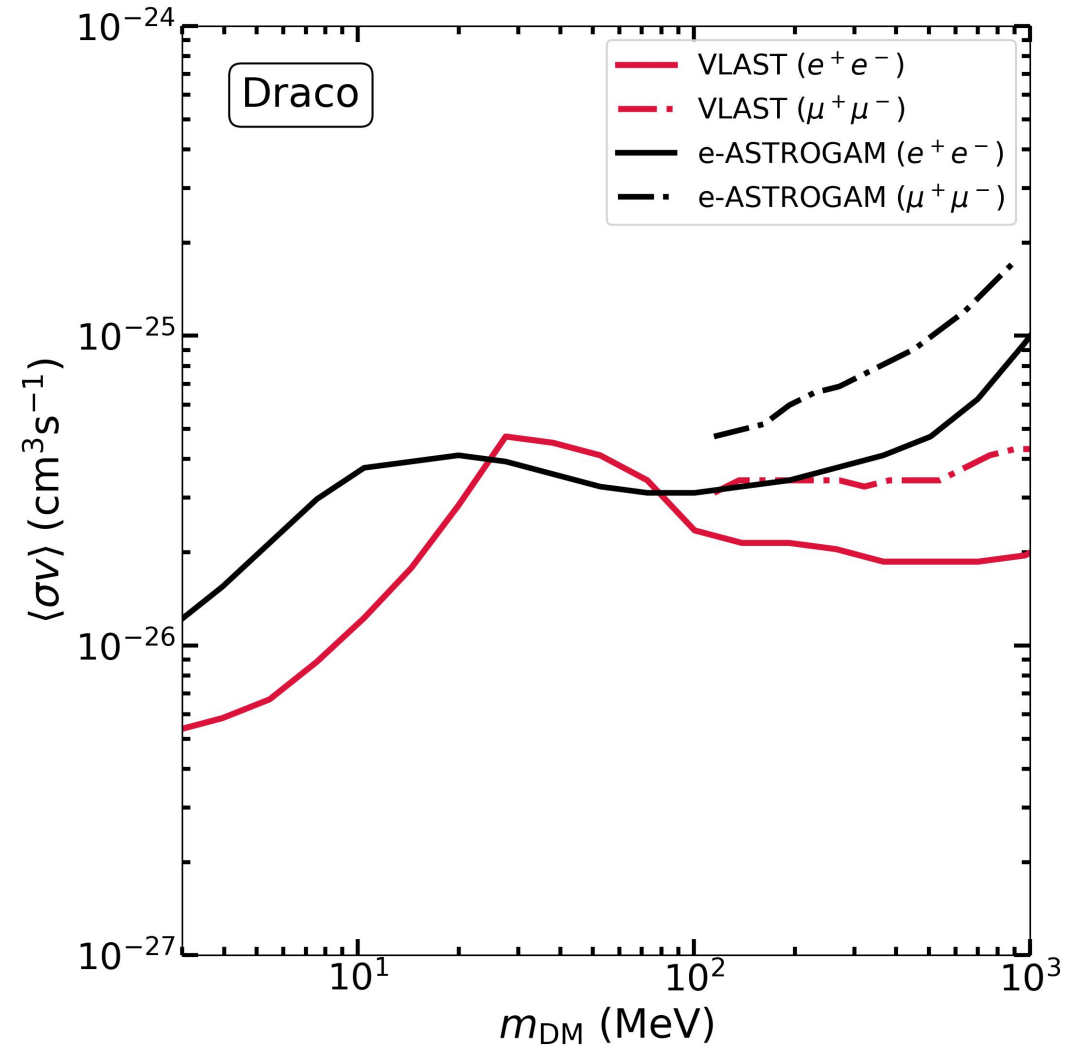
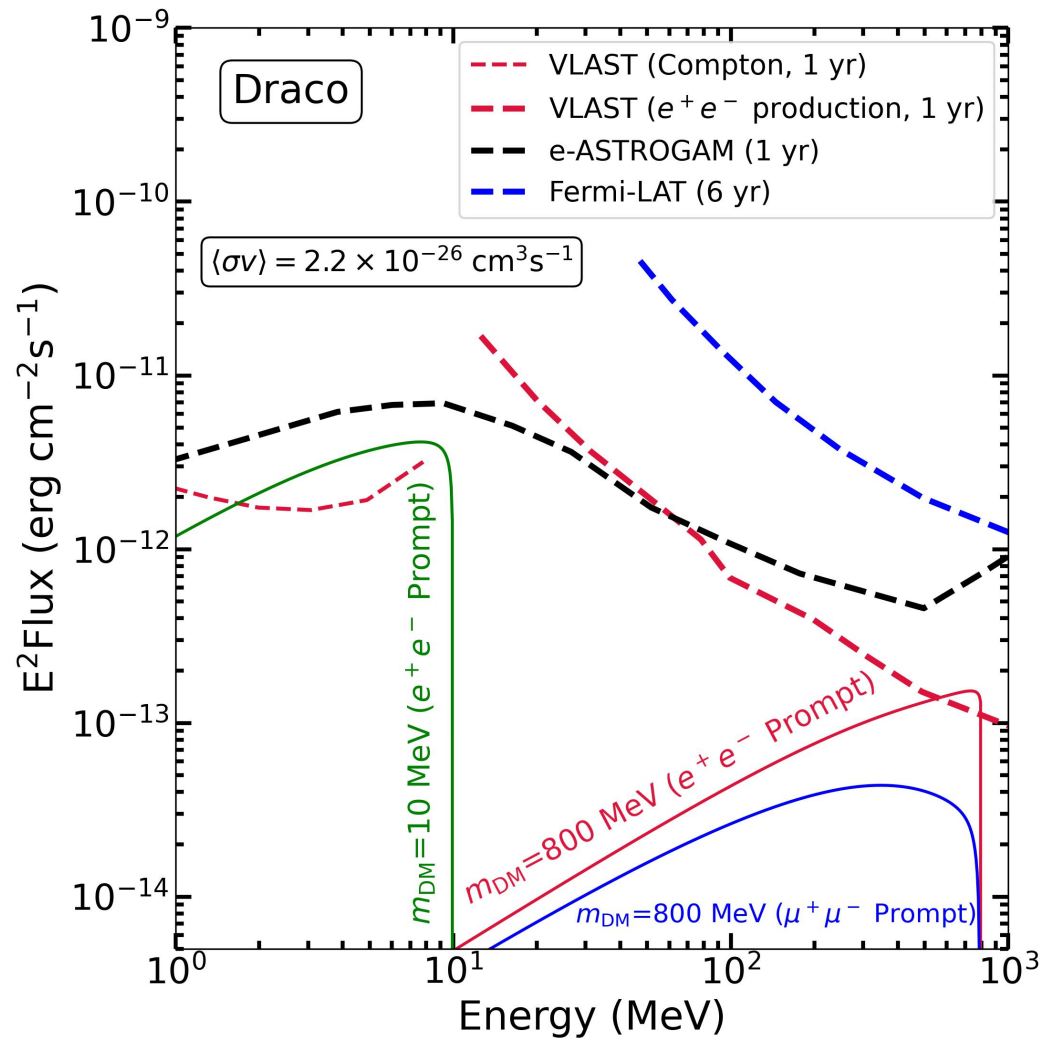
(8 中国科学院国家天文台 北京 100049)



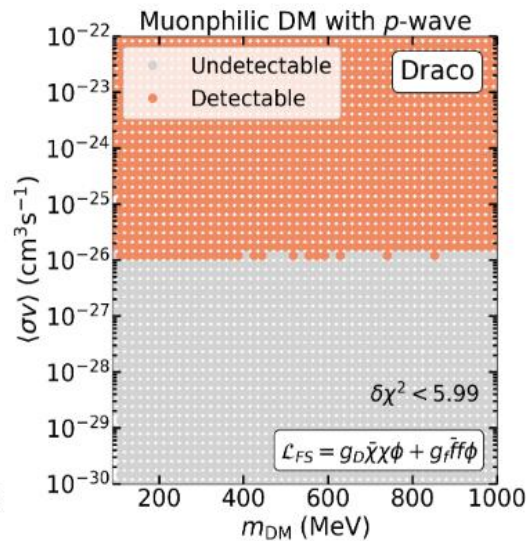
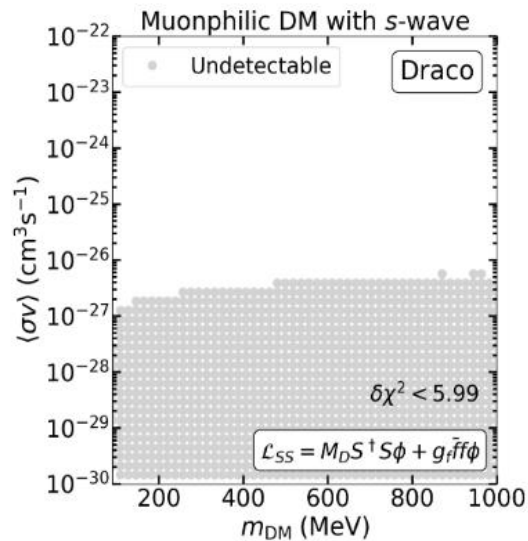
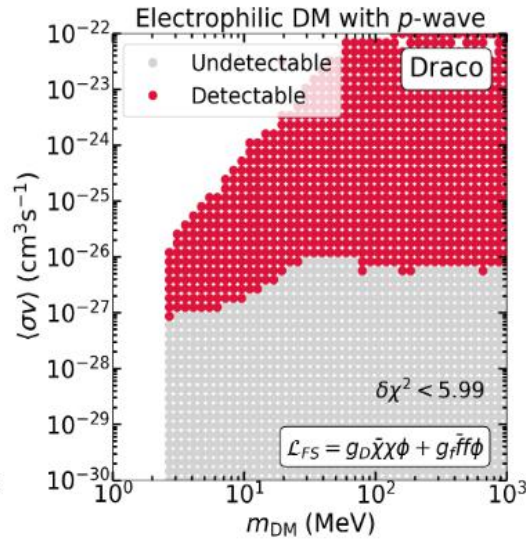
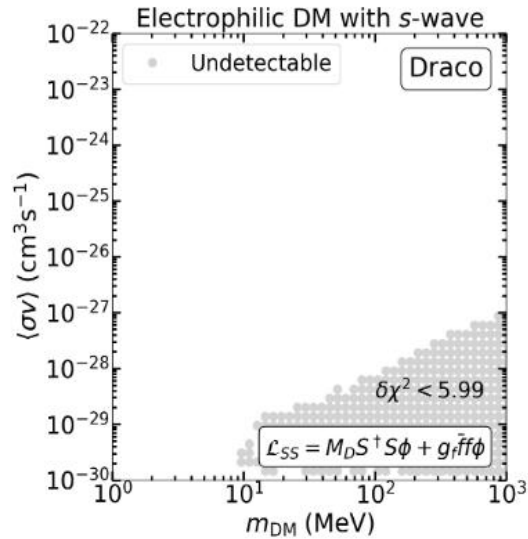
- Very Large Area gamma-ray Space Telescope (VLAST), the successor of DAMPE
- The first 10 m² sr level gamma-ray satellite (~20 tons)
- Leading the research on dark matter detection and time-domain astronomy based on MeV - TeV gamma-rays



Very Large Area Space Telescope (VLAST)



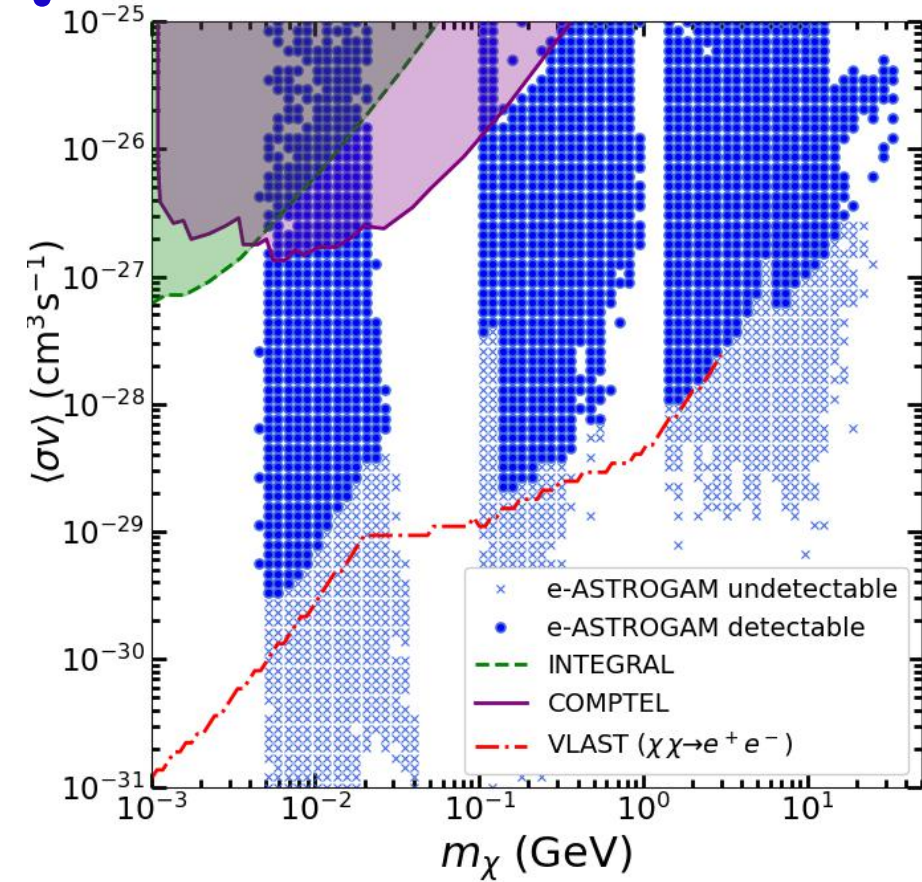
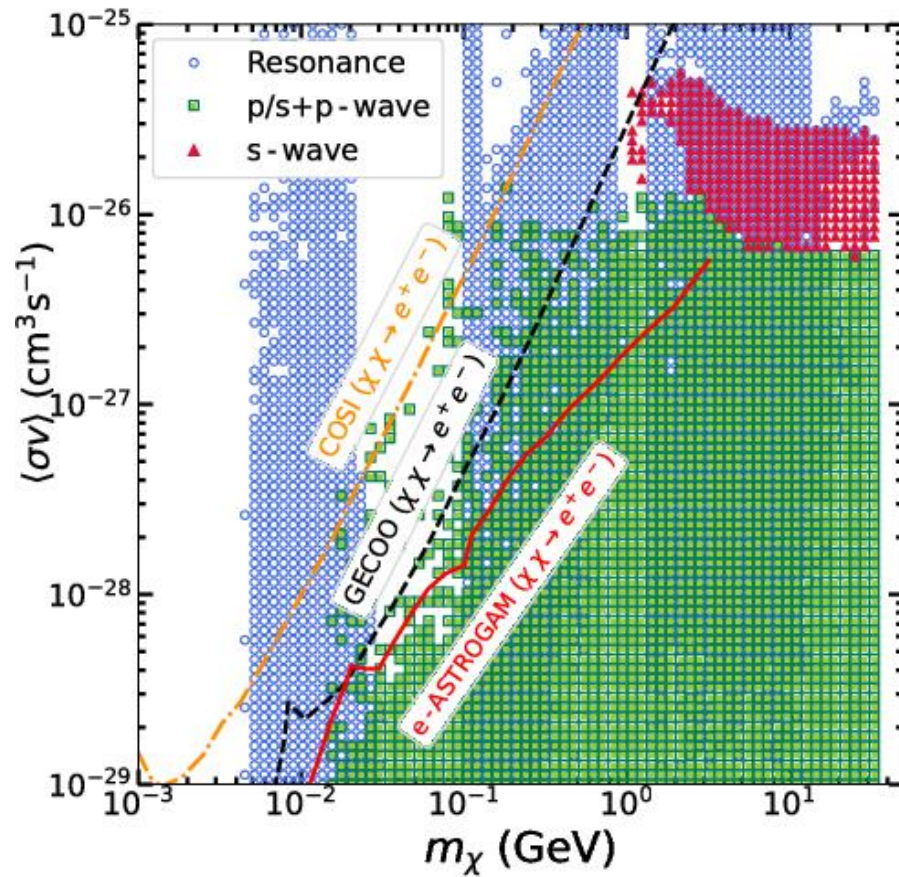
PROBING RESONANCE REGION BY VLAST (Leptonphilic)



Property	Value	Notes
Distance	$\sim 76 \text{ kpc}$	From the Sun.
Stellar Mass (M_*)	$\sim 3 \times 10^5 M_\odot$	Baryonic mass in stars.
Dynamical Mass (M_{dyn})	$\sim 5 \times 10^8 M_\odot$	Total mass (stars + dark matter).
J-factor (0.5°)	$\sim 10^{18.9} \text{ GeV}^2 \text{ cm}^{-5}$	Dark matter annihilation luminosity.
Mass-to-Light Ratio (Υ)	$\sim 1000 M_\odot / L_\odot$	Indicator of extreme dark matter dominance.

- Draco is an ideal source for DM gamma-ray detection.
- s-wave annihilation remains difficult to detect with future VLAST.

PROBING RESONANCE REGION BY VLAST (Higgs portal model)



Future DM indirect detection (like VLAST) can probe resonance DM.

Summary

- The light thermal DM has a lower mass limit around MeV.
- Direct detection can also constrain the low mass mediator mass region, but pseudoscalar can relax this tension.
- Pseudoscalar can generate s-wave annihilation which is testable in indirect detection.
- Considering CMB constraints, most of s-wave annihilation with mass below GeV is excluded, while the resonance for p-wave annihilations are still testable in future MeV gamma ray telescopes, e.g. **VLAST**.

Thank you for
listening and please
stay on VLAST!

