

# Fractionally charged particles (millicharged particles)

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# Outline

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1 Millicharge

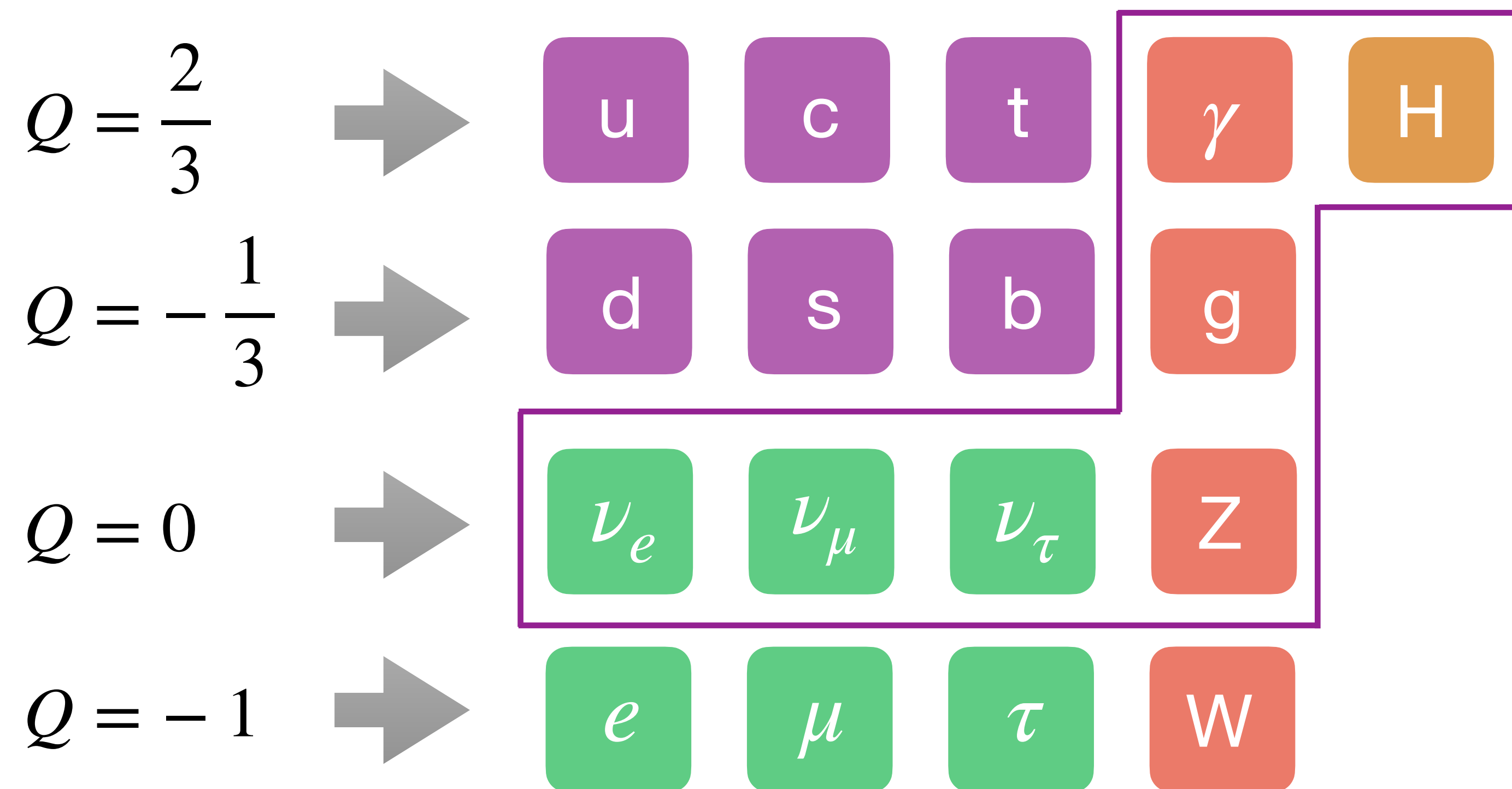
2 Theoretical models

3 Astro/cosmo probes

4 Terrestrial searches

# 1 Millicharge

# Millicharge is a very small electric charge



photon

↓

$$e Q_f A_\mu^\gamma \bar{f} \gamma^\mu f$$

$Q_f = \frac{n}{3}$  is quantized in SM

non-quantized  $Q_f \equiv \epsilon \ll 1$  → millicharge

# Magnetic monopole leads to charge quantization

Dirac's paper on magnetic monopole & charge quantization in 1931

Quantised singularities in the electromagnetic field, [from <https://inspirehep.net/>]

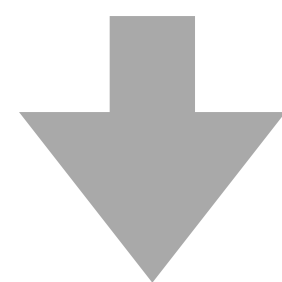
Paul Adrien Maurice Dirac (St John's Coll., Cambridge) (Sep 1, 1931)

Published in: *Proc.Roy.Soc.Lond.A* 133 (1931) 821, 60-72

 DOI  cite

 2,263 citations

magnetic monopole



charge quantization

$$e = \frac{2\pi\hbar n}{\mu}$$

$n$  is integer

$\mu$  is magnetic strength

$e$  is electric strength

# Standard model & grand unified theories

Anomaly cancellations in SM lead to charge relations

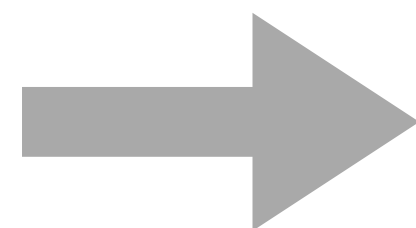
$$\sum_i Y_i = 0$$

But  $U(1)_Y$  in SM in general allows non-quantized charge

$$SU(3)_c \times SU(2)_L \times U(1)_Y$$

Embedded into a bigger gauge group? e.g.  $SU(5)$

[Georgi & Glashow, PRL.32.438, 1974]



charge quantization

# Stringent constraints on millicharge in SM

$$Q_p - Q_e < (0.8 \pm 0.8) \times 10^{-21} e$$

[Marinelli et al. 1984]

$$Q_n < (-0.1 \pm 1.1) \times 10^{-21} e$$

[Bressi et al. 2011]

$$Q_n < (-0.4 \pm 1.1) \times 10^{-21} e$$

[Baumann et al. 1988]

$$Q_\nu < 10^{-17} e$$

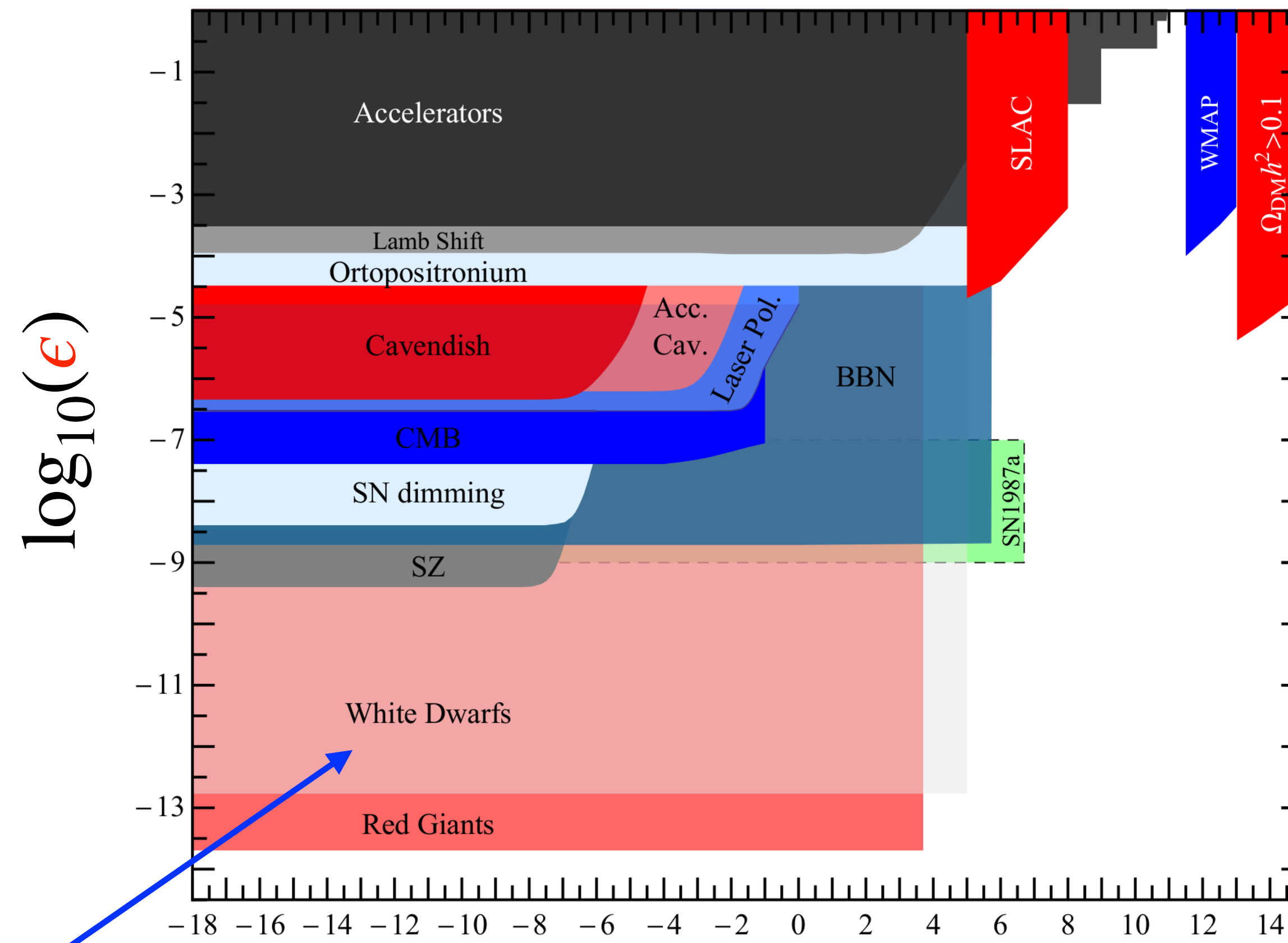
[Barbiellini et al. 1987]

[Siemensen et al., PRD 97, 052004 (2018)]

# Millicharge in BSM can be quite “large”

[Jaeckel & Ringwald, 1002.0329]

$$e \epsilon A_\mu^\gamma \bar{\chi} \gamma^\mu \chi$$



terrestrial

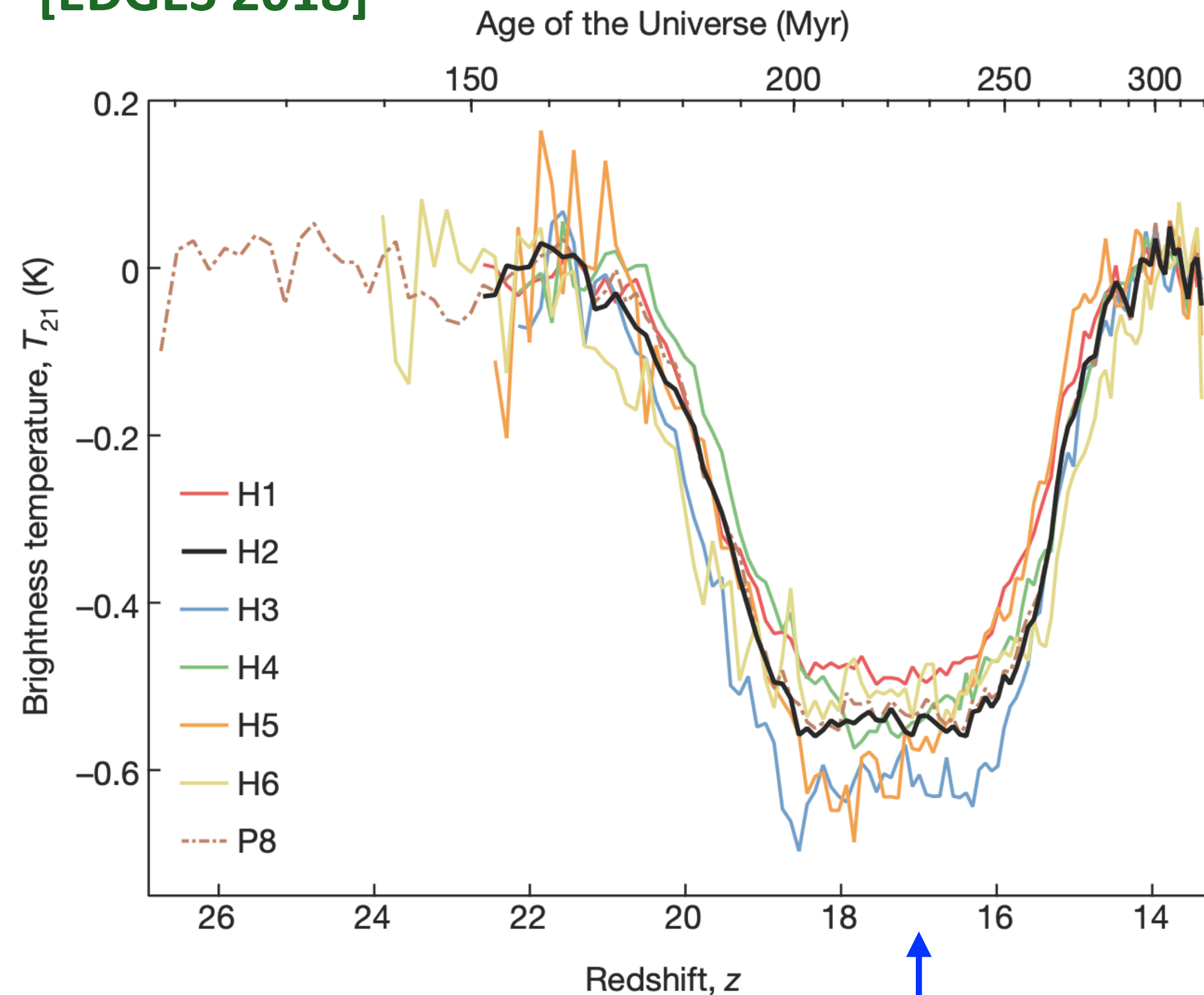
astro/cosmo

$\log_{10}(m_\chi/\text{eV})$



# Millicharged DM can explain EDGES 21 cm anomaly

[EDGES 2018]



colder gas or hotter CMB than expected

DM can cool gas

Millicharged DM xsec

$$\sigma \propto v^{-4}$$

easier to evade early universe constraints

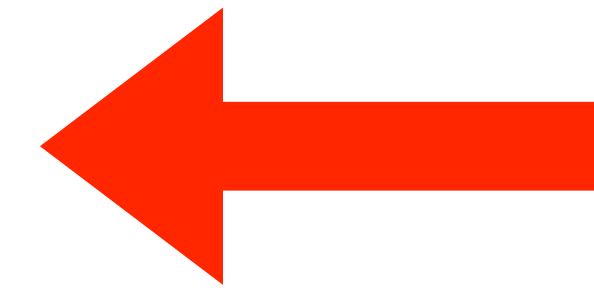
## 2 theoretical models

# Models to generate millicharged particles (MCPs)

$$e \epsilon A_{\mu}^{\gamma} \bar{\chi} \gamma^{\mu} \chi \quad \longrightarrow \quad \chi \text{ is MCP}$$

Kinetic mixing

Mass mixing



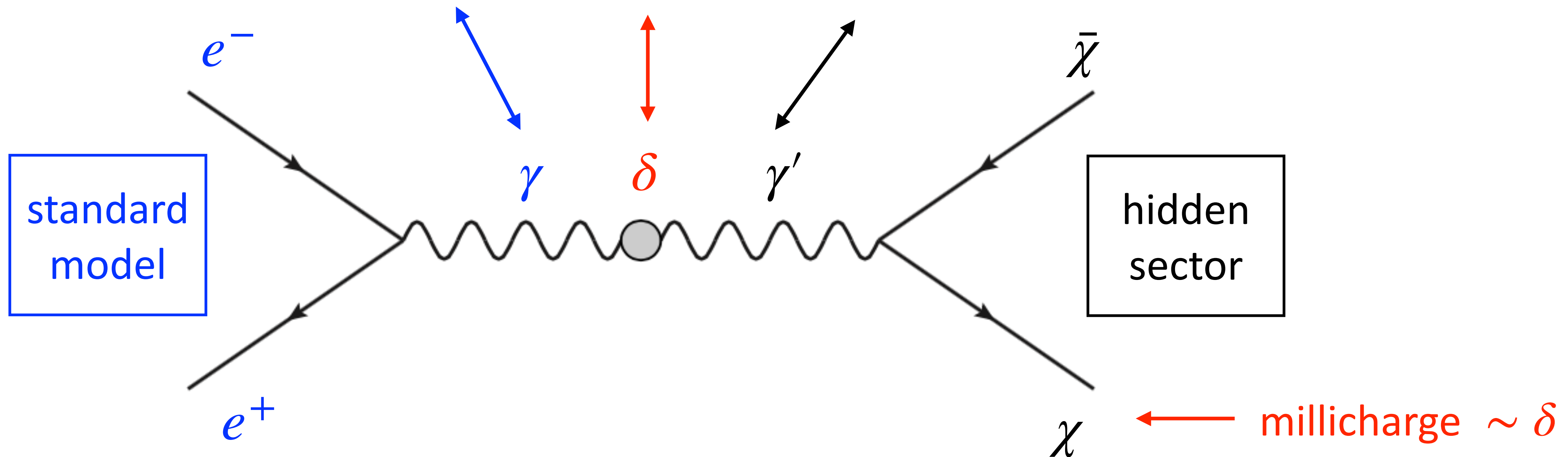
# Kinetic mixing between two U(1) fields

Kinetic mixing between two U(1) fields:  $A_1$  &  $A_2$

$$\mathcal{L} = -\frac{1}{4}F_{2\mu\nu}F_2^{\mu\nu} - \frac{\delta}{2}F_{2\mu\nu}F_1^{\mu\nu} - \frac{1}{4}F_{1\mu\nu}F_1^{\mu\nu}$$

[Holdom, Phys.Lett. 166B, 196 (1986)]

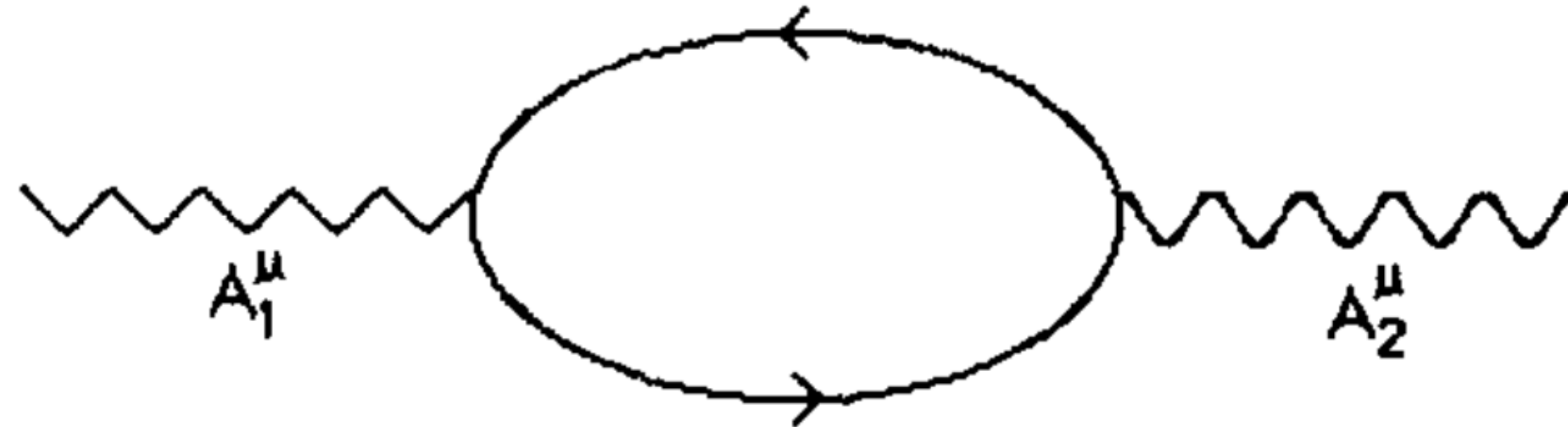
[Foot & He, Phys. Lett. 267B, 509 (1991)]



# One example to generate kinetic mixing

[Holdom, Phys.Lett. 166B (1986) 196-198]

2 high scale fermions charged under two U(1)s



$$\mathcal{L} = -\frac{1}{4} F_{2\mu\nu} F_2^{\mu\nu} - \frac{1}{4} F_{1\mu\nu} F_1^{\mu\nu} - \frac{\delta}{2} F_{2\mu\nu} F_1^{\mu\nu}$$

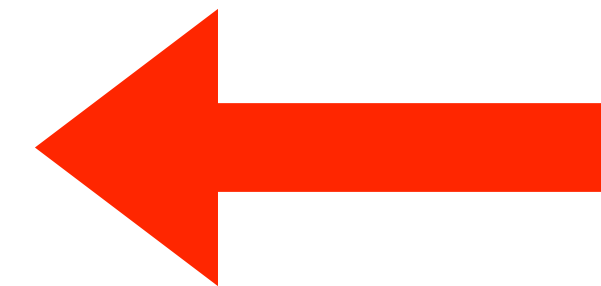
$$\delta = \frac{e_1 e_2}{6\pi^2} \ln \frac{m'_{12}}{m_{12}}$$

# Models to generate millicharged particles (MCPs)

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Kinetic mixing

Mass mixing



# Stueckelberg mass mixing between two U(1) fields

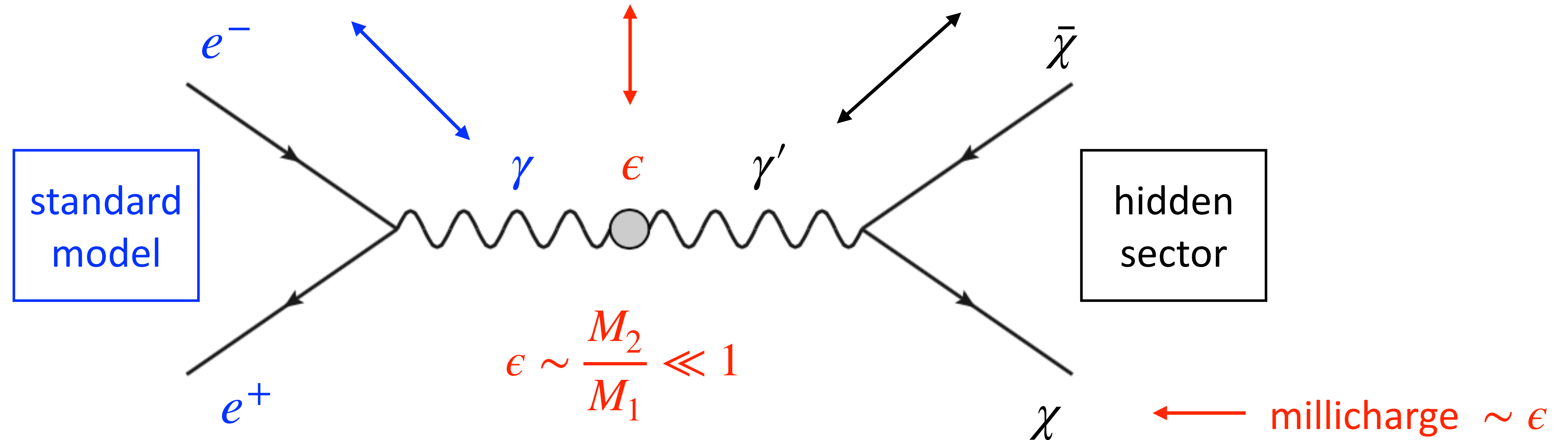
Stueckelberg mass mixing between two U(1) fields:  $A_1$  &  $A_2$

[Kors & Nath hep-ph/0402047]

[Feldman, ZL, Nath, hep-ph/0603039]

[Cheung & Yuan hep-ph/0701107]

$$\mathcal{L} = -\frac{1}{4}F_{2\mu\nu}F_2^{\mu\nu} - \frac{1}{2}(\partial_\mu\sigma + M_1A_{1\mu} + M_2A_{2\mu})^2 - \frac{1}{4}F_{1\mu\nu}F_1^{\mu\nu}$$



# “Toy” model w/ both kinetic & mass mixings

$$-\frac{1}{4}F_{1\mu\nu}F_1^{\mu\nu} - \frac{1}{4}F_{2\mu\nu}F_2^{\mu\nu} - \frac{\delta}{2}F_{1\mu\nu}F_2^{\mu\nu} + J'_\mu A_1^\mu + J_\mu A_2^\mu - \frac{1}{2}M_1^2 A_{1\mu}A_1^\mu - \frac{1}{2}M_1^2 \epsilon^2 A_{2\mu}A_2^\mu - M_1^2 \epsilon A_{1\mu}A_2^\mu$$

after field redefinition & rotation  $\longrightarrow$  canonical ( $-F^2/4$ ) & diagonal mass matrix

degeneracy

dark photon

SM photon

$$\frac{1}{\sqrt{1 - 2\delta\epsilon + \epsilon^2}} \left( \frac{\epsilon - \delta}{\sqrt{1 - \delta^2}} J_\mu + \frac{1 - \delta\epsilon}{\sqrt{1 - \delta^2}} J'_\mu \right) A_M^\mu + \frac{1}{\sqrt{1 - 2\delta\epsilon + \epsilon^2}} \left( J_\mu - \epsilon J'_\mu \right) A_\gamma^\mu$$

mass mixing  $\longrightarrow$  millicharge

[Feldman, ZL, Nath, hep-ph/0702123, 309 cites]

[see also, Fabbrichesi et al., 2005.01515, The Dark Photon [review], p 7-9]



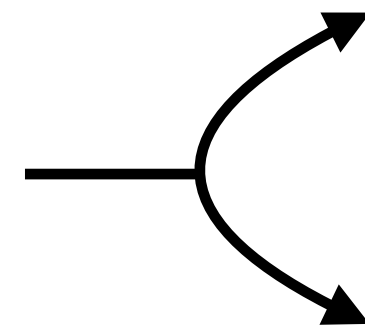
# Kinetic versus mass mixings (in “realistic” model)

common features

hidden boson mix w/ hypercharge  
(instead of photon)

$\implies$  3 by 3 mass matrix

hidden boson



photon-like (dark photon)

$$m_{A'} \ll m_Z$$

hypercharge-like

$$m_{A'} \gg m_Z$$

difference

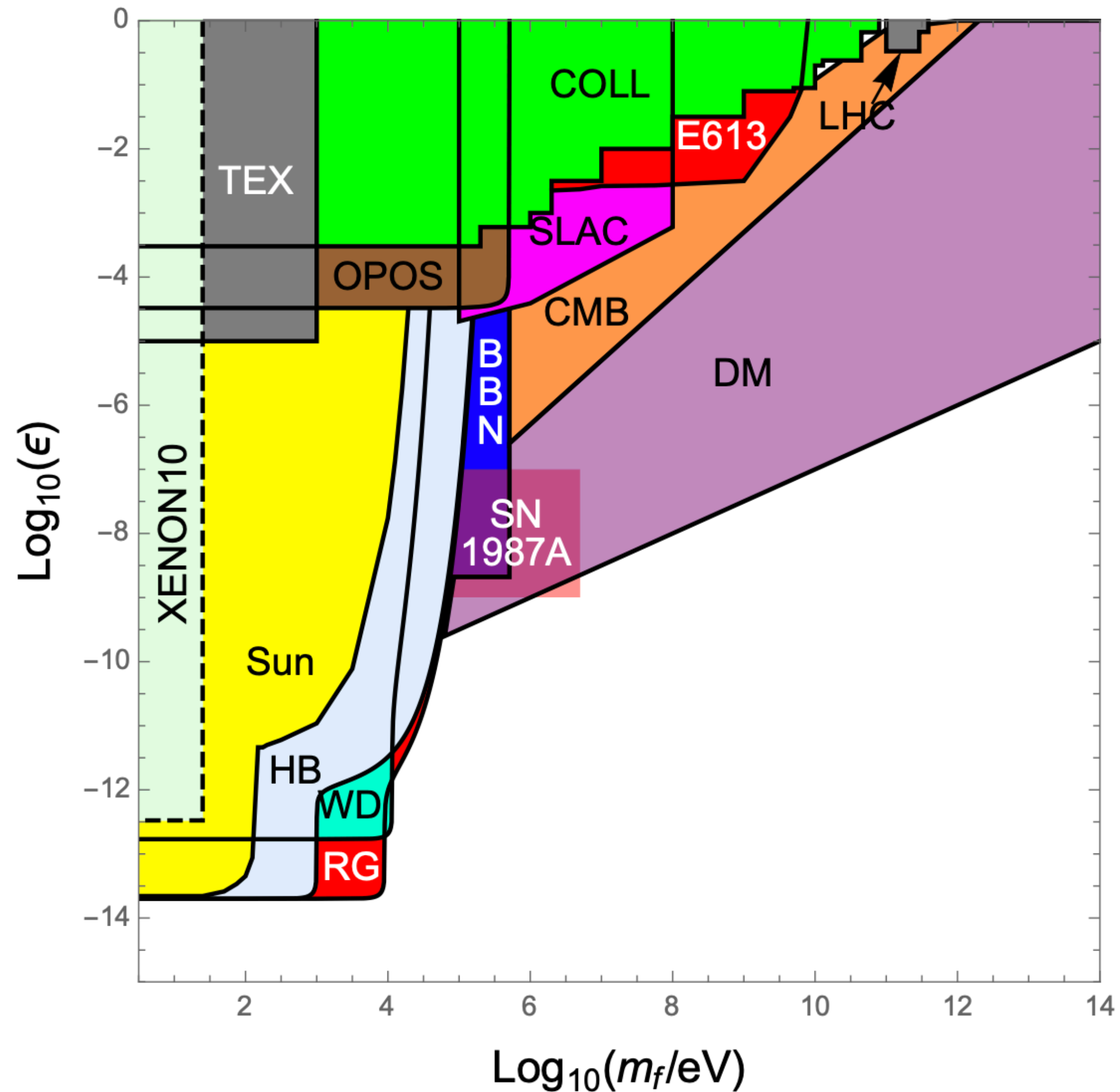
mass mixing  $\implies$  millicharge & massive dark photon

kinetic mixing  $\implies$  millicharge & massless dark photon

[Feldman, ZL, Nath, hep-ph/0702123, 309 cites]

# 3 astro/cosmo probes

# Energy loss argument in star evolutions



White dwarfs

Supernova

Red giant stars

Horizontal branch stars

Sun

[see e.g. Vinyoles, Vogel, 1511.01122]

# A small fraction of millicharged DM is allowed by CMB

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[Dubovsky et al, hep-ph/0311189v2]      WMAP data on the CMB anisotropy:  $\Omega_\chi h^2 < 0.007$

[de Putter et al, 1805.11616]

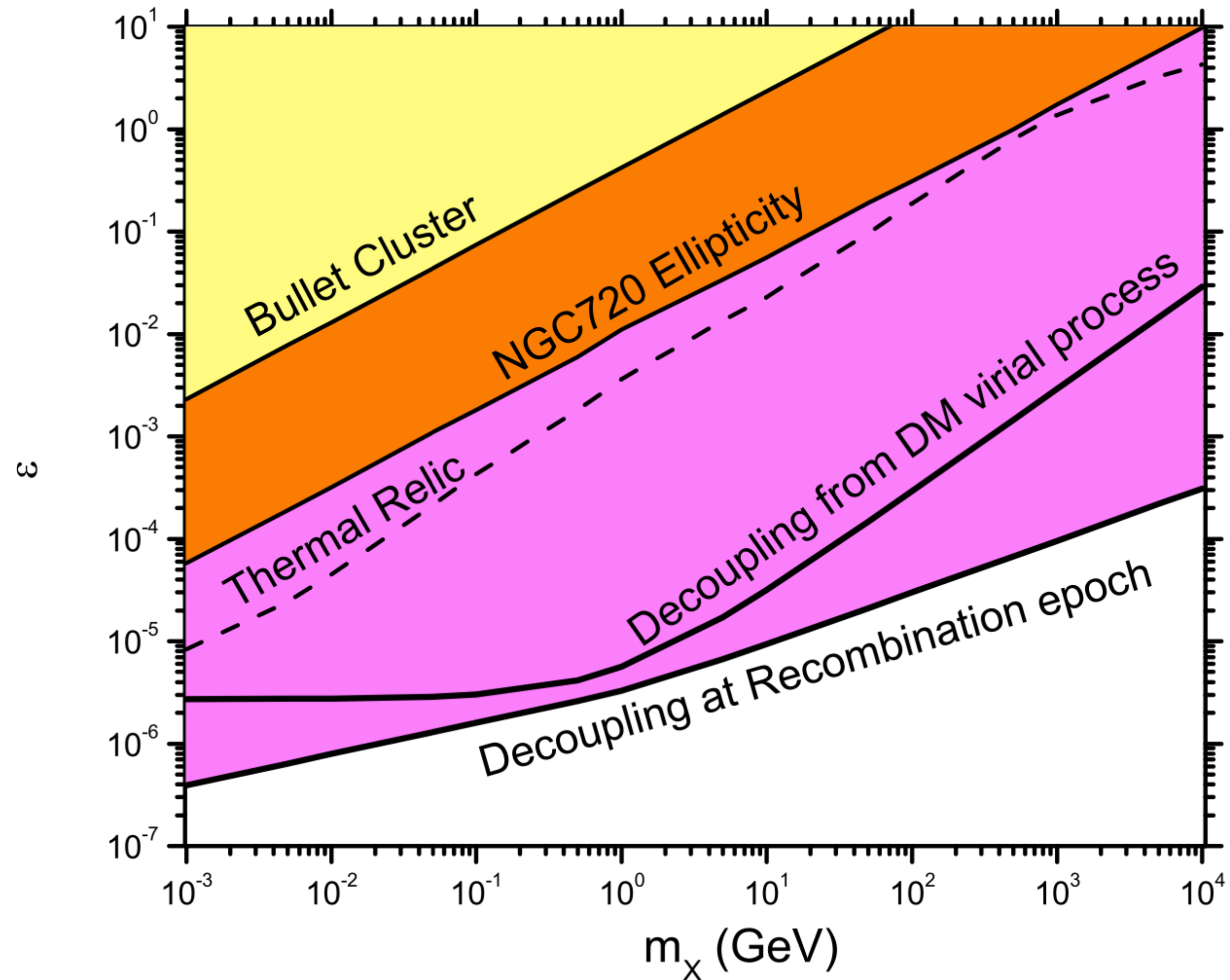
[Kovetz et al, 1807.11482]

[Boddy et al 1808.00001]      The Planck limits vanish if millicharged DM is less than **0.4%**;  
DM essentially becomes cosmologically **indistinguishable**  
from a small additional amount of **baryons**.

# Or millicharge has to be very small to not disturb CMB

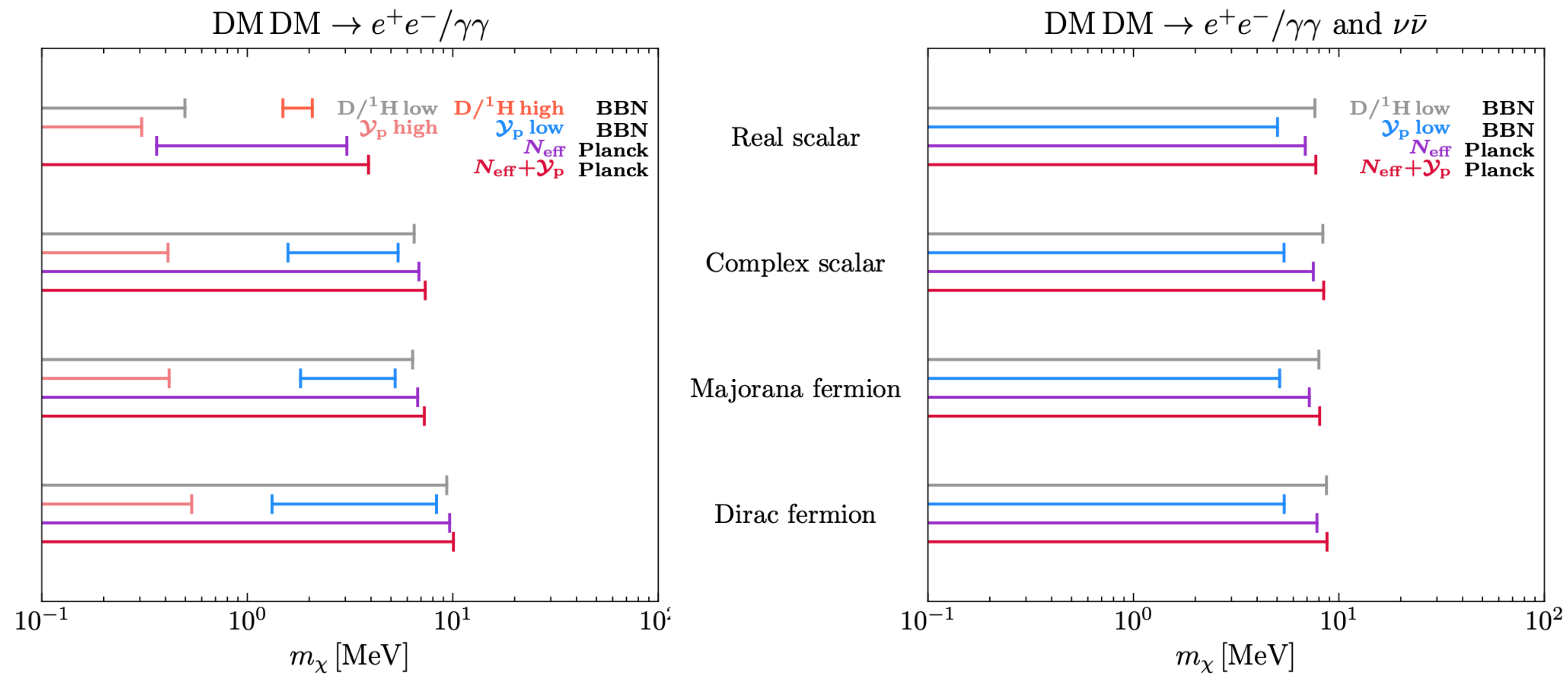
[McDermott, Yu, Zurek, 1011.2907]

completely decouple MCPDM  
from baryon at recombination



# $N_{\text{eff}}$ constraints

light MCP and/or annihilations  $\rightarrow N_{\text{eff}}$  neutrinos  $\rightarrow$  constrained by BBN & CMB



[Depta et al, 1901.06944]

4

# terrestrial searches

# Precision tests of QED

[Dobroliubov, Ignatiev, PRL 1990]

electron/muon  $g-2$  (2-loop)

Lamb shift (1-loop)

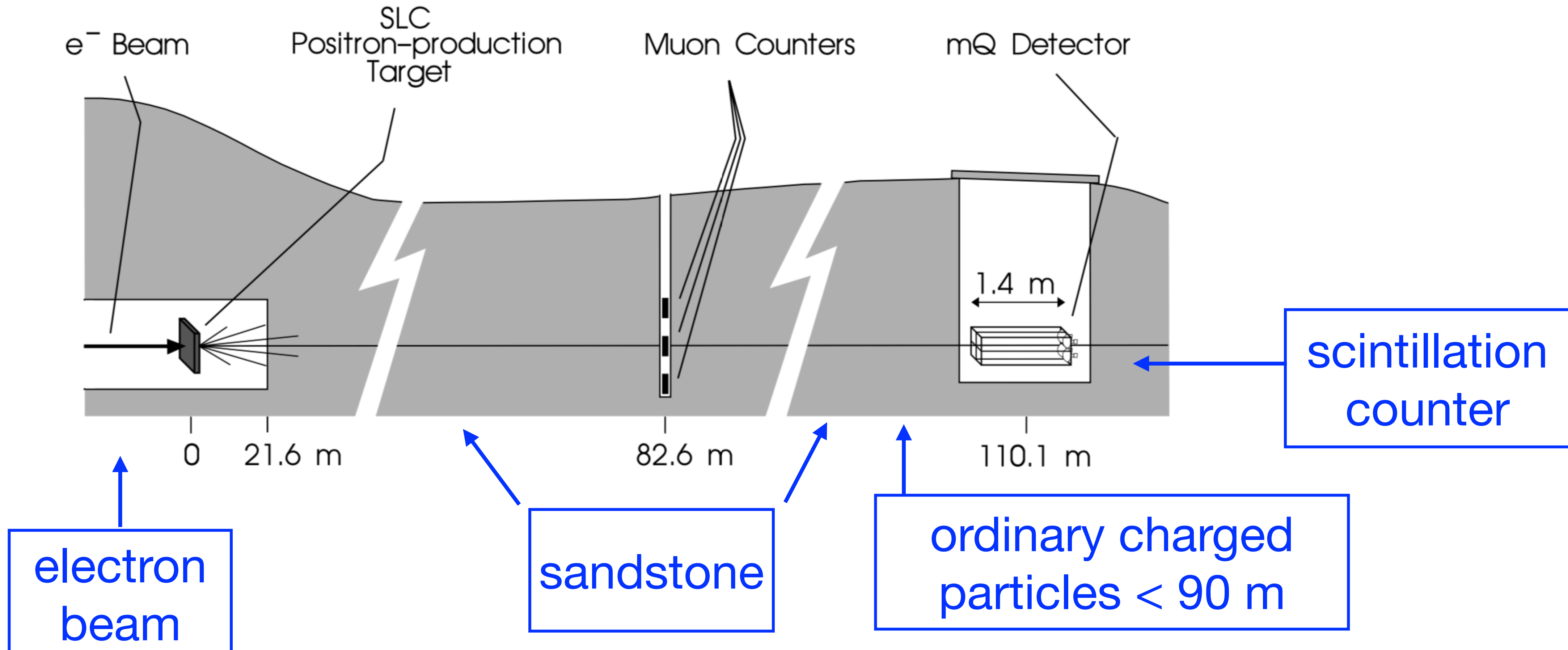
orthopositronium decay

but QED precision limits are weak compared to others



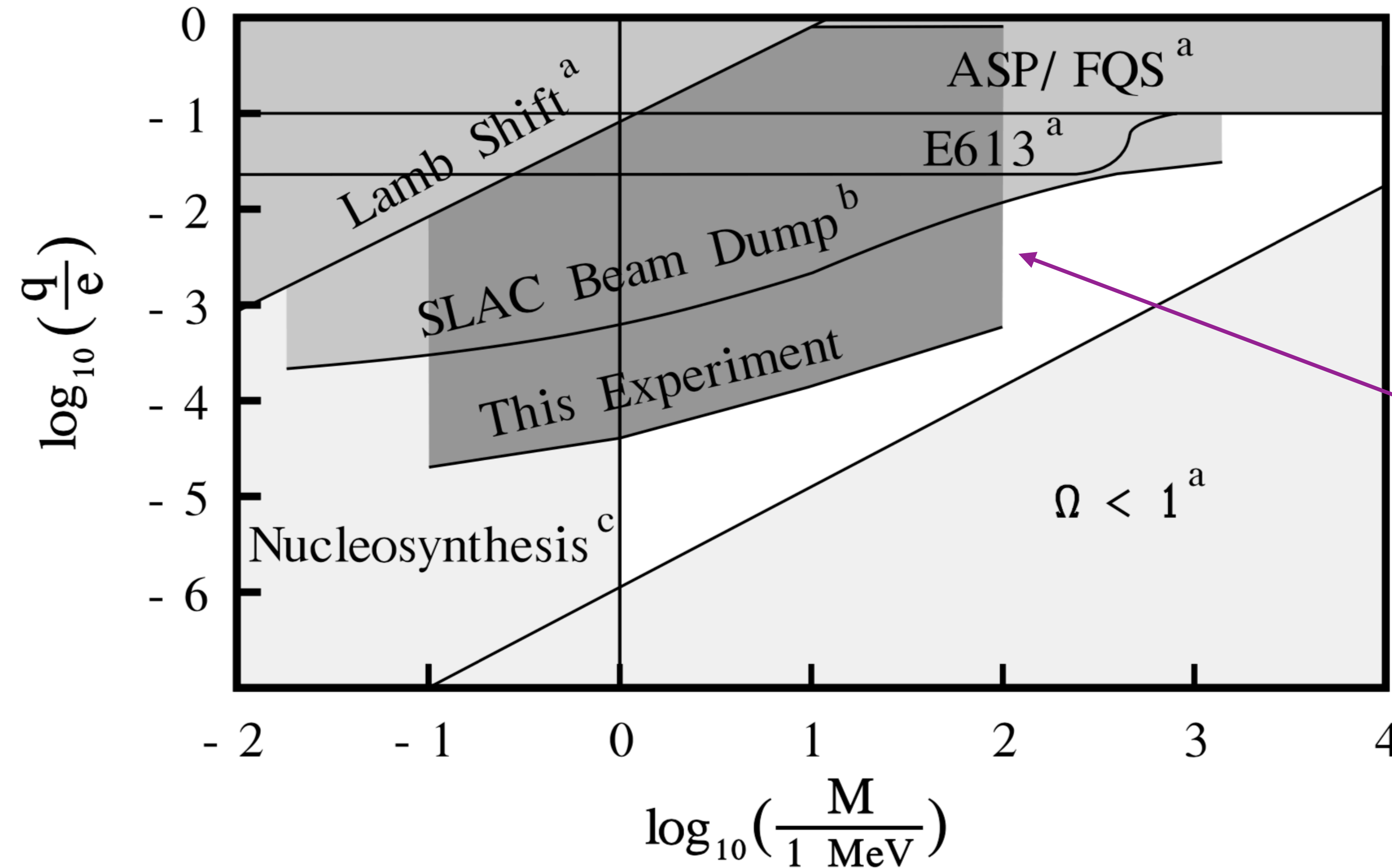
# SLAC mQ: electron beam dump experiment

[Prinz et al, PRL 81.1175, 1998]



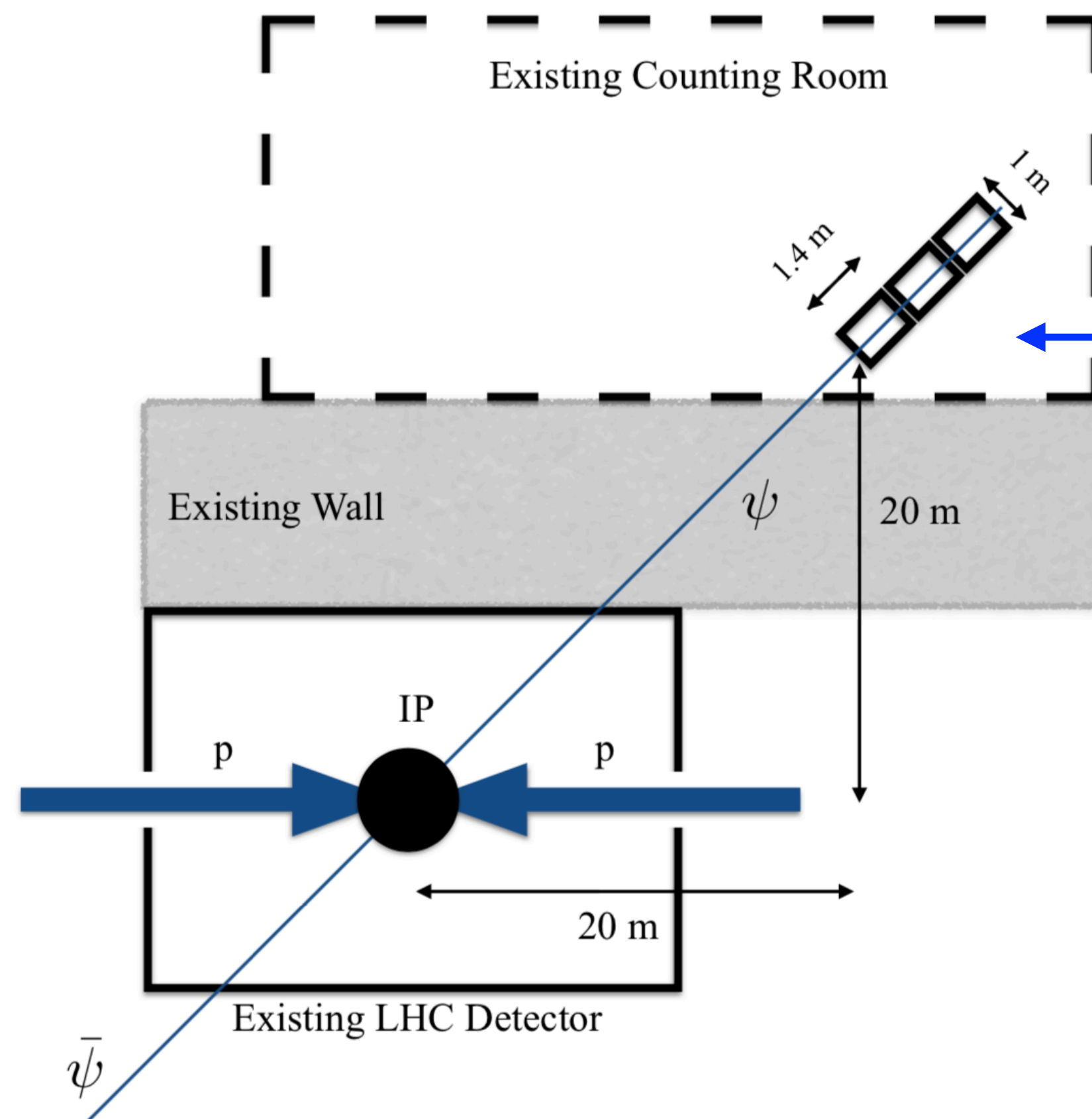
# SLAC mQ constraints on millicharge

[Prinz et al, PRL 81.1175, 1998]



$m > 100 \text{ MeV}$   
MCP yield die off

# MilliQan: proposed scintillation detector @ LHC

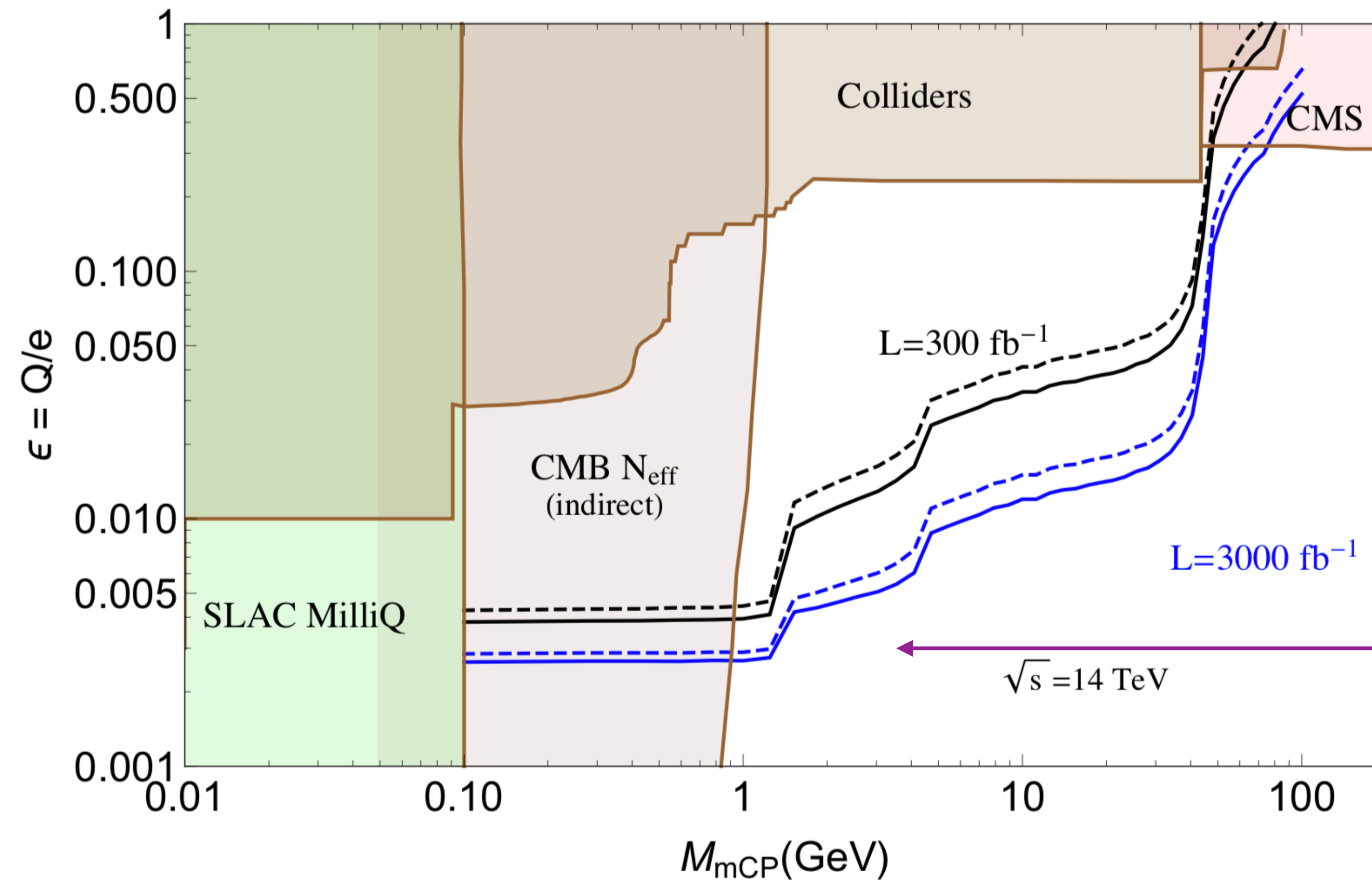


1 m × 1 m × 3 m plastic scintillators array with 3 sections (pointing to CMS IP), each containing 400 5 cm × 5 cm × 80 cm scintillator bars coupled to PMT. 33 m away from CMS IP.

[Haas et al., 1410.6816]

[Ball et al., 1607.04669]

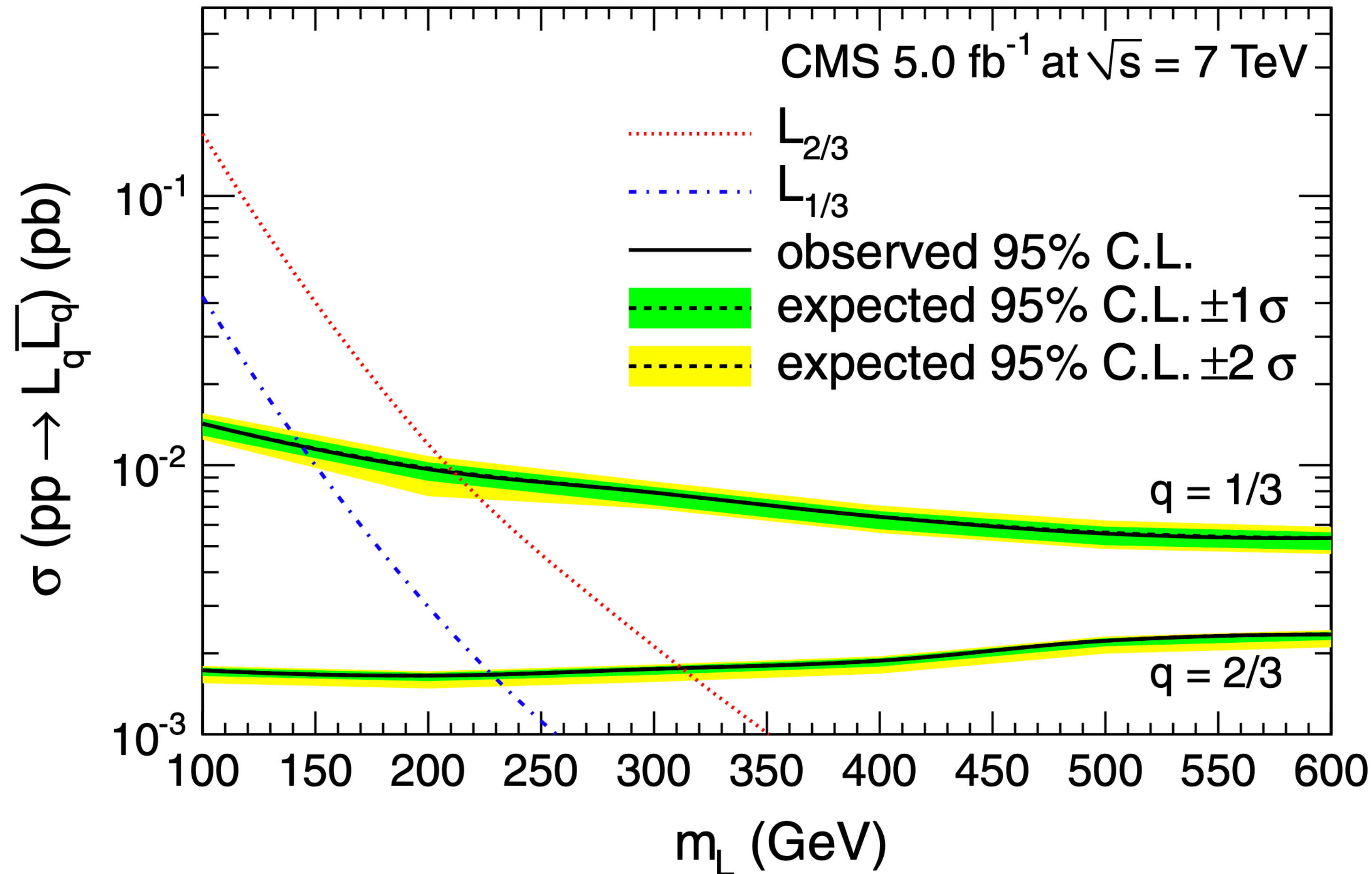
# MilliQan sensitivity on millicharged particles



[Ball et al., 1607.04669]

photoelectron drops  
below one when  
 $\epsilon \lesssim O(10^{-3})$

# CMS searches on fractional charged particles

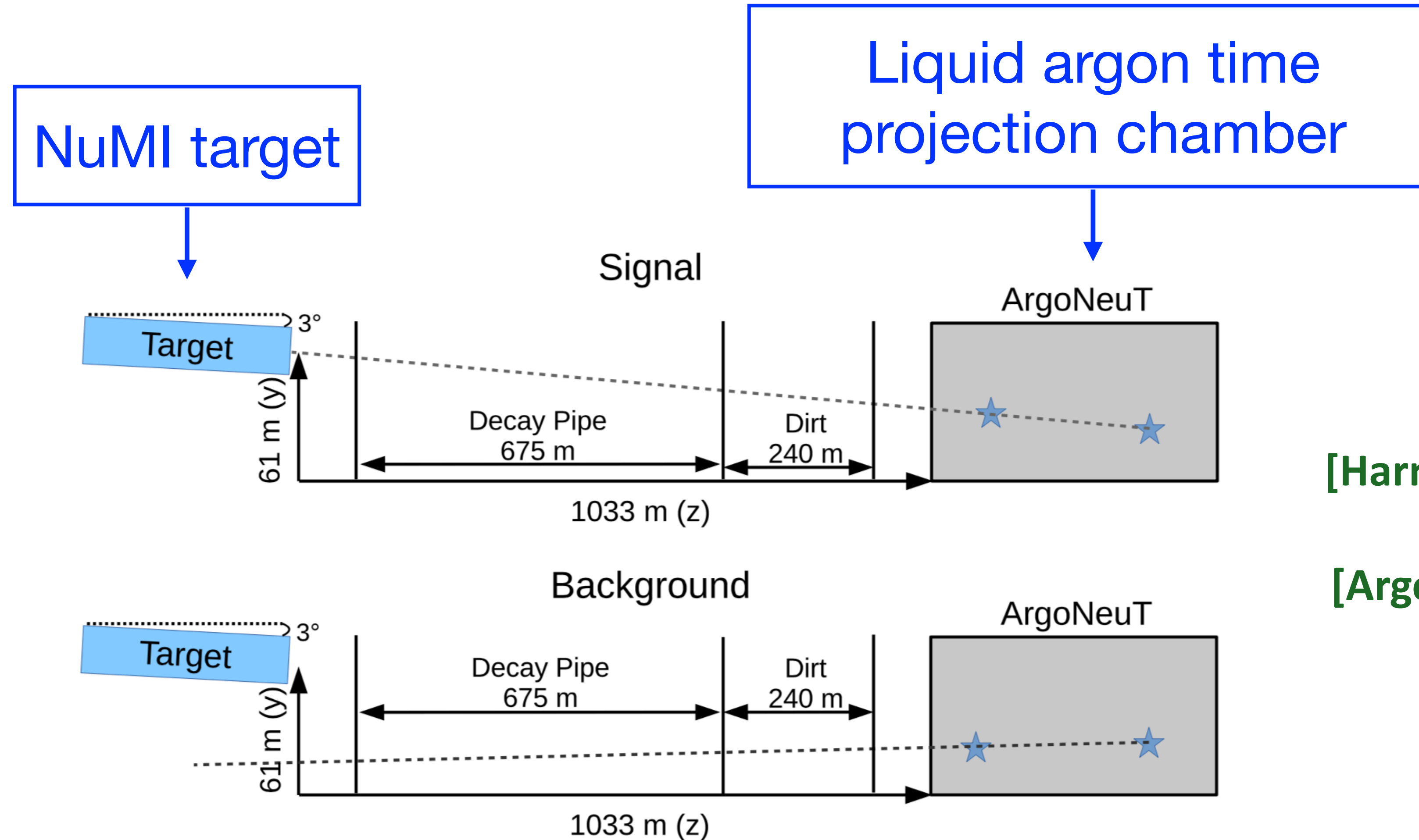


[CMS, 1210.2311]

exclude mass below 310  
(140) GeV for  $2e/3$  ( $e/3$ )

low-charge tracks in  
silicon tracking detector

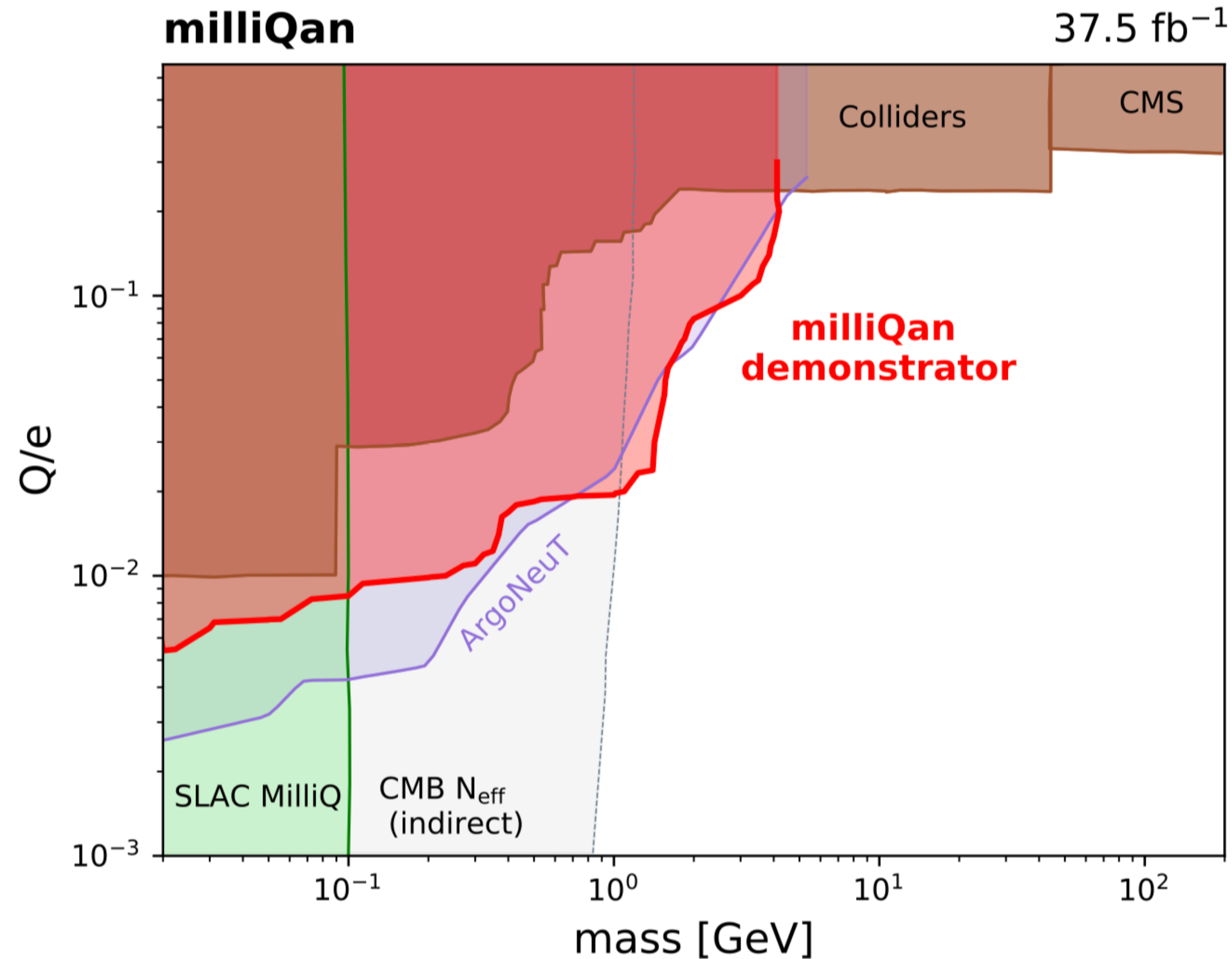
# ArgoNeuT @ FermiLab: liquid argon neutrino detector



[Harnik et al., 1902.03246]

[ArgoNeuT, 1911.07996]

# Limits from ArgoNeuT & MilliQan demonstrator (1%)



[ArgoNeuT, 1911.07996]

ArgoNeuT 10<sup>20</sup> POT

[milliQan demonstrator, 2005.06518]

1% of total detector

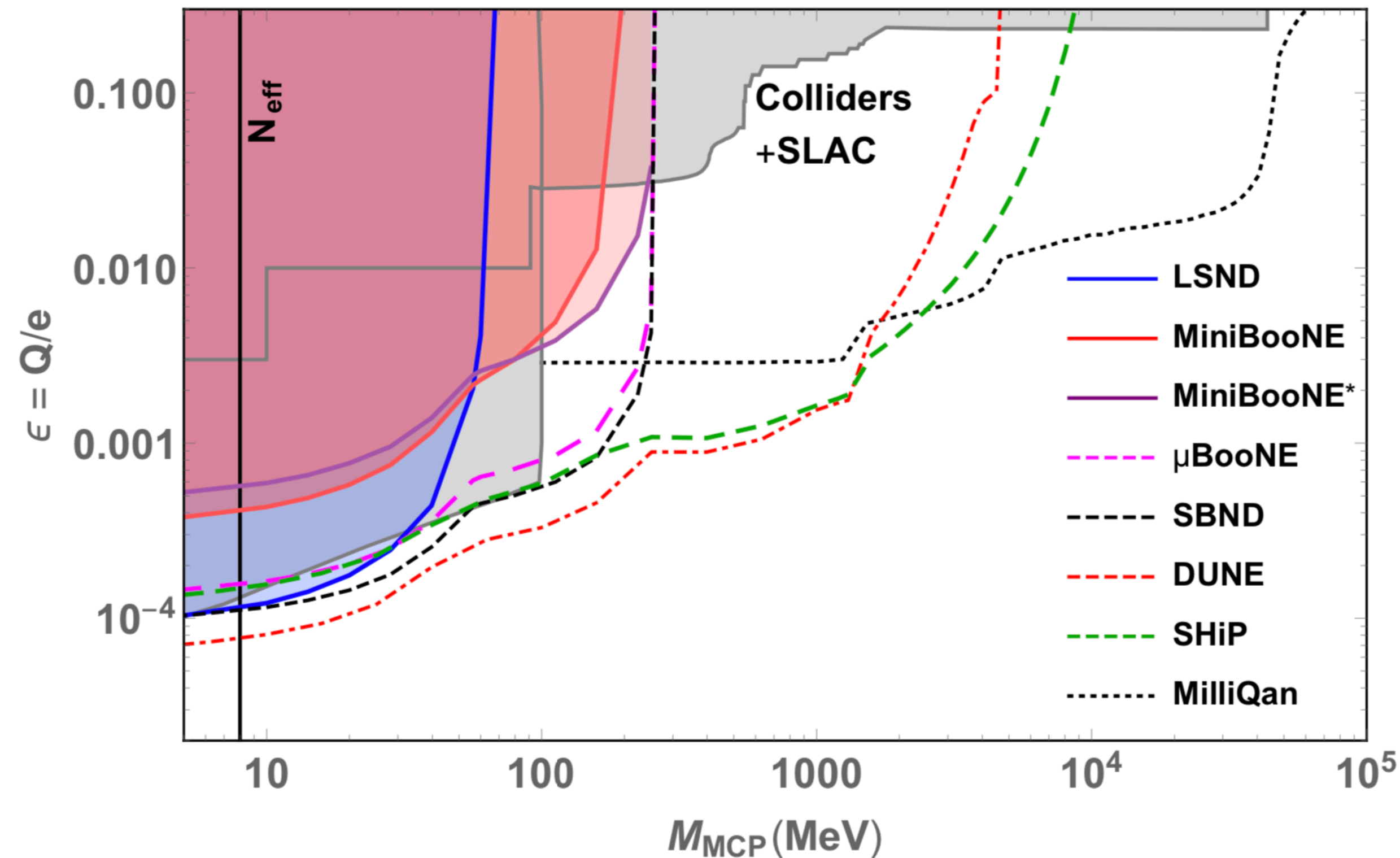
37/fb in 2018

# MCP scatters w/ neutrino experiment target

MCP produced in proton fixed-target experiments



MCP scatters w/ target (electron) in the neutrino detector



[Magill et al., 1806.03310]

solid: data  
dashed: projection



# Future experiments to probe millicharge

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LDMX: electron fixed target

Dark Shine: fixed target

NA64: muon fixed target

FerMINI: scintillators @ proton fixed target

SUBMET: scintillators @ JPARC

neutrino experiments

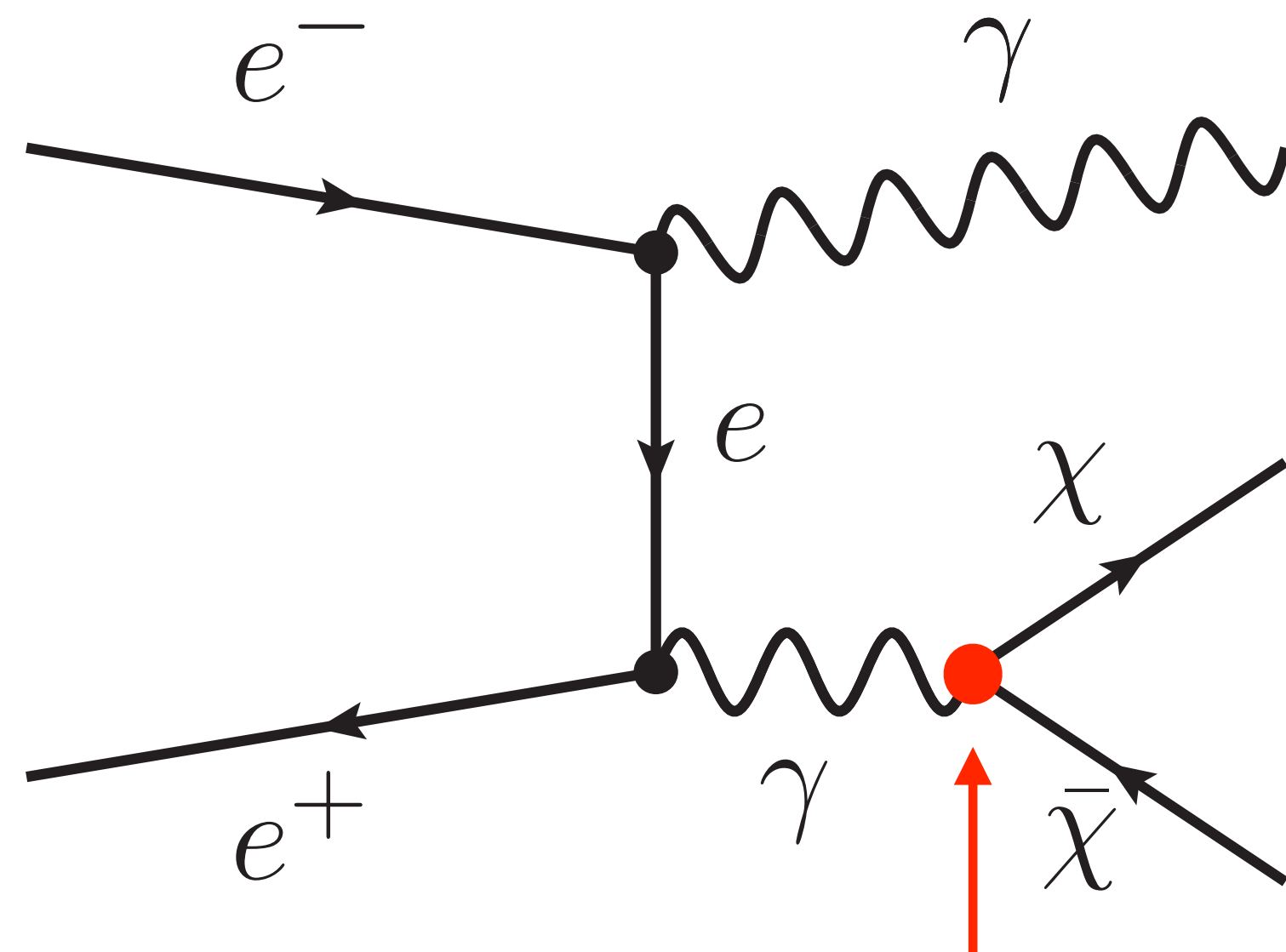
DM experiments

# Search for millicharged particles at electron colliders

very small ionization  
signal from millicharge

mono-photon @  
electron colliders

[ZL, Zhang, 1808.00983]

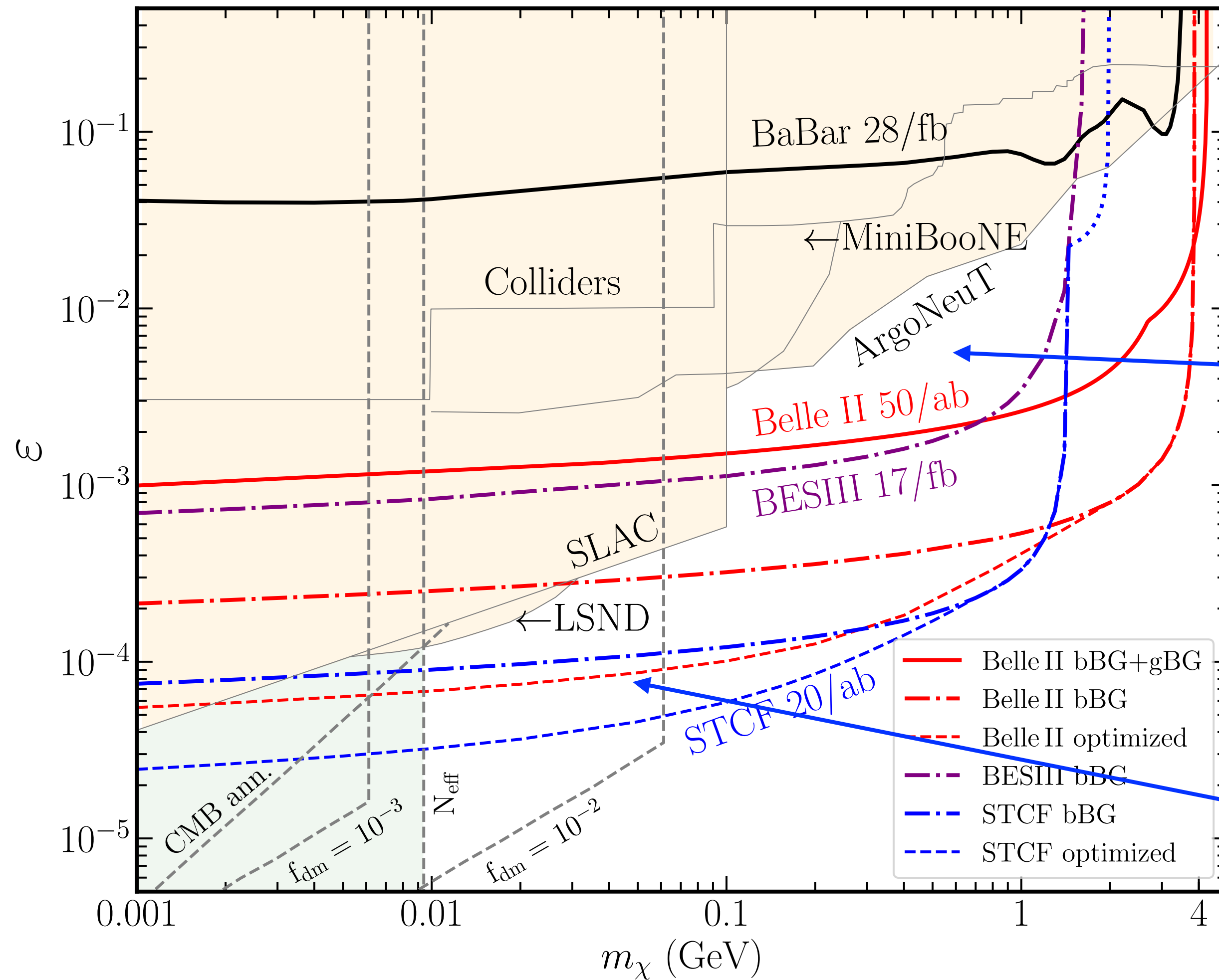


monophoton

“invisible” at  
electron colliders

millicharge vertex:  $e \epsilon A_{\mu}^{\gamma} \bar{\chi} \gamma^{\mu} \chi$

# Electron colliders can probe new parameter regions



diff BG considerations

BaBar

BESIII

new regions: 0.1-4 GeV

Belle II

STCF

$\epsilon \lesssim O(10^{-4})?$

[Liang, ZL, Ma, Zhang, 1909.06847]

# Summary

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Millicharged particles can naturally appear in [kinetic mixing](#) or [Stueckelberg mass mixing](#) models

Millicharged particles can be searched for in astrophysics/cosmology processes, precision tests of QED, and accelerator experiments

A number of [terrestrial experimental methods](#) have been carried out and recently proposed to search for millicharged particles