

Searching for fractionally charged particles with DAMPE

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On behalf of the DAMPE collaboration



Shanghai Particle and Cosmology Symposium 2021:

Emerging Frontiers of Axion,
Dark Photon,
Fractional Charged Particle
and MonoPole

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State Key Laboratory of Particle Detection and Electronics
University of Science and Technology of China

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Tsung-Dao Lee Institute @ SJTU, Shanghai, China

The 2021 Shanghai Particle Physics and Cosmology Symposium (SPCS2021)

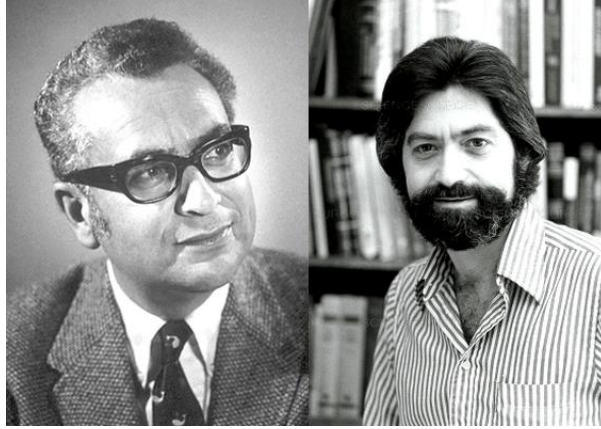
Outline

- Motivation
- Previous searches of FCP
- DAMPE experiment
- Search for FCP with DAMPE
- Summary

Motivation



Oil Drop Experiment



Quark Model

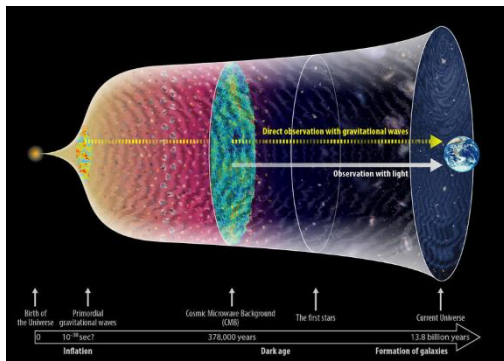


QCD Theory

- Since the oil drop experiment performed in 1909, all particles are measured as having charges of **multiples of electron charge**.
- In 1964, **quark model** for hadrons was proposed by Gell-man and Zweig.
- Due to the **QCD theory**, the quarks will not exist freely.
- Fractionally Charged Particle (FCP) is supposed to carry **any non-integer** charge.

The possible origins of FCP

The basic assumption: **FCP is a kind of heavy lepton**



Early universe



Supernova explosion



Extensive air shower

There are three possible sources of FCP in cosmic rays:

- **First**, it may be produced at the early Universe after the Big Bang.
- **Second**, it may be produced through high-energy astrophysical processes.
- **Third**, it may be produced in the extensive air shower of cosmic-rays.

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The previous experiments of large volume

For searches of **underground**, the target FCP should have energy above **hundreds GeV** to penetrate the rocks

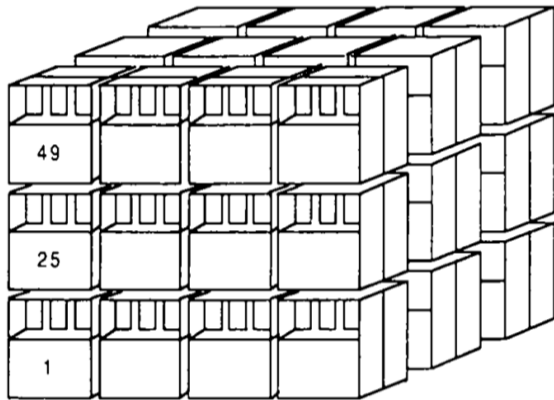


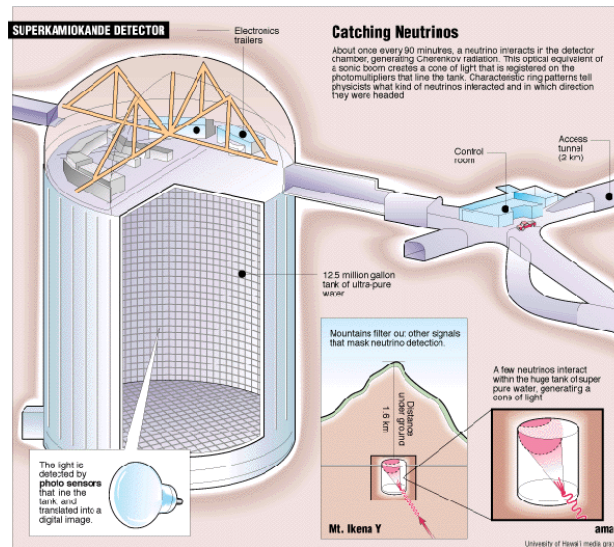
Fig. 1. The LSD experimental detector. The 72 tanks are considered as divided into 24 vertical columns (e.g. tanks 1–25–49 form the first telescope).

LSD 1800 m

$$\Phi\left(\frac{1}{3}\right) = 2.3 \times 10^{-13} \text{ cm}^{-2} \text{sr}^{-1} \text{s}^{-1}$$

$$\Phi\left(\frac{2}{3}\right) = 2.7 \times 10^{-13} \text{ cm}^{-2} \text{sr}^{-1} \text{s}^{-1}$$

2021/11/17

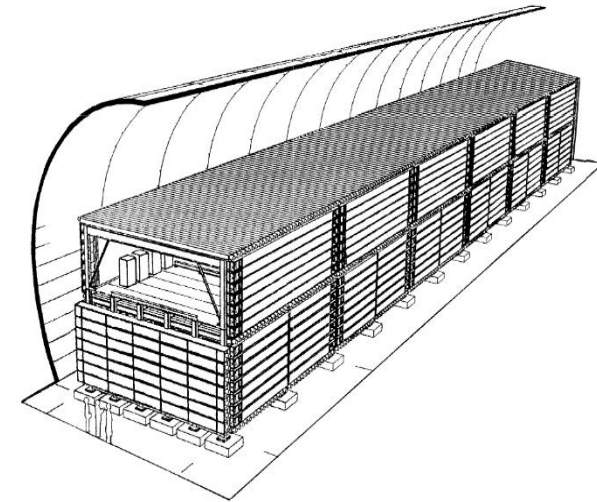


Kamiokande II 1000 m

$$\Phi\left(\frac{1}{3}\right) = 2.1 \times 10^{-15} \text{ cm}^{-2} \text{sr}^{-1} \text{s}^{-1}$$

$$\Phi\left(\frac{2}{3}\right) = 2.3 \times 10^{-15} \text{ cm}^{-2} \text{sr}^{-1} \text{s}^{-1}$$

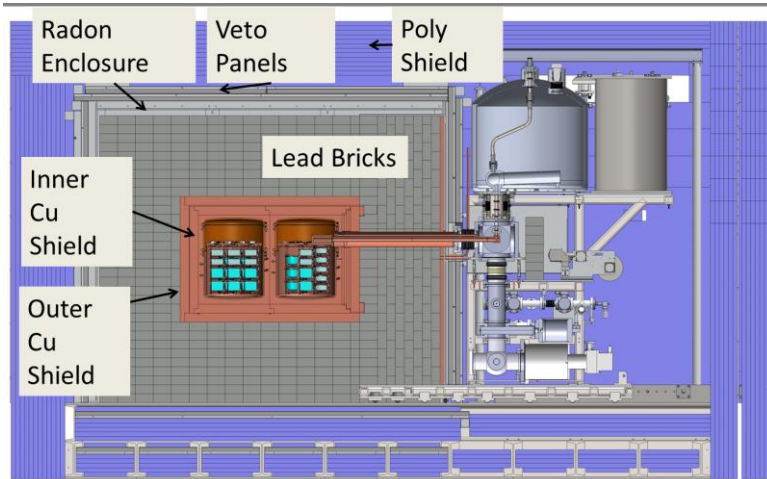
SPCS2021



MARCO 1400 m

$$\Phi\left(\frac{1}{4} \sim \frac{2}{3}\right) = 6.1 \times 10^{-16} \text{ cm}^{-2} \text{sr}^{-1} \text{s}^{-1}$$

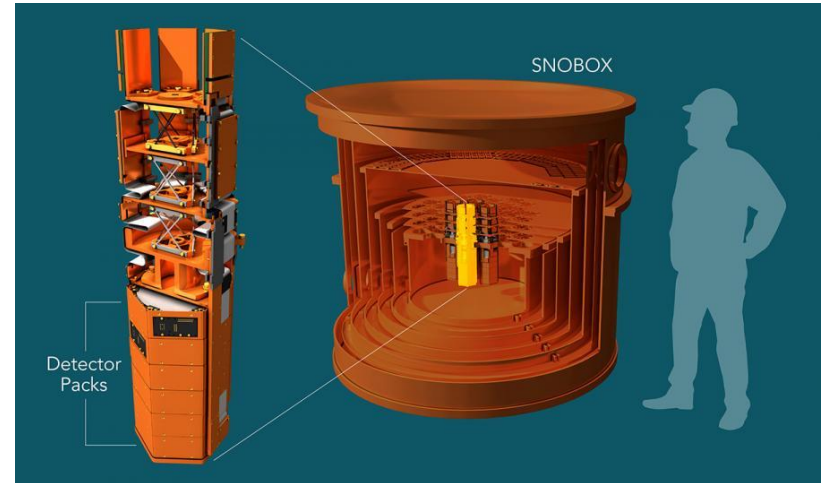
The previous experiments of charge sensitivity



MAJORANA 1600 m

$$\frac{1}{1000} < Q < \frac{1}{6}$$

$$\Phi\left(\frac{1}{6} \sim \frac{1}{30}\right) = 2 \times 10^{-9} \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$



CDMS II 700 m

$$\frac{1}{200} < Q < \frac{1}{6}$$

$$\Phi\left(\frac{1}{160}\right) = 1.36 \times 10^{-7} \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$$

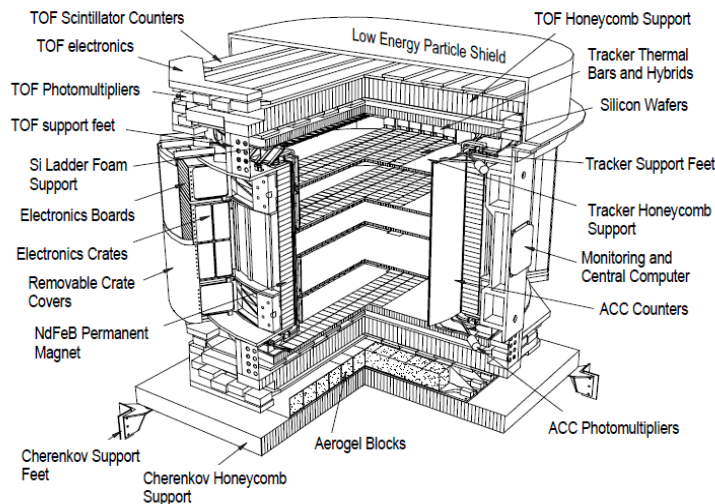
Searches for **Millicharge**, these experiments have very good capability of charge measurement

CDMSlite: $\frac{1}{1000} < Q < \frac{1}{6}$

TEXONO: $10^{-6} < Q < 10^{-3}$

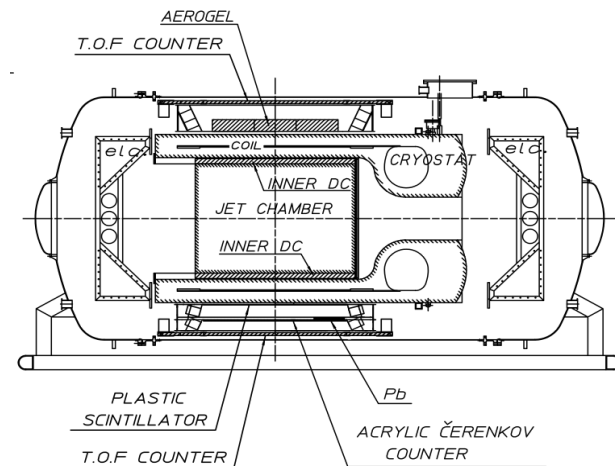
The previous experiments in space

For searches **in space**, the target FCP need not to penetrate the rocks, it can have energy **as low as ~GeV level** (the geomagnetic cutoff)



AMS01 space shuttle

$$\Phi\left(\frac{2}{3}\right) = 3.0 \times 10^{-7} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$$

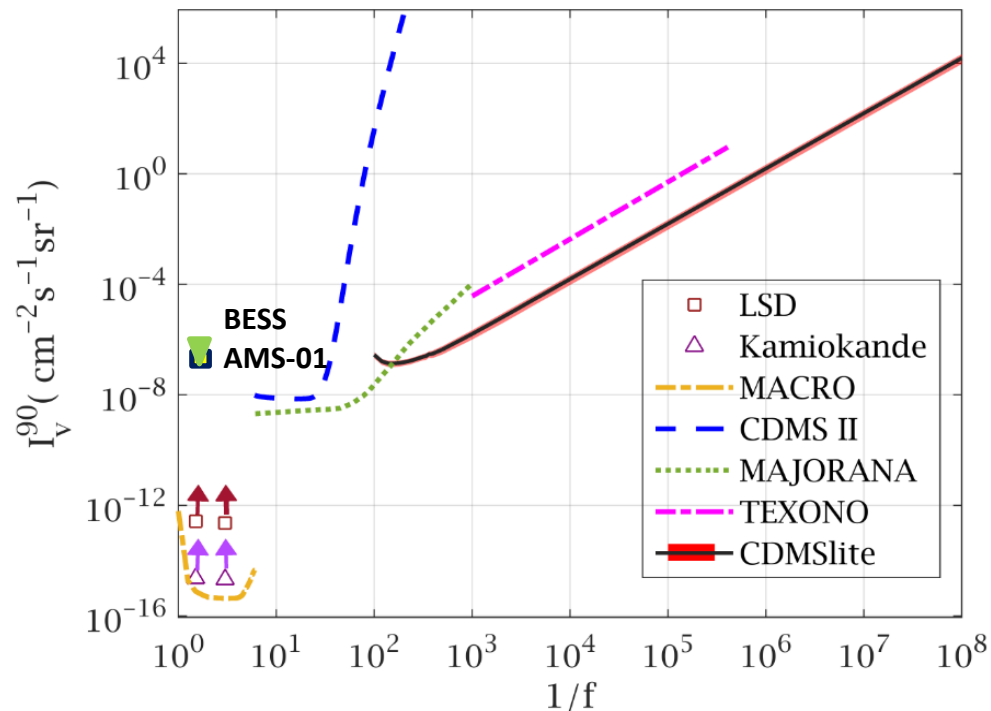


BESS balloon

$$\Phi\left(\frac{2}{3}\right) = 4.5 \times 10^{-7} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$$

The previous experiments

The flux upper limit versus the inverse charge value

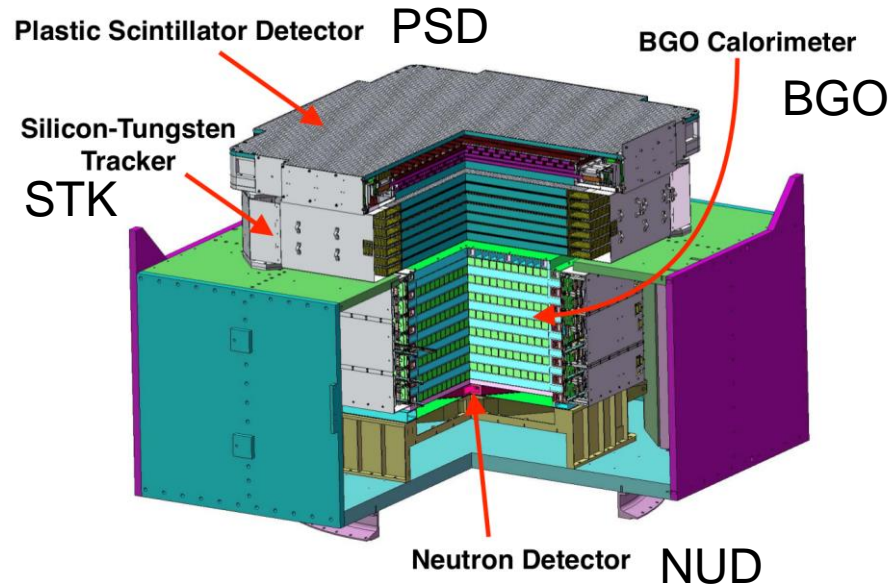


- **DAMPE** has been working stably on orbit for **nearly six years**.
- A lot of scientific data was accumulated.
- We hope to do something in searching for FCP as an on-orbit apparatus.

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DAMPE experiment



Dark Matter Particle Explorer(DAMPE)

Main Scientific Goals:

- Origin and Propagation of Cosmic-Rays
 - Dark Matter Indirect Detection
 - Gamma-ray Astronomy
-
- Orbit: sun-synchronous
 - Altitudes: 500 km
 - Period: about 90 minutes
 - 5 million events/day
 - 16 GB/day downlink
 - Launched on Dec.17th 2015
 - Life time > 5 years

Charge measurement (**PSD, STK**)
 Precise tracking (**STK + BGO**)
 Precise energy measurement (**BGO**)
 Particle identification (**BGO + NUD**)

DAMPE collaboration

CHINA

- Purple Mountain Observatory, CAS, Nanjing
- University of Science and Technology of China, Hefei
- Institute of High Energy Physics, CAS, Beijing
- National Space Science Center, CAS, Beijing
- Institute of Modern Physics, CAS, Lanzhou



ITALY

- INFN Perugia and University of Perugia
- INFN Bari and University of Bari
- INFN Lecce and University of Salento
- INFN LNGS and Gran Sasso Science Institute

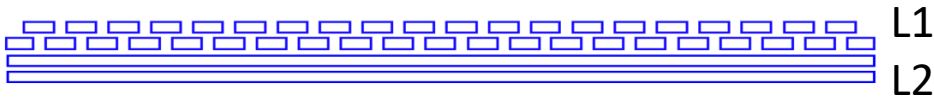


SWITZERLAND

- University of Geneva

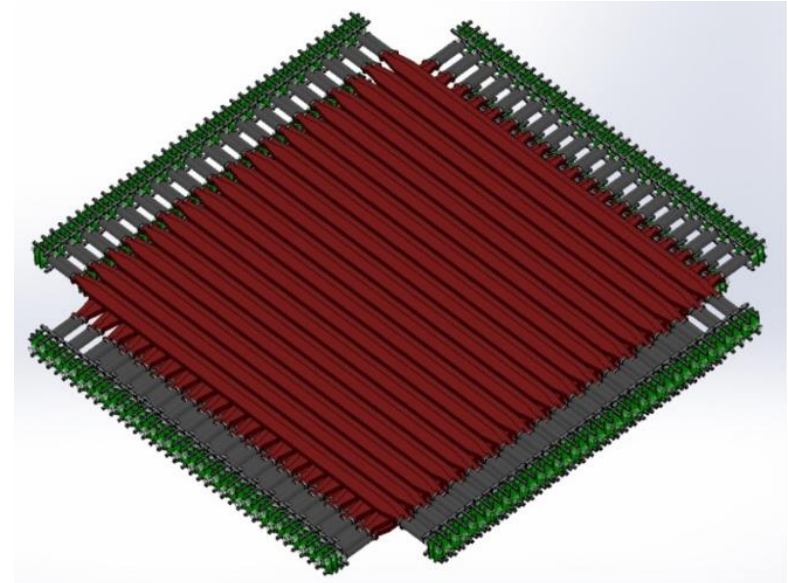
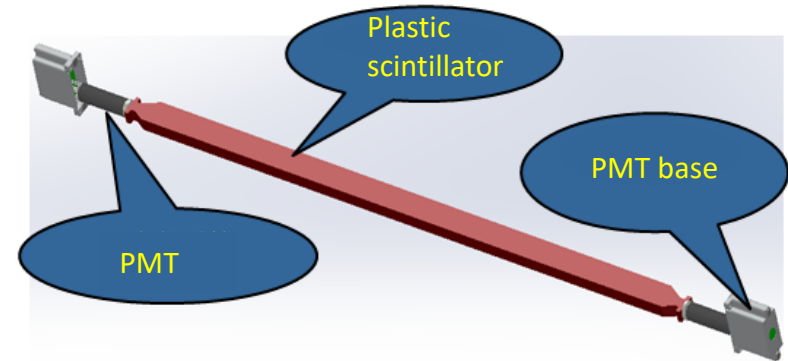


Plastic Scintillator Detector (PSD)

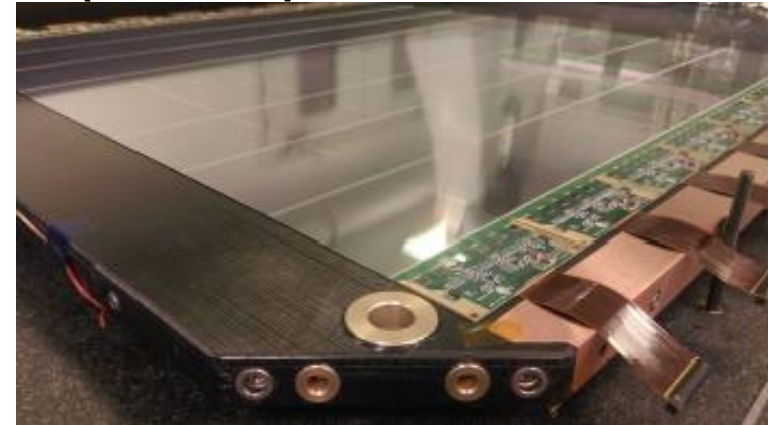
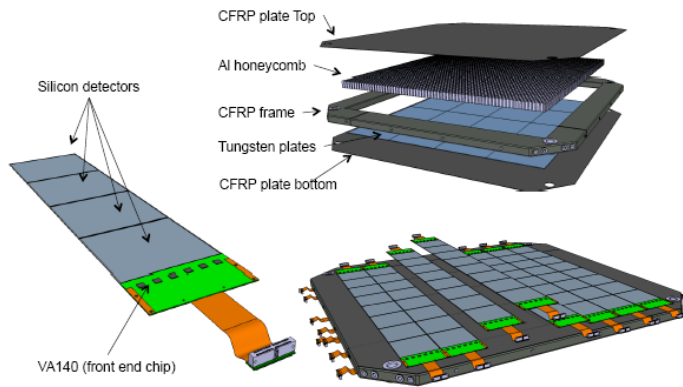


PSD is located on the top of the payload

- Active area: 82 cm × 82 cm
- Number of layers: 2
- 41 modules each layer
- A PMT at each end of strip
- Each PMT provides two signals (from Dy5 and Dy8 for large dynamic range)
- Charge resolution: 6% for $Z = 1$

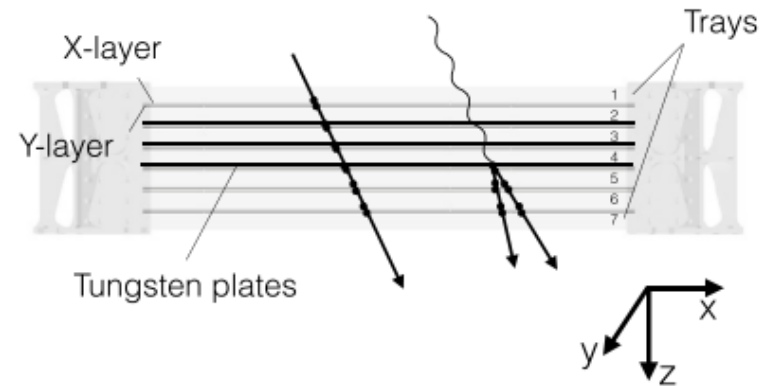


Silicon Tungsten tracker (STK)



STK is composed by:

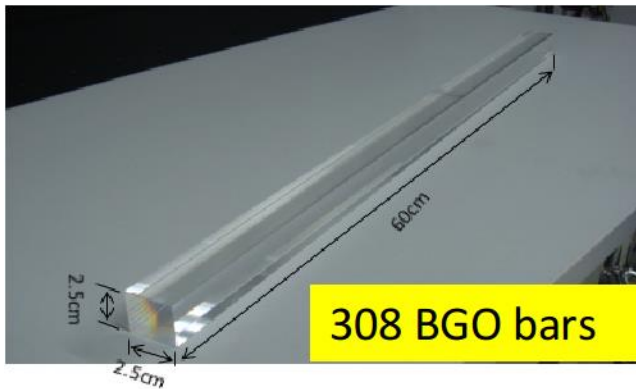
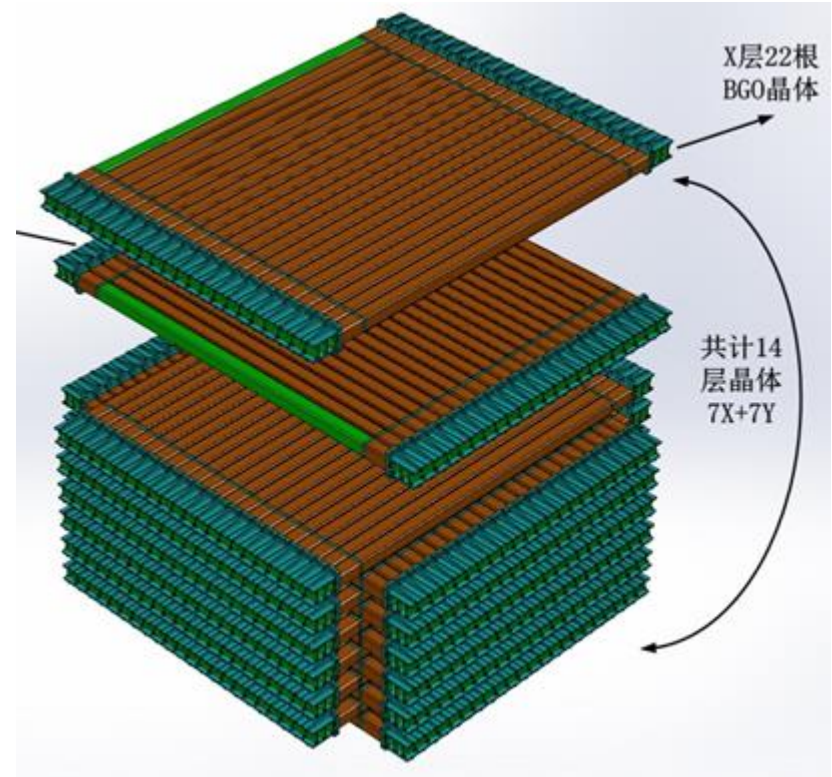
- Active area: 758 mm × 758 mm
- 7 CERP trays, 3 with tungsten plates for photon conversion
- Number of layers: 6
- 192 Si ladders. 16 on each sensitive face(12)
- each tray is orthogonal to the previous one to allow 3D tracking



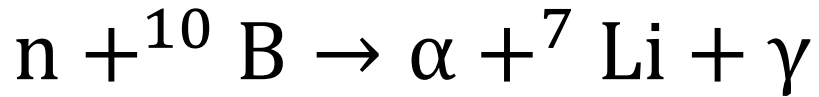
BGO calorimeter

BGO structure:

- 14 layers of 22 BGO crystals
- Dimension of BGO bar: $2.5 \times 2.5 \times 60 \text{ cm}^3$
- Hodoscopic stacking alternating orthogonal layers
- r.l: 32X0, NIL:1.6
- Two PMTs coupled with each BGO crystal bar in two ends
- Electronics boards attached to each side of module

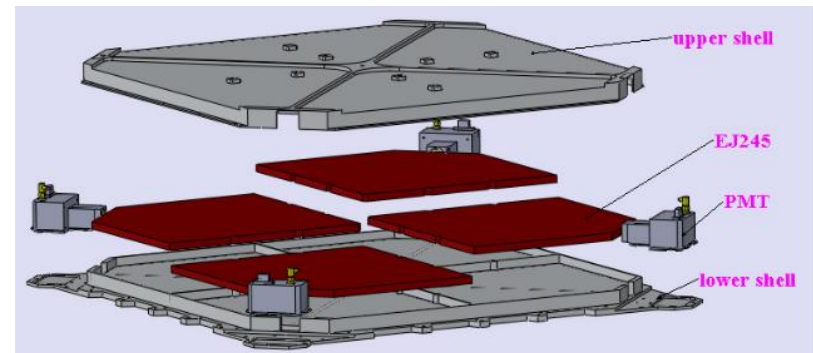


Neutron Detector

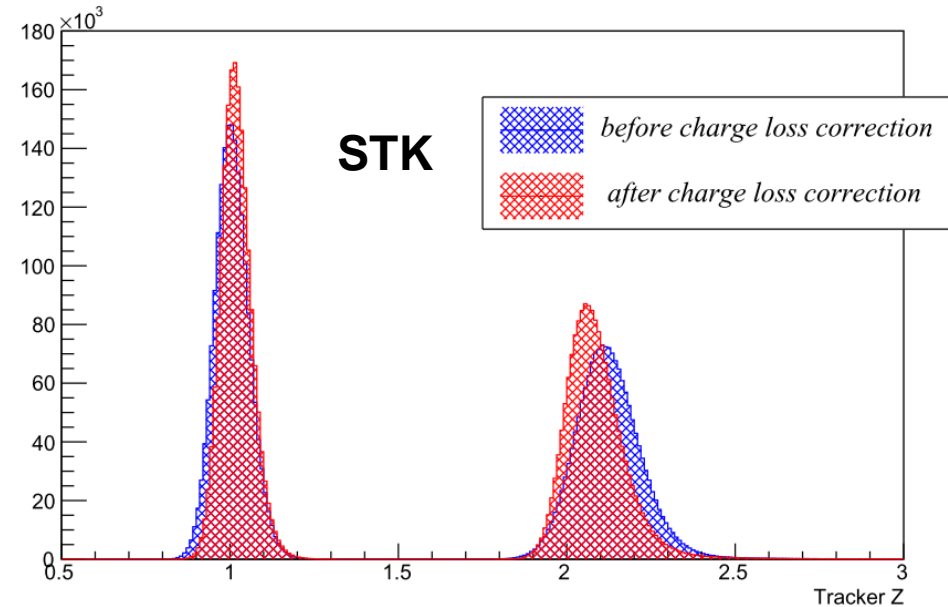
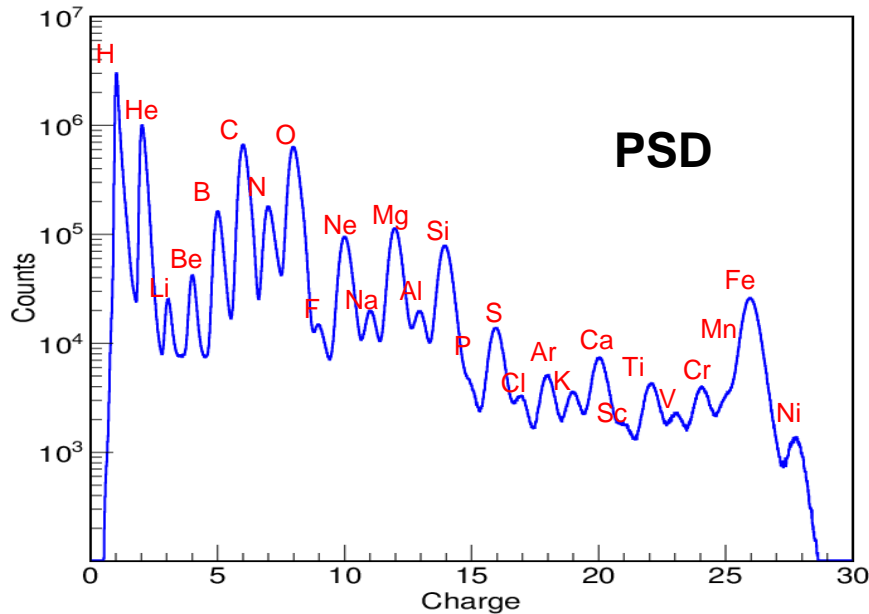


NUD designed parameter:

- 4 boron-doped plastic scintillators (30 cm × 30 cm × 1 cm)
- Active area: 61 × 61 cm²
- Energy range: 2 - 60 MeV for single detector
- Energy resolution: ≤ 10% at 30 MeV



Performance of charge measurement



Left: The charge discrimination of **PSD** in Z from 1 to 28

Right: The charge discrimination of **STK** between proton and helium

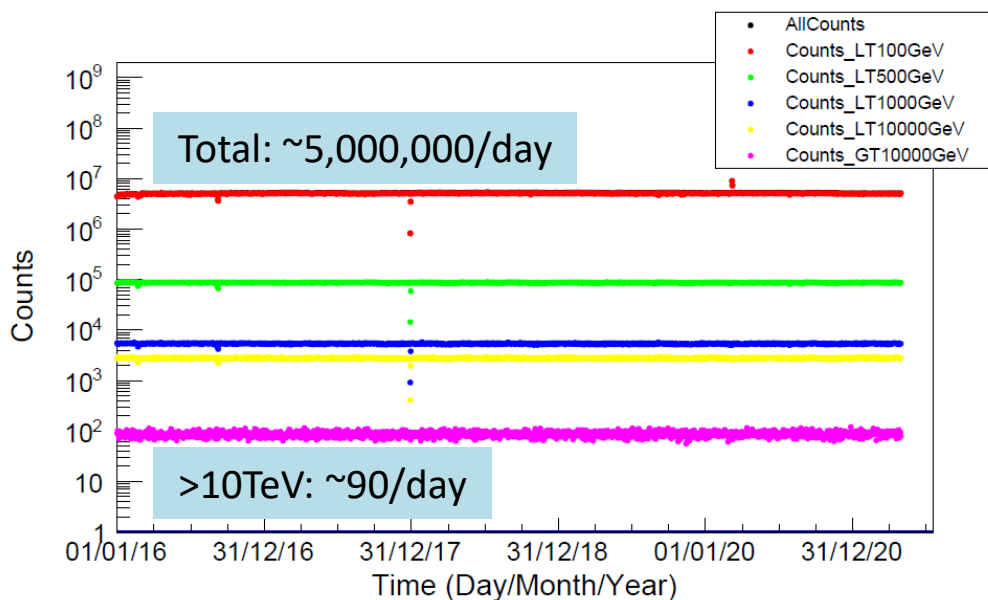
Charge resolution for singly charged particles:

PSD: **0.06e**

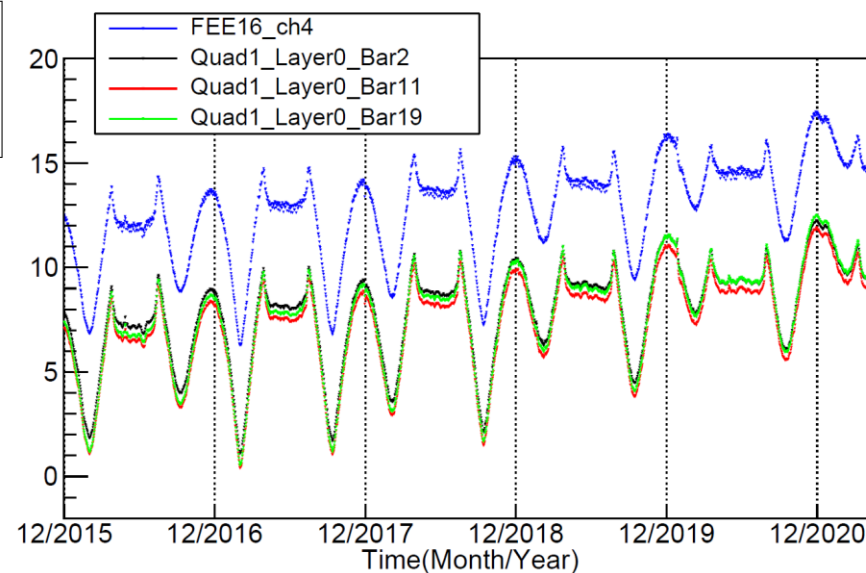
STK: **0.04e**

DAMPE status

Trigger rate per day



Daily temperature variation of BGO

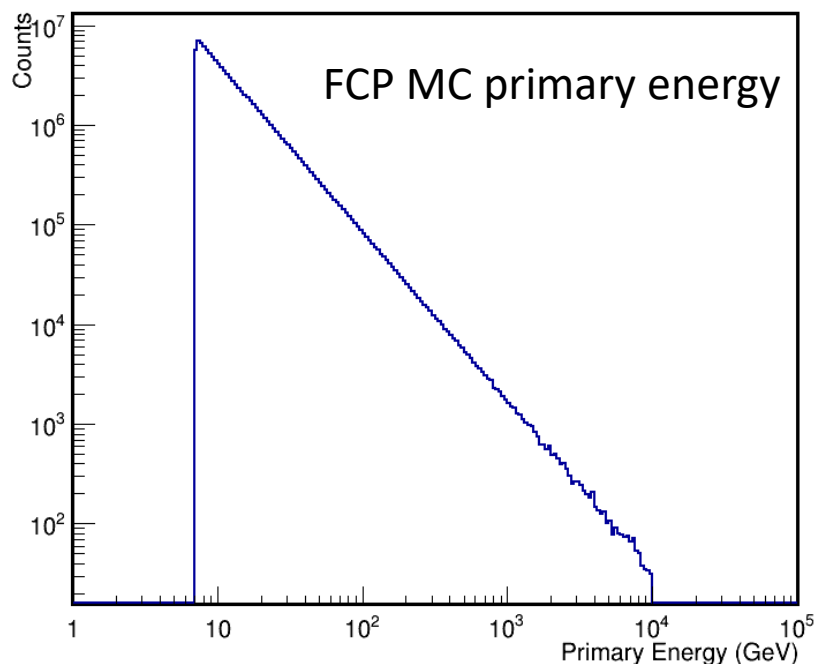


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Data samples

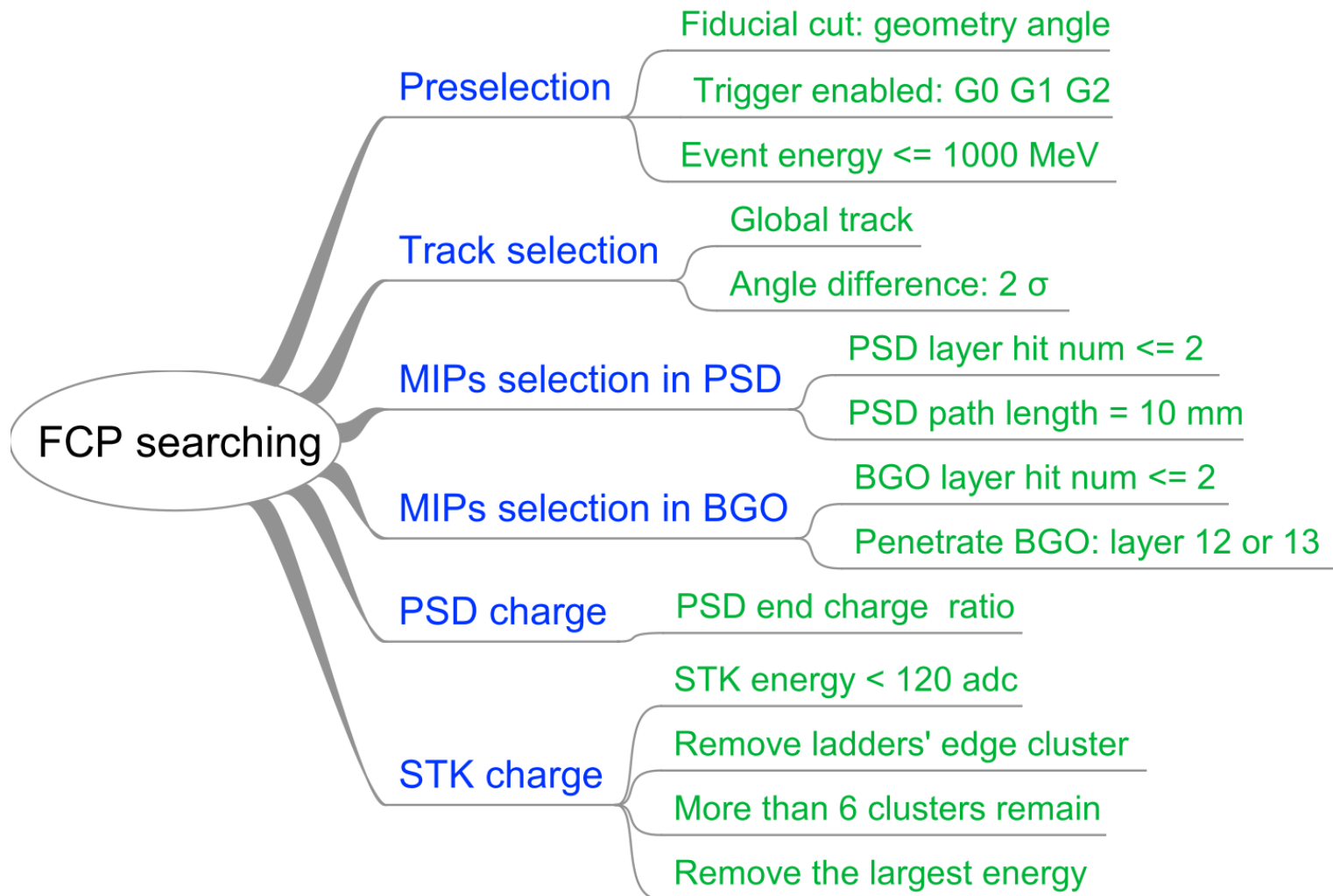
- Five years on-orbit data
- Proton simulation based on Geant4 in DAMPE software 10 GeV - 100 TeV
- FCP simulation based on Geant4 in DAMPE software 7 GeV – 10 TeV



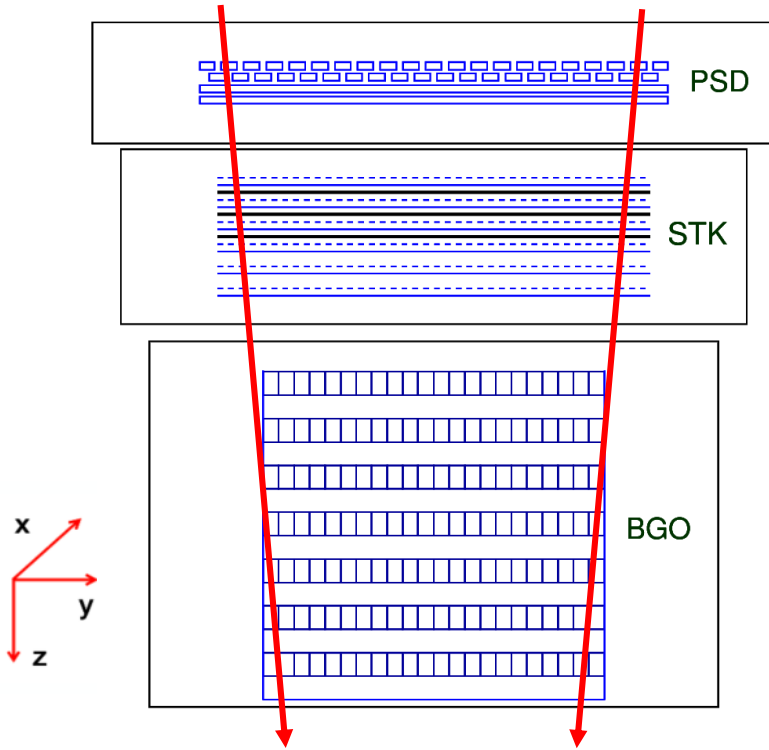
FCP simulation:

- Created a virtual particle
- Charge with $\frac{2}{3} e$
- Mass with 1.2 GeV
- Add ionization and multi scattering process
- Energy spectrum obey the E^{-3}
- Spheric particle source

Searching FCP with DAMPE



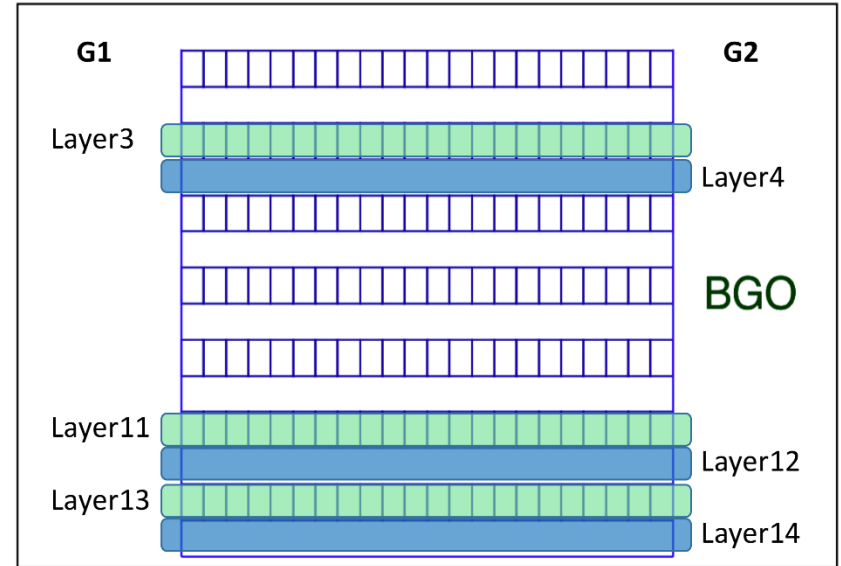
Pre-selection



Fiducial cut:

Constrain the positions of injection and ejection to maintain the event in the whole detector

MIPs Trigger: G1||G2 trigger enabled
 $-20^\circ < \text{latitude} < 20^\circ$



Since the calibrated trigger threshold of G1 G2 is ~ 0.2 MIPs, higher than the 1/3 charged particles (1/9 MIPs), we aim to search the $2/3$ charged particles.

Mips energy:

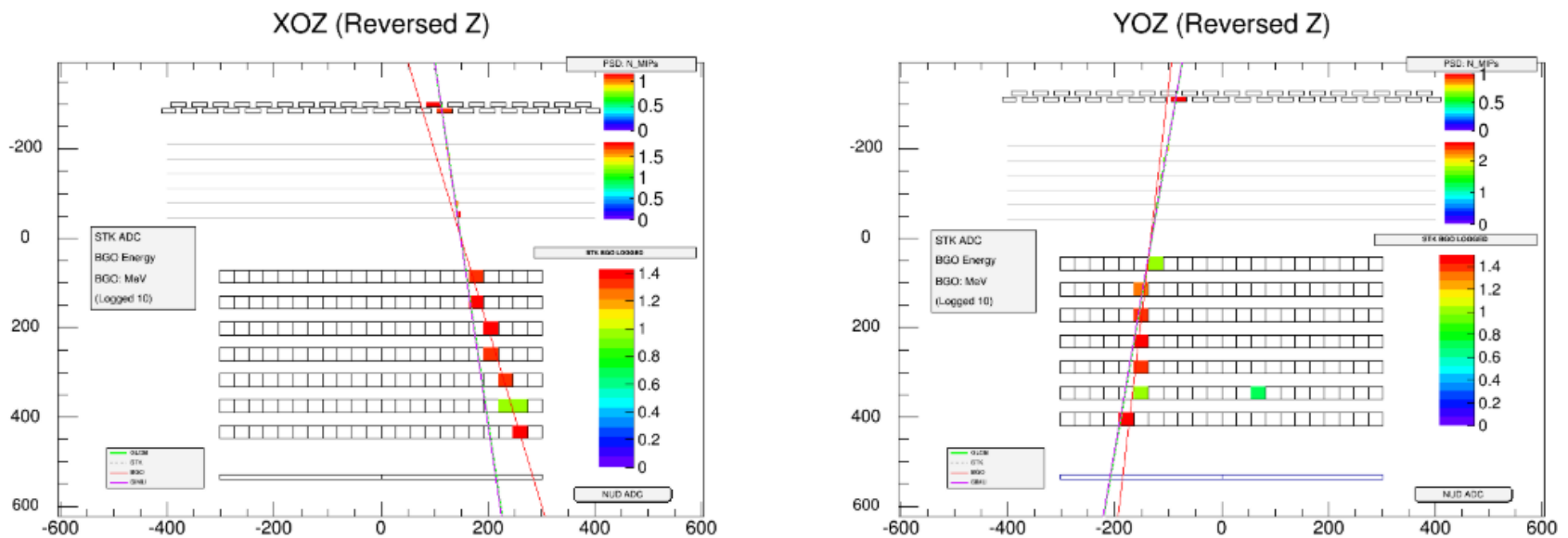
1000 MeV loose cut since singly charged particle depositing energy is about 700 MeV

Track selection

Global track:
should be reconstructed in STK.

Angle difference between Global Track and BGO Track:

The hits in BGO calorimeter can be used to reconstruct a BGO track. If global track deflects too much from BGO track, there may be scatters happened



MC_Proton event Angle > 6°

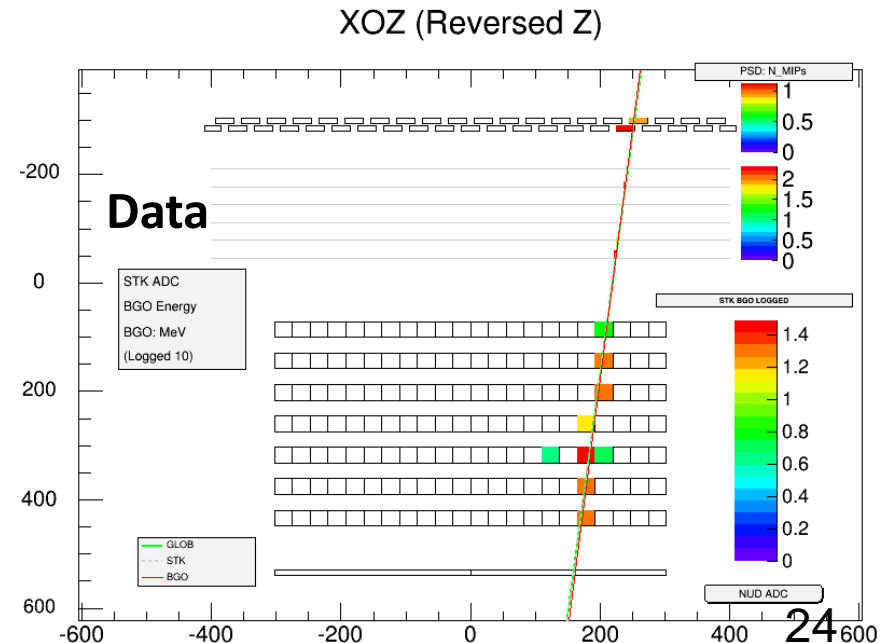
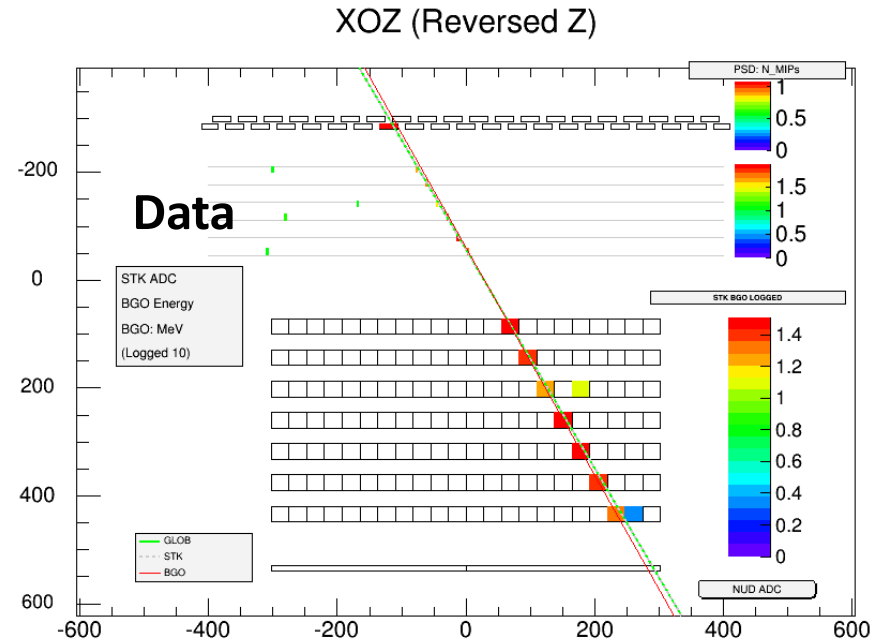
MIPs selection

MIP in PSD:

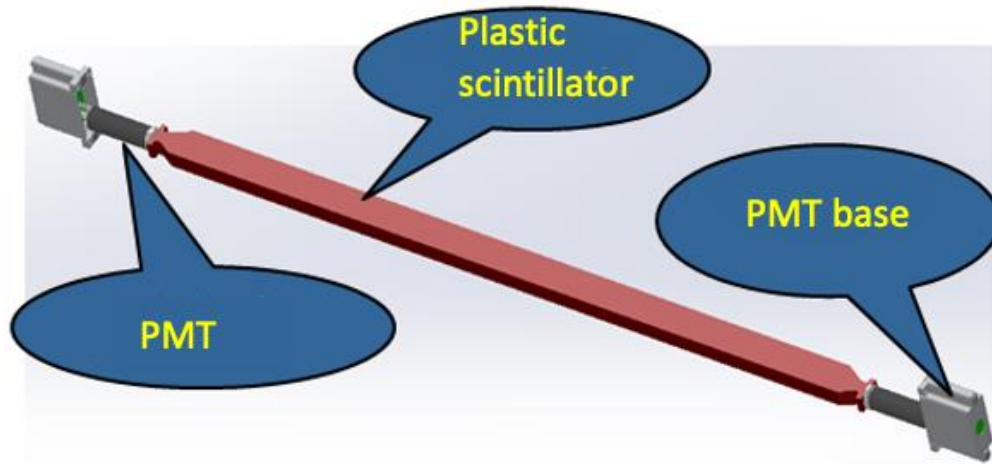
- Every Layer hit bars ≤ 2
- Path length = 10 mm

MIP in BGO:

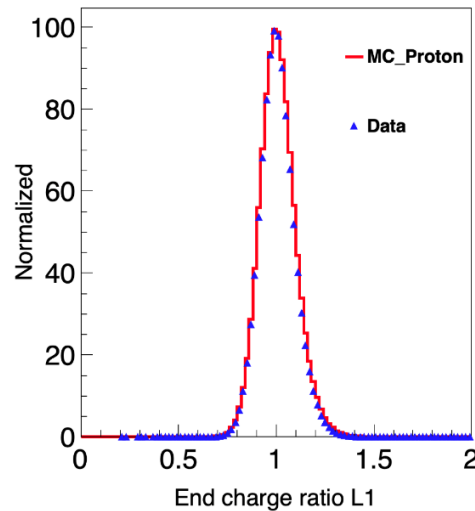
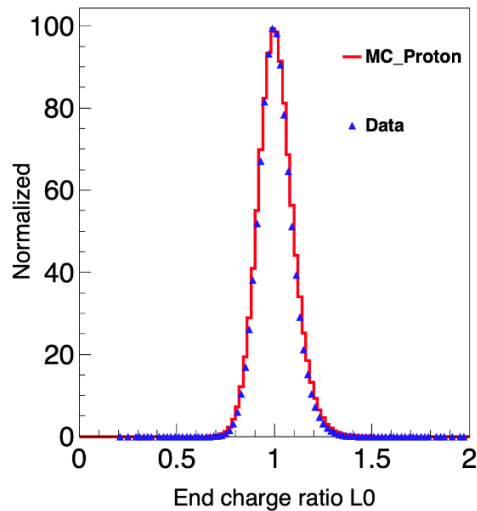
- Over-threshold(2 MeV) hits no more than 2 in one layer along the track and no more than 28 in the whole calorimeter
- Hit layers > 5 in both YOZ and XOZ, **one of last two layers** should be fired



PSD charge

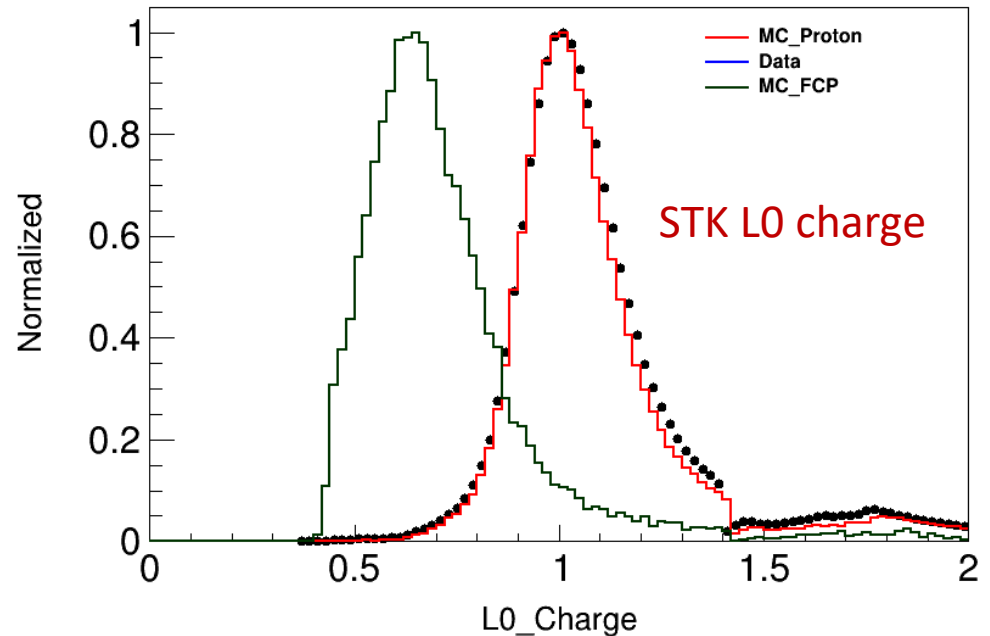
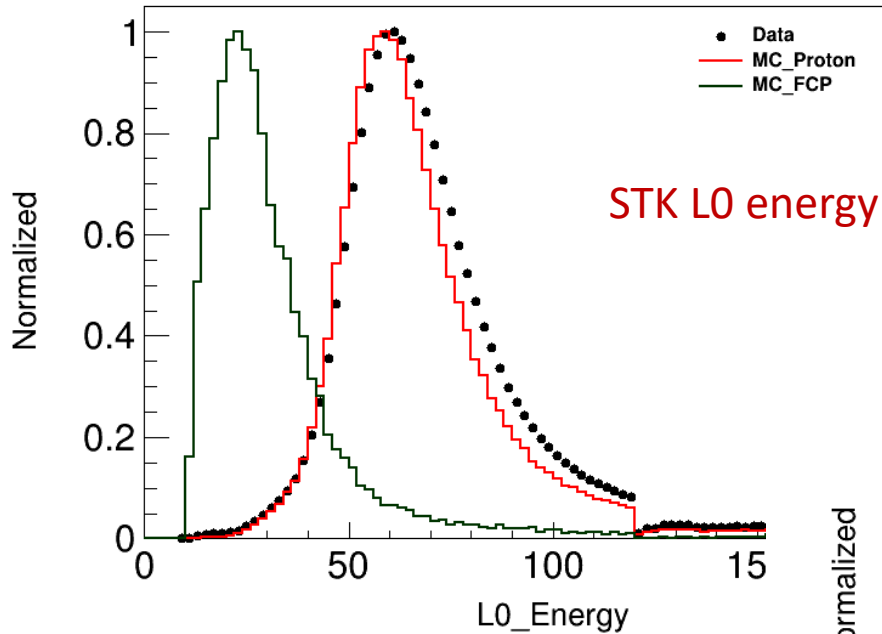


To ensure the reliability of charge reconstruction, the ratio of two ends of one strip should be consistent



STK sub-layer energy and charge

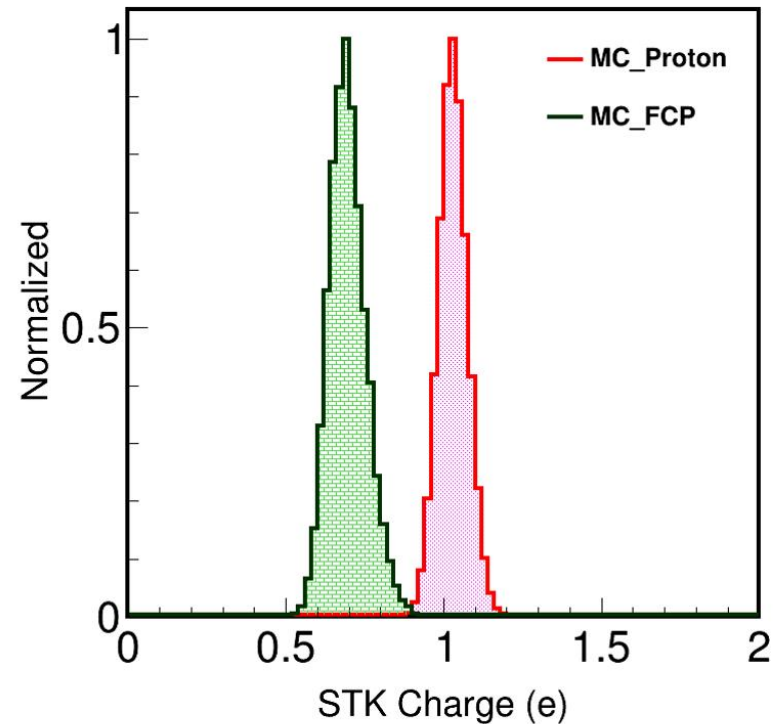
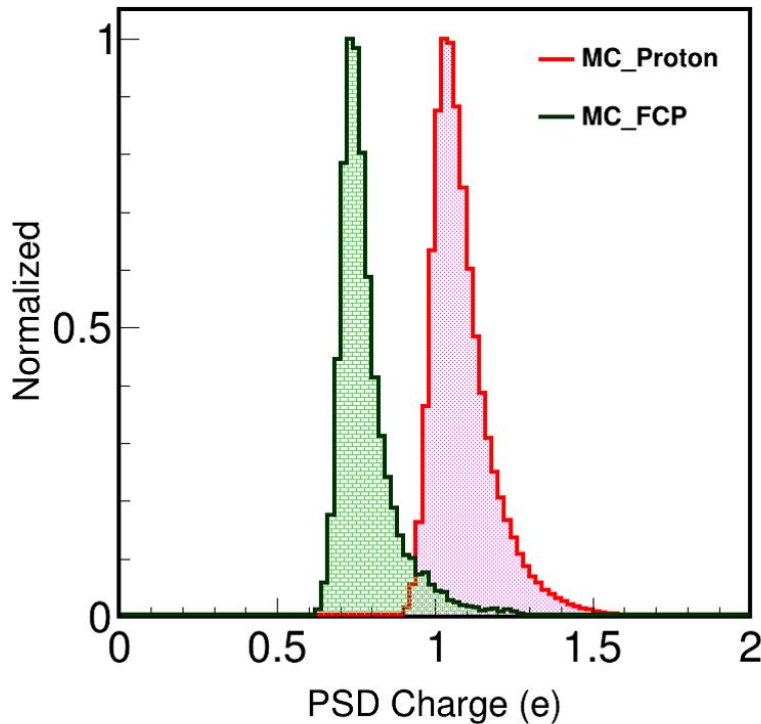
- Take layer 1 for example



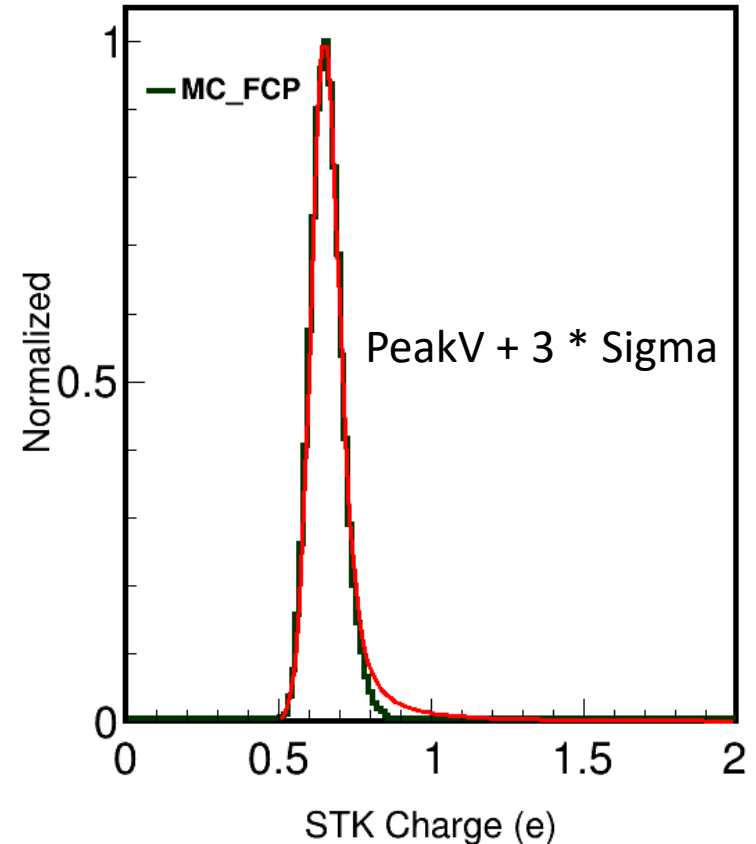
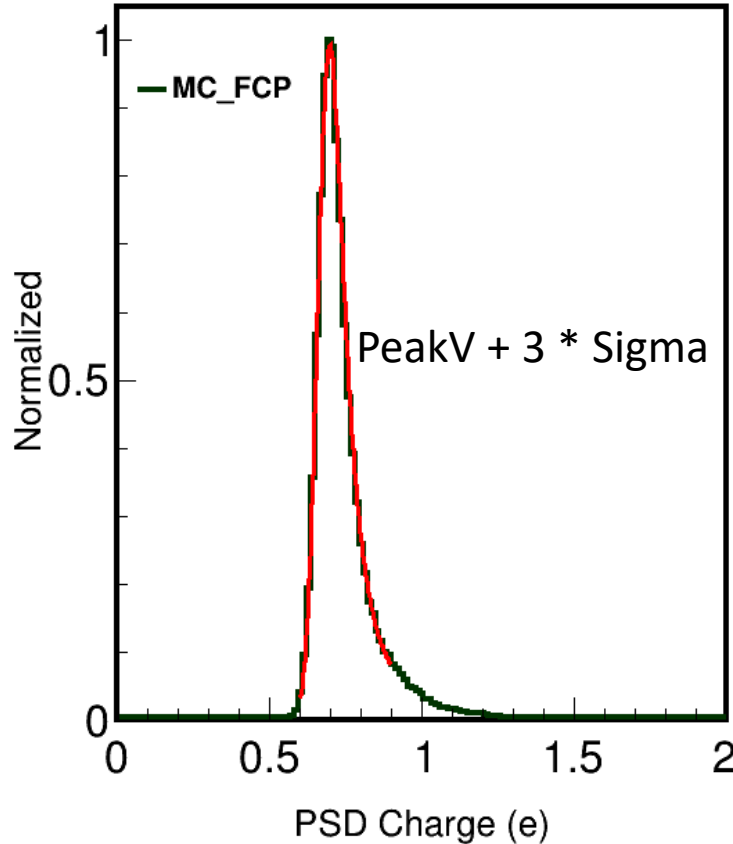
Charge reconstruction

$$Q_{PSD} = \frac{Q_0 + Q_1}{2}$$

$$Q_{STK} = \frac{\sum_{i=1}^N Q_i}{N}$$

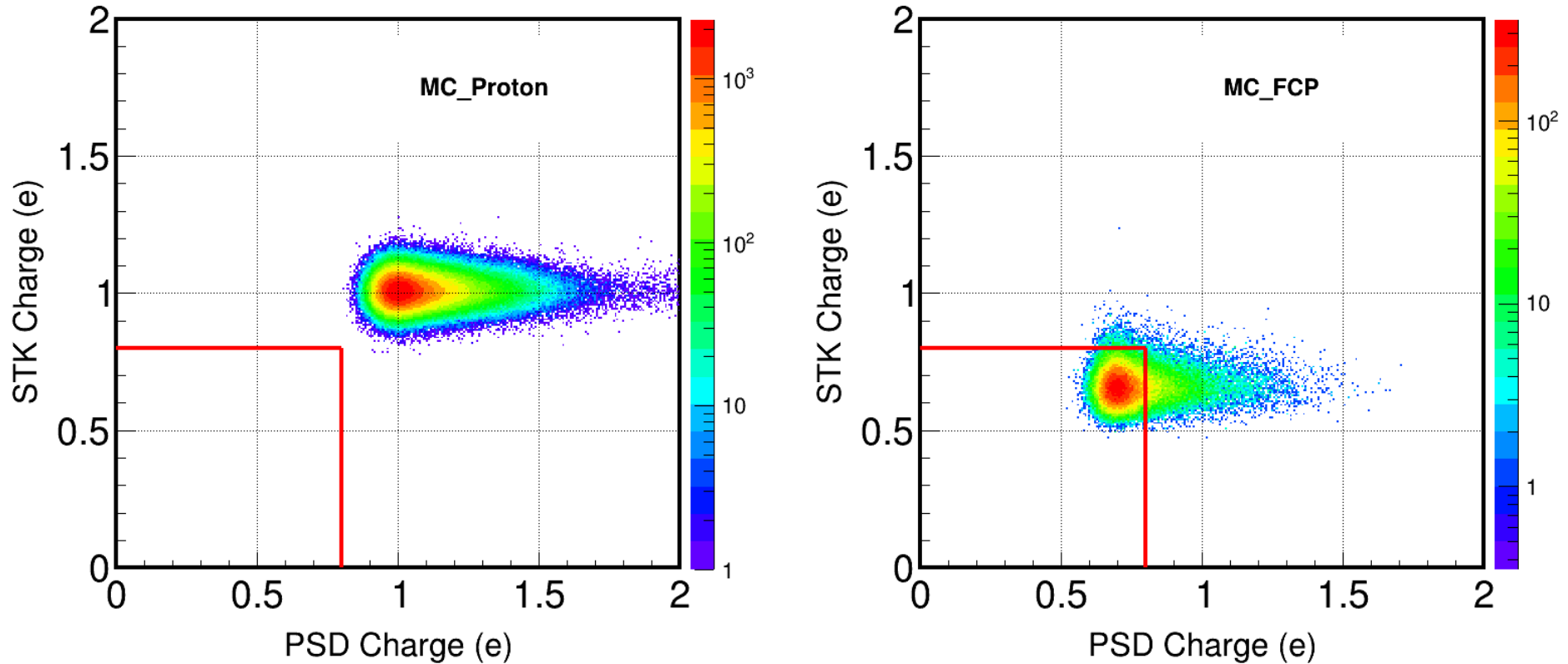


Signal region definition



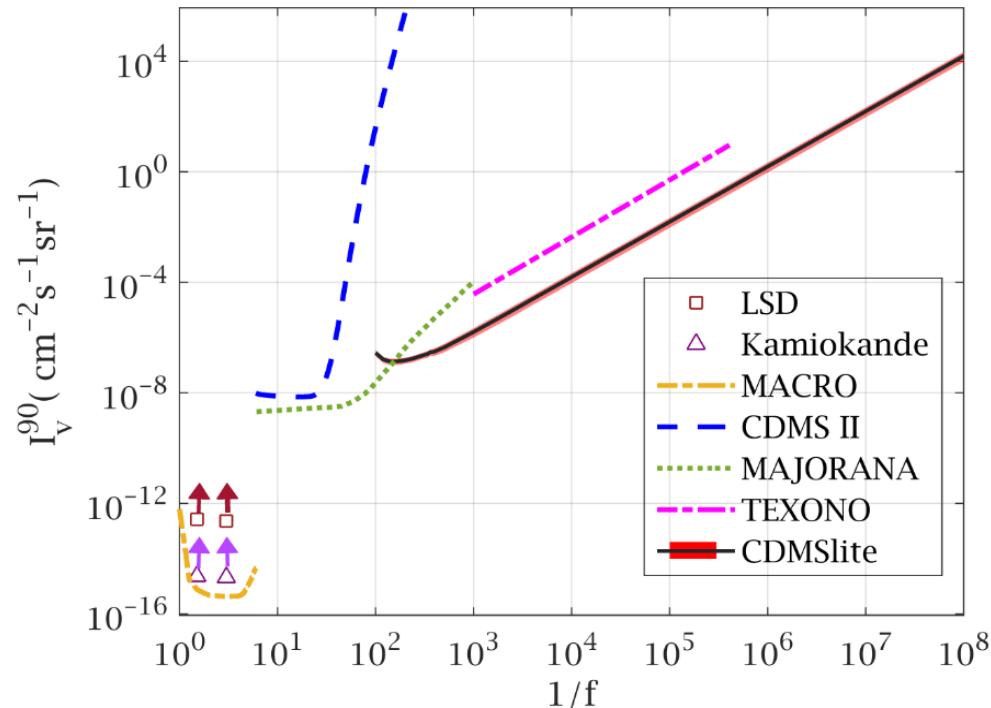
The charge distributions are fitted with the convoluted Landau and Gaussian function, the **sigma** is taken with $FWHM/2.35$.

Signal region application



A signal region can be defined by PSD and STK MC as the red lines shows, the charge of two lines are set to around $0.8e$. The signal region is effective to exclude the background and has a **86% efficiency** of covering the FCP.

Results from underground experiments



With the **large volume and long exposure time**, The MACRO released the most stringent upper limit; with **high degree of charge sensitivity**, the CDMS and others could measure very small charge.

Comparison with same type experiments

| Experiments | Geometry acceptance (cm^2sr) | Exposure time (s) | Upper limit $\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1}$ |
|--------------|------------------------------------------------|-------------------|---------------------------------------------------------|
| AMS-01 | 3000 | 3.6×10^4 | 3.0×10^{-7} |
| BESS | 1500 | 3.2×10^5 | 4.5×10^{-7} |
| DAMPE | 3000 | 2.4×10^7 | To be released |

- AMS-01 has a large acceptance, but short of the exposure time
- BESS integrates four times of flights to achieve a longer exposure time
- DAMPE has a relatively larger acceptance and the longest exposure time. We hope to release a lower upper limit in the near future.

Summary

- The history of FCP has been reviewed briefly.
- The DAMPE experiment has been introduced.
- The selection criteria to search FCP with DAMPE have been studied.
- A MC simulation has been performed and an evaluation of the detection efficiency has been carried out.
- A comparison between DAMPE and other equipment has been done. DAMPE is hopeful to release a lower upper limit in space.

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Thank you!