



李政道研究所  
TSUNG-DAO LEE INSTITUTE

SHINE



# Progress on the DarkSHINE Experiment

Zhiyu Zhao (TDLI/SJTU)

On behalf of DarkSHINE R&D Group

首届SHINE缪子源及其应用研讨会

28 November, 2025



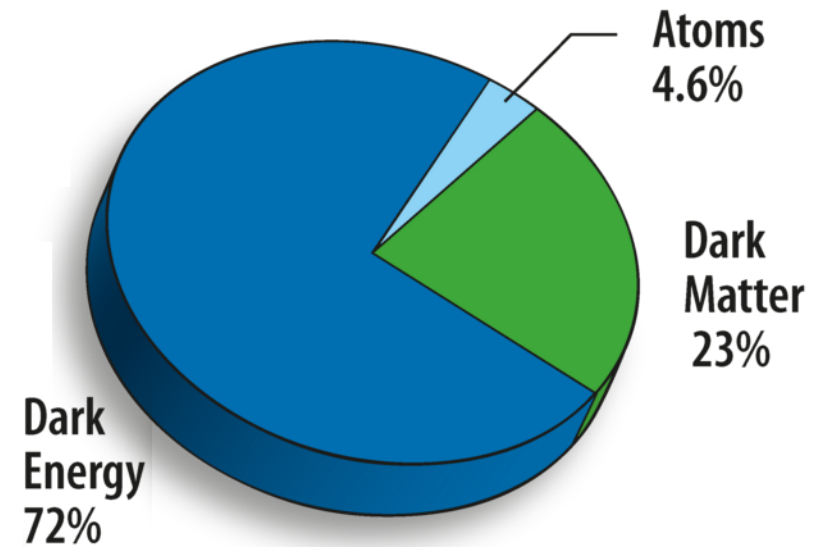
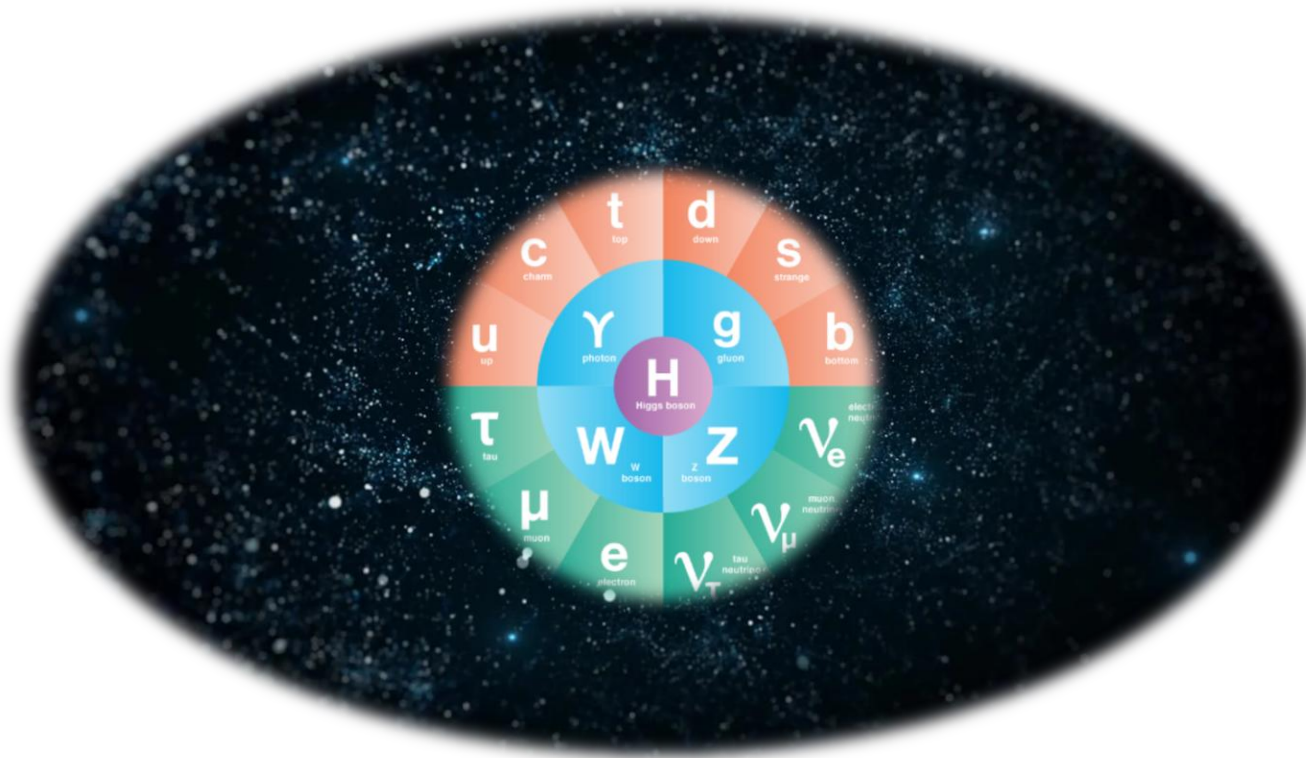
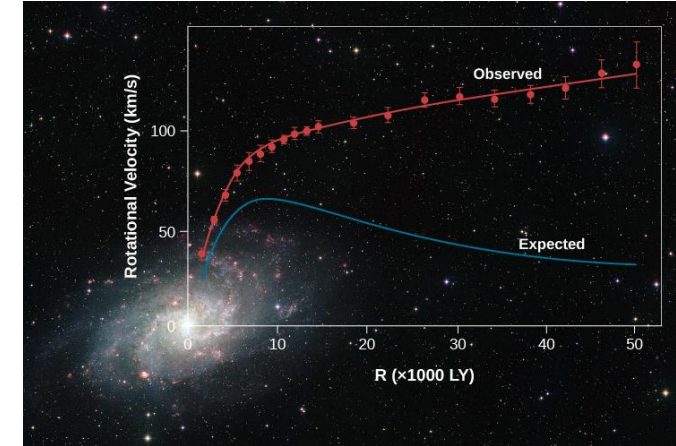
# Dark Matter



上海交通大学  
SHANGHAI JIAO TONG UNIVERSITY

李政道研究所  
Tsung-Dao Lee Institute

- Standard Model (SM): framework for particle physics
- Beyond Standard Model: **dark matter**, dark energy, neutrino mass, ...
  - Dark matter evidence from astronomical observations and gravitational effects





# Dark Matter Candidates



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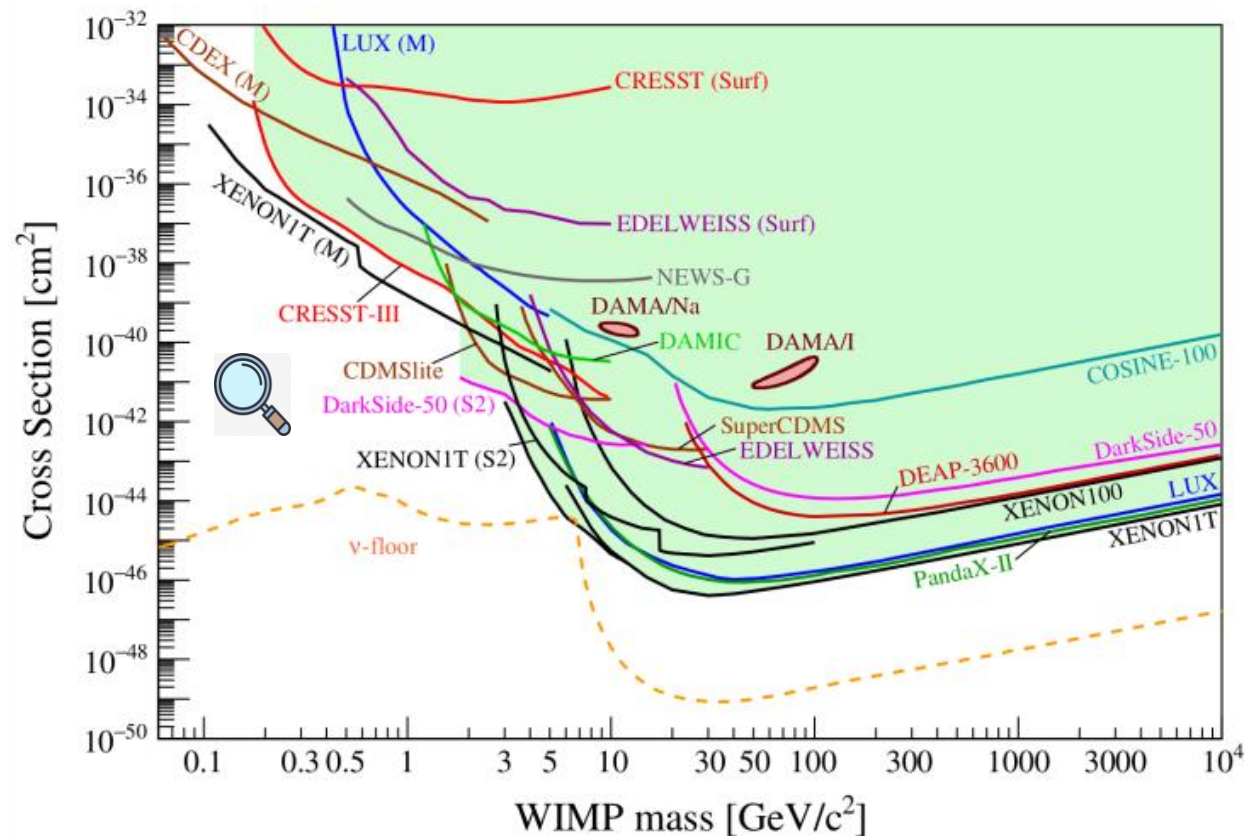
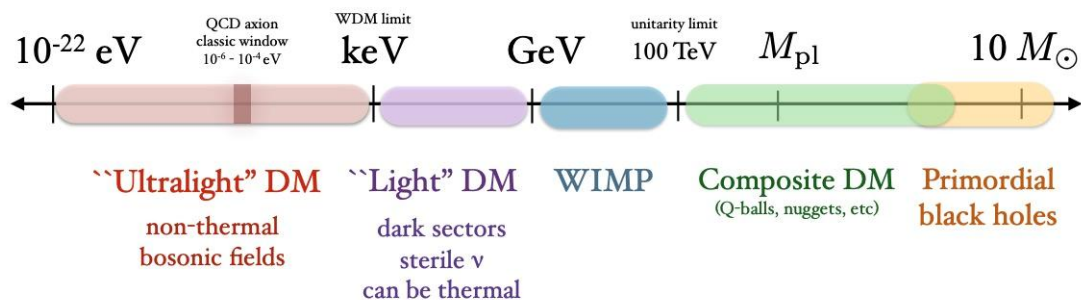
- The “freeze-out” mechanism predict the mass of dark matter is mainly distributed from **MeV to tens of TeV**
  - Weakly Interacting Massive Particles (WIMP): A large parameter space ruled out in GeV~TeV mass range.
  - Light DM: Sub-GeV mass range not fully explored yet.**

APPEC Committee Report: 2104.07634

arXiv:1904.07915

## Mass scale of dark matter

(not to scale)



# Dark Photon Theory in a Nutshell



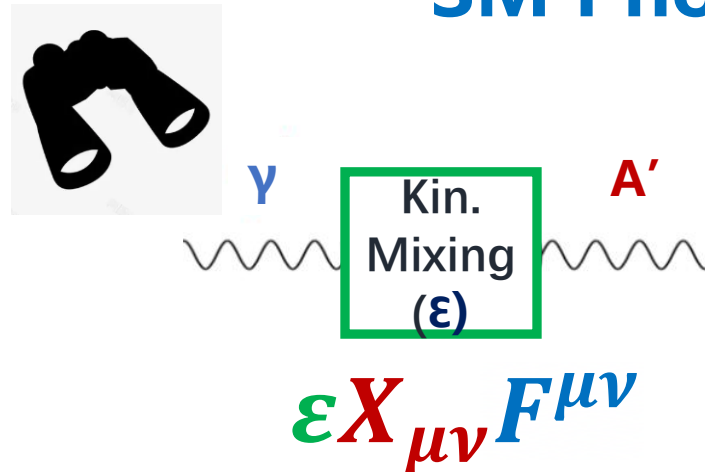
- Dark photon: one candidate of light DM, kinetic mixing with SM photon

Introduce extra  $U(1)_X$  symmetry  $\rightarrow$  New Gauge Field  $X \rightarrow$  Dark Photon Mediator  $A'$

$$\mathcal{L} = \underbrace{-\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + A_\mu j_{em}^\mu}_{\text{SM Photon } \gamma} \underbrace{-\frac{1}{4} X_{\mu\nu} X^{\mu\nu} + X_\mu j_X^\mu}_{\text{Dark Photon } A'}$$

SM Photon  $\gamma$

Dark Photon  $A'$



- $A'$  &  $\gamma$  kin. mixing
- Renormalizable and Gauge Invariant
- Straightforward for experimental search
  - Free param, kin. mixing ( $\epsilon$ ), mass ( $m_{A'}$ )

B. Holdom, Phys. Lett. B 166, 196 (1986)

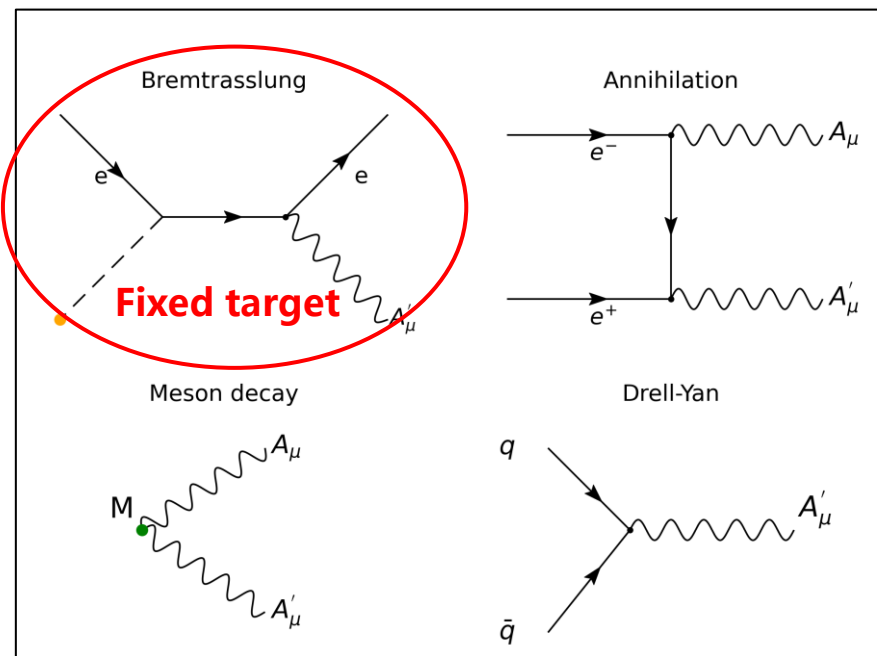
R. Foot & X.-G. He, Phys. Lett. B 267, 509 (1991)



# Dark Photon Production and Decay

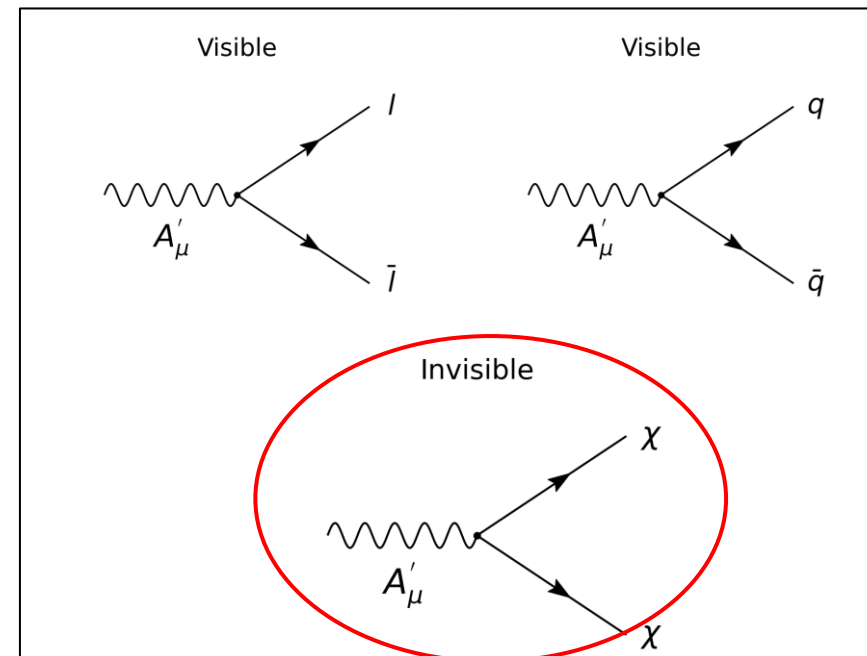


## Production



- **Bremsstrahlung,  $eZ \rightarrow eZA'$  &  $pZ \rightarrow pZA'$ , fixed-target experiment**
- Annihilation,  $e^+e^- \rightarrow A'\gamma$ ,  $e^+e^-$  collider
- Drell-Yan,  $q\bar{q} \rightarrow A'$ , hadron collider / fixed nuclear target w/ proton-beam
- Meson decay,  $\pi^0 \rightarrow A'\gamma$  or  $\eta \rightarrow A'\gamma$  (w/  $m_{A'} < m_{\pi,\eta}$ ), any experiment w/ high meson production rates

## Decay



- Visible decay  
Two interaction vertices  $\rightarrow$  highly suppressed
- **Invisible decay**  
**Better sensitivity: one interaction vertex**  
**Distinctive signatures: missing E&p**

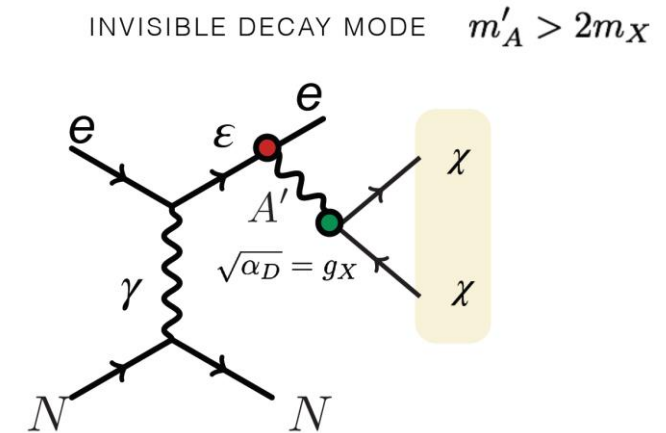
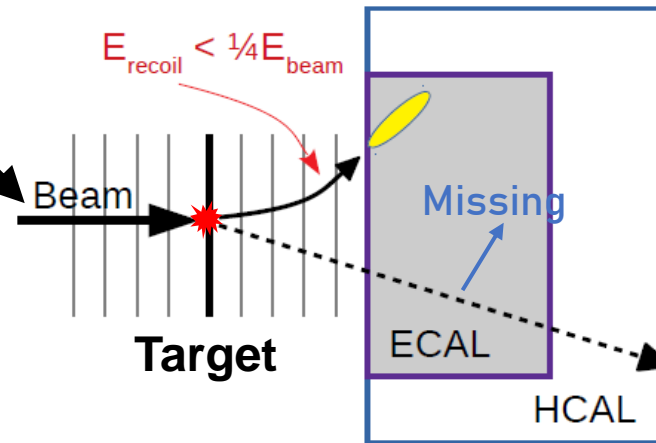
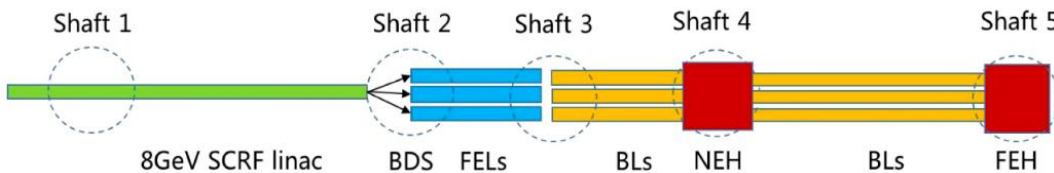
# DarkSHINE Experiment Searching for Dark Photon



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- **DarkSHINE**: search for **bremsstrahlung dark photon** and its **invisible decay**
  - Fixed-target experiment, **MHz single electron beam** (to be) provided by SHINE facility
  - Detector R&D: SJTU / FDU / USTC / SIC-CAS / IHEP

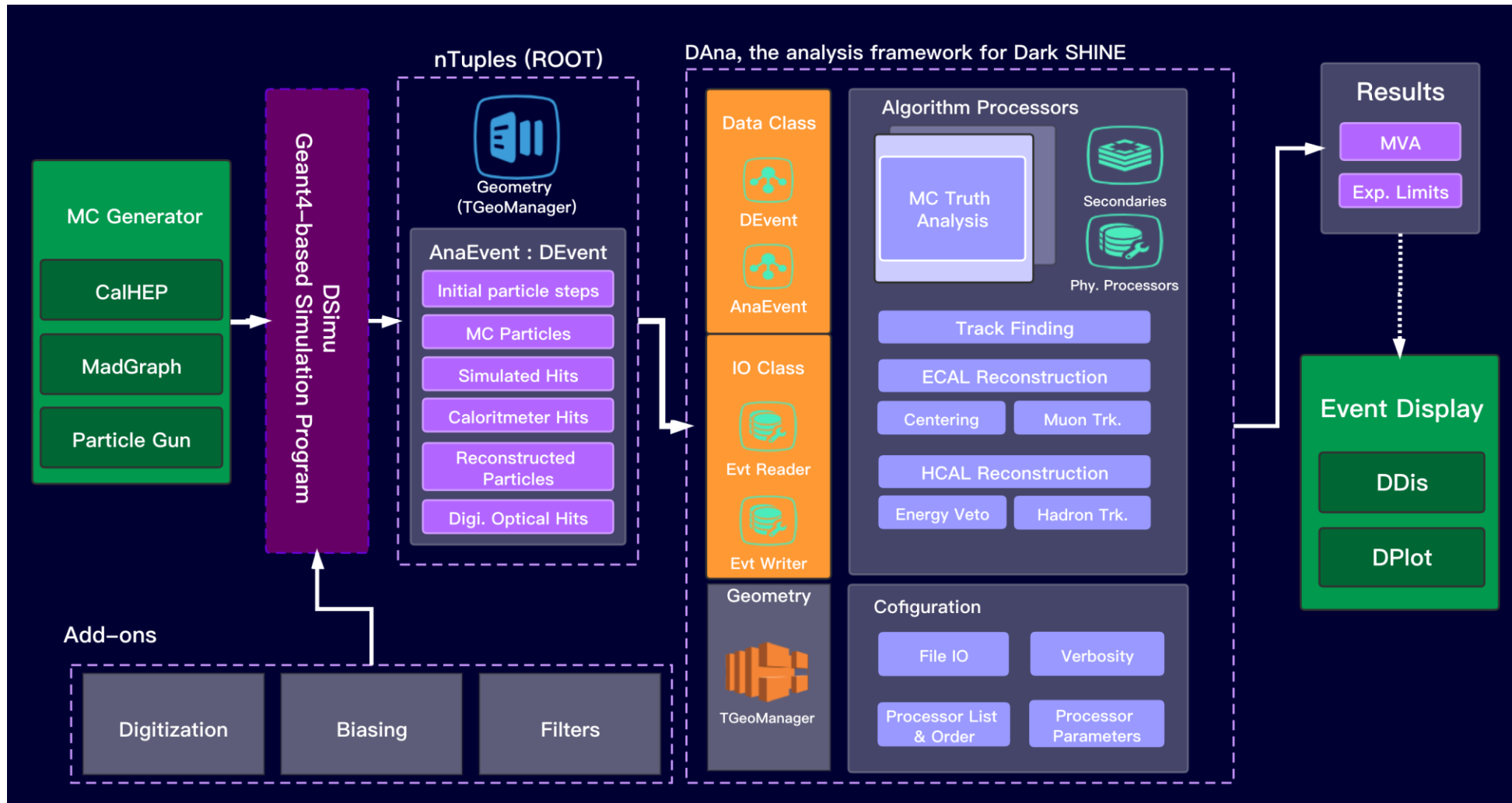


e- beam with one e- per bunch is required

# DarkSHINE Software Framework



- Integrates detector simulation, signal digitization, event reconstruction, and data analysis, etc.





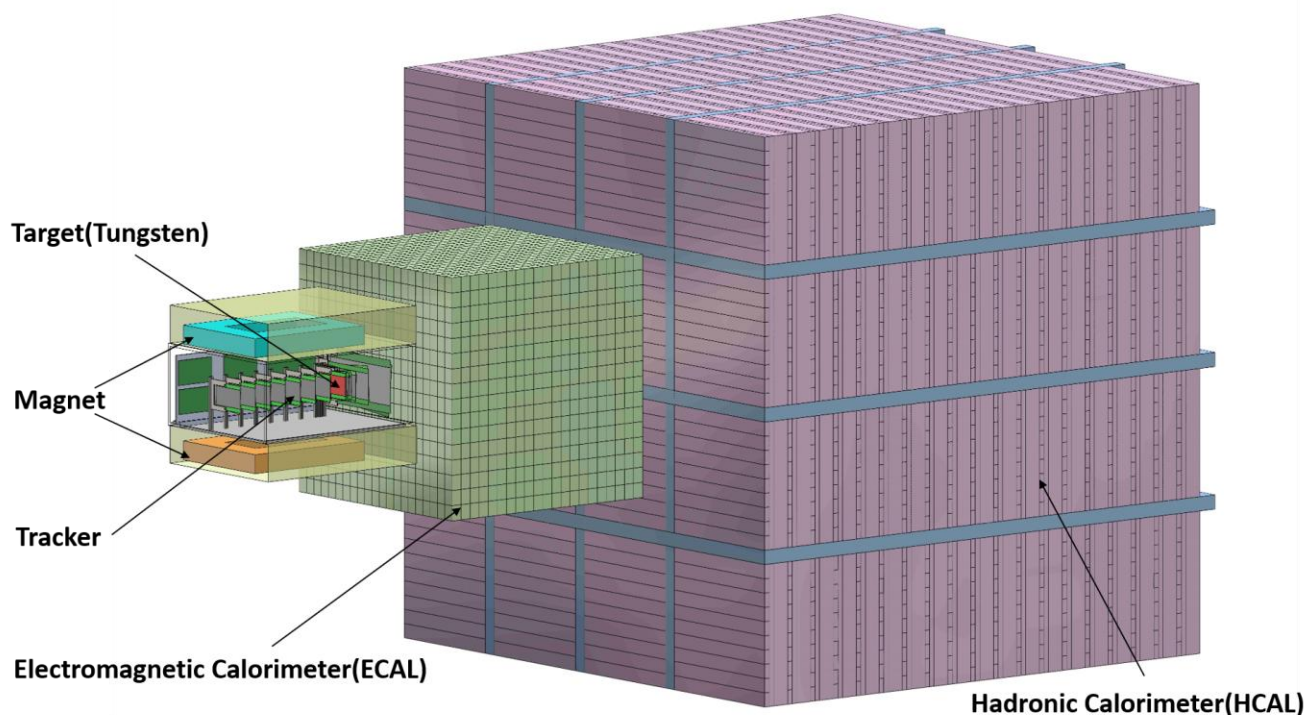
# Detector System Conceptual Design



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## DarkSHINE Baseline Design Report, arXiv: 2411.09345



**Tracker:** measure the track and  $p$  of charged particles

- Silicon strip sensors are arranged in an interleaved pattern
- Tagging tracker(7 layers) + recoil tracker(6 layers)
- Resolution:  $10\mu\text{m}$ (horizontal),  $60\mu\text{m}$ (vertical)

**ECAL:** measure energy of electrons and photons

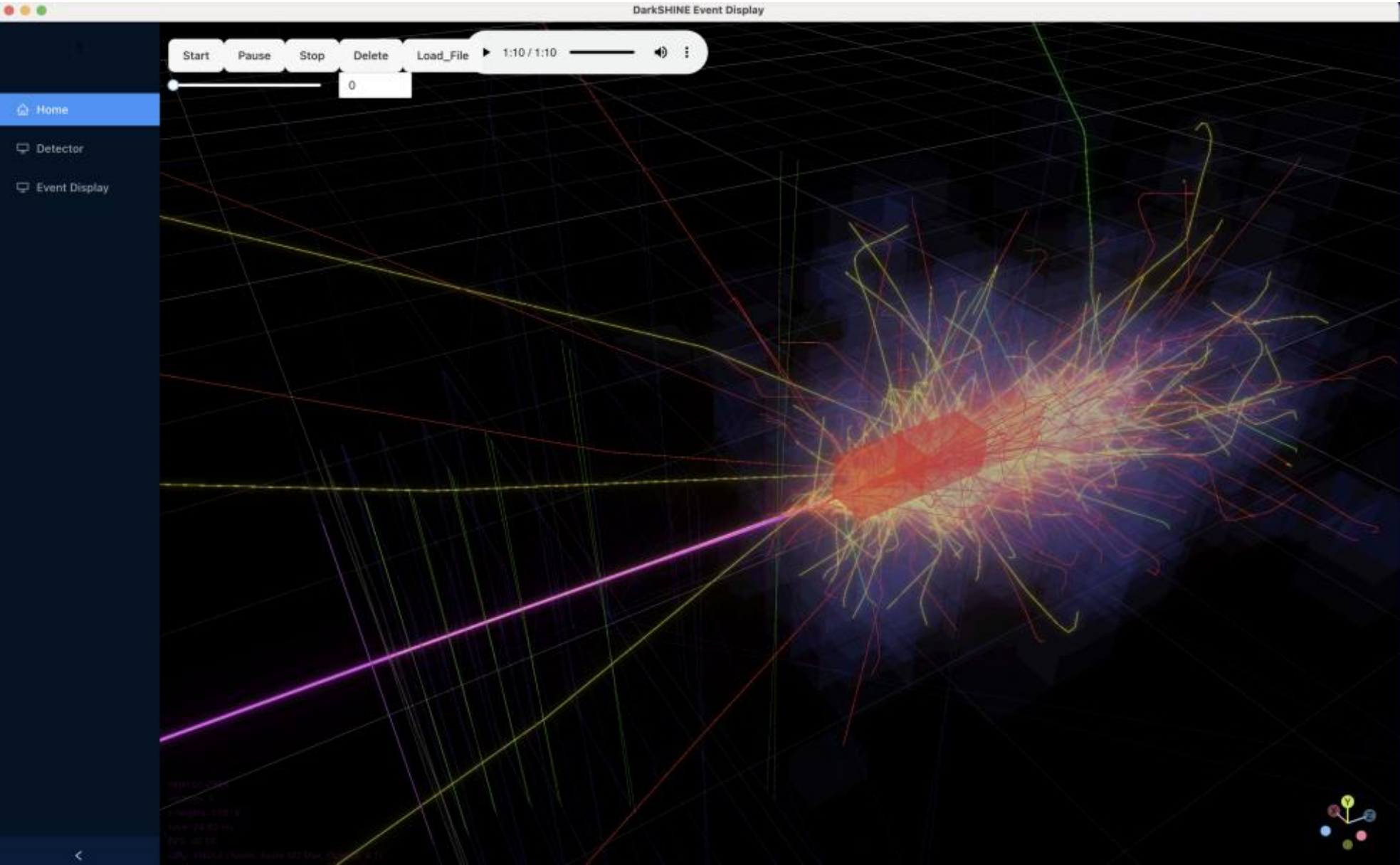
- Homogenous design with minimal EM energy leakage
- Fast response: MHz repetition rate w/o pileup
- Radiation hard: survive in  $\sim 10^7 \text{ rad}$  dose

**HCAL:** veto muon and hadron backgrounds

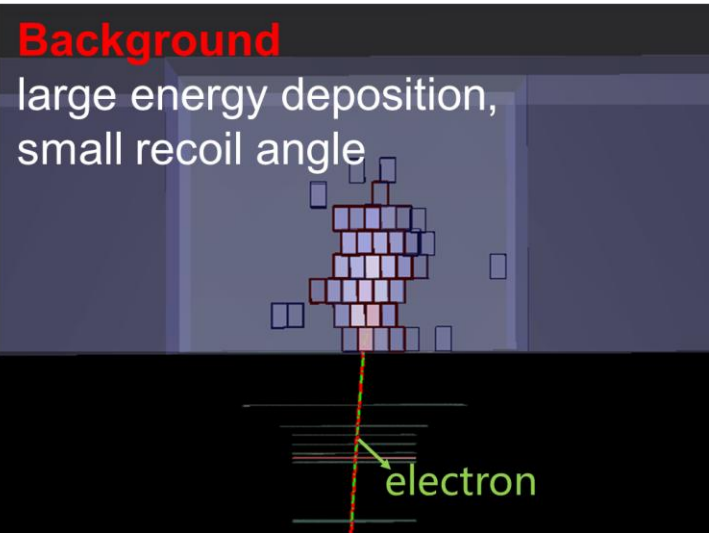
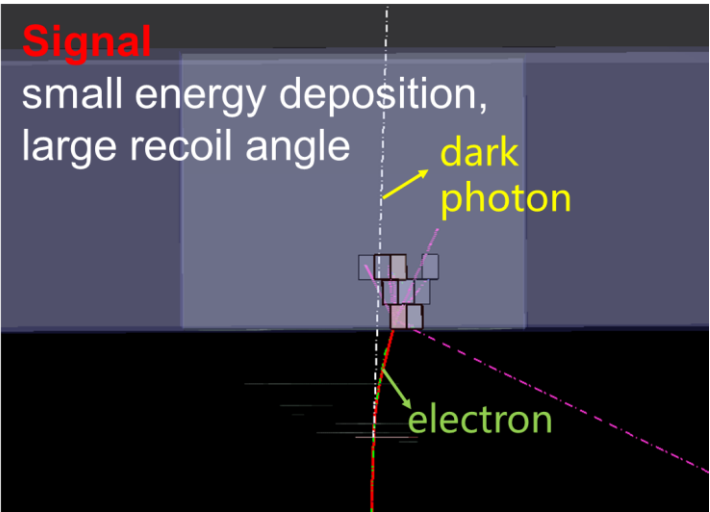
- Scintillator-absorber “sandwich” detector structure
- Small noise
- High rejection efficiency for neutron and muon

Additional system: readout electronics, trigger system, magnetic system (1.5 T), etc.

# Event Display



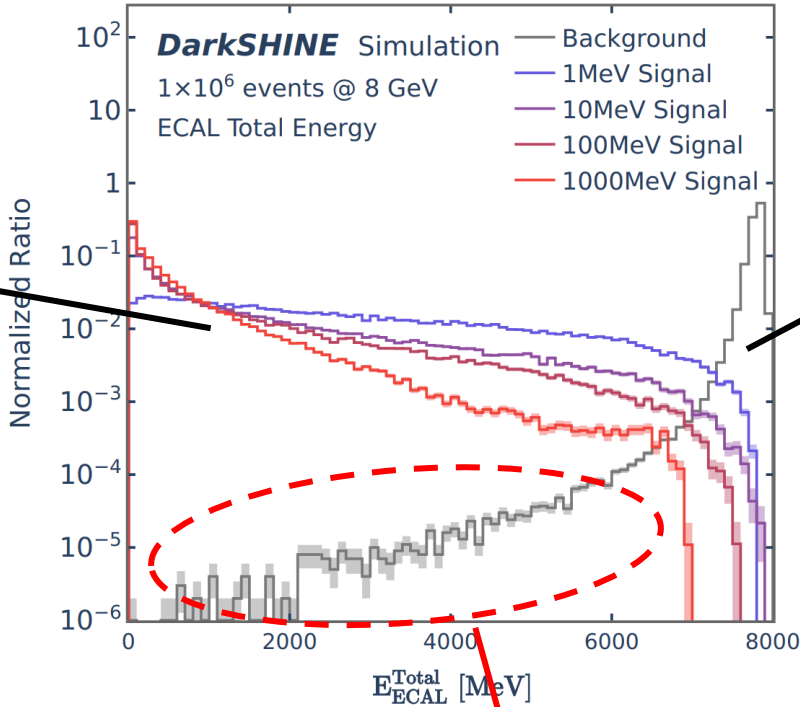
# Signatures of Signal and Background



## Energy deposition in ECAL

**Signal:**

Missing energy  
(vs. 8 GeV  $e^-$ )



**Background:**

~ 8 GeV,  
full deposition

## Background branching ratio

Inclusive	Bremsstrahlung	Electron-Nuclear	Photon-nuclear	Photon-to-muons
1.0	$6.7 \times 10^{-2}$	$3.76 \times 10^{-6}$	$2.32 \times 10^{-4}$	$1.64 \times 10^{-6}$

**Rare backgrounds are our main focus**



**Sci. China-Phys. Mech. Astron., 66(1): 211062 (2023)**

SCIENCE CHINA

Physics, Mechanics & Astronomy



• Article •

June 2022 Vol. 63 No. 1: 000000  
<https://doi.org/??>

## Prospective Study of Light Dark Matter Search with a Newly Proposed DarkSHINE Experiment

Jing Chen<sup>†b,c,a</sup>, Ji-Yuan Chen<sup>b,c</sup>, Jun-Feng Chen<sup>h</sup>, Xiang Chen<sup>b,c</sup>, Chang-Bo Fu<sup>ij</sup>, Jun Guo<sup>b,c</sup>,  
Le He<sup>f</sup>, Zheng-Ting He<sup>a,n</sup>, Kim Siang Khaw<sup>a,b,c</sup>, Jia-Lin Li<sup>b,c</sup>, Liang Li<sup>b,c</sup>, Shu Li<sup>a,b,c,d,e\*</sup>, Meng Lv<sup>g</sup>,  
Dan-Ning Liu<sup>a,b,c</sup>, Han-Qing Liu<sup>b,c</sup>, Kun Liu<sup>a,b,c\*</sup>, Qi-Bin Liu<sup>a,b,c</sup>, Yang Liu<sup>a,b,c</sup>, Ze-Jia Lu<sup>b,c</sup>,  
Cen Mo<sup>b,c</sup>, Si-Yuan Song<sup>b,c</sup>, Xiao-Long Wang<sup>ij</sup>, Yu-Feng Wang<sup>†a,b,c</sup>, Zhen Wang<sup>a,b,c</sup>, Zi-Rui Wang<sup>m</sup>,  
Wei-Hao Wu<sup>b,c</sup>, Dao Xiang<sup>k,a,l</sup>, Hai-Jun Yang<sup>b,c,a\*</sup>, Jun-Hua Zhang<sup>a,b,c</sup>, Yu-Lei Zhang<sup>†b,c</sup>,  
Zhi-Yu Zhao<sup>a,b,c</sup>, Xu-Liang Zhu<sup>a,b,c</sup>, Chun-Xiang Zhu<sup>b,c</sup>, and Yi-Fan Zhu<sup>b,c</sup>

## Highlight remarks

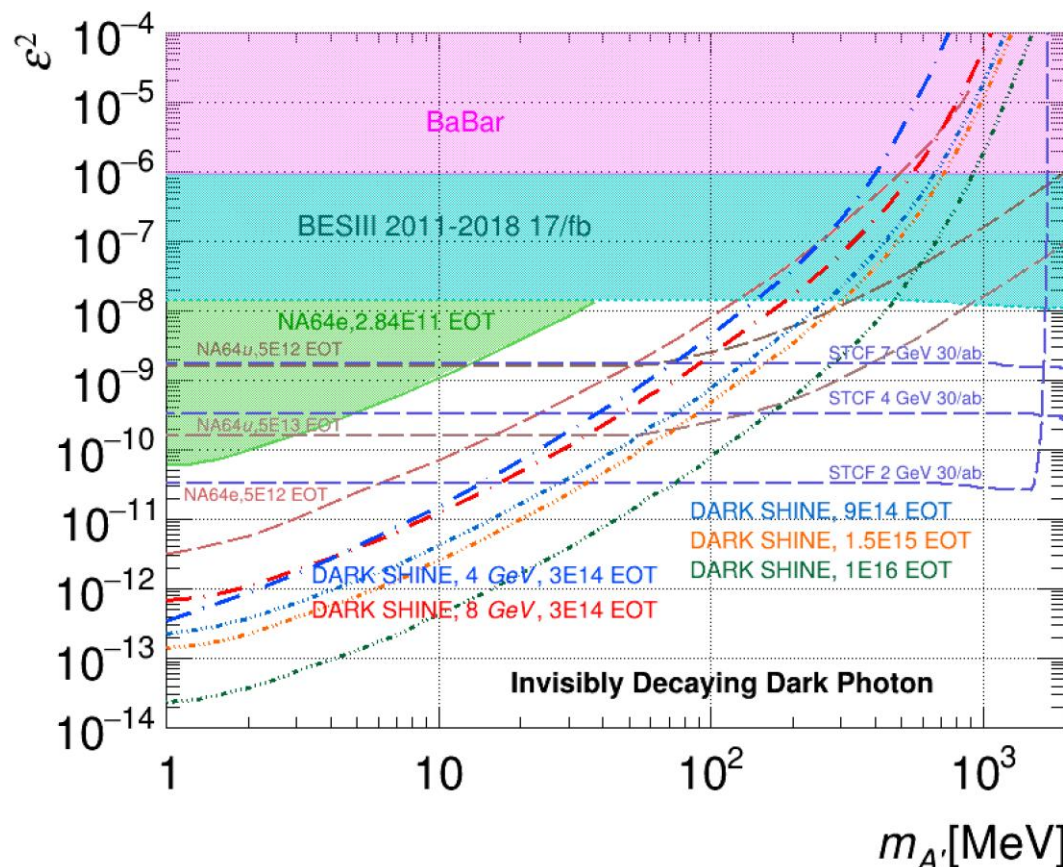
**Sci. China-Phys. Mech. Astron. 66(1): 211063 (2023)**

## “Editor’s Focus”

**Sci. China-Phys. Mech. Astron. 66(1): 211061 (2023)**

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SMS 2025



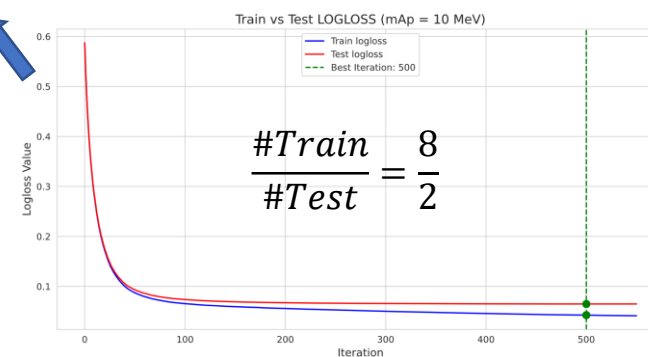
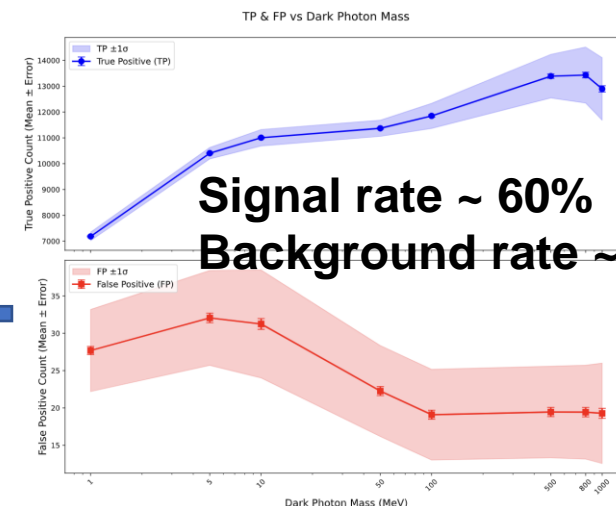
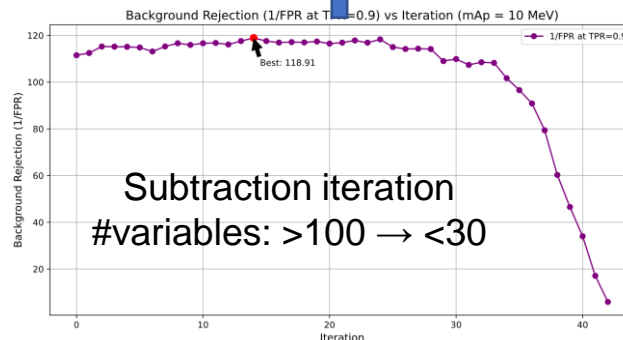
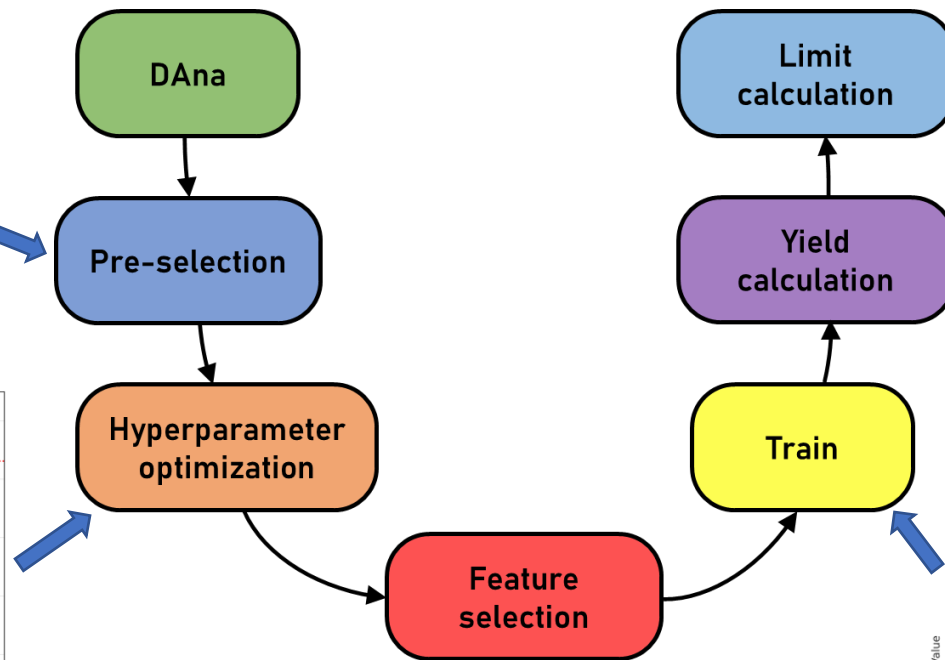
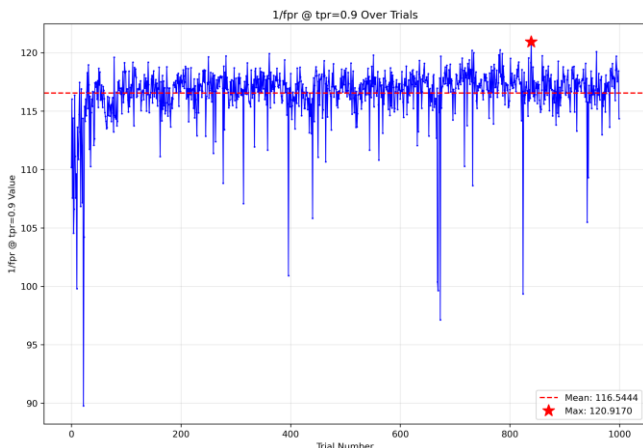
- Sensitivity competitive with other experiment internationally (e.g. NA64)
- Aim to deliver 10<sup>16</sup> EOT stat. and cover most of the sensitivity regions of interests

# Enhance Background Rejection Through Multi-Variable-Analysis

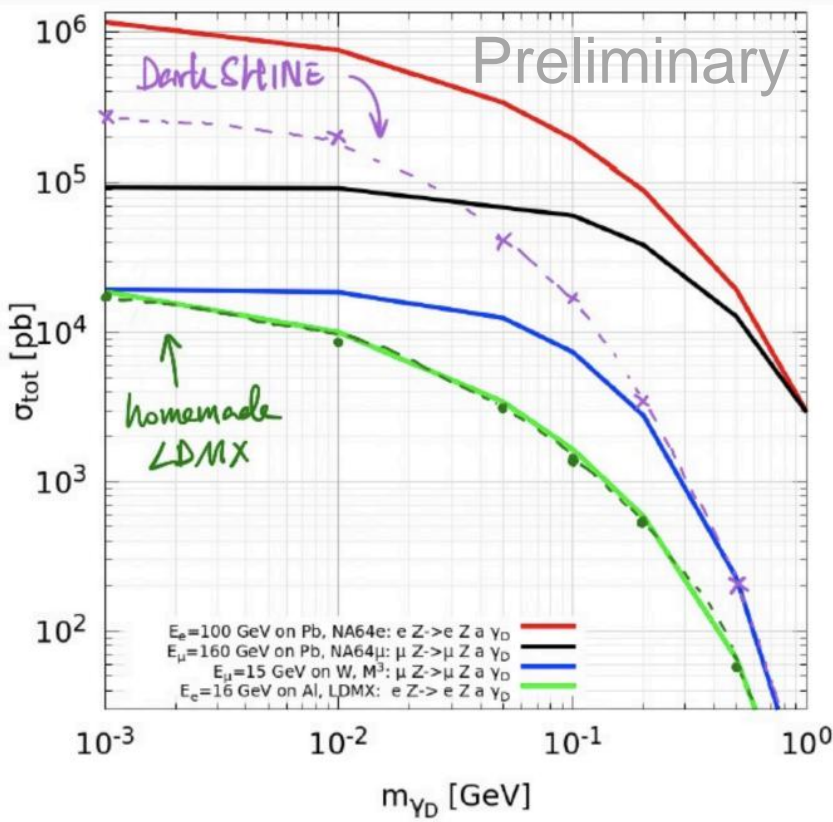
- XGBoost-based analysis pipeline → **100× reduction in bkg.** (preliminary)



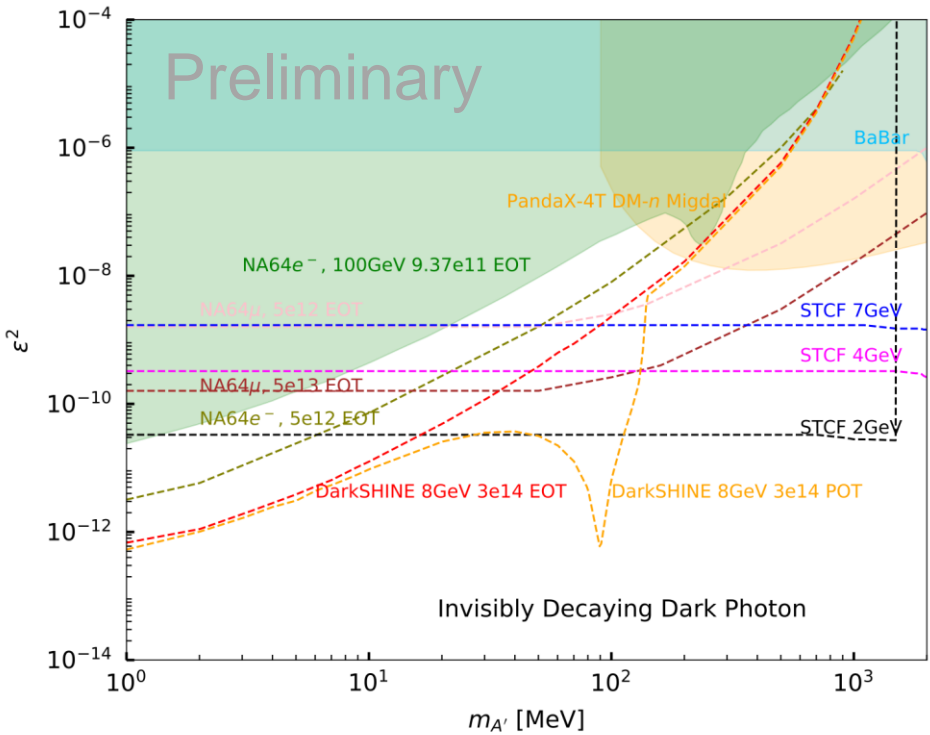
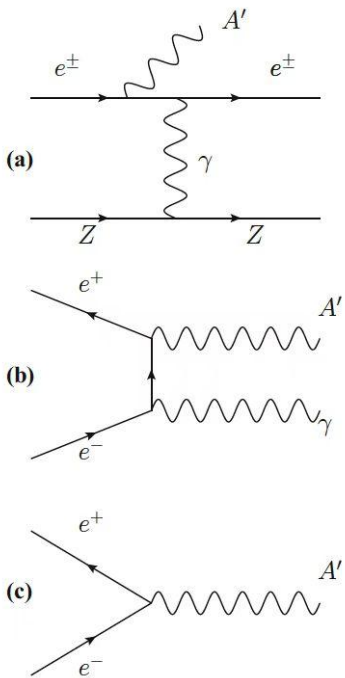
Signal  $\sim 10^5$   
Background  $\sim 10^{15}$



## Minimal dark Axion-like particle portal and Axion+DP co-existence



## Changing from electron beam to positron beam





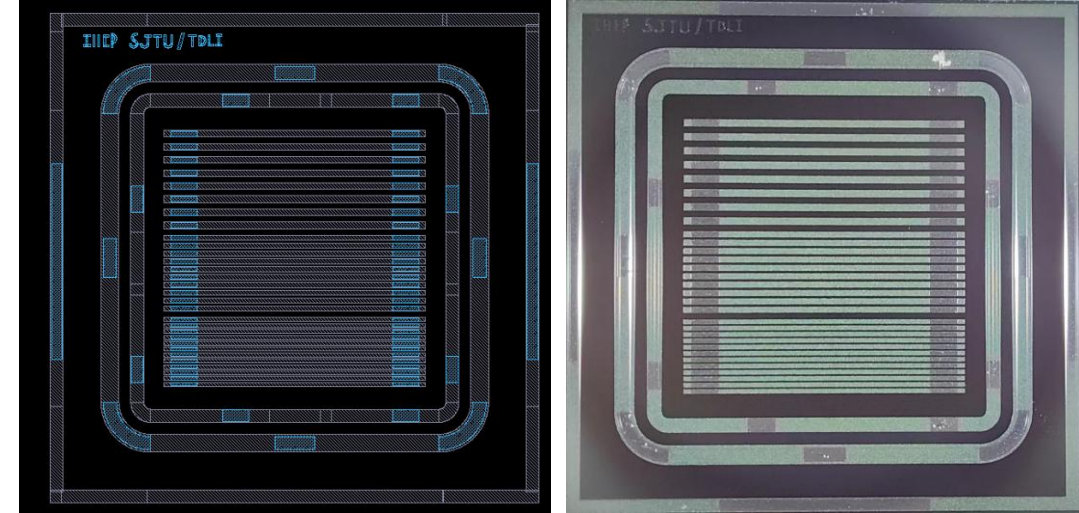
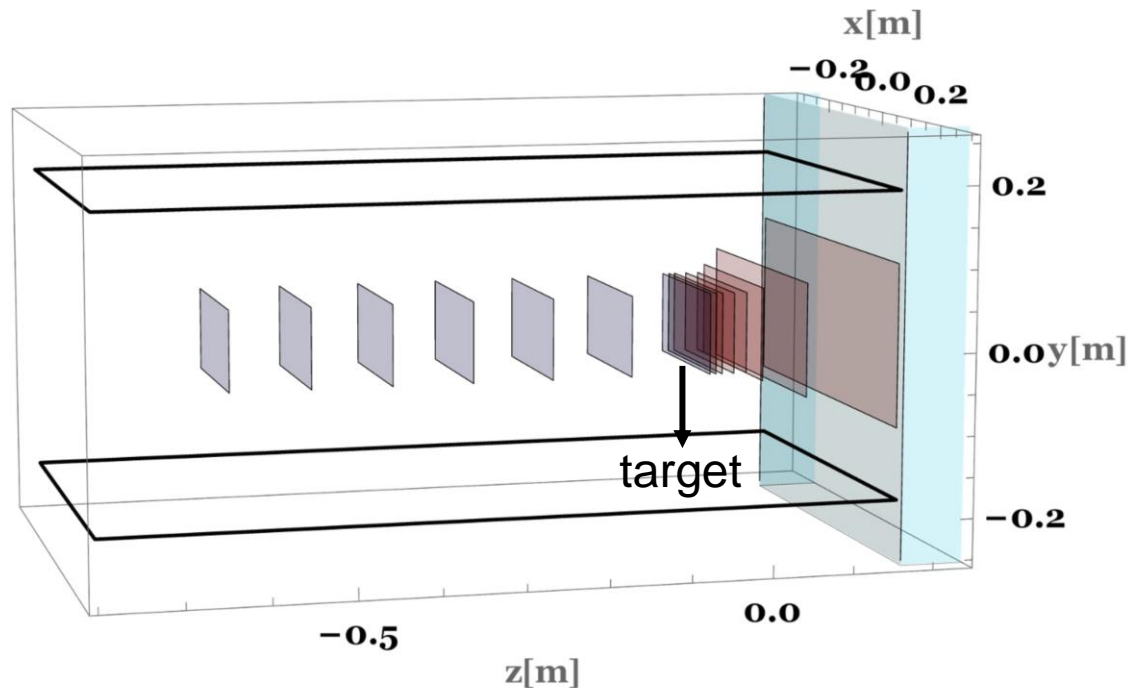
# Hardware R&D: Tracking System



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- Silicon strip detector in a strong magnetic field,  $\sim 10 \mu\text{m}$  position resolution
- 7 layers of tagging + 6 layers of recoil tracker, two silicon strips sensors w/ a small angle ( $0.1 \text{ rad}$ )

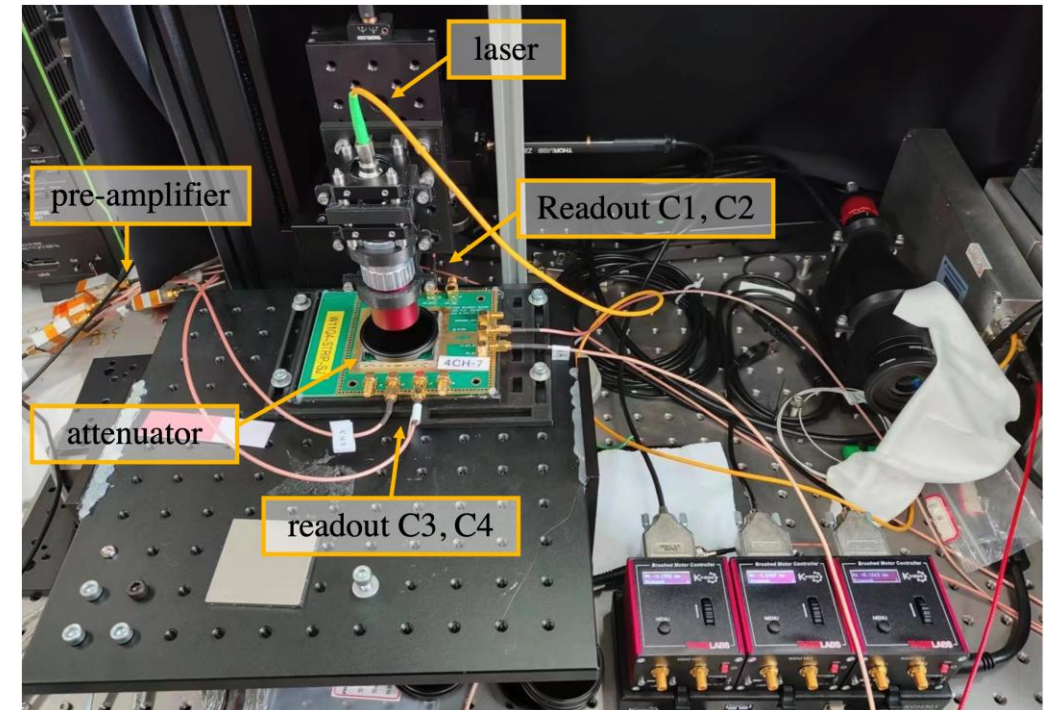
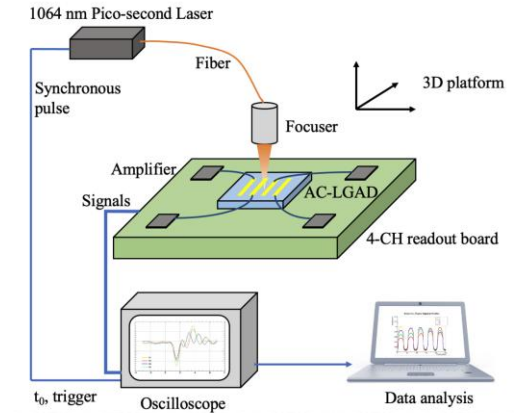
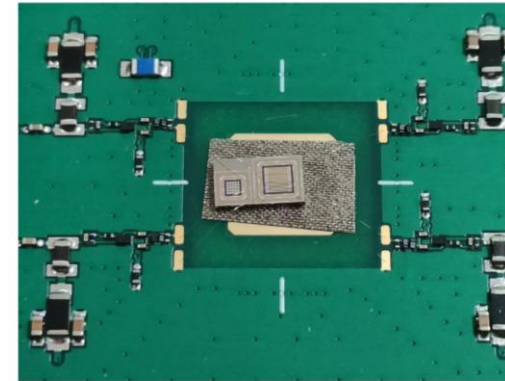
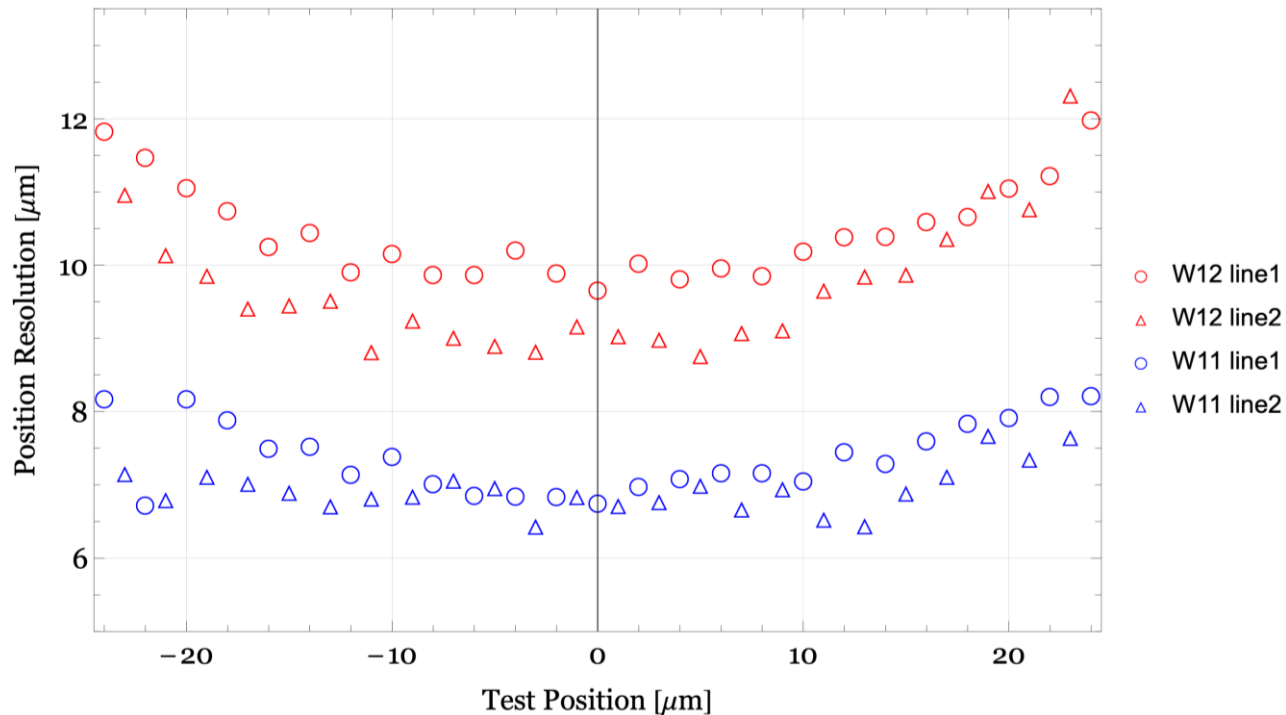


**AC-LGAD silicon strip sensor  $1 \times 1 \text{ mm}^2$  designed, in collaboration with IHEP**

# Hardware R&D: Tracking System

- AC-LGAD silicon strip sensor prototype tested using laser beam.
- Spatial resolution  $7\ \mu\text{m} \sim 12\ \mu\text{m}$  can be achieved, with good timing resolution.

*Nucl. Sci. Tech. 35, 201 (2024)*





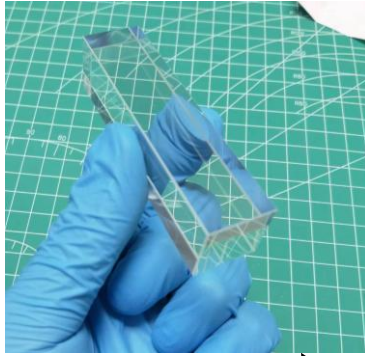
# Hardware R&D: Electromagnetic Calorimeter (ECAL)



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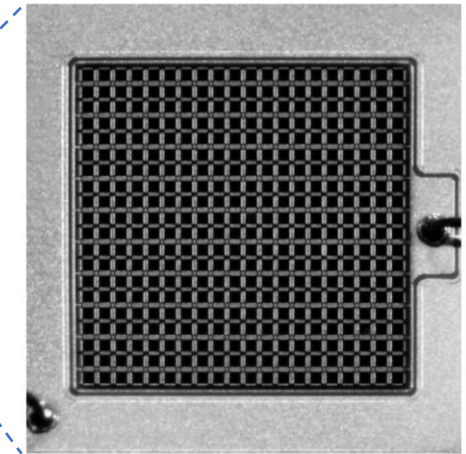
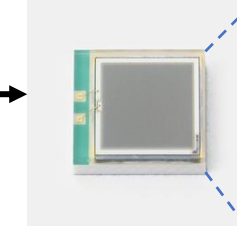
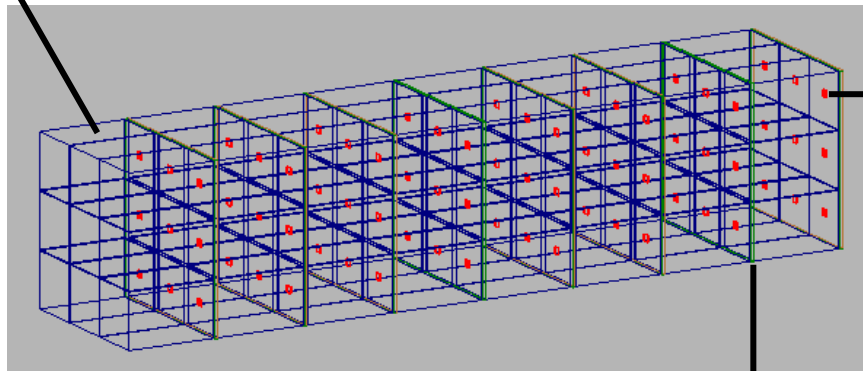
**Nucl. Sci. Tech. 36, 41 (2025)**



## LYSO crystal scintillator

- High light yield:  $\sim 30000$  ph/MeV
- Fast scintillation decay time: 40ns
- Great radiation tolerance

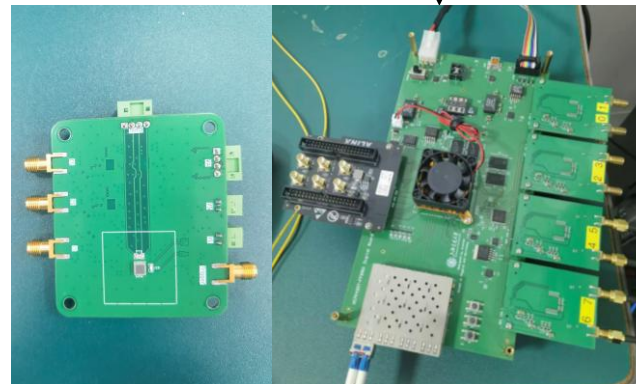
**Precise measurement  
on energy of electrons  
and photons**



Other parts: mechanics,  
cooling system, etc.

## Readout electronics

- Digitization and transmission of electrical signals
- High precision, large dynamic range, high-repetition operation



## Silicon Photomultiplier (SiPM)

- Compact size
- Single-photon sensitivity
- Great time resolution
- Low operating voltage



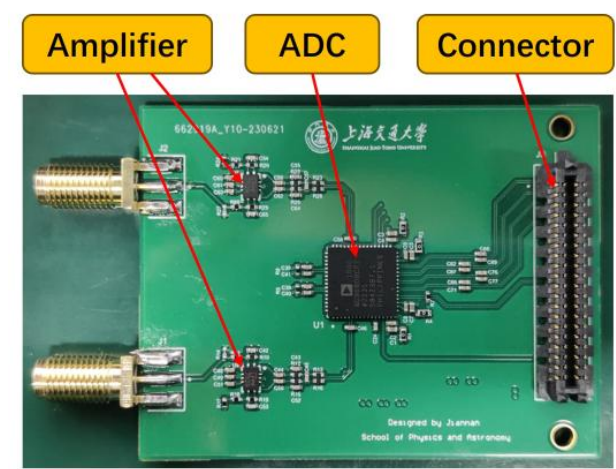
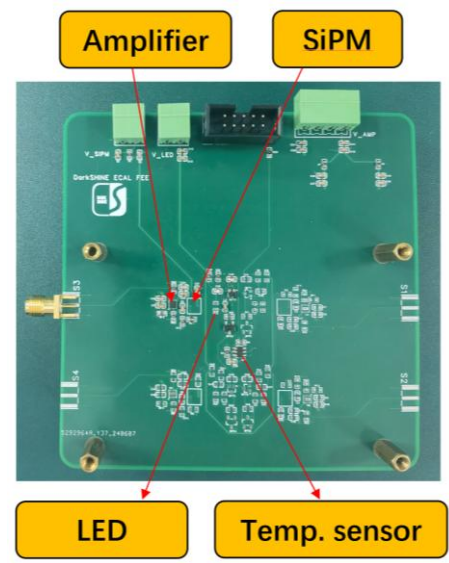
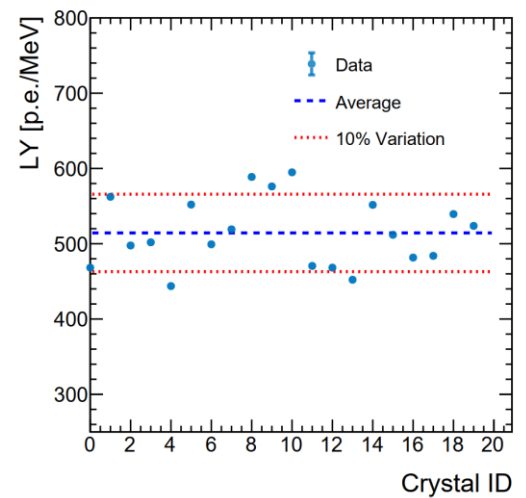
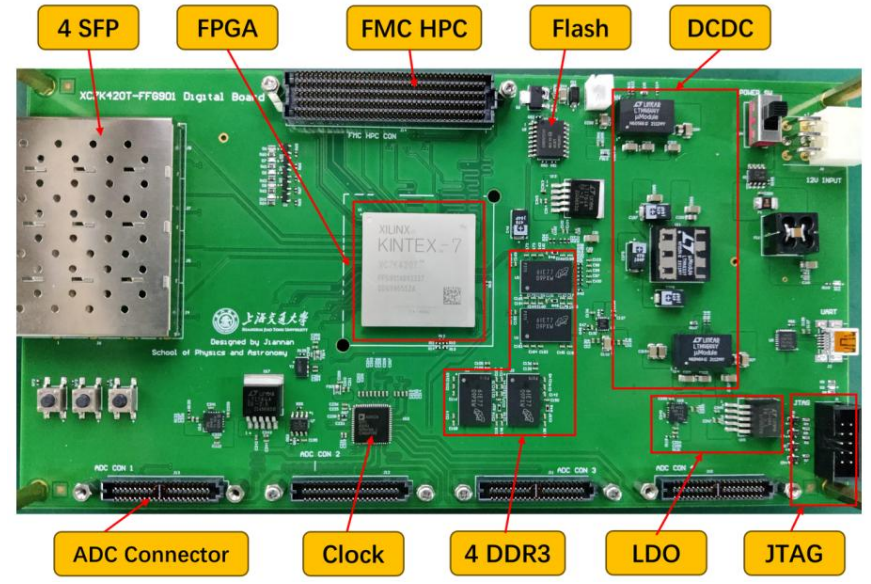
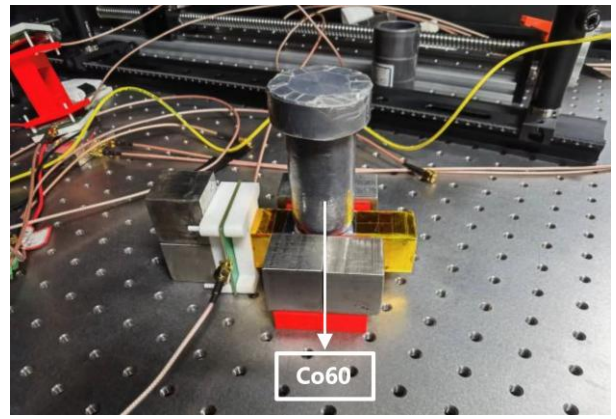
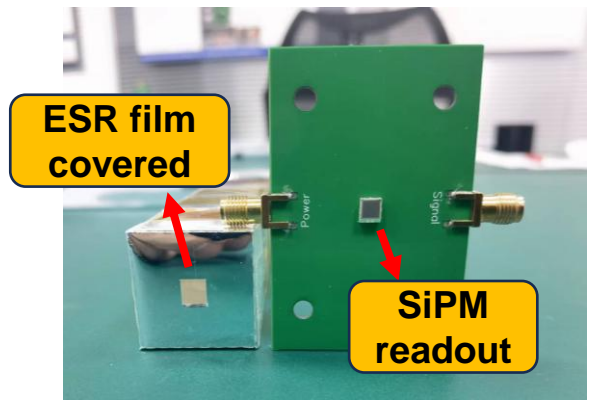
# Hardware R&D: Electromagnetic Calorimeter (ECAL)



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- Dedicated R&D works on crystal, SiPM and electronics





# Prototype Beam Test at DESY & CERN

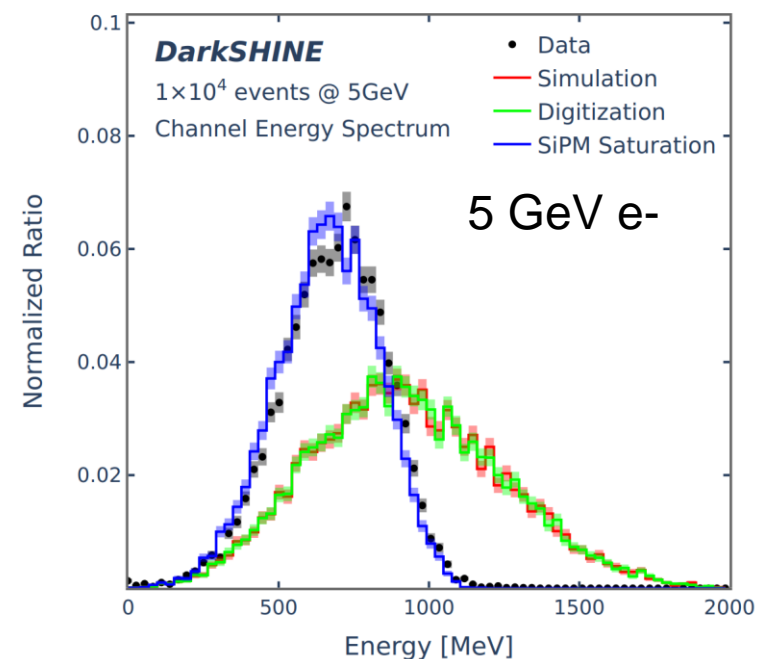
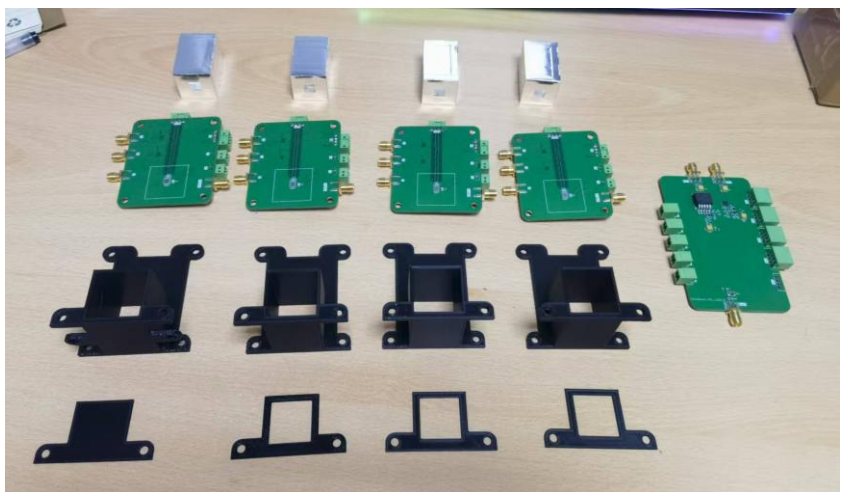


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- 4-ch crystal prototype with self-designed electronics
- **1<sup>st</sup> beam test of DarkSHINE sub-detector prototype**
  - DESY (Germany) 2024, and CERN (Switzerland) 2025

[arXiv: 2411.09345](https://arxiv.org/abs/2411.09345)



Good agreement b/w  
data and simulation

# Hardware R&D: Hadronic Calorimeter (HCAL)



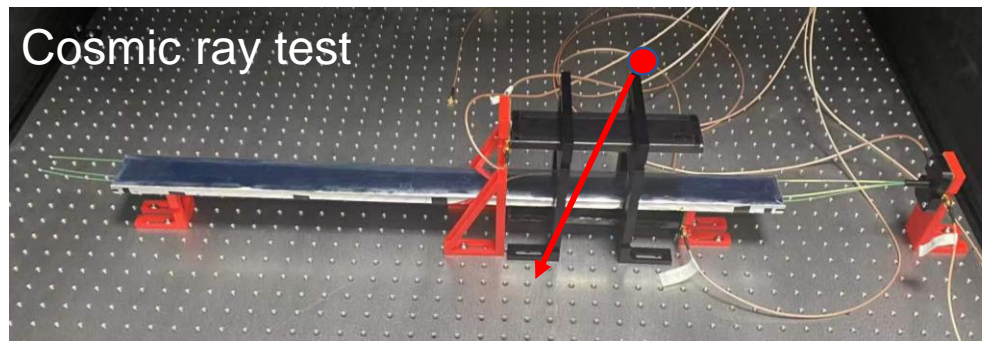
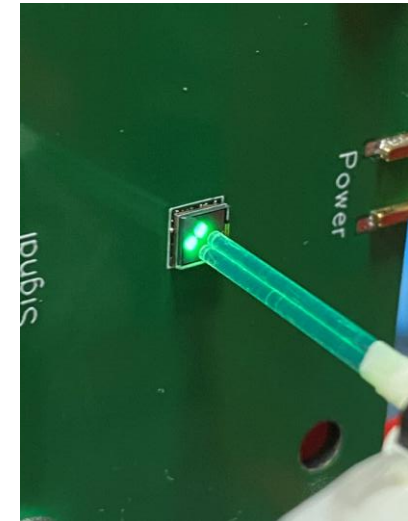
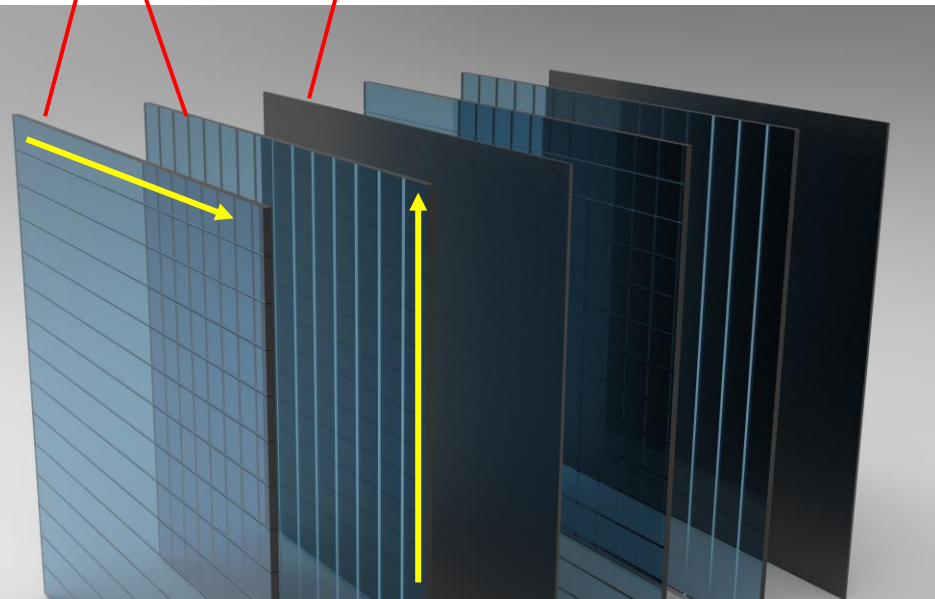
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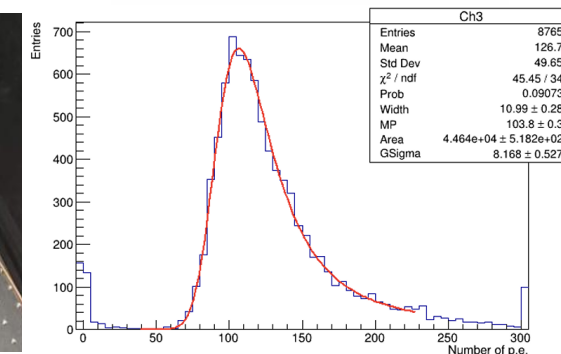
- Veto muon and hadron backgrounds
- Scintillator-Fe sampling structure
  - 1.5 m × 1.5 m × 2.5 m, ~11  $\lambda$
  - Plastic scintillator + Wavelength shift fiber + SiPM
- Good efficiency in n/p/ $\mu$  detection

**Nucl. Sci. Tech. 35, 148 (2024)**

Scintillator Absorber(Fe)



Cosmic ray test





# DarkSHINE Baseline Design Report

arXiv:2411.09345v1 [physics.ins-det] 14 Nov 2024

## DarkSHINE

### *Baseline Design Report*

### Physics Prospects and Detector Technologies



## Contents

<b>1</b>	<b>Physics</b>	<b>3</b>
1.1	Introduction to Dark Matter	3
1.2	General Picture of Worldwide Research	4
1.3	DarkSHINE Physics Program	9
<b>2</b>	<b>Detectors</b>	<b>11</b>
2.1	Introduction to DarkSHINE Beamline	11
2.2	Overview of Detector System	12
2.3	Magnetic Field System	13
2.4	Tracking System	14
2.4.1	Tagging Tracker	15
2.4.2	Recoil Tracker	16
2.4.3	Tracking detector module and expected performances	17
2.4.4	Tracking Electronics	18
2.5	Target System	19
2.6	Electromagnetic Calorimeter System	22
2.6.1	Introduction	22
2.6.2	Software configuration	23
2.6.3	ECAL expected performance	25
2.6.4	Radiation damage	27
2.6.5	ECAL Electronics	28
2.6.6	Measurements	31
2.7	Hadronic Calorimeter System	37
2.7.1	Introduction	37
2.7.2	HCAL Conceptual Design	38
2.7.3	HCAL Expected performance	40
2.7.4	Electronic	44
2.7.5	HCAL Sensitive Unit and Performance Test	44
2.8	Mechanical System	48
2.8.1	Constrain and Requirements	48
2.8.2	Beamline and Detector Interface	48
2.8.3	Support Structure of Tracker System	48
2.8.4	Support for Electromagnetic Calorimeter System	48
2.8.5	Support for Hadronic Calorimeter System	49



## Dark SHINE Seasonal Workshop

22 May 2025, 09:00 → 24 May 2025, 17:10 Asia/Shanghai

Tsung-

Descrip

## DarkSHINE New Physics Opportunity Mini-Forum

Saturday 2 Aug 2025, 19:30 → 21:55 Asia/Shanghai

Tsung-Dao

Haijun Yang

Liang Li (Sh

Description

09:00 → 09:30

09:30 → 10:00

10:00 → 10:30

## DarkSHINE Winter Mini-Tutorial

Saturday 8 Nov 2025, 10:00 → 18:10 Asia/Shanghai

Tsung-Dao Lee Institute/N4F-N400 - meeting room (Tsung-Dao Lee Institute)

Haijun Yang (Shanghai Jiao Tong University (CN)) , Shu Li (TDLI, SJTU) , Kun Liu (TDLI / SJTU) , Yifan Zhu , Xuliang Zhu ,  
Haidar Mas'ud Alfanda (TDLI, SJTU)

Description Tencent meeting ID: 719270255

Password: 295188



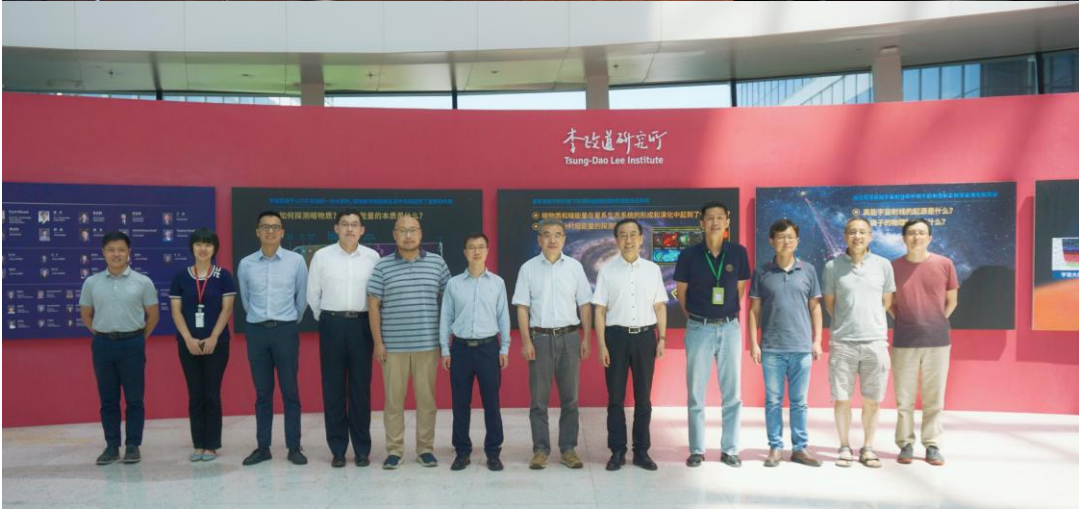


# Collaboration with SHINE



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- DarkSHINE: a fixed-target experiment searching for dark photon bremsstrahlung and invisible decay via missing  $E/p$
- Sensitivity competitive with other international experiments.
- Extensive R&D was carried out on the detector's critical technologies
- Conceptual design report completed — submission underway
- Future: explore more physics opportunities and develop large-scale detector prototypes

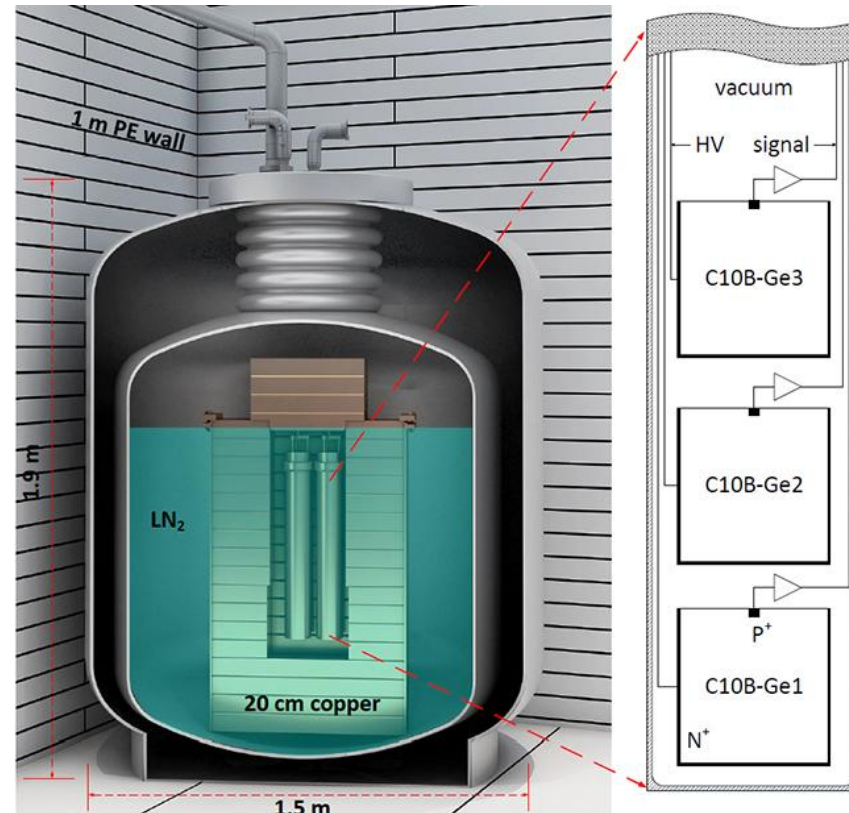
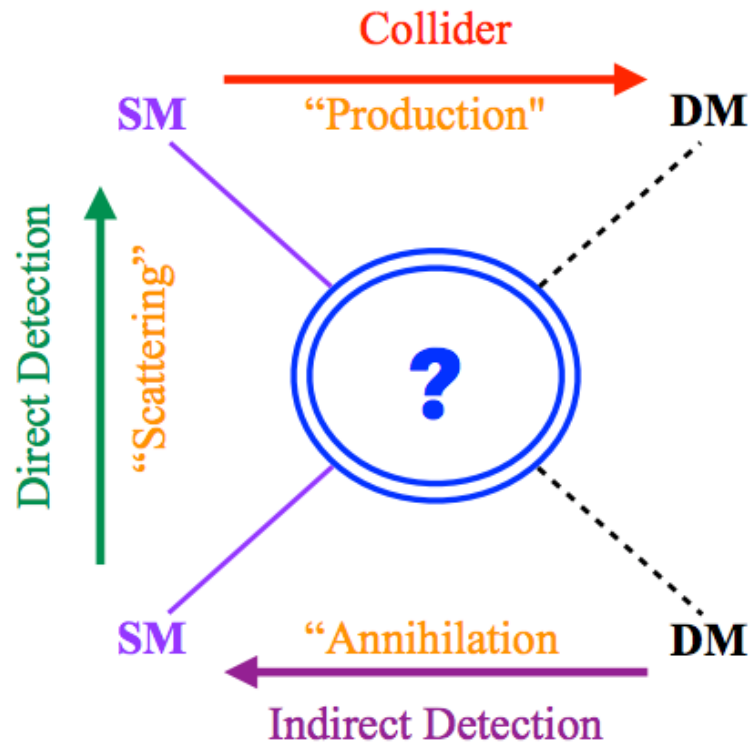
Thanks!





## ● Dark Matter Direct Detections

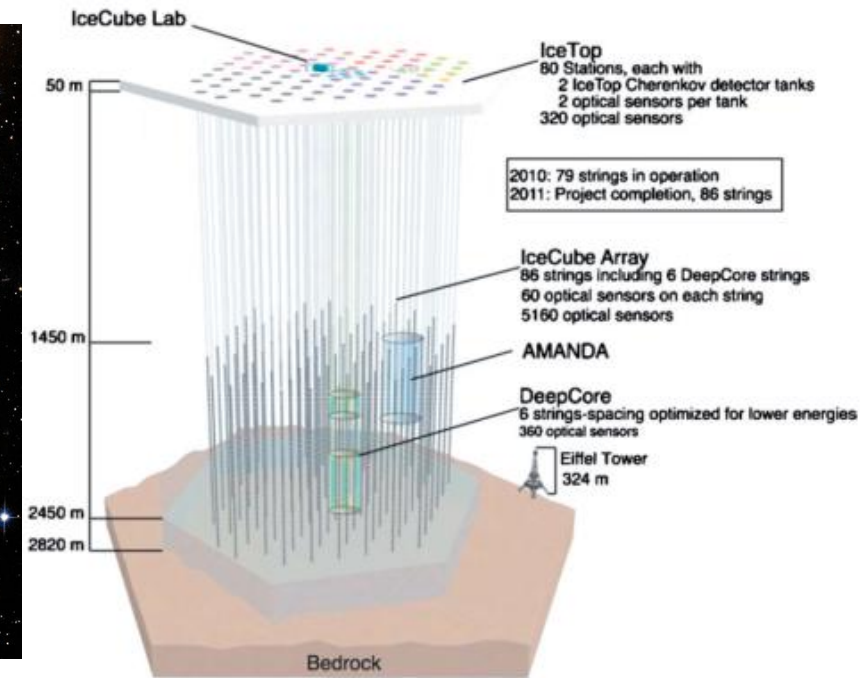
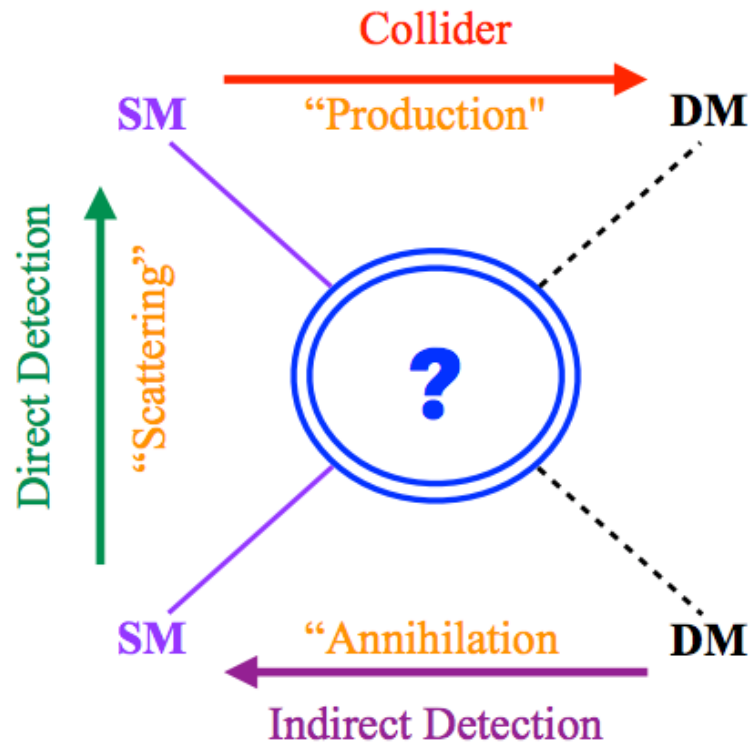
- Direct Detection (DD): nuclear recoils from DM-nuclei scattering (XENONnT, LZ, CDEX, PandaX, ...)
- Indirect Detection (ID): products from DM annihilation (DAMPE, HESS, IceCube, ..)
- Colliders: DM production in high-energy collisions, focusing on the productions of a SM particle(s) (X) with large missing  $E_T$





## ● Dark Matter Indirect Detections

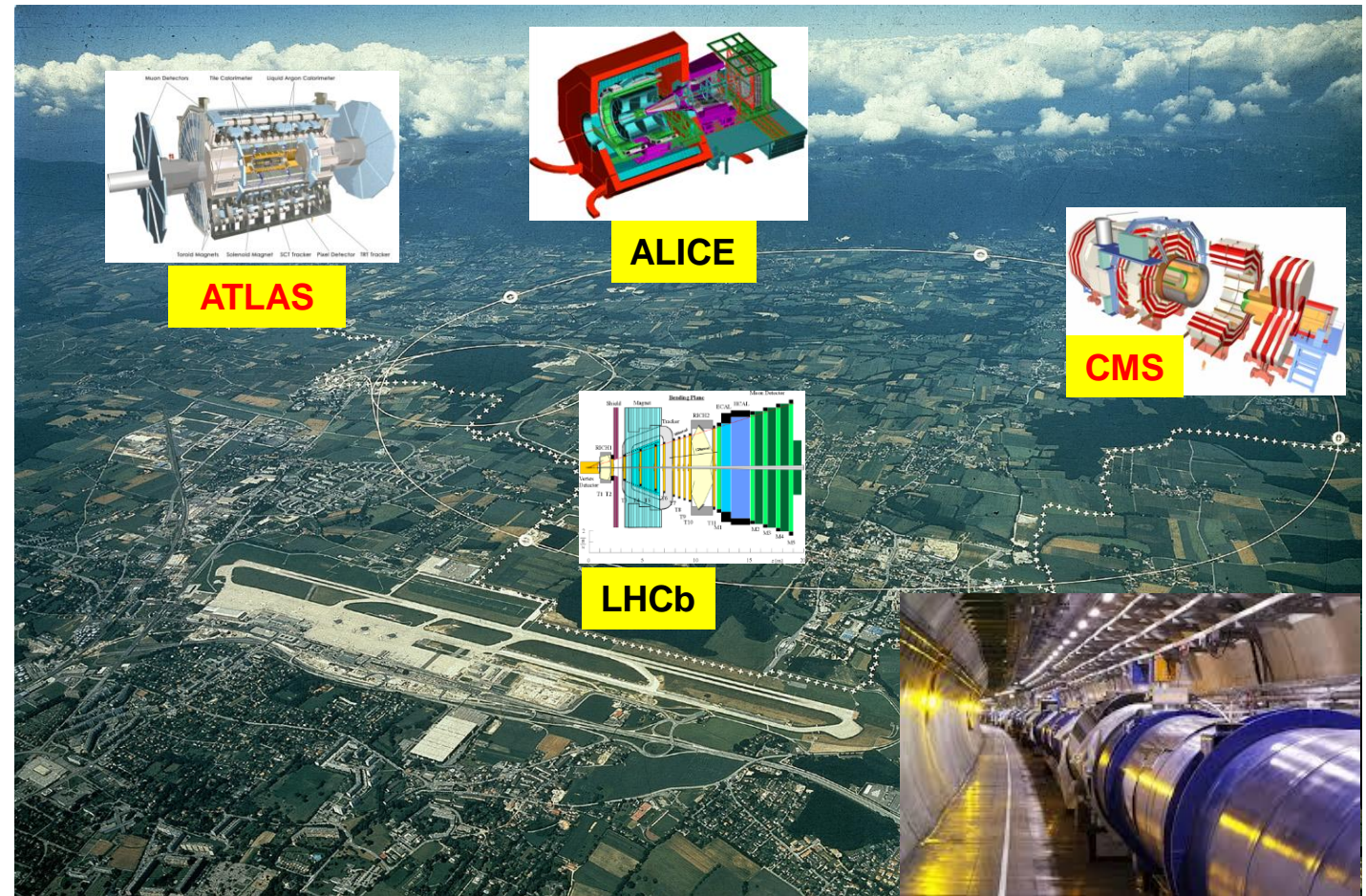
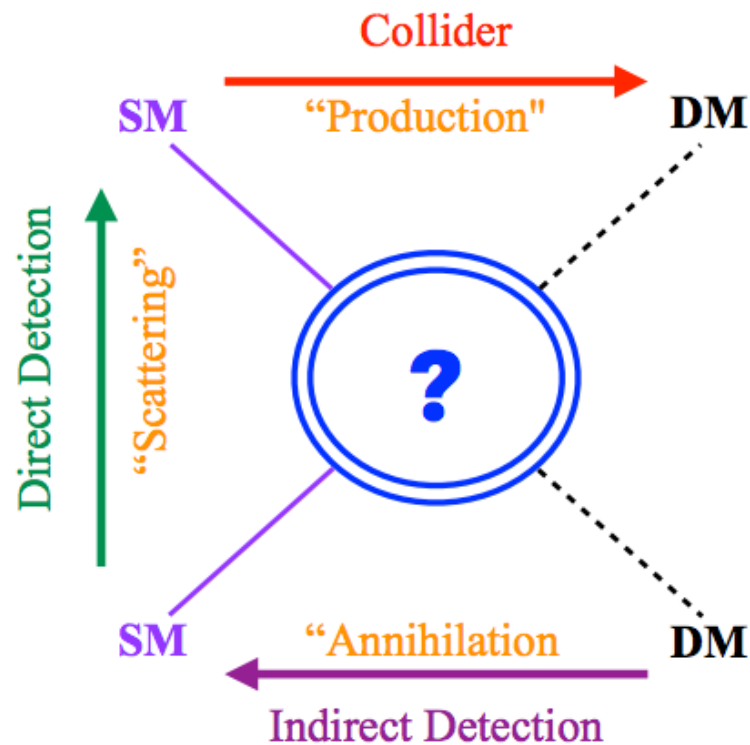
- Direct Detection (DD): nuclear recoils from DM-nuclei scattering (XENONnT, LZ, CDEX, PandaX, ...)
- Indirect Detection (ID): products from DM annihilation (DAMPE, HESS, IceCube, ..)
- Colliders: DM production in high-energy collisions, focusing on the productions of a SM particle(s) (X) with large missing  $E_T$





# ● Dark Matter Collider productions

- Direct Detection (DD): nuclear recoils from DM-nuclei scattering (XENONnT, LZ, CDEX, PandaX, ...)
- Indirect Detection (ID): products from DM annihilation (DAMPE, HESS, IceCube, ..)
- Colliders: DM production in high-energy collisions, focusing on the productions of a SM particle(s) (X) with large missing  $E_T$

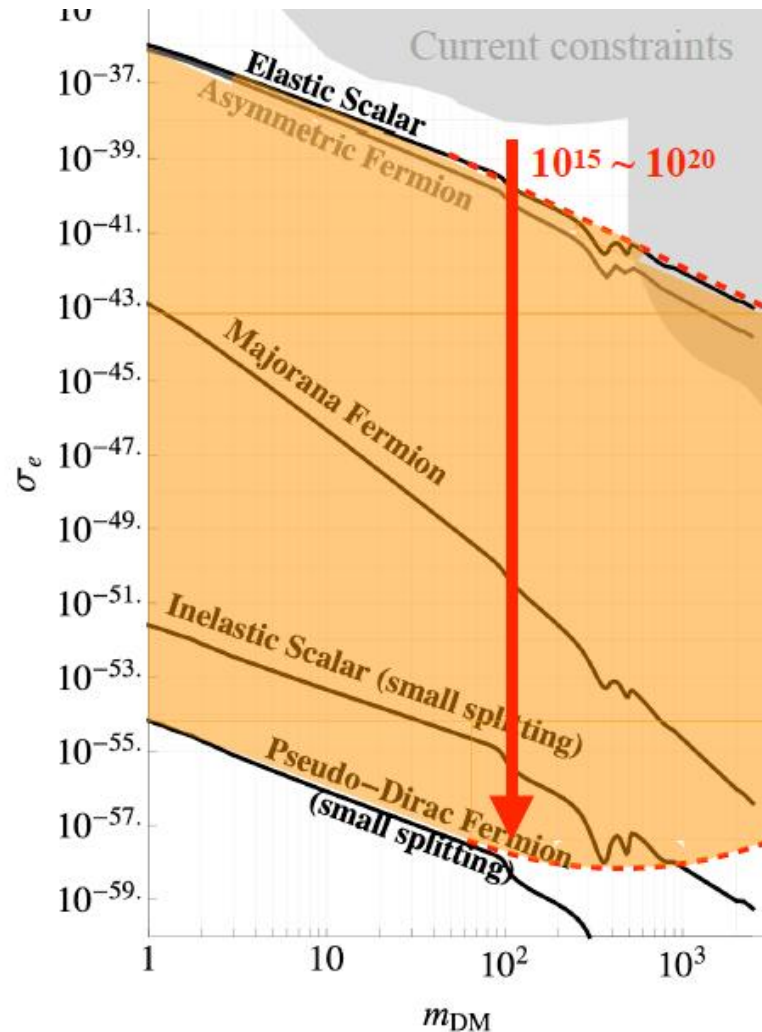




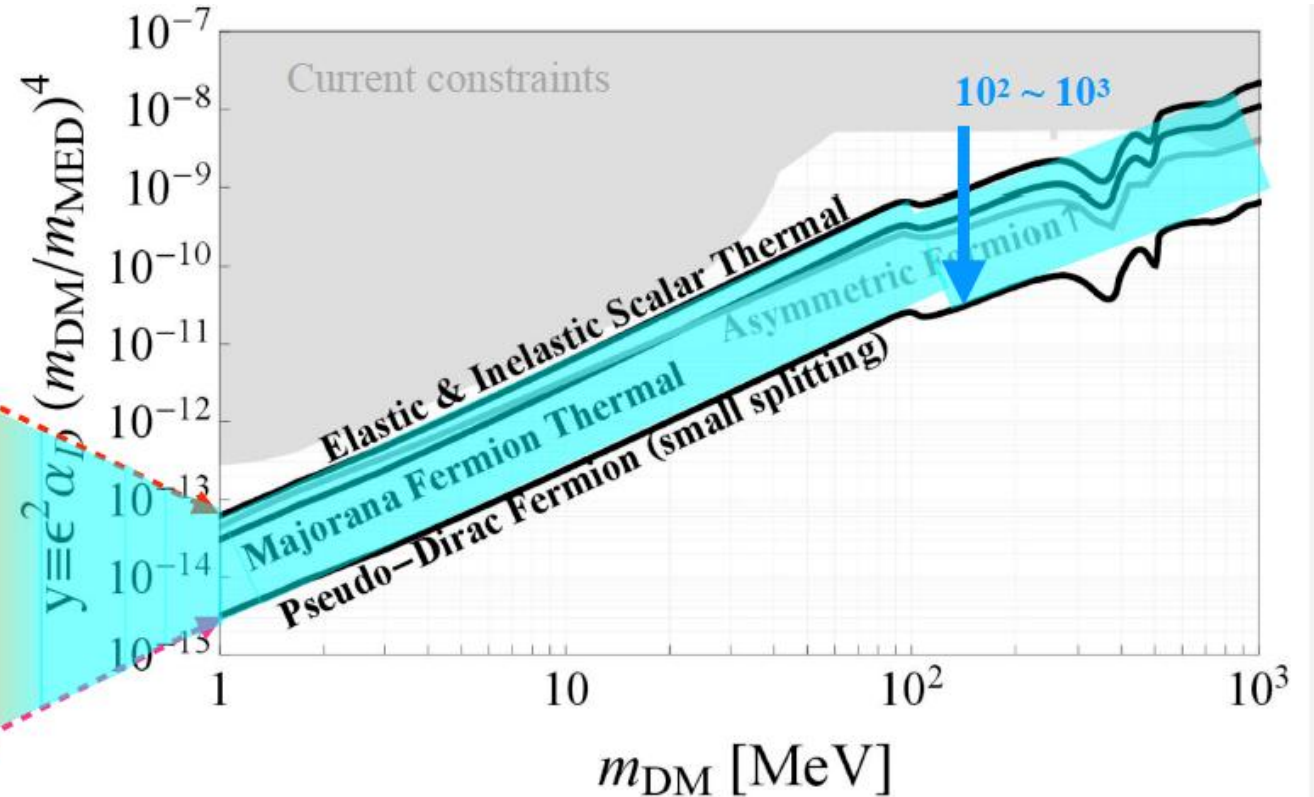
# Dark Matter search at Accelerator Experiments



## Direct Detection

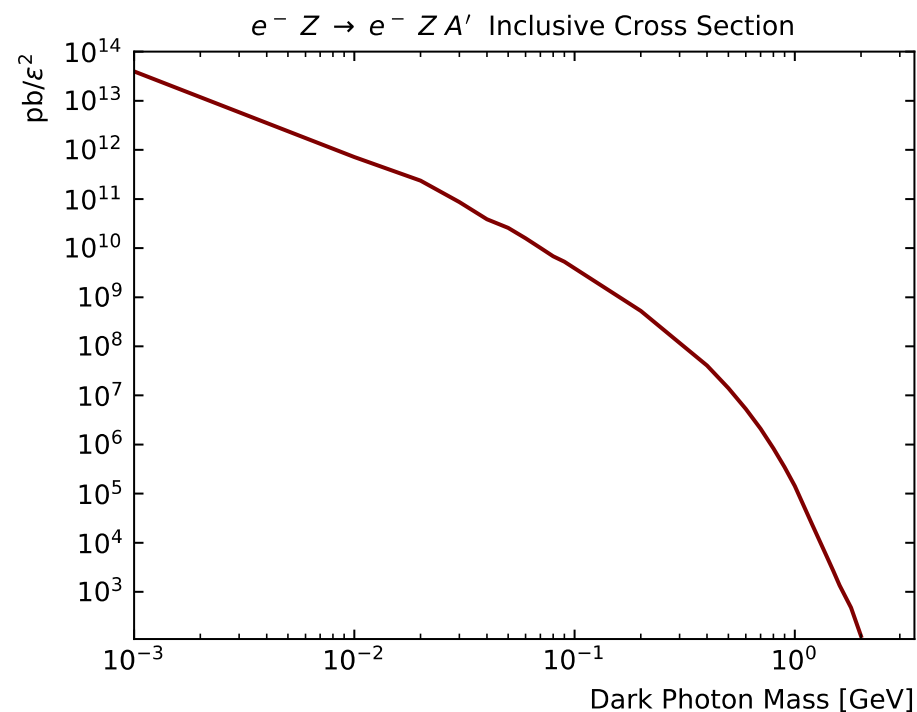


## Accelerator based experiments



- In accelerator-based experiments, this difference can be reduced to  $10^2 \sim 10^3$  orders of magnitudes, due to the fact of insensitive to DM's mass and spin in its production.





Inclusive cross-section of dark photon bremsstrahlung from electron interacting with W target, assuming  $\epsilon = 1$ .

# Signal & Background Features

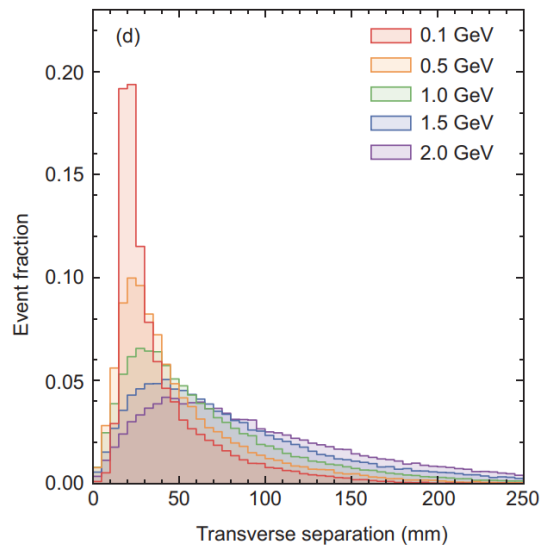
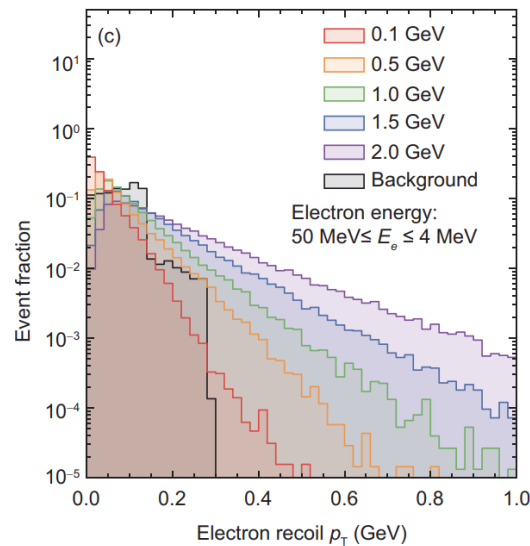
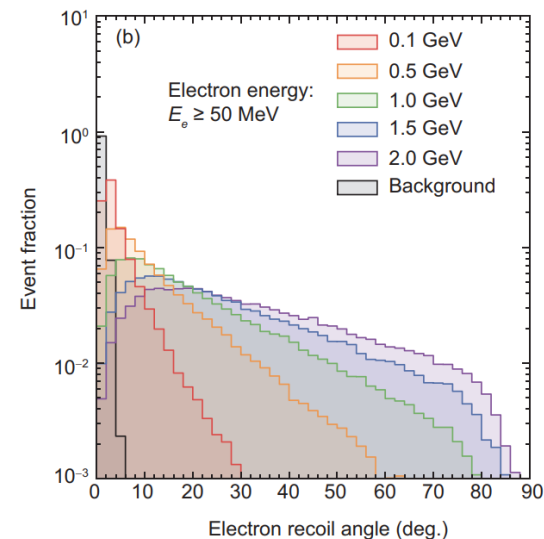
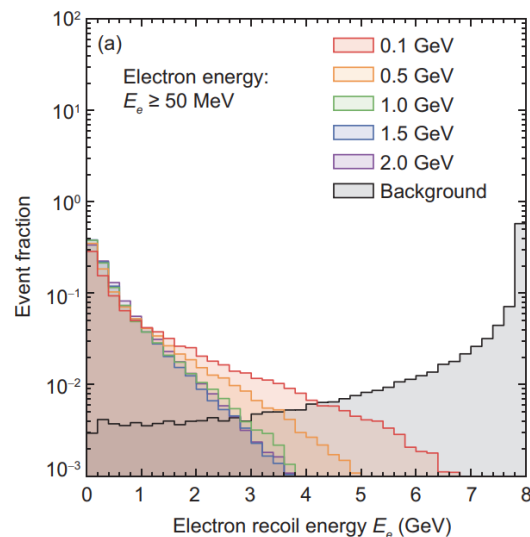


## Signal

- Low recoil energy
- Large recoil angle and recoil  $p_T$

## Background

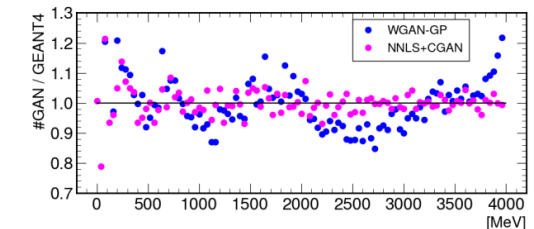
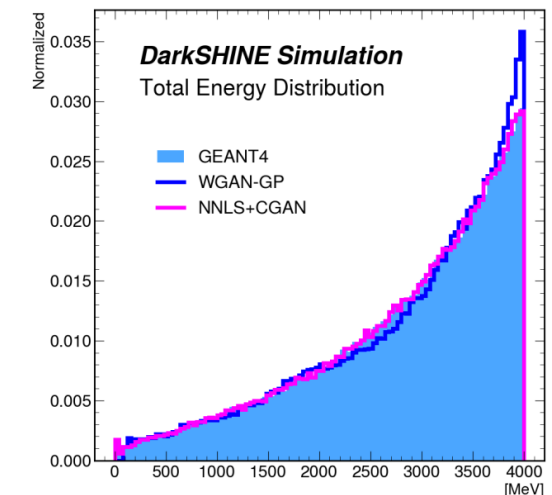
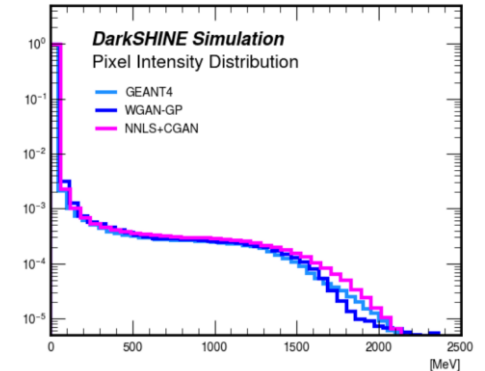
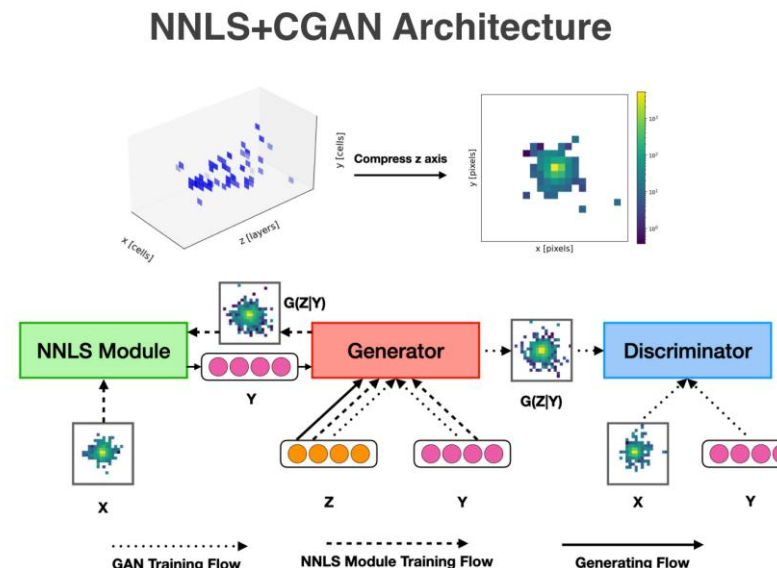
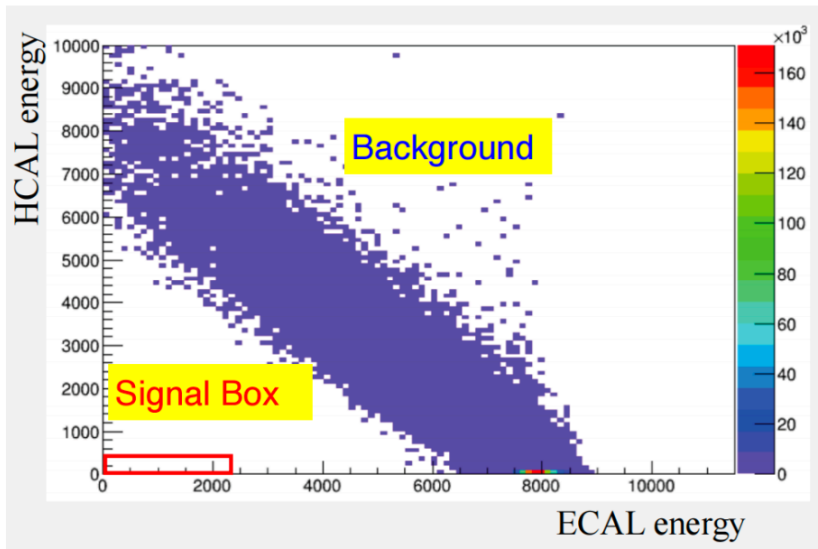
- Large recoil energy
- Small recoil angle and recoil  $p_T$



# Fast Simulation with Machine Learning



- Motivation: As we increase the simulation statistical quantities, it tends to be more background events closing to signal region
  - Geant4 simulation for  $3 \times 10^{14}$  EOTs will consume a massive amount of resources
    - ~5ms per inclusive event.  $3 \times 10^{14}$  EOTs ~  $10^4$  CPUs run for 5 years
  - Generation with deep learning network could save a lot of time/money
- NNLS+CGAN can work for ECAL energy data synthesis
  - $\times 10^4$  speed-up could be achieved
  - Upgrade to provided shower shape for detailed validation

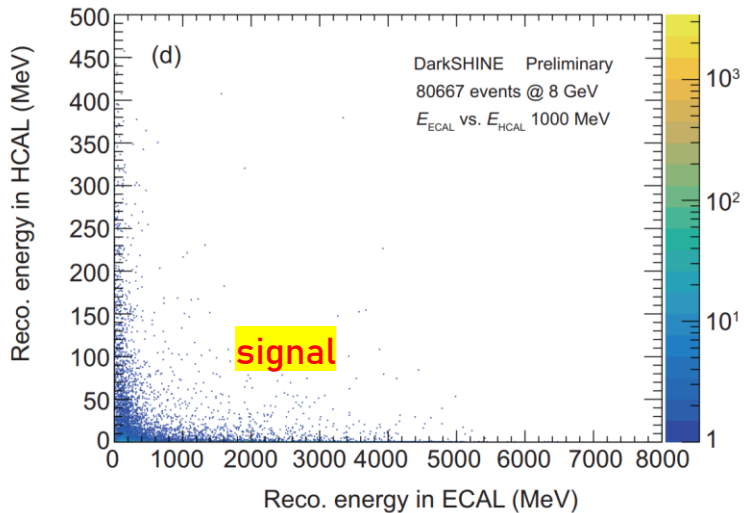
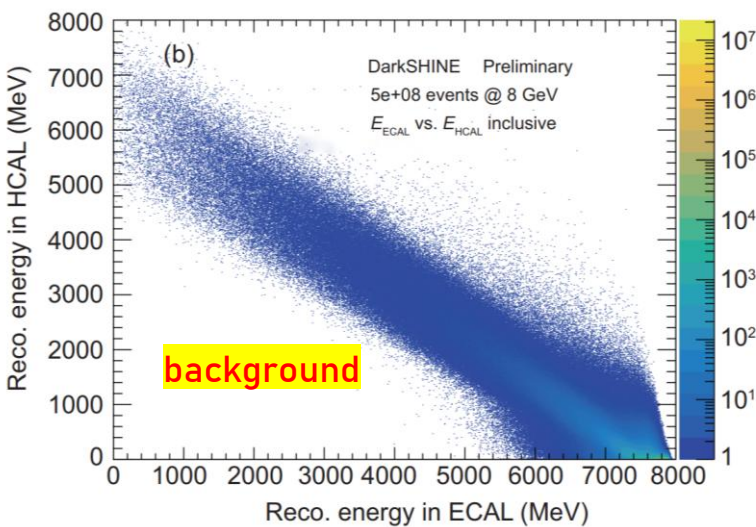
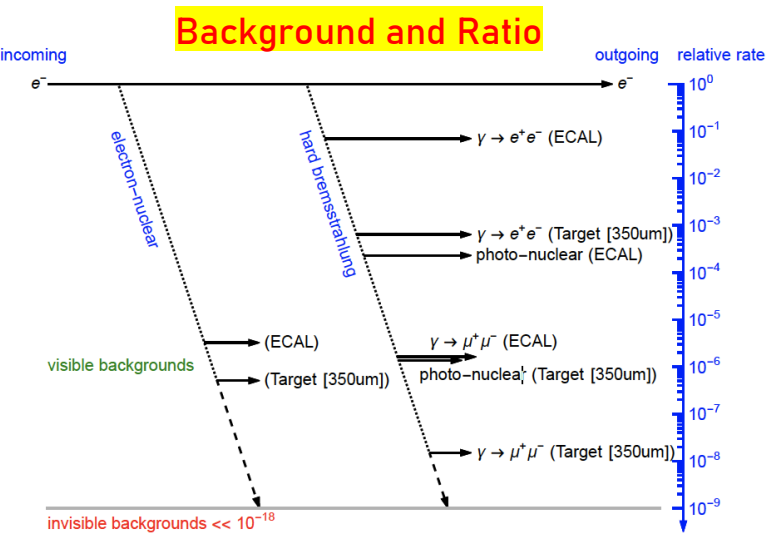


[1] Mirza M, Osindero S. Conditional generative adversarial nets[J]. arXiv preprint arXiv:1411.1784, 2014.

[2] Chen D, Plemmons R J. Nonnegativity constraints in numerical analysis[M]//The birth of numerical analysis. 2010: 109-139.



# Background Rejection



Signal region

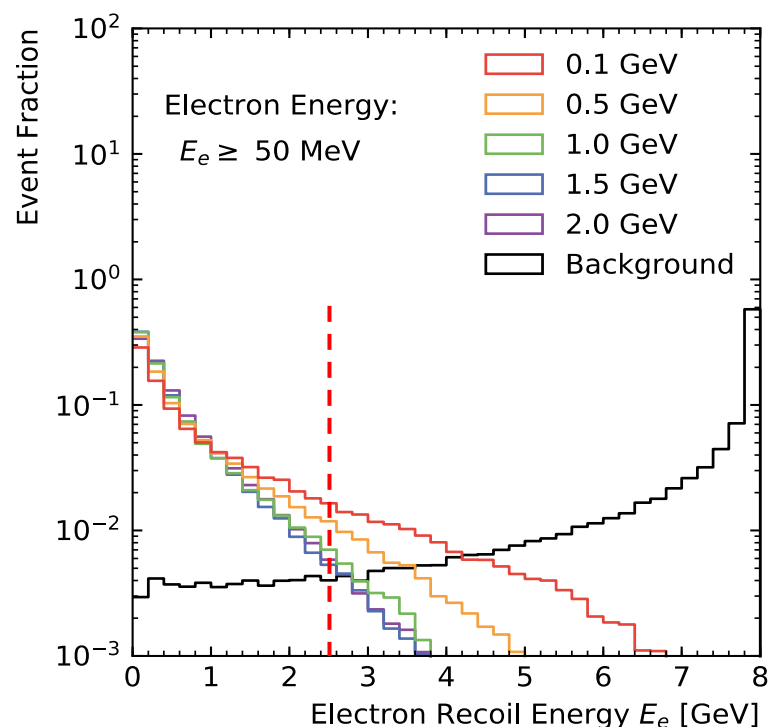
	EN_ECAL	PN_ECAL	GMM_ECAL	EN_target	PN_target	GMM_target	Hard_brem	Inclusive
Total events	100	100	100	100	100	100	100	100
Only 1 track	58.87	70.48	87.36	5.85	5.88	< 10 <sup>-3</sup>	78.73	84.40
$p_{\text{tag}} - p_{\text{rec}} > 4 \text{ GeV}$	0.0044	0.0033	0.0041	5.58	5.46	< 10 <sup>-5</sup>	70.49	4.80
$E_{\text{HCAL}}^{\text{total}} < 100 \text{ MeV}$	< 10 <sup>-3</sup>	< 10 <sup>-3</sup>	0	0.30	0.72	0	69.61	4.76
$E_{\text{HCAL}}^{\text{MaxCell}} < 10 \text{ MeV}$	< 10 <sup>-3</sup>	< 10 <sup>-3</sup>	0	0.13	0.27	0	65.00	4.48
$E_{\text{HCAL}}^{\text{MaxCell}} < 2 \text{ MeV}$	< 10 <sup>-3</sup>	< 10 <sup>-3</sup>	0	0.058	0.095	0	58.14	4.04
$E_{\text{ECAL}}^{\text{total}} < 2.5 \text{ GeV}$	0	0	0	0	0	0	0	0

Zero background can be achieved, for  $2.5 \times 10^9$  inclusive EOTs and  $\sim 10^{12}$  rare EOTs!

# Dark Photon Efficiency



- After applying all cuts in signal region, a **signal efficiency of around 60%** can still be achieved.



Signal loss due to the cut box

